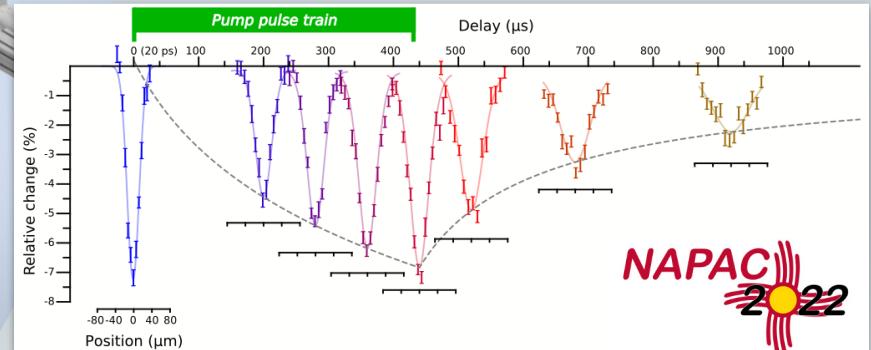
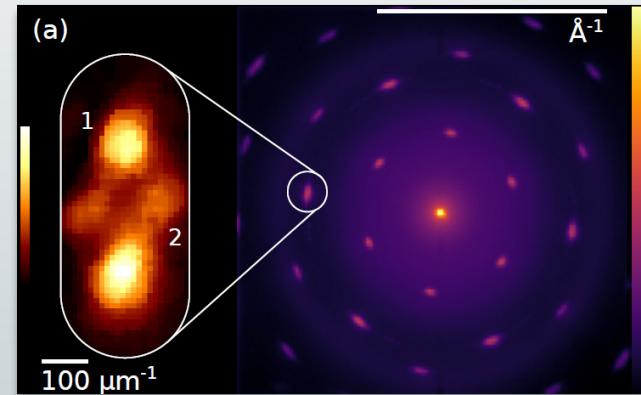


Ultrafast Electron Diffraction with Low Emittance Photocathodes

Jared Maxson, Cornell University



CLASSE
Cornell Laboratory for Accelerator-based Sciences & Education

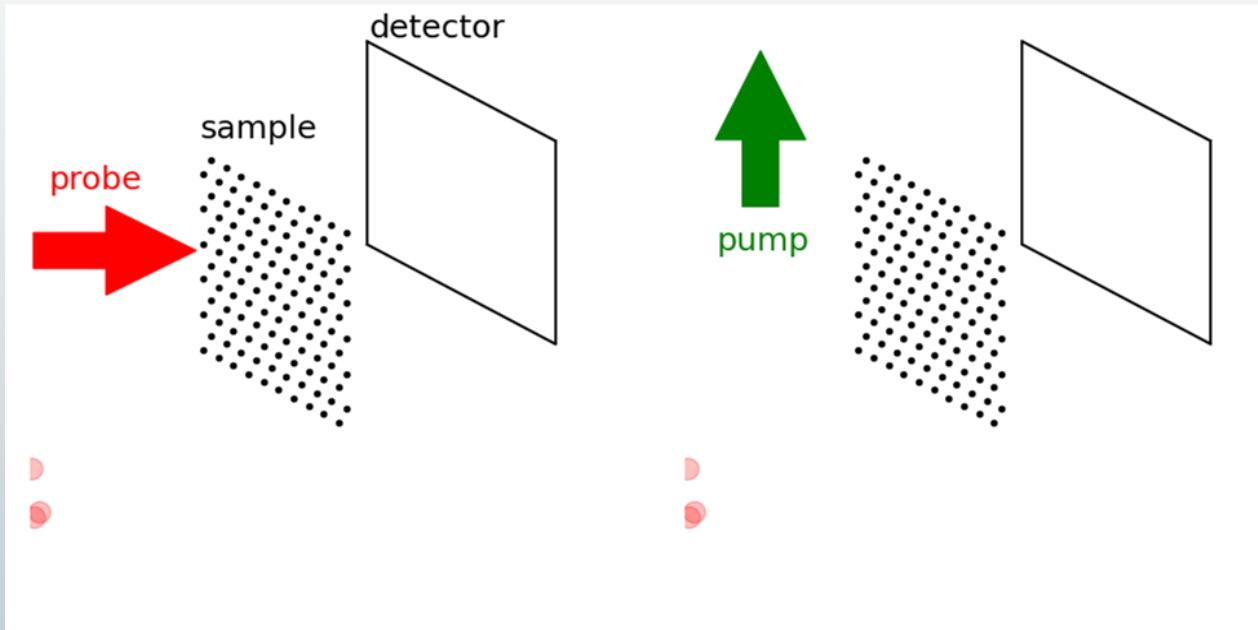


U.S. DEPARTMENT OF
ENERGY

Office of
Science

NAPAC 2022

Introduction to UED: The need for *brightness*



We want a **small transverse probe size** → some samples are hard to make with large dimensions

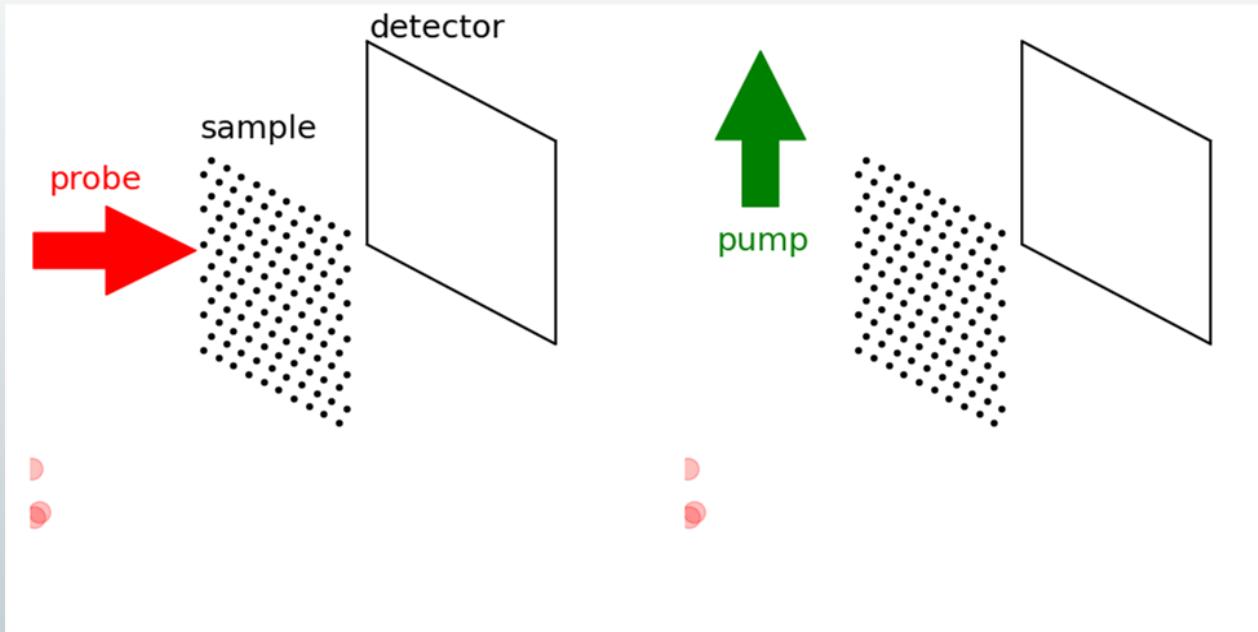
We want a **small transverse momentum spread** → need high coherence to see small features in k-space

We want a **short bunch duration** → natural timescale for atomic motion is fs-ps

We want as **many electrons as possible** → large signal to noise for subtle diffraction features

High source brightness is critical for UED!

Introduction to UED: The need for *brightness*



We want a **small transverse probe size** → some samples are hard to make with large dimensions

We want a **small transverse momentum spread** → need high coherence to see small features in k-space

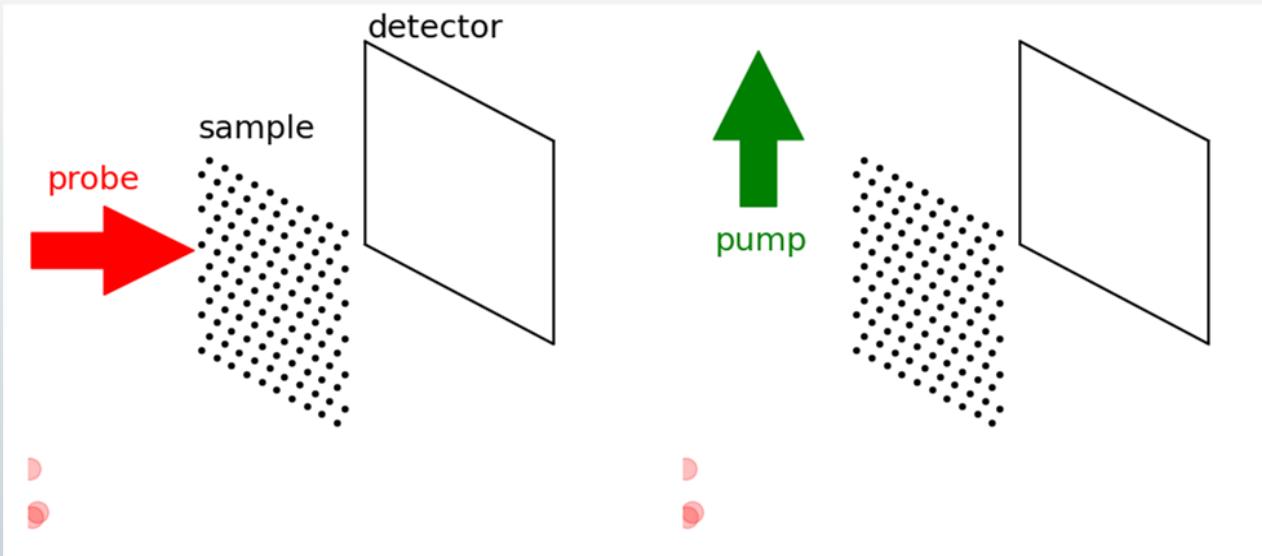
We want a **short bunch duration** → natural timescale for atomic motion is fs-ps

We want as **many electrons as possible** → large signal to noise for subtle diffraction features

High source brightness is critical for UED!

Much previous UED work

Introduction to UED: The need for *brightness*



Our device focuses on these

We want a **small transverse probe size** → some samples are hard to make with large dimensions

We want a **small transverse momentum spread** → need high coherence to see small features in k-space

We want a **short bunch duration** → natural timescale for atomic motion is fs-ps

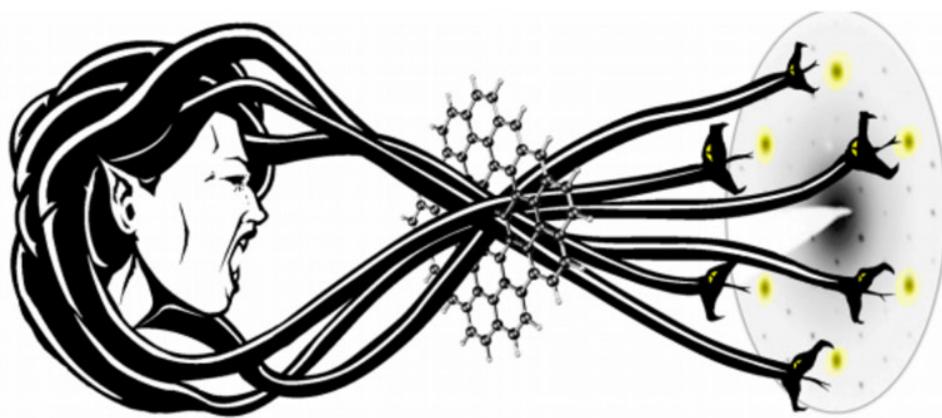
We want as **many electrons as possible** → large signal to noise for subtle diffraction features

High source brightness is critical for UED!

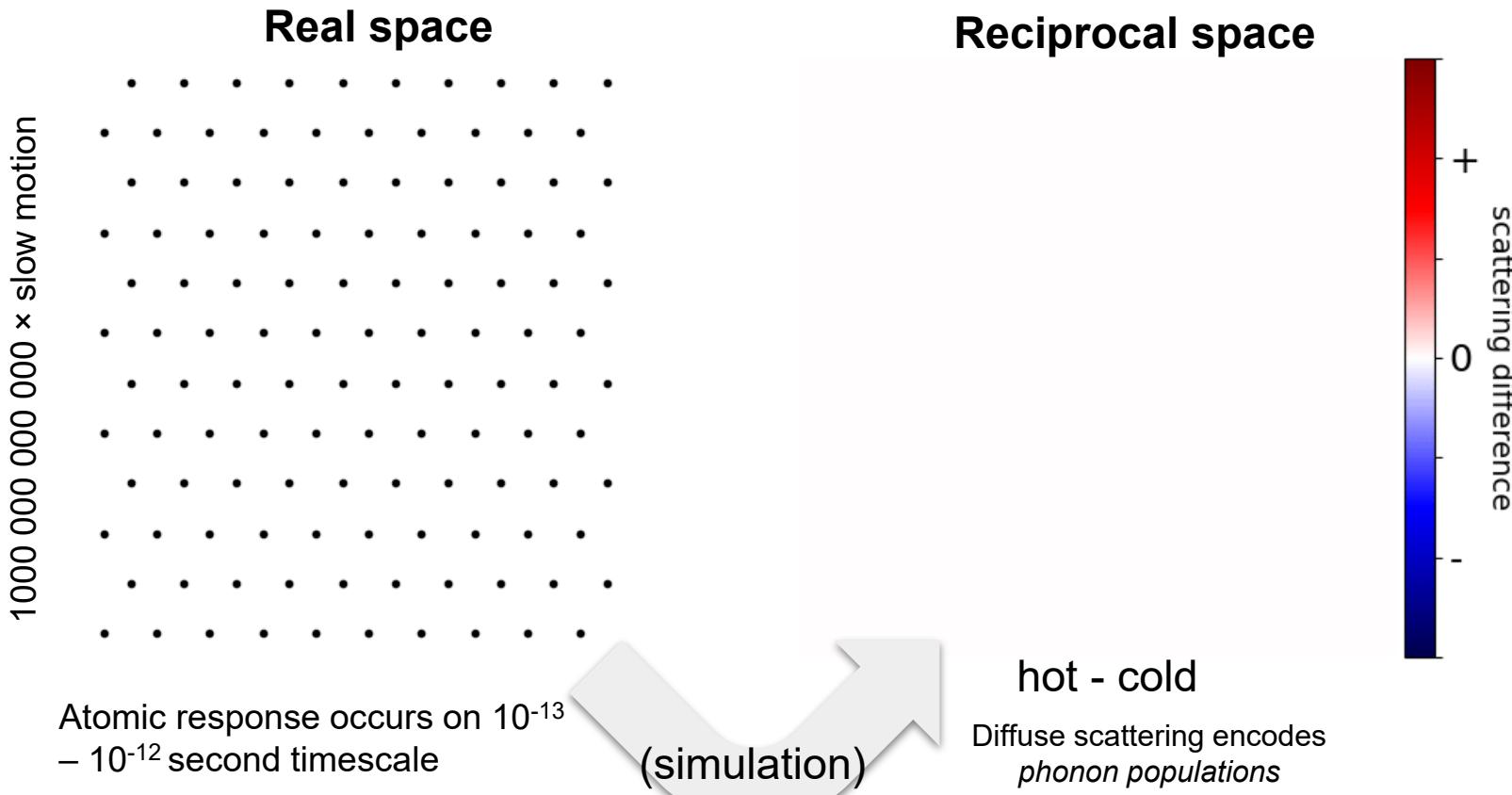
Much previous UED work

We focus on transverse probe size and coherence,
(hence the name)

MICRO
ELECTRON
DIFFRACTION FOR
ULTRAFAST
STRUCTURAL
ANALYSIS



Not just about Bragg Peaks

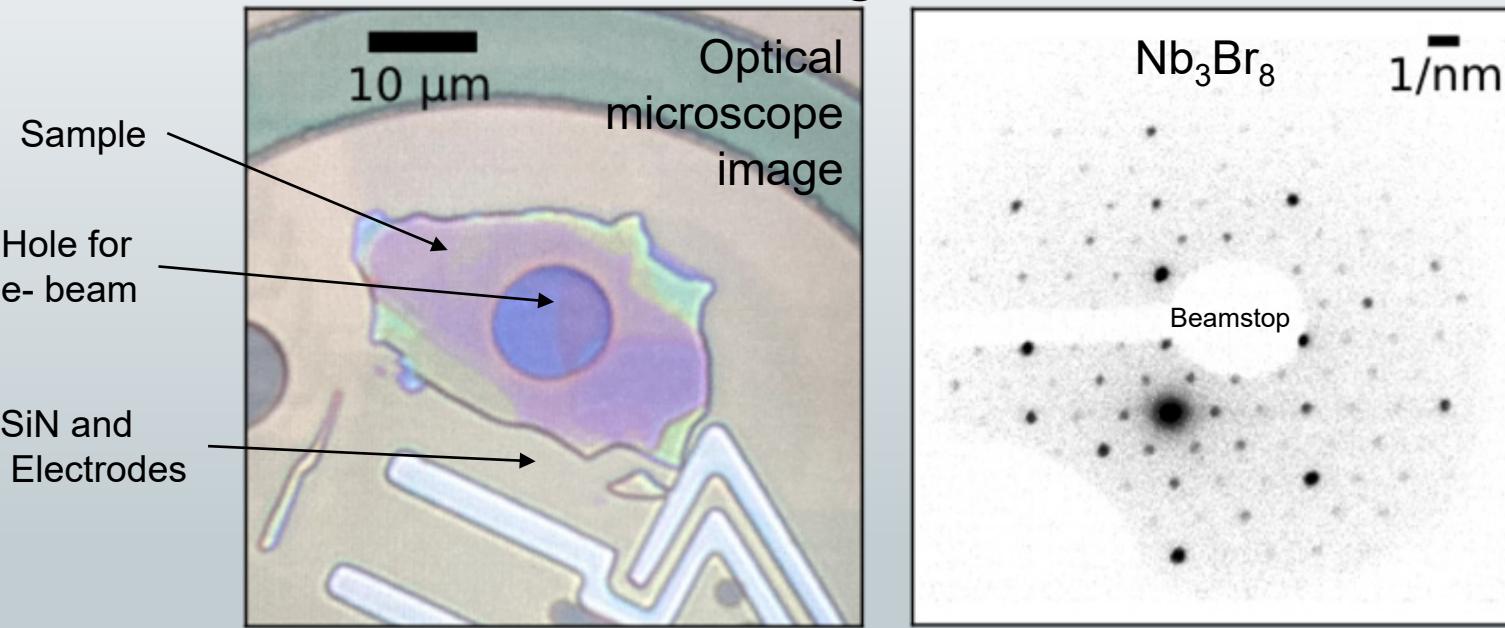


The need for small probe sizes

Preparing large films of quantum materials for UED can be challenging.

Example: Nb_3Br_8 , thin film flakes, exhibits periodic lattice distortion

UED @ MEDUSA



The rich k-space means that one cannot merely focus strongly– need small divergence too.

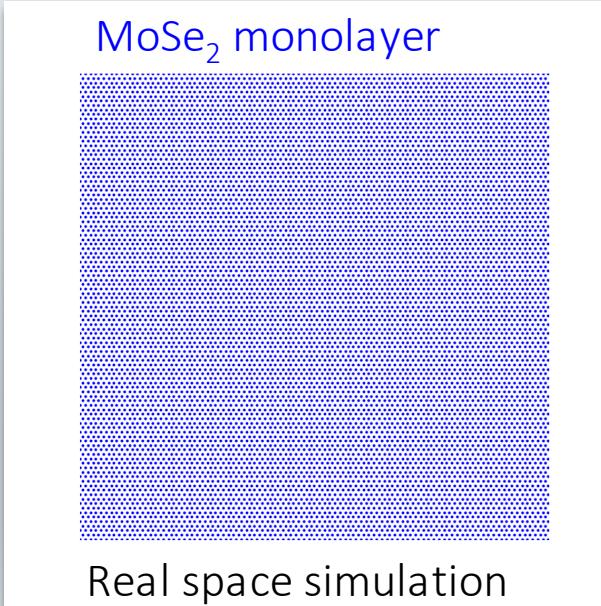
The need for reciprocal space resolution *Case Study: Moire Materials @ MEDUSA*

- Overlapping two monolayers with a small twist can yield remarkable new materials physics: moire materials.

The need for reciprocal space resolution

Case Study: Moire Materials @ MEDUSA

- Overlapping two monolayers with a small twist can yield remarkable new materials physics: moire materials.

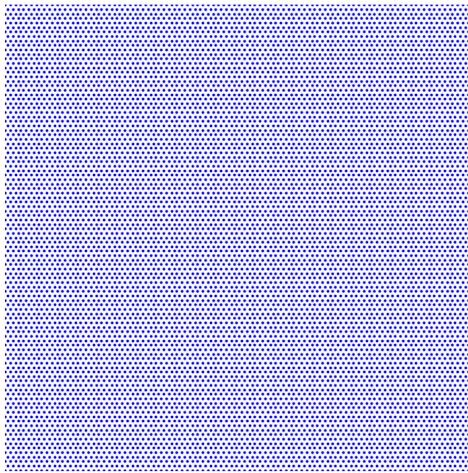


The need for reciprocal space resolution

Case Study: Moire Materials @ MEDUSA

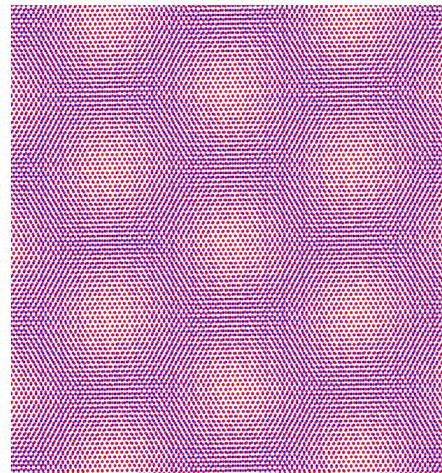
- Overlapping two monolayers with a small twist can yield remarkable new materials physics: moire materials.

MoSe₂ monolayer



Real space simulation

WSe₂ rotated 2° / MoSe₂ monolayer

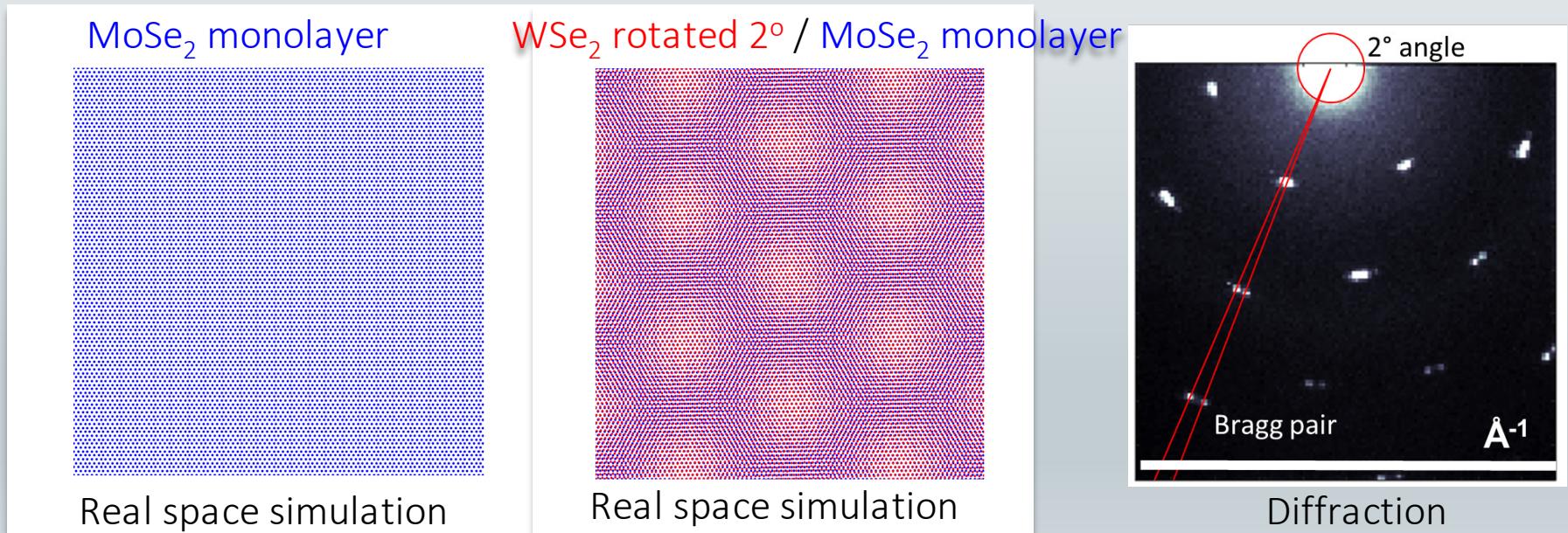


Real space simulation

The need for reciprocal space resolution

Case Study: Moire Materials @ MEDUSA

- Overlapping two monolayers with a small twist can yield remarkable new materials physics: moire materials.

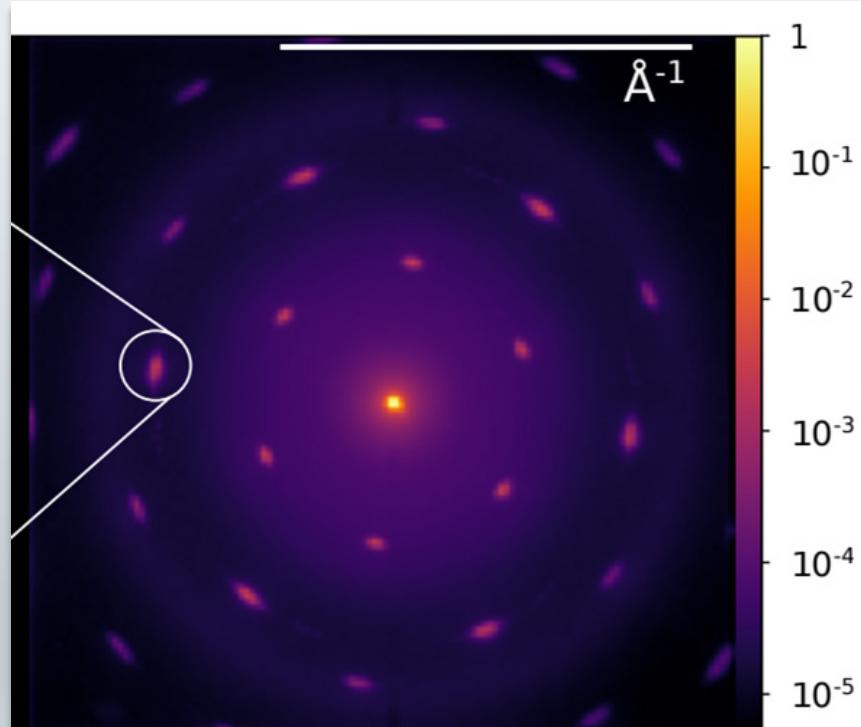


Samples prepared @ Stanford by Fang Liu and Helen Zeng

The need for reciprocal space resolution

Case Study: Moire Materials @ MEDUSA

Interesting physics at multiple scales:



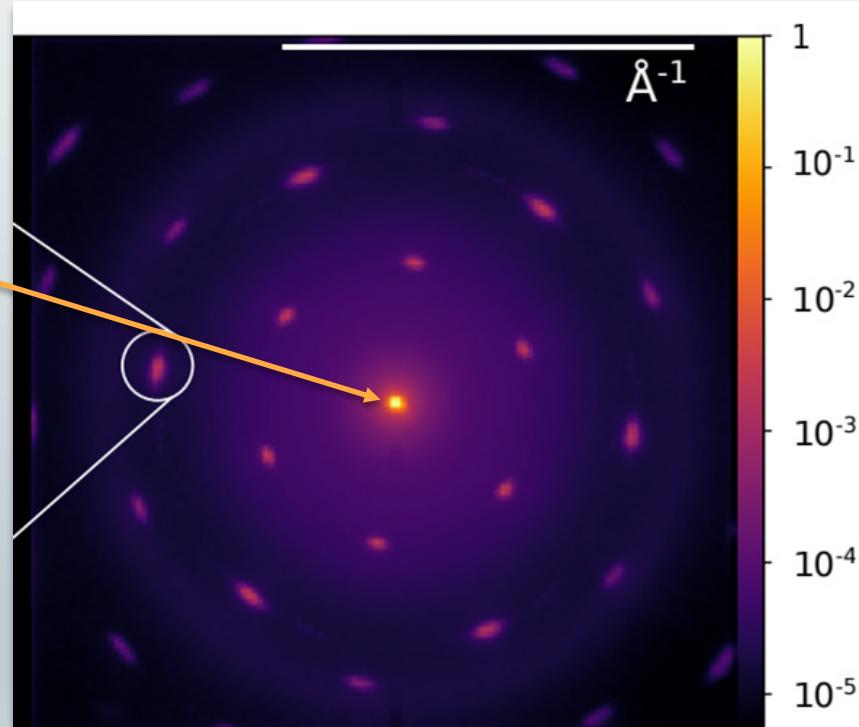
WSe₂ rotated 2° / MoSe₂ monolayer

The need for reciprocal space resolution

Case Study: Moire Materials @ MEDUSA

Interesting physics at multiple scales:

Undiffracted beam



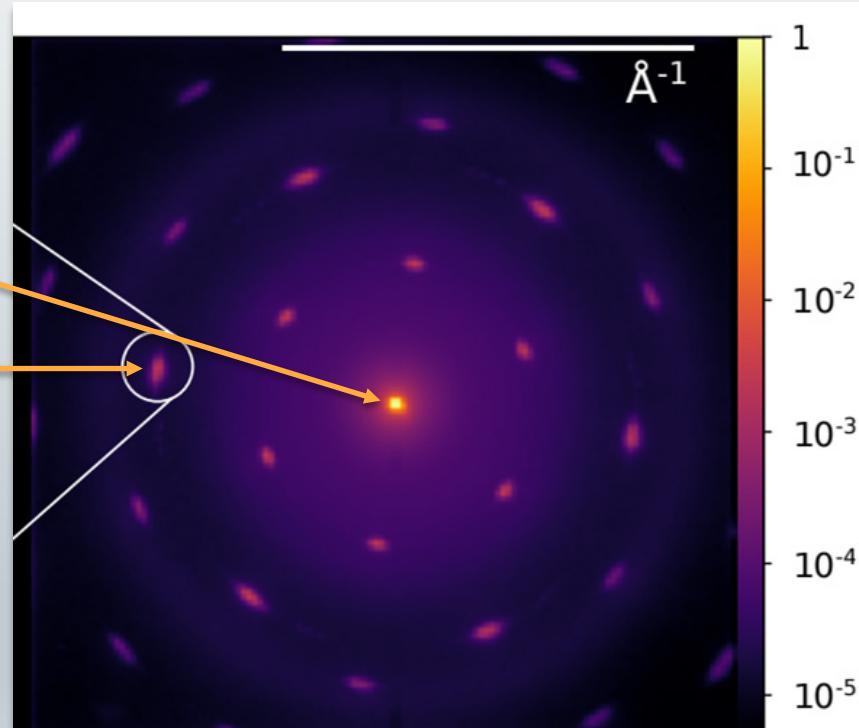
WSe₂ rotated 2° / MoSe₂ monolayer

The need for reciprocal space resolution

Case Study: Moire Materials @ MEDUSA

Interesting physics at multiple scales:

Undiffracted beam
Bragg Peaks: note “smearing” due
to moire twist)



WSe₂ rotated 2° / MoSe₂ monolayer

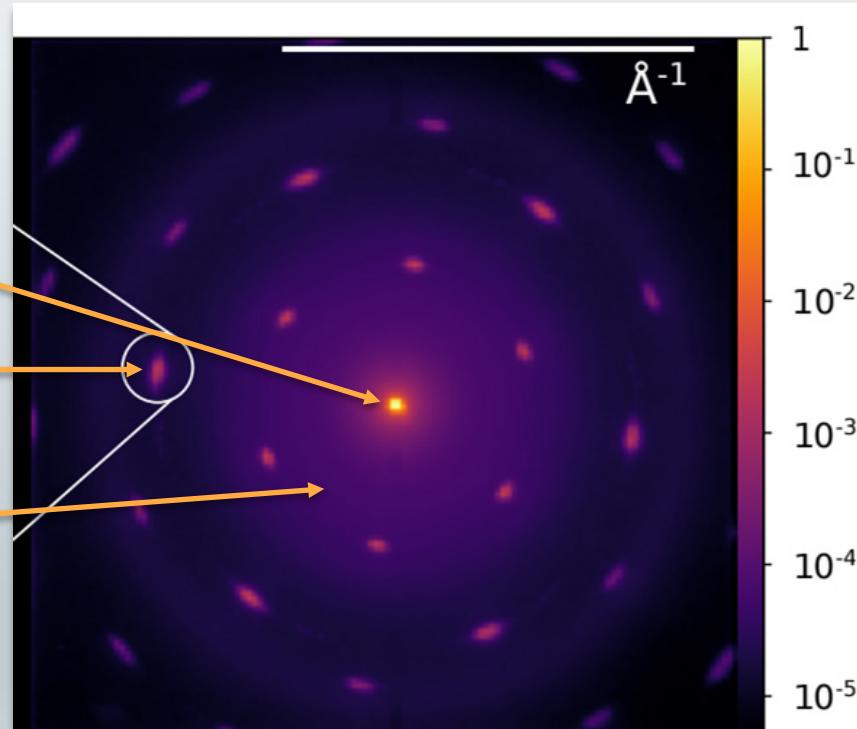
The need for reciprocal space resolution

Case Study: Moire Materials @ MEDUSA

Interesting physics at multiple scales:

Undiffracted beam
Bragg Peaks: note “smearing” due
to moire twist)

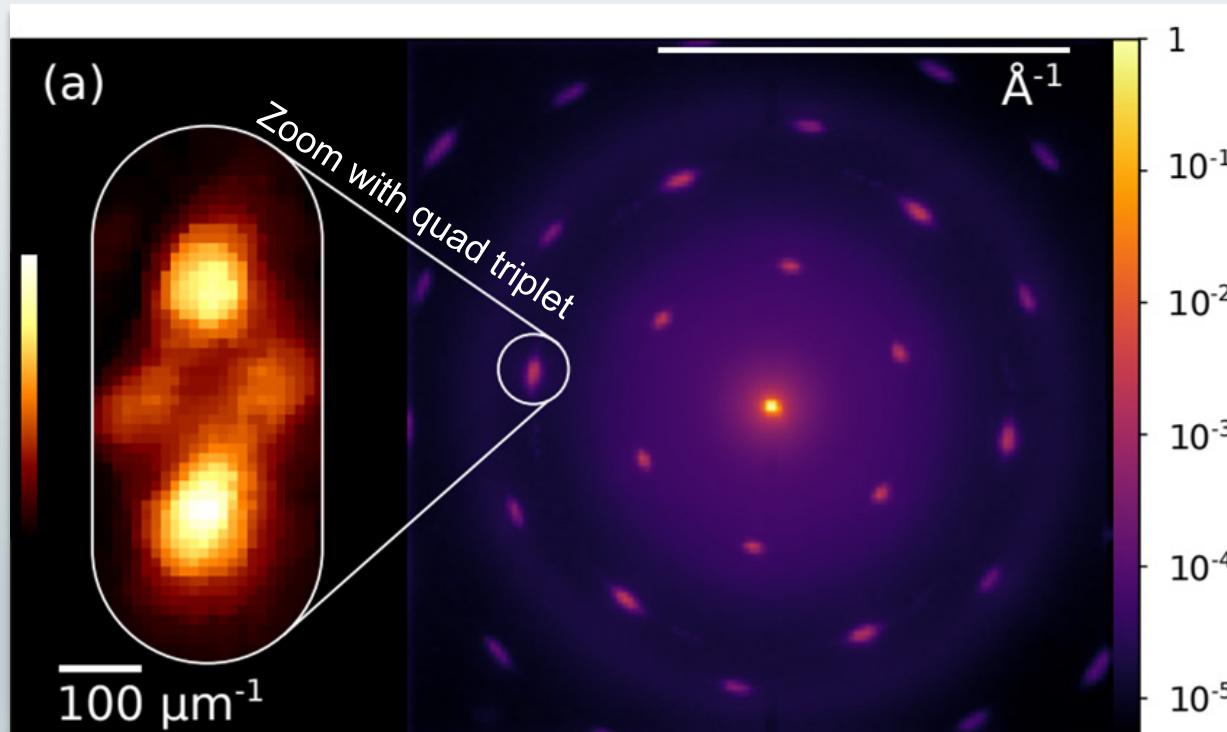
Diffuse Scattering
Only two monolayers thick!



WSe₂ rotated 2° / MoSe₂ monolayer

The need for reciprocal space resolution

Case Study: Moire Materials @ MEDUSA



First observation of moire atomic reconstruction with any ultrafast probe!
Requires very high coherence (≥ 10 nm coherence length)

MEDUSA Strategy for source brightness: Seek Low MTE

- MTE is rms transverse photoelectron momentum spread expressed in energy units:

$$\epsilon_{n,cath} = \sigma_{laser} \sqrt{\frac{\text{MTE}}{mc^2}}$$

$$B_{n,max} \propto \frac{E_{acc}^n}{\text{MTE}}$$

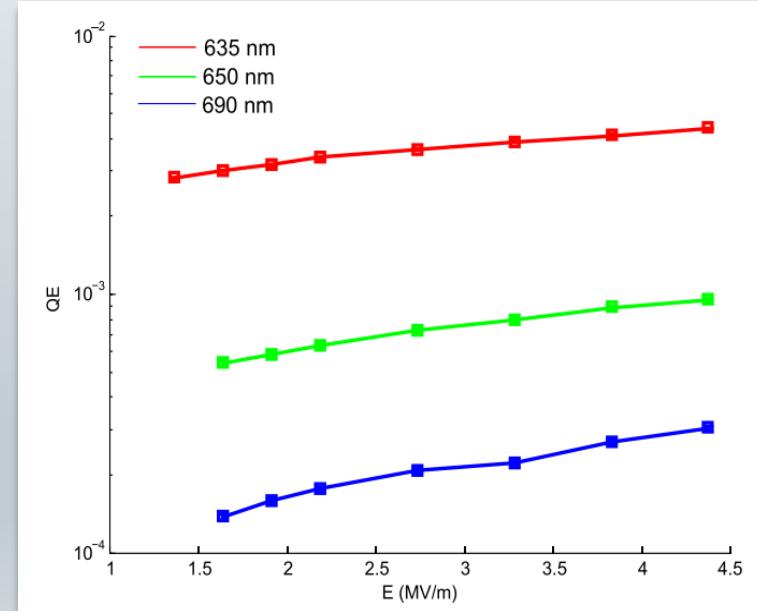
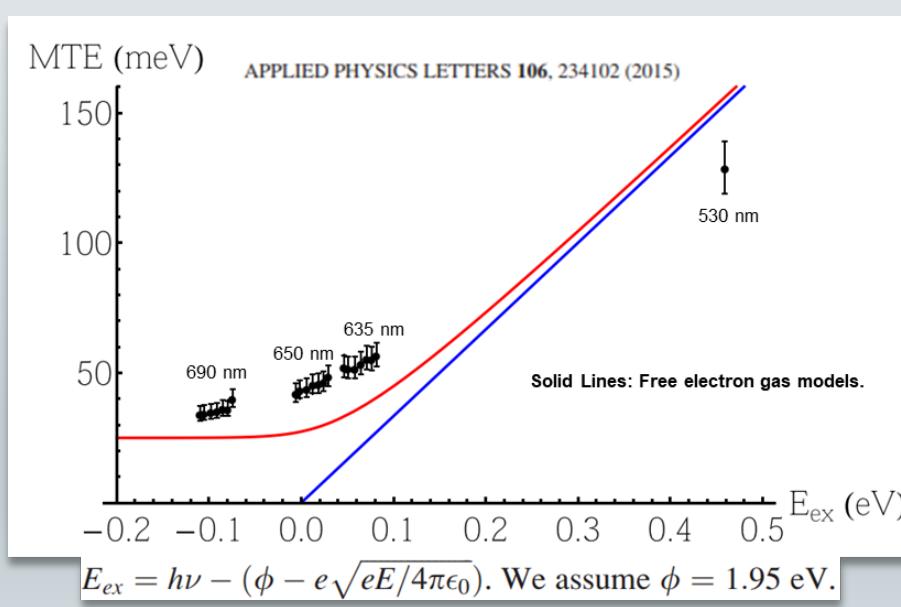
where $1 < n < 2$ depends on bunch aspect ratio.

MTE's of several hundred meV are common (Cu, Cs-Te).

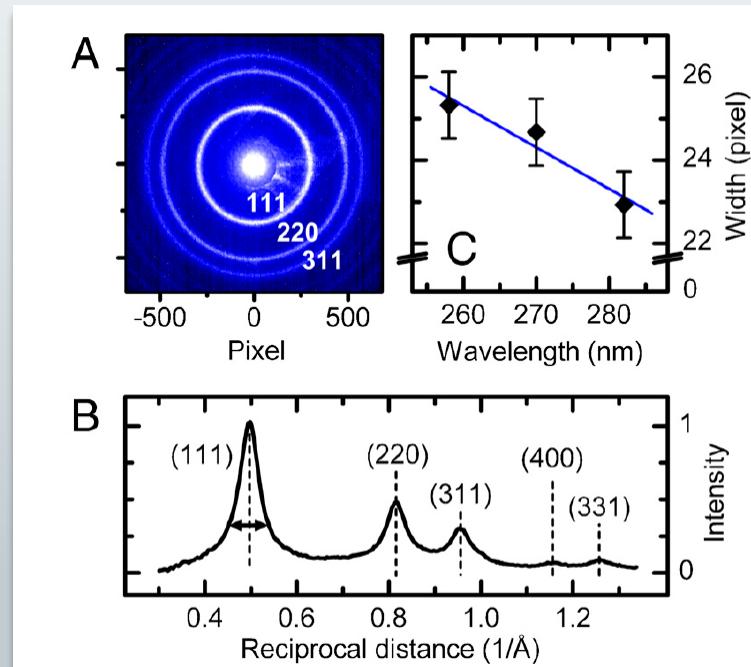
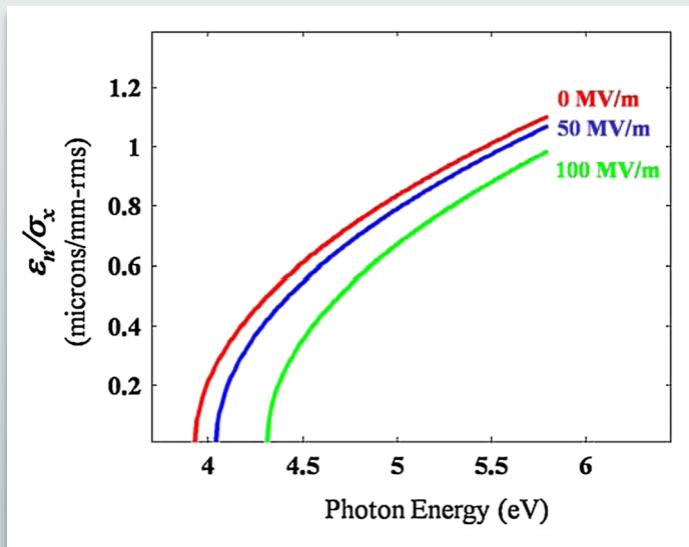
- But does low MTE actually matter when space charge is present? We will get back to this.

Excellent Low MTE Candidate Materials: Alkali Antimonides

- By reducing the excess energy of photoemission, one can *trade* quantum efficiency for lower MTE.
- Alkali antimonides achieve as low as ~30 meV (shown below: Na-K-Sb, min MTE of 35 meV) with photon energy tuning.
- >10x max. brightness as compared to Cs-Te or polycrystalline Cu in traditional operation.
- Lower QE must be balanced against increased laser energy—ultimate limit is multiphoton photoemission, which spoils low MTE.



Threshold photoemission for low MTE is not a new idea

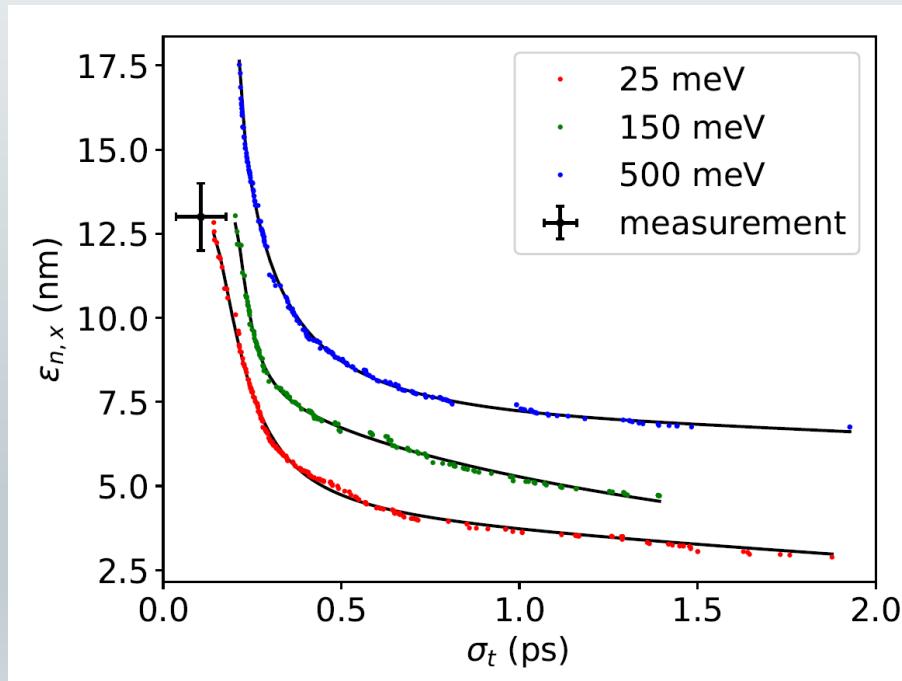


Dowell and Schmerge, PRSTAB 12, 074201 (2009)
See also: J. Feng, APL 107, 134101 (2015)

M. Aidelsburger, PNAS 107 46 (2010).

But even *very low MTE* values remain useful

MOGA optimization: MEDUSA Beamline. DC Gun (140 kV) + RF buncher, 100,000 e/bunch
(space charge very much not negligible)



Lower MTE provides access to:

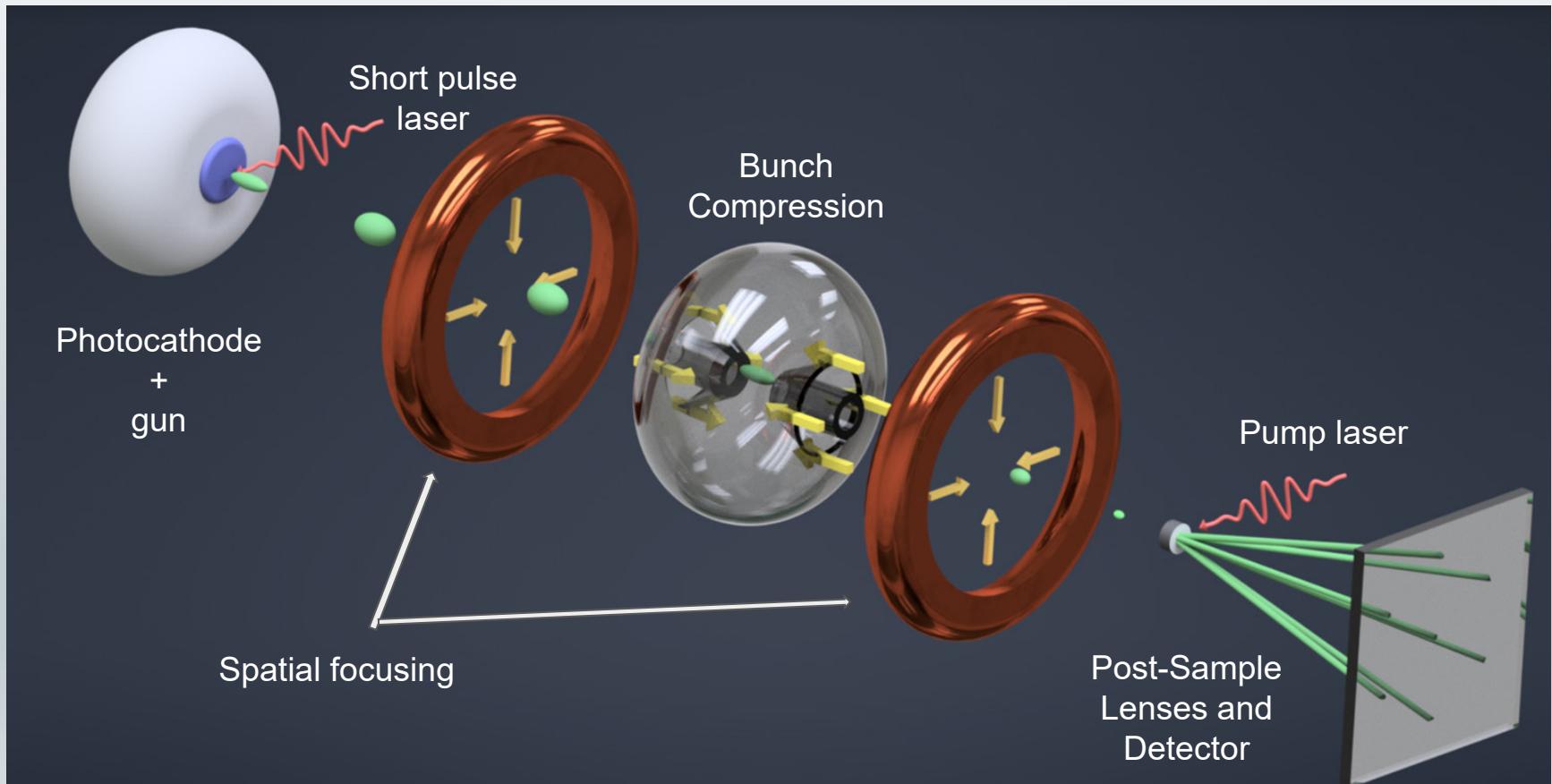
Smaller bunch length for a given emittance, or vice versa

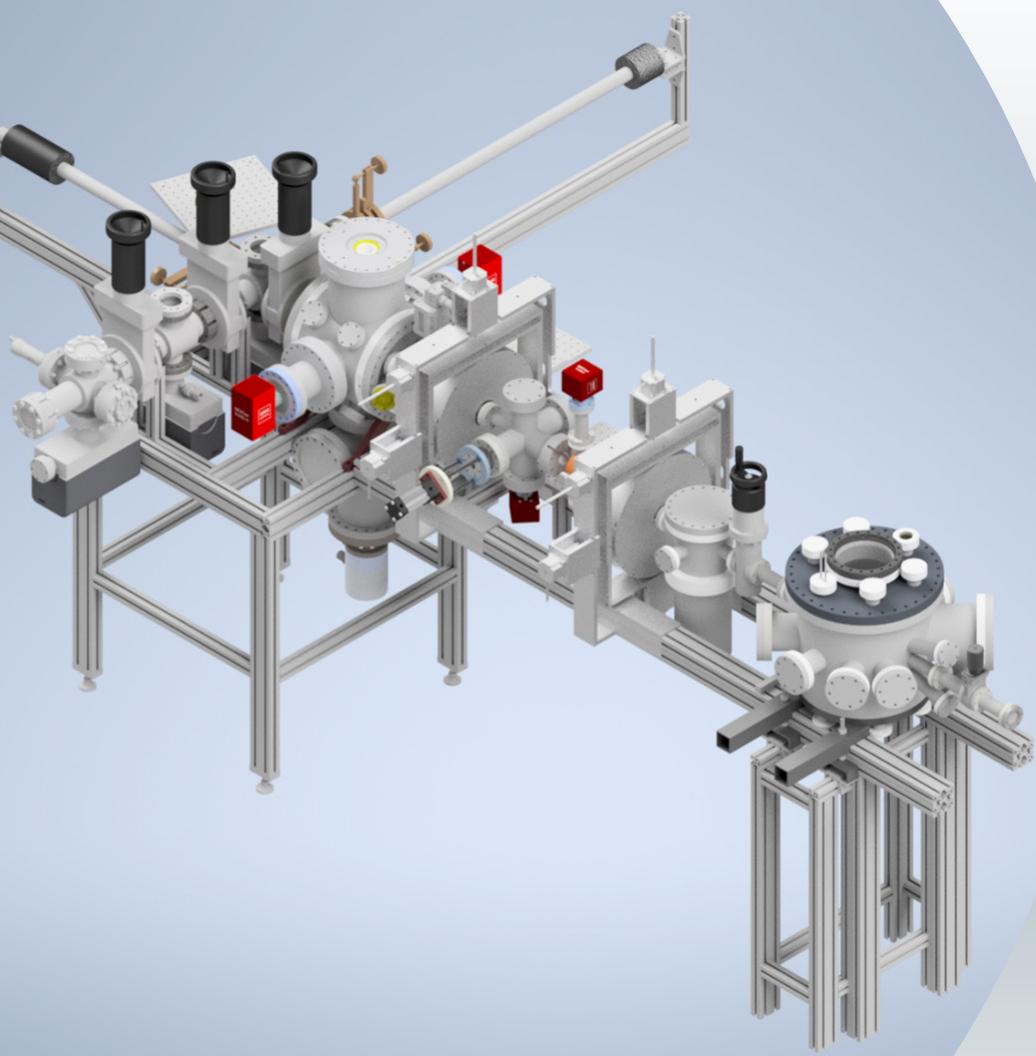
An important effect even below 100 meV MTEs

We regularly measure emittances of 12-14 nm depending on transverse optics, and bunch length between 100-200 fs rms.

Infer MTE significantly less <100 meV

Cartoon of MEDUSA: The critical components





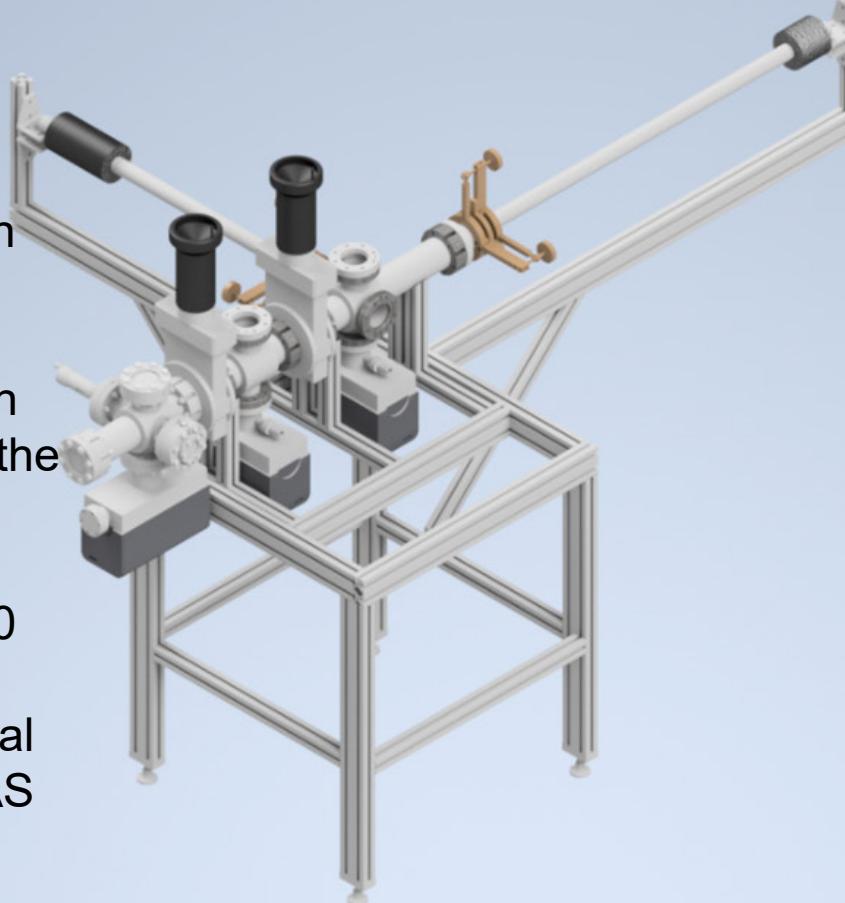
Our device: piece by piece.

Beamline

High brightness,
semiconductor electron
source (Na-K-Sb)

650 nm photo-emission
wavelengths matches the
cathode bandgap

50 W @ 250 kHz, 1030
nm Yb fiber laser (AS
Tangerine) feeds optical
parametric amplifier (AS
Mango)

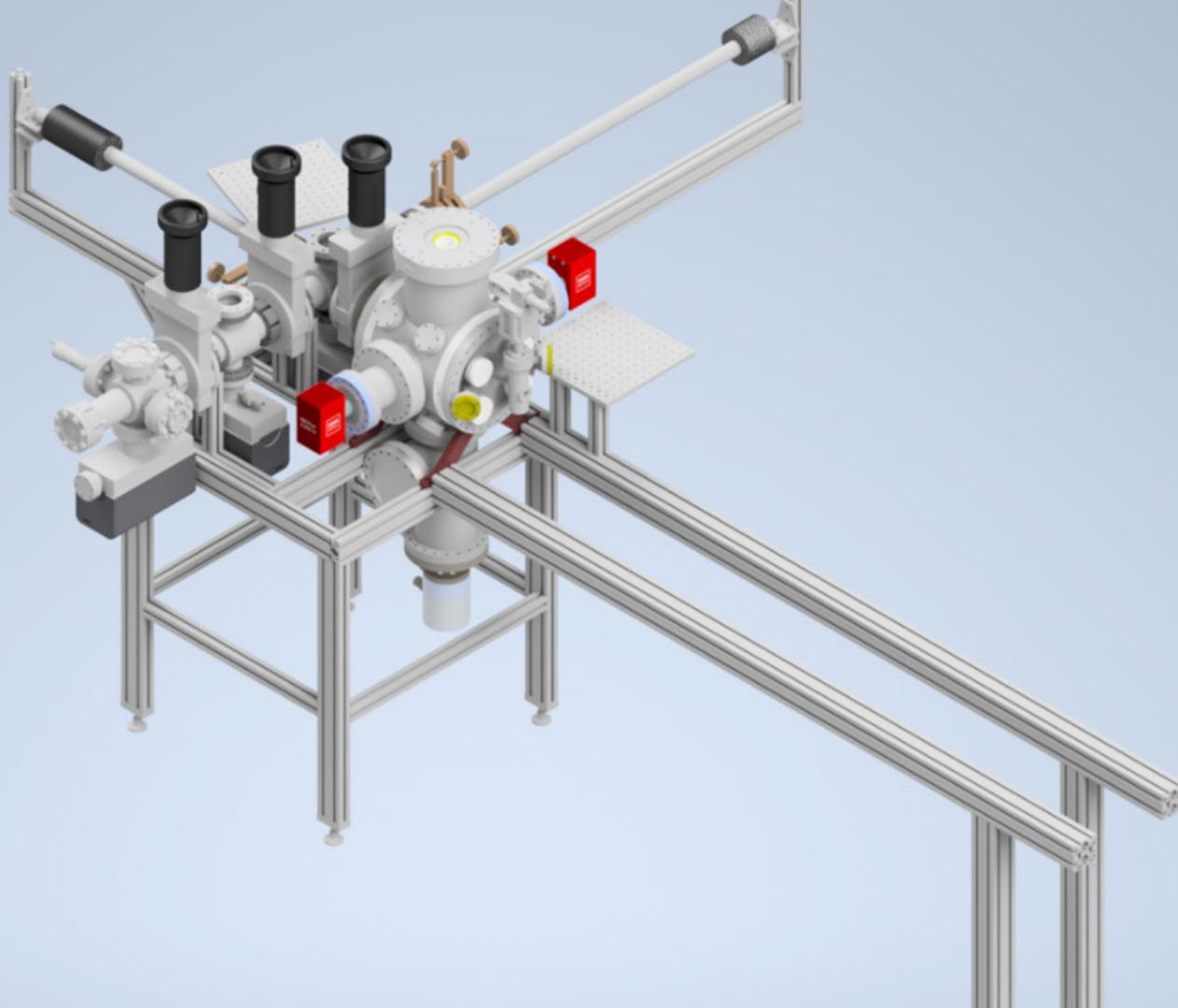


Source is extremely
vacuum sensitive (XHV)

Beamline

150 kV DC gun

Base pressure:
 8×10^{-12} Torr



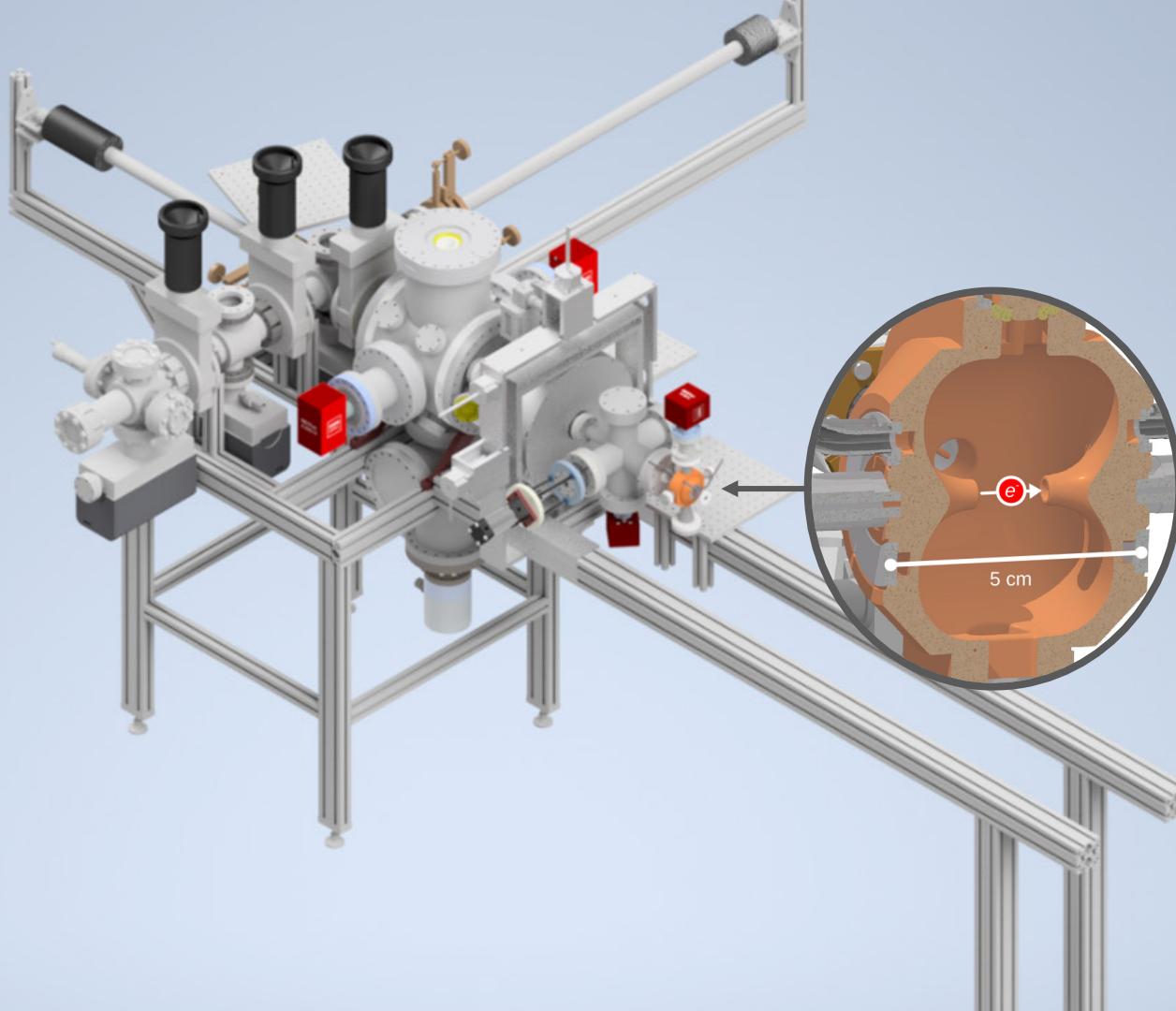
Beamline

3 GHz bunching cavity

Long pulse length at cathode: ~10 ps

Bunching after acceleration mitigates space-charge

~100 fs rms bunch-length at sample



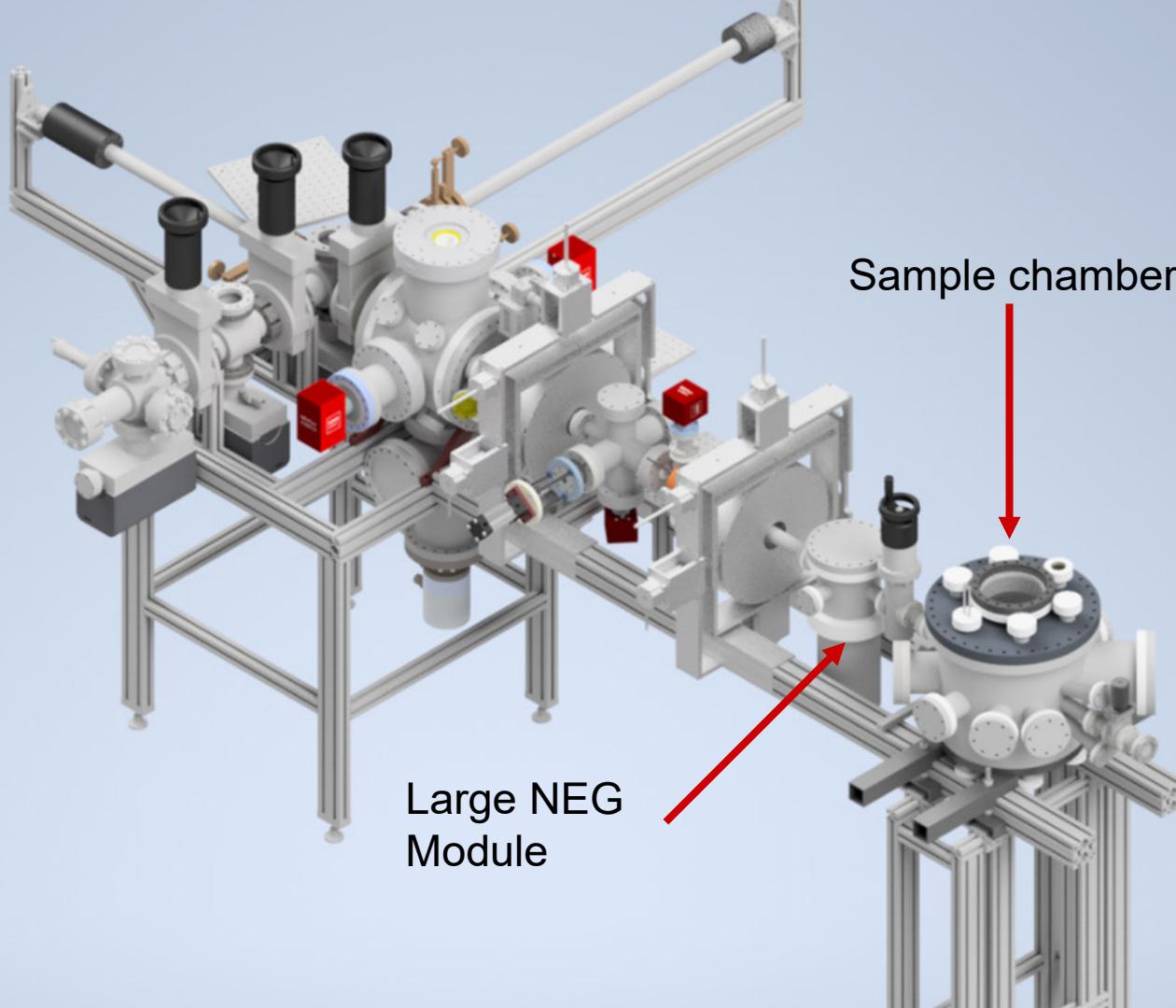
Beamline

Sample chamber

Pressure $10^4 \times$ higher
than gun

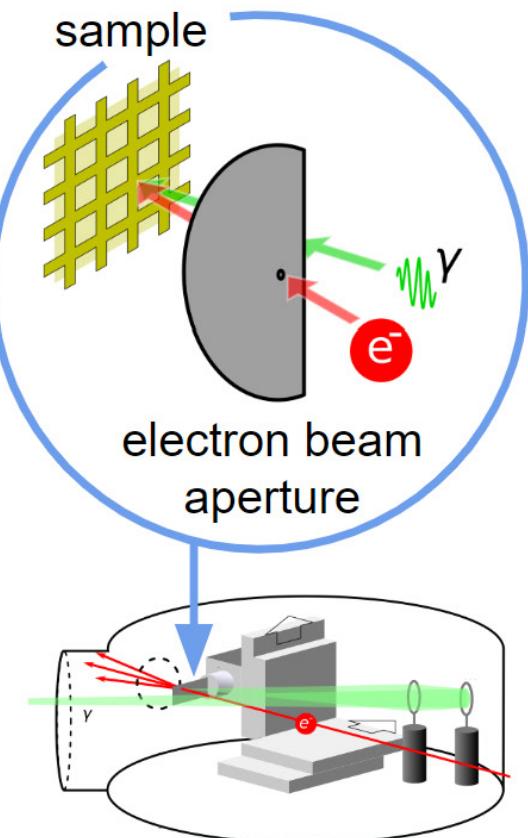
515 nm pump pulses,
second-harmonic of
Tangerine

(future upgrade:
Visible and NIR OPA)



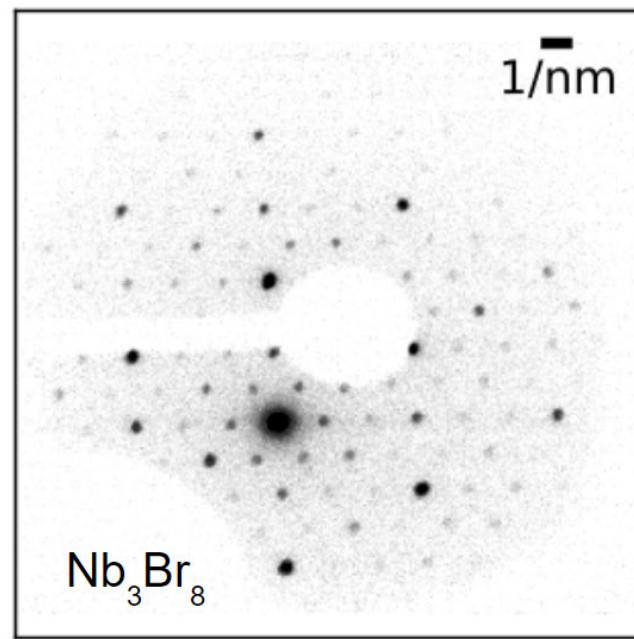
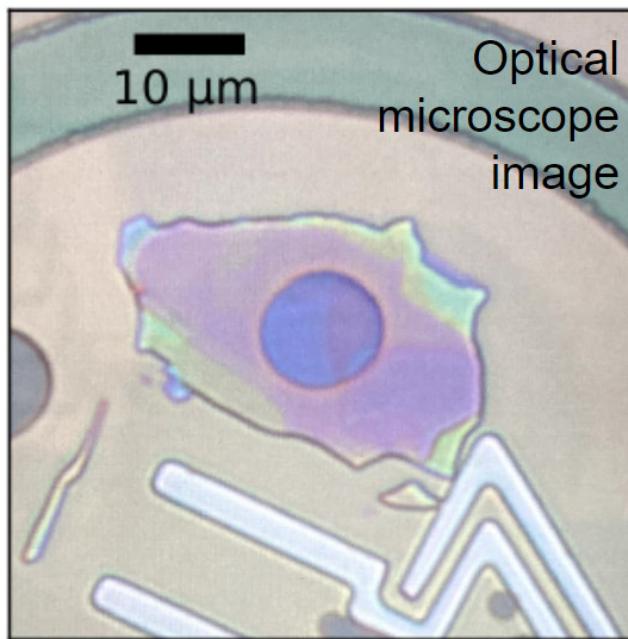
Spatial Resolution

Microdiffraction probe



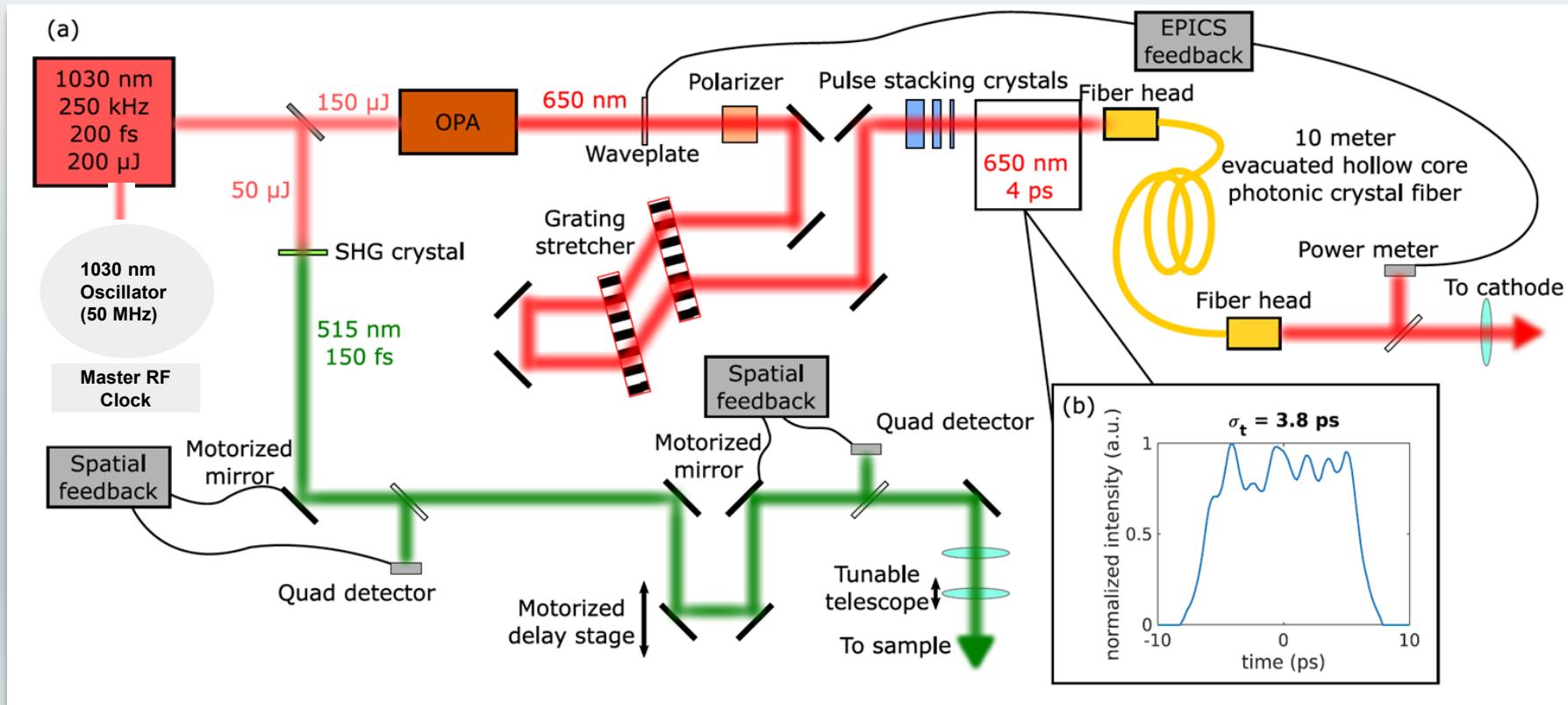
Spot size on sample (r.m.s.):
Electrons on sample per pulse:

With aperture	No aperture
3 μm	40 μm
500	10^5



Thanks to Lena Kourkoutis and Elisabeth Bianco for providing the Nb₃Br₈ sample

The UED Laser system: “Much ado about the photocathode drive laser”



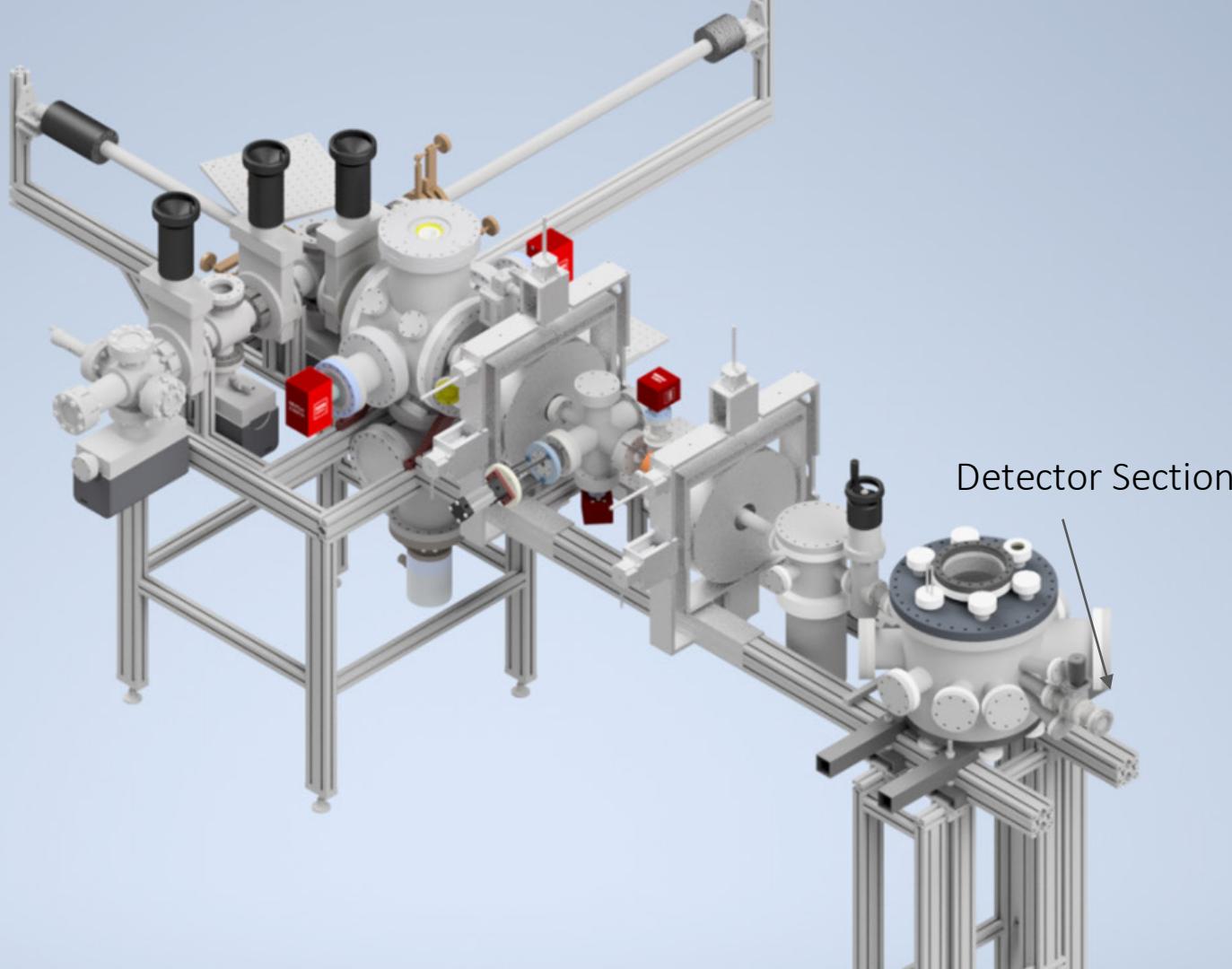
Beamline

Sample chamber

Pressure $10^4 \times$ higher
than gun (UHV)

Direct Electron
Detector
(EMPAD) deployed in
collab. with
Gruner/Thom/Muller

Quad triplet post-
sample not shown

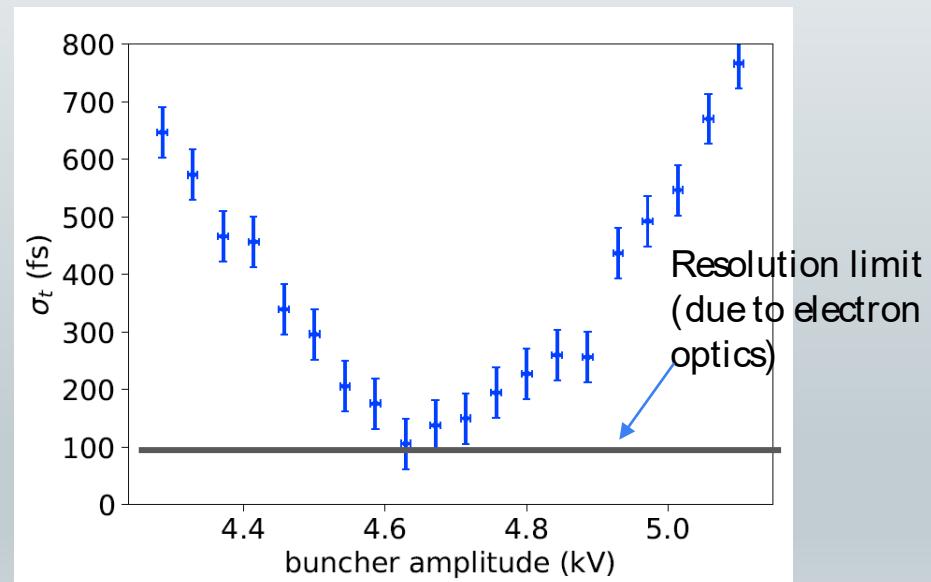


Bunch length measurements: compact RF deflector cavity

3 GHz *insertable deflection cavity manufactured by Dr. X. Works, Eindhoven.*

Drops in just upstream of sample location

Looking through laser entrance window
Note 1" light optics for scale

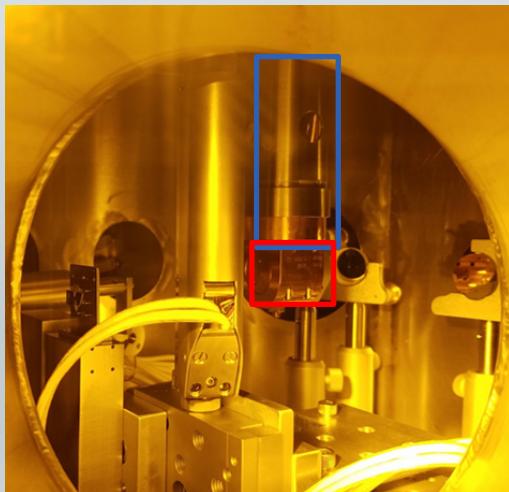


Bunch length measurements: compact RF deflector cavity

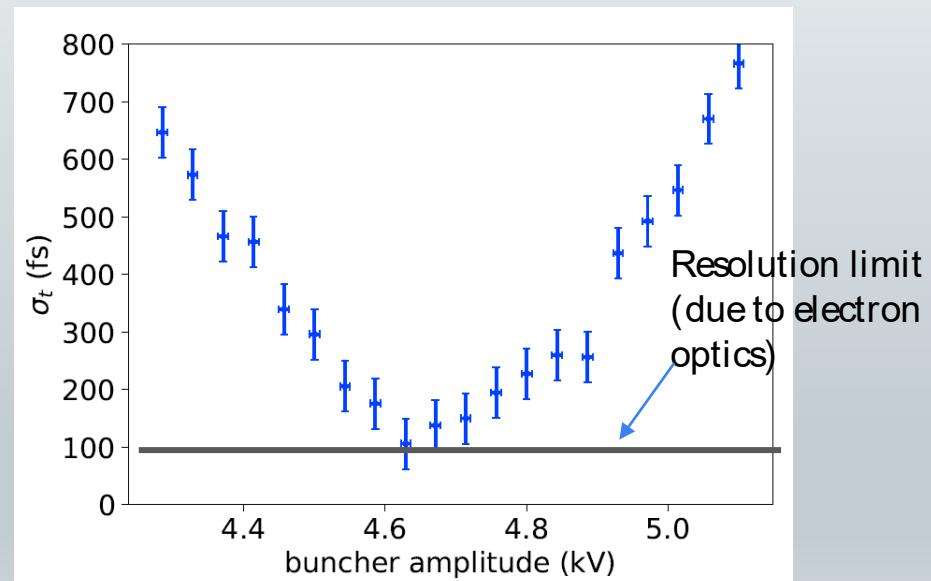
3 GHz *insertable deflection cavity manufactured by Dr. X. Works, Eindhoven.*

Drops in just upstream of sample location

Looking through laser entrance window
Note 1" light optics for scale

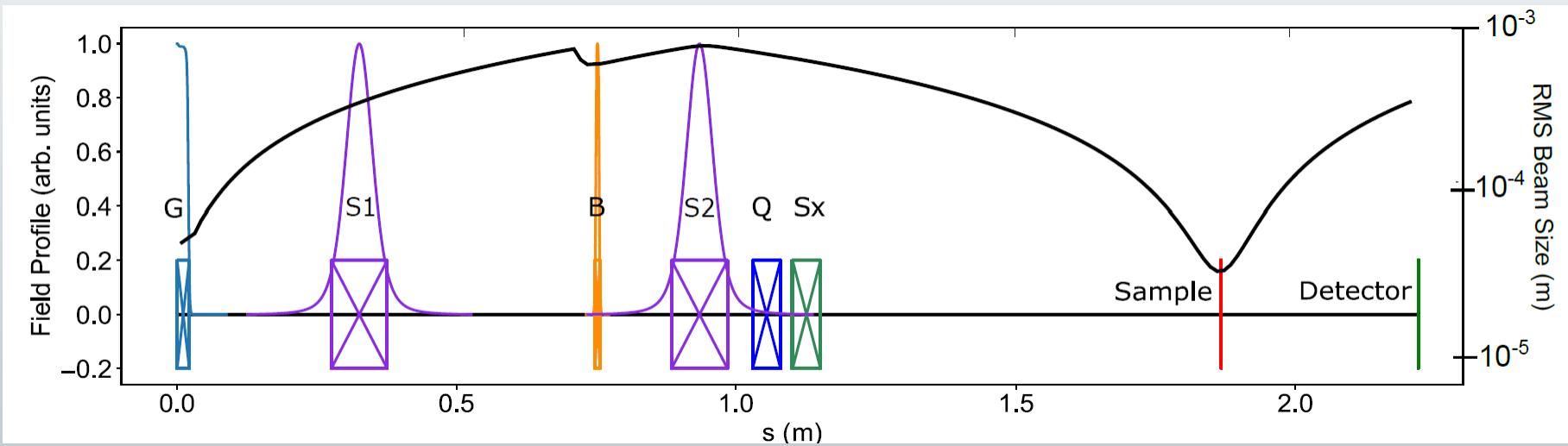


Mounting Pole
Cavity



Aberrations: an important obstacle

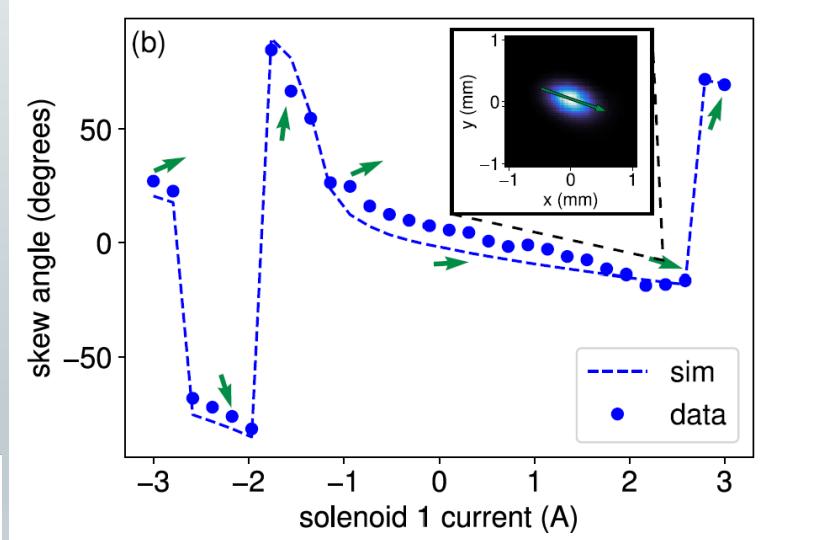
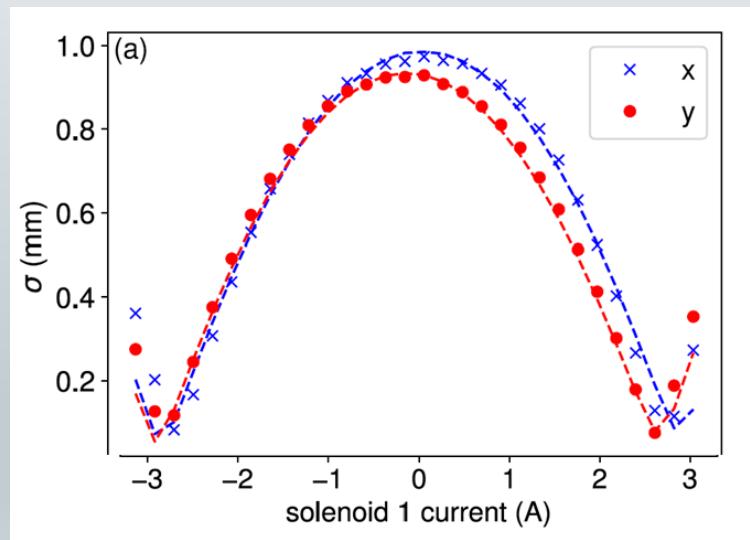
- Our microdiffraction optics rely on large changes in beam size:



- This leaves us *very vulnerable to field aberrations*.
- We have dedicated correction magnets for quadrupole, skew quadrupole, and sextupole moments.

Aberrations: Normal and Skew Quadrupole

- Erroneous quads are found in our solenoids *and* due to the coupler kick of the bunching cavity where beam size is large.



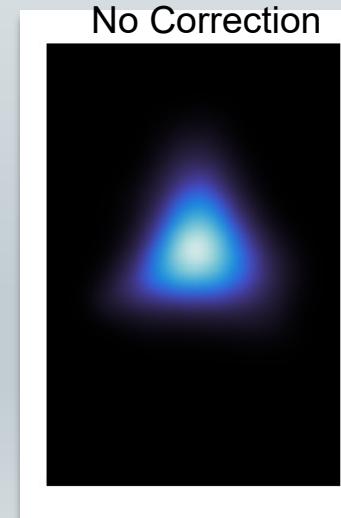
Lines: simulation, dots: measurement

Aberrations: Sextupole

- We use quadrupole correctors just downstream of our second solenoid → as previously demonstrated [L. Zheng, PRAB **22**, 072805 (2019)] very effective in removing transverse coupling.
- Sextupole moment primarily arises from buncher coupler
- Once we do that, our beam looked like this on our diffraction detector:

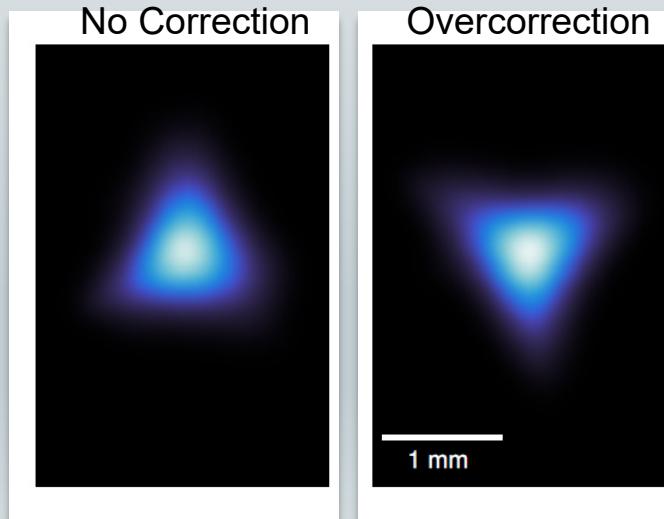
Aberrations: Sextupole

- We use quadrupole correctors just downstream of our second solenoid → as previously demonstrated [L. Zheng, PRAB **22**, 072805 (2019)] very effective in removing transverse coupling.
- Sextupole moment primarily arises from buncher coupler
- Once we do that, our beam looked like this on our diffraction detector:



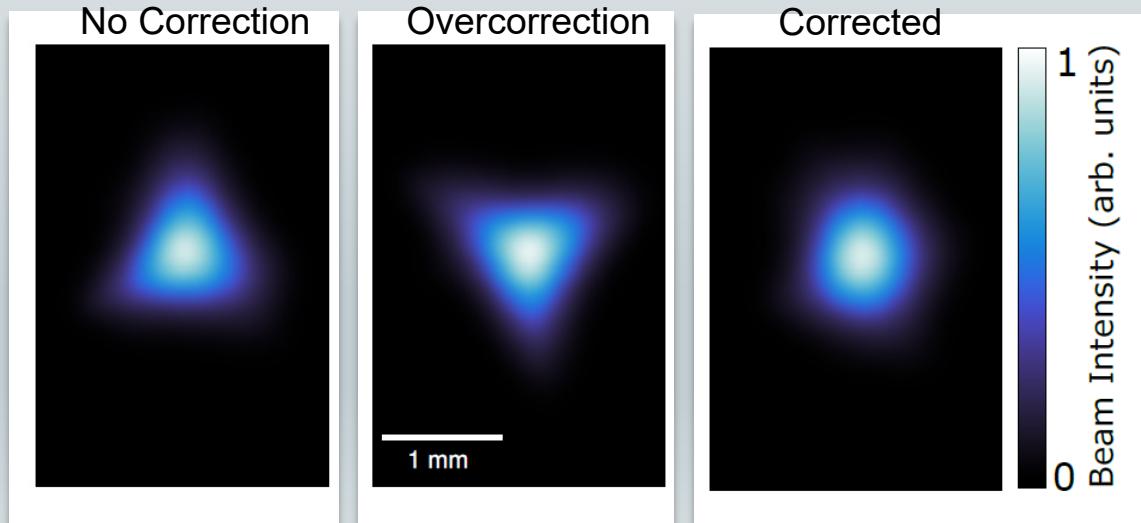
Aberrations: Sextupole

- We use quadrupole correctors just downstream of our second solenoid → as previously demonstrated [L. Zheng, PRAB **22**, 072805 (2019)] very effective in removing transverse coupling.
- Sextupole moment primarily arises from buncher coupler
- Once we do that, our beam looked like this on our diffraction detector:



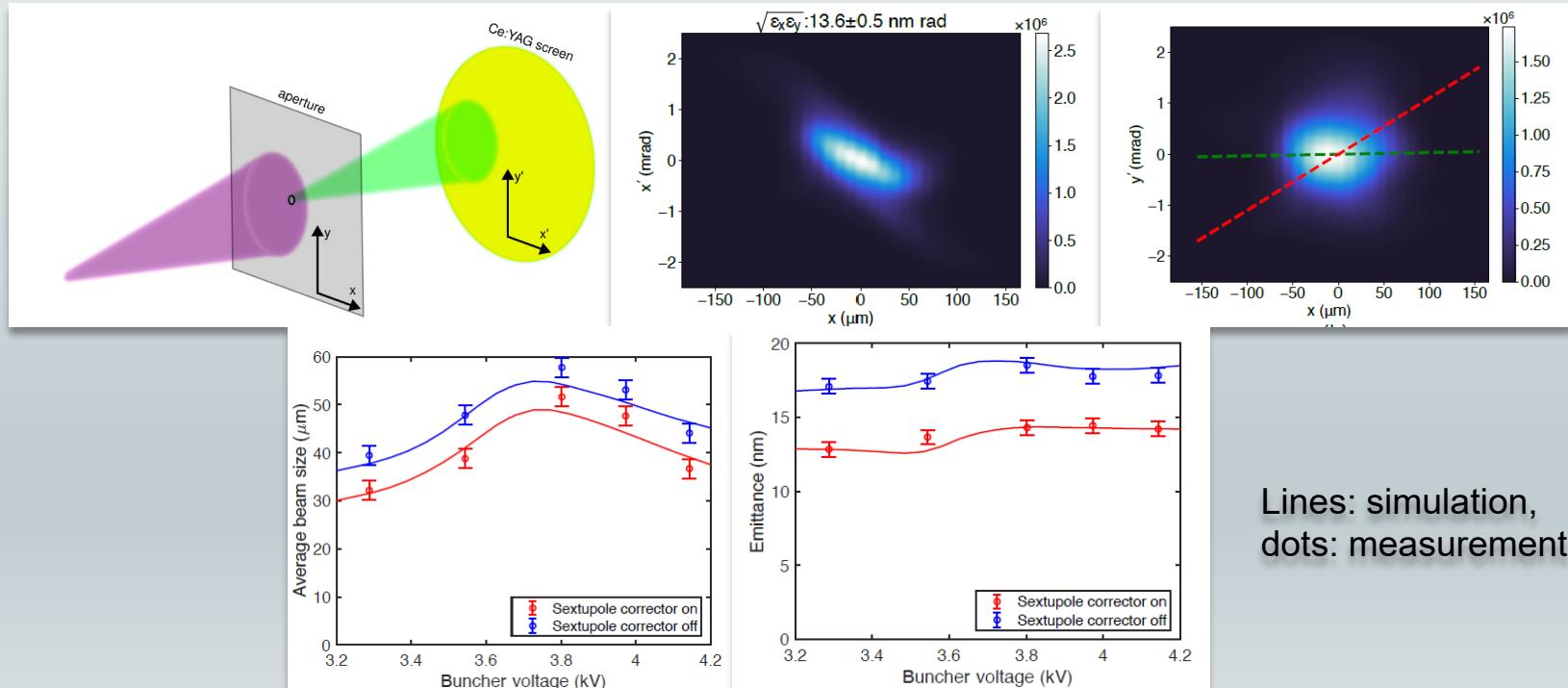
Aberrations: Sextupole

- We use quadrupole correctors just downstream of our second solenoid → as previously demonstrated [L. Zheng, PRAB **22**, 072805 (2019)] very effective in removing transverse coupling.
- Sextupole moment primarily arises from buncher coupler
- Once we do that, our beam looked like this on our diffraction detector:



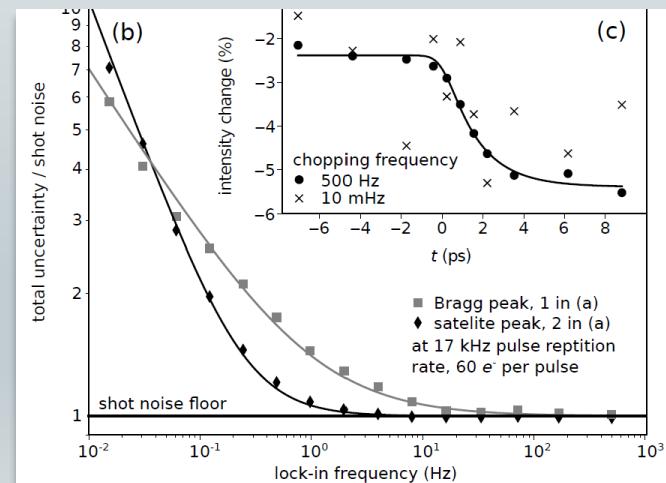
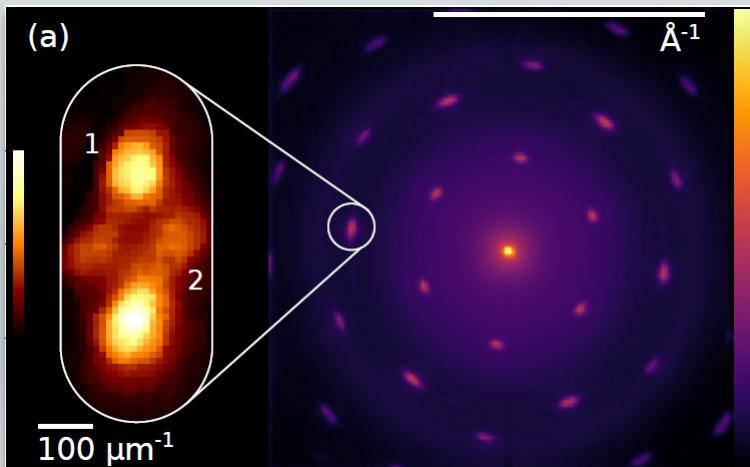
4D Phase Space Measurements

- We use 4D transverse phase space mapping to ensure cancellation of skew quadrupoles and sextupole moment



Critical Step Forward: Direct Electron Detection

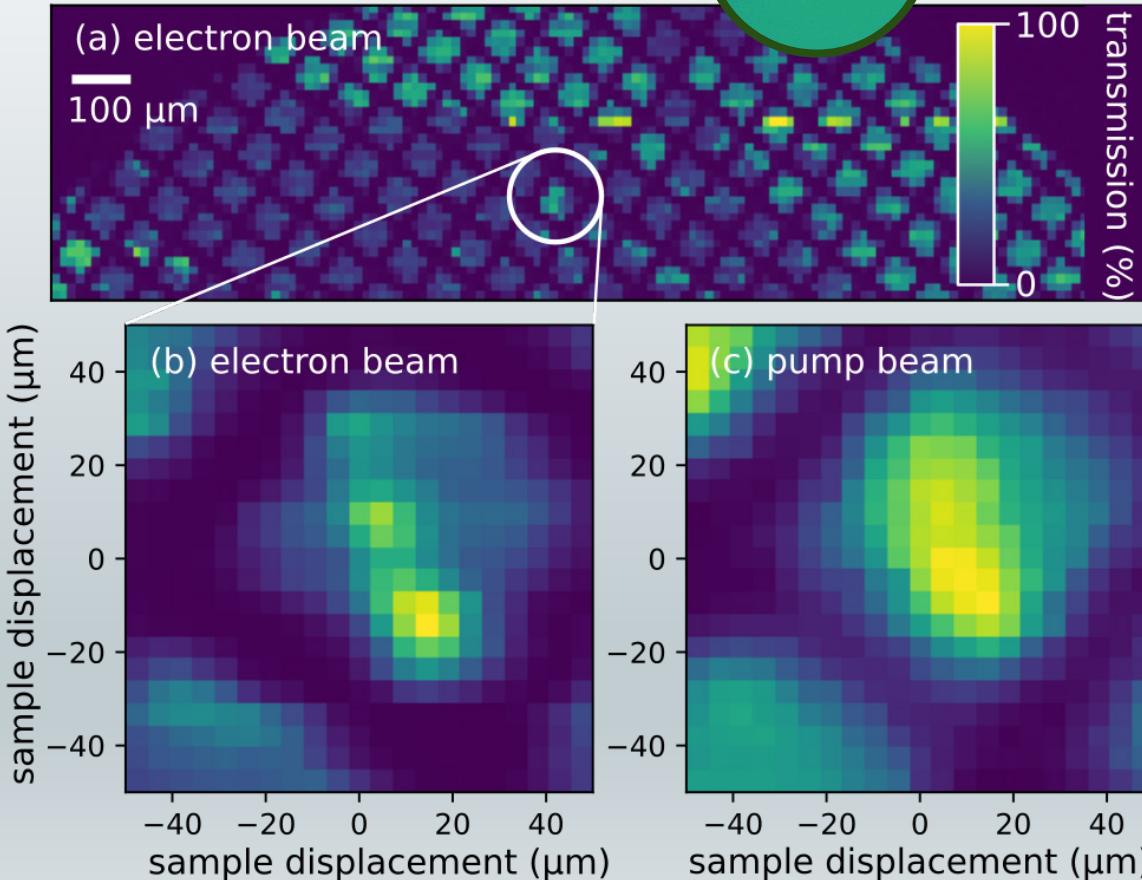
- Brightness in UED is only as good as your detector.
- A direct electron detector called the EMPAD, has been a huge step forward for MEDUSA. Collab with Gruner, Thom-Levy, and Muller at Cornell.
- Single Particle sensitivity (SNR ~ 100 per electron), and very high dynamic range (10^6)
- Images up to **1000 frames/second** → outrun noise!



To Conclude: Some examples of what you can do with a
UED microprobe

Commissioning UED experiment

Mosaic gold film, $\sim 20\text{ nm}$ thin



Why gold?

Responds strongly to temperature changes that are small compared to the melting point, 100 K vs 1300 K

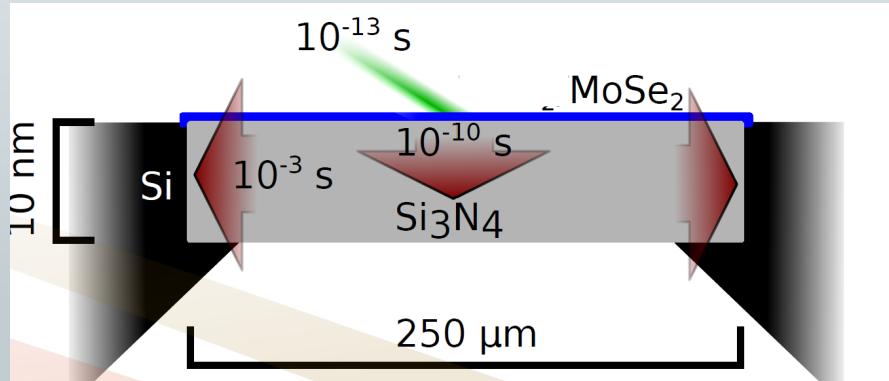
We can pick out individual grains of a mosaic material

We can make our pump beam very small! (10 micron rms)

→ Reduces average power needed for pumping!

Mapping the full life cycle of optical excitation

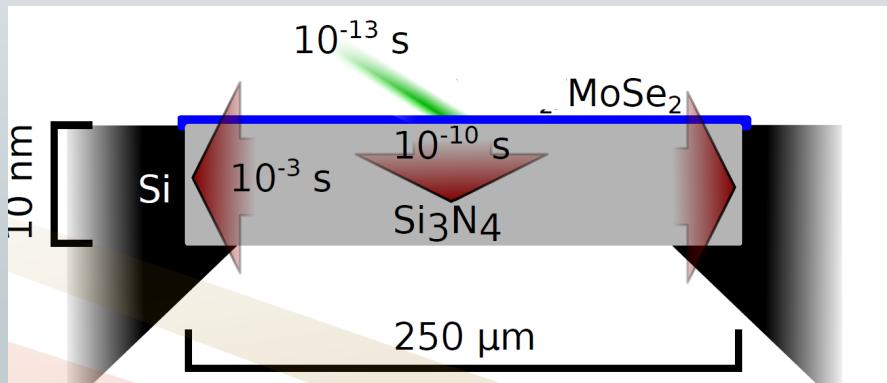
- With a high rep-rate laser *and detector* need both, you can use pulse-picking to extend UED delays out to microsecond-millisecond-second timescales.
- This allowed us to watch the full “life cycle” of optical excitation in thin films.
- Example 1: Monolayer MoSe₂ atop SiN



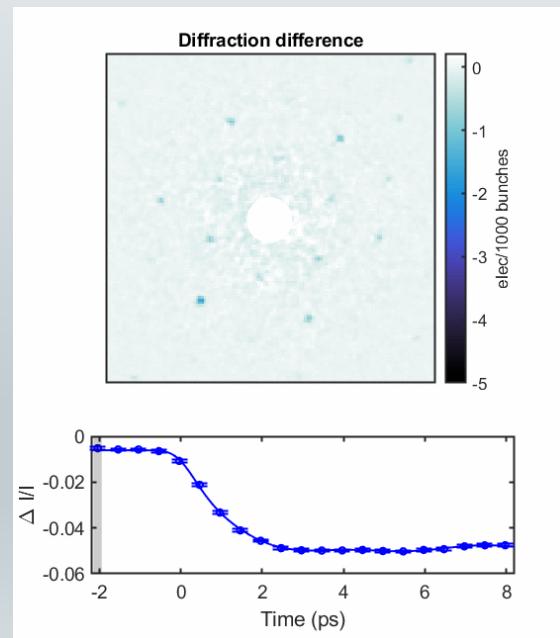
Step 1: light absorption and e-ph scattering, few ps

Mapping the full life cycle of optical excitation

- With a high rep-rate laser *and detector* need both, you can use pulse-picking to extend UED delays out to microsecond-millisecond-second timescales.
- This allowed us to watch the full “life cycle” of optical excitation in thin films.
- Example 1: Monolayer MoSe₂ atop SiN

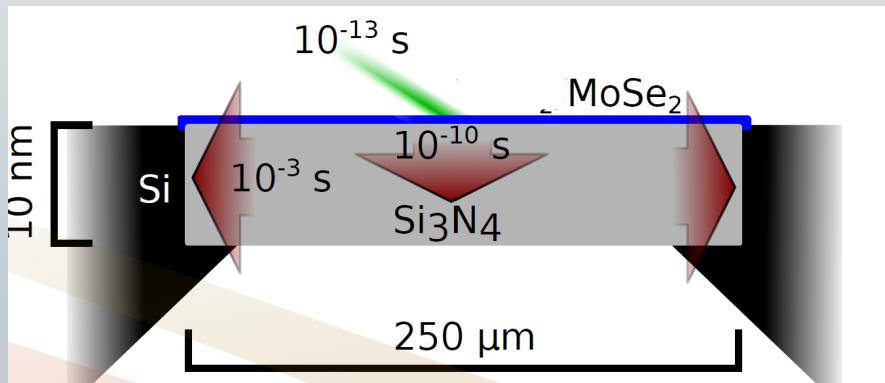


Step 1: light absorption and e-ph scattering, few ps

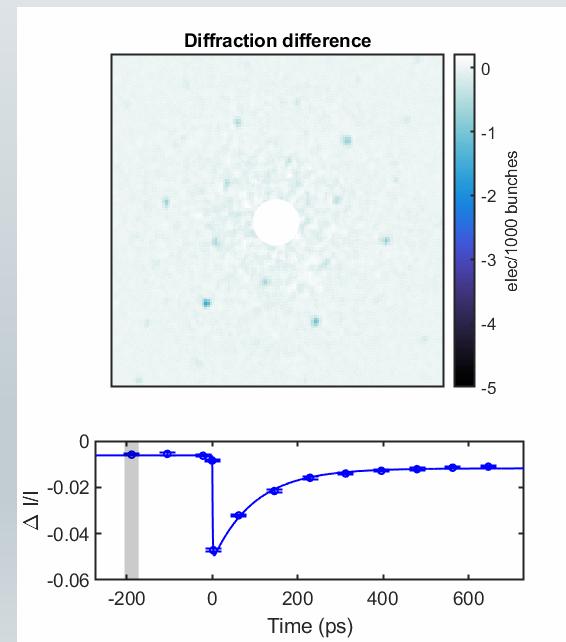


Mapping the full life cycle of optical excitation

- With a high rep-rate laser *and detector* need both, you can use pulse-picking to extend UED delays out to microsecond-millisecond-second timescales.
- This allowed us to watch the full “life cycle” of optical excitation in thin films.
- Example 1: Monolayer MoSe₂ atop SiN

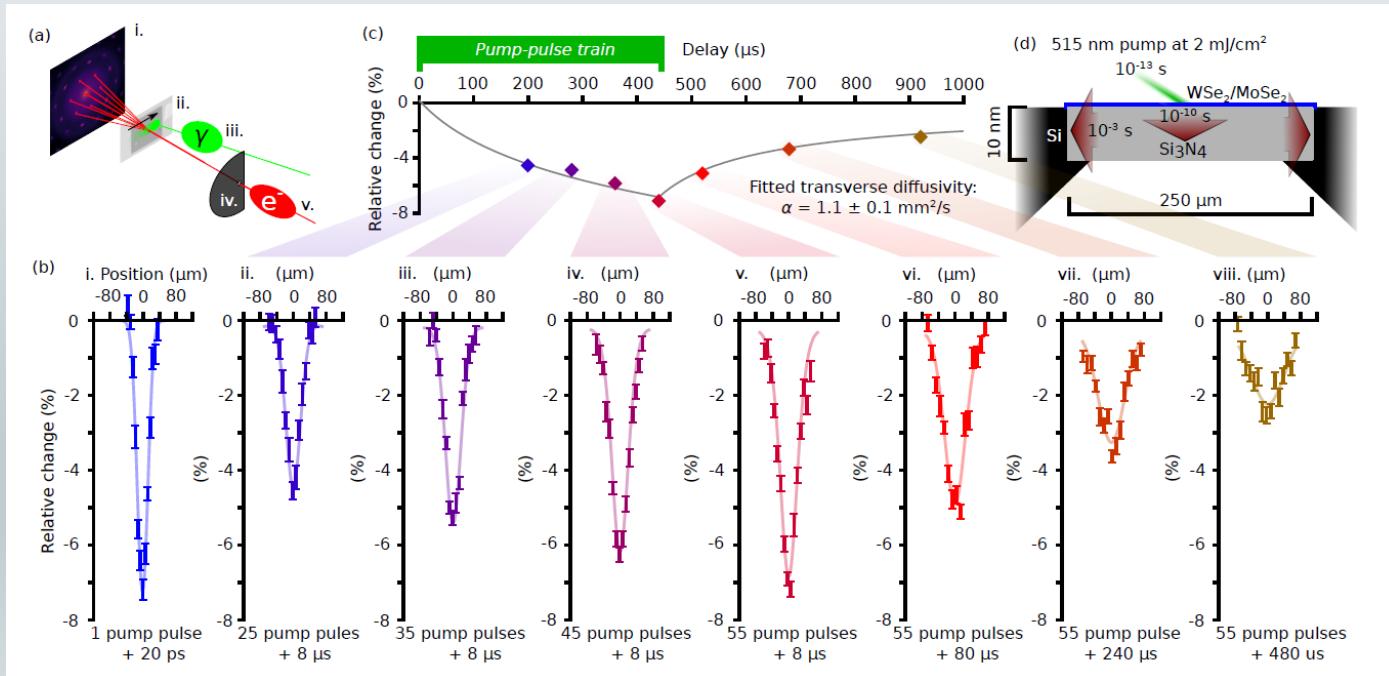


Step 2: *Longitudinal* heat transfer to SiN (100s of ps)



Mapping the full life cycle of optical excitation

- Example 2: Monolayer **WSe₂/MoSe₂ moire bilayers** atop SiN
- We use pulse picking, small spot sizes, and high rep rate to track heat transfer out to millisecond timescales, and 10s of microns in space.



Conclusions

- Operating alkali antimonides at threshold gave us dramatic improvement in beam quality in UED.
- Diagnosis and correction of aberrations out to sextupole order was critical.
- Transversely small, coherent ultrafast electron probes are very useful, particularly for UED on quantum materials.
- Our electron source, coupled with a state-of-the-art direct electron detector, enabled a novel study of moire materials.

Papers this talk draws from:

- W. H. Li et al., Structural Dynamics **9**, 024302 (2022)
- M. Gordon et al., PRAB, Accepted (2022) [<https://arxiv.org/abs/2207.13634>]
- C.J. R. Duncan et al., in review [<https://arxiv.org/abs/2207.13634>]

Acknowledgements

Photo-emission Microscopy Detectors

Cameron Duncan



William Li



Michael Kaemingk



Adam Bartnik



Alice Galdi



Young-Kee Kim



U. Chicago

Luca Cultrera



Chad Pennington



Sofia Gruner



Matt Gordon



Cornell

Ivan Bazarov

Yu-Tsun Shao



Jai Kwan Bae

Chris Pierce
Elisabeth Bianco



Matt Andorf

David Muller



Lena Kourkouffis

Lopa Bhatt
Julia Thom-Levy



Mark Tate

Juan Pablo



Stanford

Fang Liu



Helen Zeng



Aaron Lindenberg



Thank you!