

Stable and Brilliant Self-Seeded XFEL at SACLA

Ichiro Inoue

RIKEN SPring-8 Center/SACLA

On behalf of SACLA accelerator and beamline teams



Contents

Stripped Brief introduction of SACLA

Background & concept of reflection self-seeding A

Early commissioning results

Recent progress & future perspectives

SPring-8 Angstrom Compact free-electron LAser "SACLA"



High gradient C-band accelerator (acceleration gradient: 35 MV/m)



In-vacuum short period undulator (λ_u=18 mm)



XFEL amplification can be achieved at a relatively short distance (~700 m)

sub-ångström region

Hitoshi Tanaka and Makina Yabashi et al.

Accelerator and undulators of SACLA

Accelerator & undulators









XFEL properties of SACLA (BL2&BL3)

Photon energy & Photon number (Pulse energy)	4—15 keV >10 ¹¹ photons/pulse (max. 600 µJ)
Pulse duration	<10 fs
Peak power	>60 GW @ 10 keV
Rep. rate	Max. 60Hz
Band width: Pink	~5x10 ⁻³ plane mirrors
Monochromatic	~1x10 ⁻⁴ Si(111)
Coherence length	Almost the same as beam size



several %

/ position H position

Stable Spectrum

Stable Operation







Fluctuation of center of photon enargy: 10 eV (STD) Photon energy spread of each pulse: ~15 eV (STD)

Pulse

Contents

Stripped Brief introduction of SACLA

Background & concept of reflection self-seeding A

Early commissioning results

Recent progress & future perspectives

Principle of single-path FELs



Amplification processes start from undulator radiation with broad bandwidth→ XFEL beam has relatively large bandwidth (∠E/E~0.3%)

Why we need narrow-band XFELs ?

Reason I:

some experiments cannot be performed with pink beam (⊿E/E~0.3%)

Example 1: Wide angle diffraction

Example 2: Spectroscopy



For performing these experiments, SASE-XFEL beam should be monochromatized at the cost of considerable loss of photon flux

Why we need narrow-band XFELs ?

Reason II:

Crystal optics are used everywhere at XFEL facilities, and most of them utilize Bragg reflection



Crystal optics can only reflect X-ray beam with a certain photon energy (The bandwidth of the reflected X-ray beam is a few eV or less): most pulse energy of SASE-XFEL beam is lost in Bragg reflection

Generation of narrow-bandwidth XFELs by self-seeding



J. Amann *et al.*, *Nat. Photon.* (2012). R. Lindberg & Yu. Shvyd'ko, *PRST* (2012).

Transmission self-seeding at SACLA BL3



T. Inagaki et al., Proc. FEL 2014.

The e-bunch/XFEL at SACLA has some tails



I. Inoue, Phys. Rev. Accel. Beams 21 (2018).

Two Problems:

Broad SASE background

→ Probably because monochromatic wake and transmitted SASE tail was temporally overlapped and their intensities are comparable

Transmitted SASE makes machine tuning difficult

We cannot measure qualities of the seed, such as, intensity, profile, pointing etc.

Reflection self-seeding using channel-cut crystal



I. Inoue et al., Nature Photonics (2019)

Purely mono-XFEL beam is delivered to downstream IDs

•No SASE contamination in seed pulse

In the second second

 $\begin{array}{l} \textbf{(Seed power)} & = & \sim 3 \times 10^{-2} \text{ for reflection seeding (Si 111 channel cut)} \\ \textbf{(Input SASE power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm)} \\ \textbf{(Seed power)} & = & \sim 5 \times 10^{-3} \text{ or less for transmission seeding (C400, 100 μm

Difficulty in realizing reflection self-seeding



Maximum e-bunch delay: ~300 fs Optical delay by Si 111 CC-crystal



5-m compact chicane enabling large e-beam delay

Hara, *Nature Commun*. (2013). Inoue, *Proc. Natl. Aca. Sci. USA* (2016).



100-µm gap Si(111) channel-cut crystal





Rocking curve measured with 10-keV SR monochromatized Taito Osaka by Si(111) DCM (RIKEN/SACLA)





cf. Theoretical value@peak: ~80%

Photon energy: 5 keV or higher Aperture: 50 μ m@10 keV Optical delay: 120 fs@10 keV X-ray beam offset: 180 μ m@10 keV

T. Osaka et al., arXiv: 1811.0860 (2018).

Installation of µ-CC crystal (2017.12.24.)

Ohashi-san

(JASRI)

Hasegawa-san (TOYAMA) & Maki-san (SACLA)

Dec.23-25 2017

Early commissioning results (Jul. 2018)

Parameters of e-beam

E-beam energy: 7.8 GeV E-bunch duration: ~10 fs E-beam charge: 270 pC K-value: 2.1 Photon energy: 9.85 keV

The same e-beam parameters for normal SASE mode of SACLA

cf. Average spectrum of SASE-XFEL

Average spectrum of seeded-XFEL beams

I. Inoue et al., Nature Photon. (2019).

Single-shot spectrum of seeded-XFEL beam

Central photon energy
 of seeded-XFEL is
 slightly lower than that of seed

Bandwidth of seeded-XFEL is larger than that of seed

I. Inoue et al., Nature Photon. (2019).

Effect of energy chirp on the spectrum

If there is non-negligible energy chirp of e-beam, intervals of micro-bunches gradually changes along the undulators.

e.g. FEL simulation of seeded-XFELs with/without energy chirp of e-beam

Towards narrower and brighter seeded-XFELs

smaller energy chirp in the e⁻ beam → difficult to be realized immediately
 narrower bandwidth of the seed → use of higher order diffraction (Si 220)

Diffraction in higher order indices is more sensitive to lattice strain. much distorted wavefront

T. Hirano (Osaka U) treated the Si(220) crystal with a plasma etching technique (PCVM).

Nearly ideal reflection profile & rocking curve

Takashi Hirano (Osaka U)

Installation of Si(220) crystal (2018.8.22.)

<u>Si(111)</u>

Energy range: >5 keV (in design) Optical delay: ~120 fs @10 keV Si(220)

Energy range: >6.5 keV (in design) Optical delay: ~200 fs @10 keV

Seeded-XFELs generated with Si (220) crystal

Typical gain of spectral brightness by Si (220) seeding with respect to normal SASE mode is ~7

cf. typical gain with Si (111) seeding: ~6

If the electron beam condition is good (unfortunately we cannot routinely achieved this good condition), the bandwidth of seeded-XFEL and seed becomes almost the same

Osaka et al., in preparation.

Early user experiments (since June 2018, # of experiments: 8)

Typical tuning time

- Tune normal SASE with full IDs including spectrum & optical axis
- 2. Open the downstream IDs & adjust μ CC 1~2 h
- Close the downstream IDs & adjust the optical axis ~0.5 h
- Set delay and offset for e- beam to calculated values, then adjust the optical axis of SASE to the seed position ~0.5 h
- Optimize some parameters of the downstream IDs by monitoring pulse energies after the Si(111) DCM (e- beam offset both in Ver. & Hor., K value, taper of IDs, delay etc.)
 3~4 h

Just realizing seeded-XFEL: ~4 hours Optimizing seeding conditions: ~8 hours

Not short but straightforward.

(All processes can be completed by Accel. operators without help of scientists)

Long term stability

XFEL intensity after Si (111)DCM@ User experiments(Dec. 2018)

Spectral brightness kept its original value over 3 days (=typical beam time for single user)

Self-seeding is officially released for user experiments

From the current run (2019A: Apr.-Aug. 2019),

seeded-XFEL is released for user experiments as a standard operat

SACLA User Information	
SACLA Guide	Call for 2019A Proposals at SACLA
Programs/Call for Proposals	Closed
 Current Calls for SACLA Research Proposals 	The Japan Synchrotron Radiation Research Institute (JASRI) is pleased to announce the call for 2019A proposals for research to be carried out at SACLA. Please follow the following quidelines and instructions to apply
 User Operation Starts at the SACLA 	
 Reference Documents 	
• For Prospective Users	[Updates] • Self-seeded XFEL
Proposal Application	A reflection self-seeding system is available at BL3. A bandwidth of self-seeded XFEL can be much narrower than that of SASE XFEL, while an average pulse energy is
Arrival/Experiment	comparable. A typical setup time for seeded XFEL is 1 shift, which will be included <i>in a user's beamtime</i> . Please also note that it takes more time to change wavelength of
• After Experiment	seeded XFEL than in the standard SASE case. Those who plan to use self-seeded XFEL should contact the XFEL Utilization Division (sacla-bl.jasri@spring8.or.jp) 🖻 in
Proposal System and Fees	advance of the proposal submission to obtain the detailed information of the operation conditions.
• List of Proposals	
	 Experiments at BL2
	The following experiments are basically conducted at BL2:
SP8/SACLA Guest House	– Senar remosecond Crystallography (SFA) – Fixed-target Protein Crystallography (FPX)
 SPring-8 Guest House 	- Coherent Diffractive Imaging (CDI)
	• Feasibility-Check Beamtime (FCBT)
Search	We may accept a request of a feasibility-check beamtime (FCBT, max 0.5 shifts) for
Jearch	sample screening before the main experiment. Please note that the feasibility-check
 Publications Entry (Login Required) 	can be performed under the following conditions during FCB1 for the time being: – Beamline & experimental hutch: BL2 EH3
	– Experimental system: DAPHNIS (J. Synchrotron Rad. 22, 532-537, 2015)

Homepage of call for 2019A Proposals at SACLA

~25% of all scheduled experiments at SACLA BL3 will utilize seeded-XFEL beam

On-going projects

Fast photon energy scan (continuously rotating seed Xtal)

0.90

0.85

-1.0

-0.5

0.0

Delay (ps)

1.0

1.5

2.0

333 reflection measured w/

third harmonic (27 keV)

Tailoring electron-beam for self-seeding (not started yet)

b.w. of 13 eV

@27 keV

 Parameter tuning of accelerator based on spectral brightness and/or new RF deflector combined with machine learning

Summary

- ☆ A new seeding scheme using a micro channel-cut crystal, called *"reflection self-seeding"*, was implemented at SACLA and works pretty well.
- The gain of spectral brightness compared with SASE-XFEL is typically 7-8, but sometimes reaches ~10 (Si 220).
- Tuning time of self-seeding is ~8 hours and seeding condition keeps over 3 days.
- Seeded-XFELs are now provided for user experiments as a standard operation mode of SACLA.
- ☆ Fast photon energy scan, evaluate and use of harmonics of seeded-XFEL, and further improve of spectral brightness by tailoring electron beam qualities are our on-going projects.

Thanks to...

Especially,

Taito Osaka, Toru Hara, Takashi Tanaka, Takahiro Inagaki, Shunji Goto, Yuichi Inubushi, Haruhiko Ohashi, Mitsuhiro Yamaga, Kenji Tamasaku, Tetsuya Ishikawa, Makina Yabashi, and Hitoshi Tanaka

Kazuto Yamauchi, Yasuhisa Sano, Takashi Hirano, Yuki Morioka, and Shotaro Matsumura

Aymeric Robert, Diling Zhu, Yiping Feng et al.

Kwang-Je Kim, Yuri Shvyd'ko

Gianluca Geloni

Thank you for your attention!