



The Future of X-ray FELs

Hans-H. Braun Paul Scherrer Institut

3rd International Particle Accelerator Conference New Orleans May 20-25, 2012



IPAC'12 talks on X-ray FELS

IPAC12 Conference Synoptic Table: May 21-25, 2012

	Monday, May 21 2012				Tuesday, May 22, 2012				Wednesday, May 23, 2012				Thursday, May 24, 2012				Frida	Friday, May 25, 2012								
	Hall B				Room 209			Room 218			Room 209			Room 218		Room 209			Room 218		Room 209	Room 218				
Time		Chair: Victor Sulle	r, CAMD, LSU	6		Chair: Thomas Roser, BNL		BNL	Chair: Bob Hettel, SLAC		Chair: V	Vladimir Shiltsev	FNAL	Chair	Chair: Kevin Jones, ORNL		Chair	Yu-Juian Chen, LLN	L	Chair: Vladim	ir Litvinenko	BNL	Chair: Lia Merminga, TRIUMF	Chair: Andrew Hutton, Jlab	e.	
09:00						Status of the J-Par JAEA/J-PARC	re Facility, Shoji I	Vagamiya,	Progress Towards Sources, Michael	Ultimate Storage F Borland, ANL	Ring Light	Project X: Resear Collaboration Effi	ich, Design, Devel lorts for an Ultra H	opment and ligh Power	Recent Advances Visualization of U	and New Techniqu Jura-short Relativi	istic Electron	Backgrounder: P rays, Mark Sybe	roducing Medical Isoto de Jong, CLS	pes using X- Suppre Broad	ession of Head-Ti band Feedback, 1	ul Instability (long Ho Chin,	Using a KEK	Marrying Lasers and Beams, Luco Serafini, IN Milano	N Symplectic Tracking and Compensation of Field Integrals in Complex Undulator Struct	if Dynamic uctures,
09:10		Welcome from	the Chairs									Superconducting	RF Linac, James	Kerby, Fermilab	Beam Bunches, L	And Allang, SLAC									Johannes Bahra, HZB	
09:20																										
09:30						The Upgrade Plan Roland Garoby, C	u for the LHC Inj 'ERN'	ector Complex,	Review of ERL P World, Norio Nai	rojects at KEK and tomura, KEK	Around the	Development of E Vasily Parkhomch	Electron Coolers in hak, BINP	n Novosibirsk,	Diagnostics for H Edda Gschwendt	ligh Power Targets ner, CERN	and Dumps,	Operation and Pa Erminia Bressi, (itient Treatments at CN CNAO Foundation	AO Facility, Experi Cohere	imental Demonstr ent Synchrotron B	ration of Supp Radiation, Vita	ressing dy Yakimenko,	Recent Developments in Lasers for use in Accelerators, Andreas Tinnermann, Friedrich	Femtosecond Electron Guns for Ultrafast E Diffraction, Jinfeng Yang, Osaka Universit	Electron
09:40	The Pi	uture of X-ray Science	e, Joachim Stok	r, SLAC																BNL				Schuler Chryerstaat, Jena		
09:50																										
10:00						Increasing the AG Jumps, Vincent Sc	S Beam Polarizat horfer, BNL	ion with 80 Tune	Overview of Pres Sources, Ying K.	ent and Future Con Wu, FEL/Duke Unit	apton Photon versity	Accelerator Physi Michael James Sy	ics and Technolog ophers, MSU	y for FRIB,	Protecting Accele of Sophisticated (erator Control Syste Cyber Attacks, Ster	ettas in the Face ven M. Hartman,	Accelerator Syste Steven Lidia, LB.	ems for Heavy-Ion Iner NL	ial Fusion, Beam Proton	and Spin Dynam EDM, Richard J	ics in an All-el dichael Talma	lectric Ring fo n, CLASSE	Overview of Recent Progress on High Repetitio Rate, High Brighmess Electron Guns,	Review of Microwave Schottky Beam Dia; Ralph James Pasquinelli, FNAL	Agnostics,
10:10	First Two	o Years of LHC Oper	ation, Steve Mys	ers, CERN											UKAL									r ernanao sannibale, LBNL		
10:20						_											_									
10:30											Coffee Break	10:30-11:00														
		Chair: Stuart Hen	derson, FNAL	S		Chair:	Oliver Brüning,	CERN	Chair: Ge	org Hoffstaetter, U	. Cornell	Chair:	Katsunobu Oide	, KEK	Ch	air: In Soo Ko. PA	AL	Chair	Akira Noda, U. Kyo	•	Chair: Valer	i Lebedev, F?	NAL	Hall B Chr	: Jeff Corbett, SLAC	
11:00						Kesearch and Dev Collider, Katsaya	elopment Toward Yonehara, Fermi	lab	Proton Beam Acc Laser Pulses, Xuo	eseration with Circi aging Yan, PKU/IHI	uar Polarized IP	CLIC Design and CERN	K&D Status, Stel	nar Stapnes,	Laser (SACLA), Centre	pstrom Compact Pr Hitoshi Tanaka, RI	IKEN SPring-8	igh field Magne KEK, Ibaraki)	et Læveloptient, Tatsus	and Ex	-beam Effects in I speriment, Kazah	ito Ohmi, KEI	ers: Theory		an and the second second	
11:10	Accelerator	e Driven Systems, Dir	rk Vandeplassch	e, SCE-CEN																				The Future of X-ray I	Ls, Hans-Heinrich Braun, PSI	
11:20						Owner of a			1	Turner of One 1	and a	Cardina of Th	n Charl Barry D					adama Tadaad								
11:30						Colliders, Visitm	Phitsyn, BNL	uection.	Accelerator Struc Joel England ST	tures from Silicon : AC	and Silica,	Future Damping F CLASSE	Rings at CestTA, 0	Gerald Dugan,	Paul J. Emma, SI	LAC	OFEL.	interaction X-ray	Source, Fred V. Harto	nan, LLNL Electro	on Cooling, Gang	Wang, BNL	ouetent			
11:40	State of the Art and F	Putate Prospects in RJ	F Superconducti	ivity, Kenji Saito, I	EK.												_							Accelerators for Intensity P	ntier Research, Paul Derwent, FNAL	
11:50												-														
12:00	Davies and Grandeder of					HEP Beijing	PTOSPECT OF BEPC	ια, Quis Qui,	Yoshiharu Mori,	KURRI	KIS,	Maria Enrica Bia	igini, INFN/LNF	1	Experiments at JI	LAB, David Dough	as, JLAB	Intensity ISOL R Pierre Gerard Br	iges and Puttice Direct IB Production, ricoult TRIUMF	ons in High Conec	lex, Elias Métral,	CERN	Injector			
12:10	Design and Similation of	Sergei Nagaitse	ry, Fermilab	ne Optics Test Aco	elerator,																			Physics Results at the LHC and Implicati	is for Puture HEP Programmes, Rolf Heart, CE	ERN
12:20																										
12:30		1									Lunch Brea	ak 12:30 - 2:00												Closing Remarks: Jeff C	bett, SLAC, IPAC12 SPC Chair, NAP IPAC13 OC Chair	
	Room 209	Room 2 Chair: Andrzei	218 Wolkh, U.	Room	221	Room	1 209	Roon	1 218	Room	221	Roon	n 209	Roon	n 218	Room	1 221			Hall B				Larving Larv,		
	Chair: Won Namkung, PAL/Paul Schmor, TRIUMF	Liverpoo Cockcroft/Joseph Wisconsin-M	ol & Bisognano, U. Madison	Chair: Paolo Pi Milan/Mats Lin	erini, INFN adroos, ESS	Chair: Vadi BNL/Byung Ho	im Ptitsyn, Choi, KAERI	Chair: Stev LBNL/Chan	e Gourlay, Joshi, UCL	CERN/Caterina Frase	Biscari, INFN tati	DESY/Chuang Beij	Zhang, IHEP	Chair: Davi Maryland/Ale	d Sutter, U. x Chao, SLAC	Special Session fo Chair: Alan M.M	or Industry I. Todd, AES		c	Awards Sessio hair: Stan Schriber	n r, IEEE					
14:00	Comell ERL Photoinjector	A Proton-driven Pla Accelerator Expe	asma Wakefiel eriment with	The European 2	FEL LLRF	3-Dimensional Electron Clouds	Modeling of in Non-uniform	Timing and Syn	thronization for	Challenges of the	FAIR Vacuum	Injected Beam In 3 with a Diaira	aging at SPEAR	New Results fr	om the EMMA	Future Governmer Accelerator Projec	nt-funded rts,	presented	by Jeff Corbett, SLAC.	Student Poster Pri SPC Chair and Kay	izes Winenburg, DE	SY, SPC Mem	iber (10)			
14:10	Progress, Bruce Dunham, Class	CERN SPS E Patric Mugg	Bunches, gli, MPI	System, Julien Br	anlard, DES	Magnetic Fields, 1	Seth Veitzer, Tech	Project, Frank	Lenkszus, ANL	System, Andreas	Eraemer, GSI	Hao Zha	ng, UMD	STFC/D	L'ASTeC	Zheniang Zhao, Si	INAF		New	by Elected IEEE-NF Presenter TBD (1	PS Fellows 10')					
14:20	Instrumentation and Diagnostic for High Repetition Rate Linac	First Results from	the Electron	Superconducting B	RF Accelerator	Focusing Charged Using Multipole	d Particle Beams e Magnets in a	Silicon, Diamo	nd and Liquid	Development of I	HTS Magnets,	Beam Profile I Highly Brillin	Diagnostics of mt and Highly	Photocathode F	L&D at Cornell	Future Medical Ac Riyoshi Yasuoka, I	ccelerators, University of		IEEE/NPSS Part	Internation Sci Presenter TBD (1	ience and Tech An 10')	wards				
14:30	driven FELs, Stefano De Santi: LBNL	FACET, Enk A	dli, U. Oslo	Test Facility a Jerry Liebfri	r Fermilab, iz, FNAL	Beam Transport L	ine, Yosuke Hari. TARRI	Measurement Christoph Ku	s at 2 Kelvin, fuerst, CERN	Eichigt Hatan	aka, RCNP	Energetic Ele Gero Kul	ctron Beams, be, DESY	CLA	ISSE	13BARCA										
14:40	Fast Feedback Strategies for Longitudinal Beam Stabilization Doe Biotic	FACET Firs Commissioning, S	st Beam Steven Yocky,	Superconducting Development for	the FRIB and	Tests of Low Ex Techniques for Su	aperB at SLS and	Five Years of Experience	Coperation e at HIT,	Aperture 11 Demonstrator Di	T Nb3Sn tpole for LHC	Development of Monitor in J-PAI	the Deam Halo RC 3-GeV RCS,	An Update on a Photonic Band	Superconducting Gap Structure	Accelerators, Pete DTRA	r Zielinski,		IEEE/NPSS I Awar	ied by Ilan Ben-Zv	n, BNL (5')	93				
14:50	DESY	SLAC	c	Alberto Facco,	INFN/LNL	INFN:	LNF	Andreas P	eters, HIT	Upgrades, Alexa FNA	ander Zlobin, L	PA	RC	Evgenya Sin	akov, LANL		1									
15:00	Thorium Evergy Futures, Robert Cywinski, U.	Transverse-to-k Emittance-excha	ongitadinal ange with an ed Beam	Electron Linar J Driver for the F	tare Isotope	Summary of Ferr Electron Cooler	nilab's Recycler Operation and	A European Pr Compton Gamm ET LNP Count	oposal for the a-ray Source of	Experimental Ver CLIC Two-beam 1 and Outlook Pro-	Scheme, Status	Construction P RHIC Elect	Progress of the ron Lenses,	Measurement of Spread of Electro	the Slace Energy n Beam at SDUV CHG-based	angle Mayanced A opplications, Wim BNL	Leemans,									
15:10	Huddersfield	Jayakar Thango	araj, FNAL	Chin Chao,	TRIUMF	Studies, Lionel	Prost, FNAL	INFN	-LNF	CER Small-Bera Cr	EV Slimation at	Wolfram Fi	ischer, BNL	Method, Chao	Feng, SINAP	Future Accelerator	r-based Material		esorgen contranging :	Presenter TBD (20)	syster Award				
15:20	Perspectives of the HIE- ISOLDE Project at CERN,	Probe to Explore 1 Storage Rings	Wakefields in Rec Biology	Cryomodule Dev	elopment for	Operation of New for RIKEN PL	scioning and w Linac Injector beam Eactorr	Commissioning	of the PLS-II,	SuperKEKB to S Scattered Particle	top Beam-Gas is and to Avoid	Tevatron End-of-I Experiments, Ale	Run Beam-Beam exander Valishev,	Laser-Compto	n X-ray Using	Processing, Wendy UMAN	y Ruth Flavell,									
15:30	llacine Kadi, CERN	PhLAMCE	ERCLA	Kensei Umen	IOTI, KEK	Kazunari Yan	nada, RIKEN	Jennyhind	week, t Ok	Instability, Higgs	de Coupling ski Nakavama, surements of e-	FN	(AL	Kazuyuki Si	ikaue, RISE	Punue Accelerator	r-based		s	pecial Dwited Prese	mtation					
15:40	High-power Coherent THz Sources and TH2-TDS System, with Chamber Shielding, A Multi-purpose X Band Accelerating Structure,		se X Band Scructure,	Compariso Superconduc	on at SNS	Status of the Fi Project, Mich	ERMI@Elettra ele Svandritit.	Cloud Mitigation Electrodes in t	using Clearing the DAFNE	Beam Tests of a Gas-Filled Cav	High Pressure	Radiation in Pha	ase Space Using	Instruments, Olive	r Kester, GSI		LIGO, the Laser In	Rainer Weiss M	y-wave Observat	ary (30')						
15:50	Masafumi Kumaka, RISE	Demin Zho	W, KEK	Micha Dek	ear, PSI	Andrei Shis	hlo, ORNL	ELE	784	Collider, Day	na Alesini, LNF	Coulider, Thomas	Schwarz, FNAL	Ivan Bazari	W, CLASSE			1								
16:00	Canal Decatur	Poydras	Royal	Bourbon	-	Canal	Decatur	Poydras	Royal	Bourbon	-	Canal	Decatur	Poydras	Royal	Bourbon	-	Canal	Decatur	Poydras B	Royal B	ourbon	-	4		
	Colliider Hadron	SR and FEL	Distriggent	Colluder	-	Beam Dyn	Sources	SR and FEL	Colliider	Beam Dyn	-	Tech	Tech	Sources	Beam Dyn	Sources	-	Tech	Tech	Hadron Ins	trument 1	Hadron	-	-		
	Beam Dyn _	-	-	Instrument	-	-	-	-	-	SR and FEL	-	-	-	Instrument	-	Tech	-	-	-	- App	lacations App	plications	-	-		
17:50		-	-		-	-	-	-	-		-	-	-	-	-	-	-	-	-	-	-	Tech	-	-		
		Chairman's]	Recption					Industrial	Reception				Wome	n m Science and	Engineering Rec	reption				Banquet				1		
			,																							

MC7: Technology

MC8: Applications

+	many	good	posters	on the	subject
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MC8: Instrumentation

MC5: Dynamics

Content

- Recent developments
- X-ray FEL* user community
- R&D issues
- Near future of X-ray FELs
- Mid term Future
- Long term future

*Definition used:

X-ray FEL is a FEL with smallest $\lambda_{\text{photon}}{<}5\text{ nm}$

Many thanks to

Paul Emma/LBNL Bart Faatz/DESY Katja Honkavaara/DESY In Soo Ko/PAL Subrata Nath/LANL Jörg Rossbach/Univ. Hamburg Siegfried Schreiber DESY Michele Svandrlik/ELETTRA Hitoshi Tanaka/Spring-8 Dong Wang/SINAP and my colleagues at PSI

for providing information for this talk

The young history of X-ray FELs WEYB02 First Demonstration of Hard X-ray Self Seeding at LCLS Michele Svandrlik Profile Monitor XPP:OPAL1K:1 12-Jan-2012 13:11:36 8.3 keV diamond OUT TUOBB03 20 eV FERMI FEL: un di materi in terent Light Source (LCLS) is transforming the face ine than 40 years, SLAC's two-mile long linear accel-FERMI 80 – 4 nm SEEDED HGHG aelettra Seeded SASE (240 µJ) L -0.25 x (eV) 🔶 0.45 eV insert Pofile Monitor XPP:OPAL1K:1 12-Jan (5×10⁻⁵) SASE diamond IN (300 µJ) n energy (eV) A well seeded -0.45 eV pulse (not typical)seeded The World's First Hard X-ray Free-Electron Laser SASEhas chicane OFFF 2009 2010 2011 2012 FLASH First Lasing at 4.45 nm on June 6/7 (with 3rd harm.) SACLA 10 June 2011 6 June 2010 🕅 R I K E N Announcement SAC A Lased 1200 Lasing < 5 nm 2010 Fixed gap undulator: 1000 K Lasing at 6.5 nm 10/2007 (MeV) 2 27 800 Energy (Lasing at 13 nm 4/2006 At 16:10 on June 7 2011, we accomplished "Lasing" with SACLA, our newest X-Ray Free Electron Laser Facility Lasing at 25 nm 12/2005 Construction of SACLA began in 2006 as part of Japan's Key Technology of National Importance program. 600 ε Ne appreciate your support in helping us to achieve this milestone. We will do our best to live up to your expectations. Bea Lasing at 32 nm 1/2005 Electron 400 FEL at TTF 1 (upgraded into FLASH in 2003) Hitoshi Tanaka Proof-of-Principle for SASE in the VUV First Lasing 2/2000, Saturation 9/2001 200 **WEYB01** 20 40 60 80 100 120 140 160 180 400 800 FEL Wavelength (nm)

Paul Emma

X-ray FEL users

- who are they ?
- what instruments are they used to?





 \Rightarrow X-ray FEL allows for flash images on time scale of fastest chemical processes



\Rightarrow FELs have the monopoly for intense coherent X ray beam

X-ray FELs vs. synchrotron

Worldwide >50 synchrotron for photon science with ≈1000 user stations total

• wide range of applications in life science and material science attractive science portfolio of high brightness X-ray sources

Advantages X-FELs vs. synchrotron

- + orders of magnitude higher peak brightness
- + order of magnitude higher average brightness
- + orders of magnitude better time resolution
- + clean polarization and coherence

Difficulties X-FELs vs. SR storage rings

- much less integrated photon flux
- less beamlines/facility

World map of SR storage rings

- less stable beam conditions
- much higher cost per user station



vast scientific potential in uncharted territory



"New Science" session organized by Jochen Schneider/DESY Every speaker was asked for a XFEL parameter-wishlist!

Monday, August 22

13:00 ~ 15:00	New Science
13:00	Multi-photon processes and two-color studies in atomic and molecular systems
	M. Meyer, European XFEL GmbH
13:30	Overview of Warm Dense Matter experiment at the FLASH and LCLS Free Electron
	Lasers
	B. Nagler, SLAC
14:00	Femtosecond protein X-ray nanocrystallography
	H. N. Chapman, DESY
14:30	XFEL interaction with correlated electron materials
	H. A. Dürr, SLAC

FEL 2011

FEL

FEL parameter wish list for

Research on Atoms and Molecules

Michael Meyer (European XFEL)

i) Non-linear processes	> 10 ¹⁵ W/cm ²				
Direct processes	"short" pulses: < 10 fs				
Sequential processes	"long" pulses: 100 – 300 fs				
Variable pulse duration	& pulse energy				
ii) Time-resolved experiments	"Synchronization"				
Pulse duration	< 10 fs (< 1 fs)				
Temporal stability	< 10 fs				
iii) "Complete" experiment	Coincidence techniques				
repetition rate	> 1 kHz ("MHz")				
iv) Soft X-ray regime	C(1s) = 290 eV, N(1s) = 410eV,				
"tunable"	O(1s) = 560eV				
v) linear / circular polarization	Dichroism				
2011 FEL parameter wish list for					
Macromolecular Single-Shot Coherent Imaging					

Henry N. Chapman (DESY-CFEL)

- The key metric is photon power. Ideally ~10 TW
 - This gives about 10²⁰ W/cm² (with 1 micron focus, assuming beamline and focusing efficiency)
 - Only the first 10 to 30 fs of the pulse usefully contributes
- Wavelength range: 4 keV to 14 keV (to cover elemental edges from S to Se)
 - Also: 300-500 eV for water-window imaging of cells, viruses
 - Up to 30 keV for time-resolved imaging of nanoparticles
- · Repetition rate: As high as possible
 - Need to match detector capabilities. 1 kHz repetition could be feasible
- · Bandwidth: As high as possible
 - 1 to 10% bandwidth would allow structure determination with about 1% of the required pulses.
 i.e. structure determined in <1000 shots

FEL 2011

FEL parameter wish list for

Matter in Extreme Conditions Research

Bob Nagler (SLAC-LCLS)

: 10 fs 00 – 300 fs	 Short pulse lengths: isochoric heating with inertial confinement beat hydrodynamic expansion (~10ps) beat electron-ion equilibration time (~1ps) beat inner shell recombination rate (1-100fs) beat electron-electron equilibration time (sub fs)
les z") 1s) = 410e∨,	 High Photon Flux High energy deposition, to increase heated sample size and temperatures (10¹³-10¹⁴ photons) Photon hungry probe techniques (10¹³-10¹⁴ photons in small bandwidth 10⁻⁴ - 10⁻⁶) Bandwidth Thomson scattering probe: 10⁻⁴ - 10⁻⁶ X-ray absorption near edge spectroscopy : 1%-5%
	 Repetition rate 10Hz, limited by optical laser, target refresh rate
g	FEL 2011 FEL performance wish list for Condensed Matter research (especially correlated materials) Hermann Dürr (SLAC-Pulse/SIMES)
sing efficiency)	Femtosecond Two Color X-Ray Holography
rom S to Se)	fs optical pump fs optical pump fs optical pump fs optical pump fs optical pump fs optical pump fs optical pump fs optical pump fs optical fs o

X-ray FEL dream machine wish-list summary from FEL'11 "New Science" session

	Shortest Pulse length	Max Peak Power	rep. rate	Polarisation control	band width
Research on Atoms and Molecules	<10 fs	0.1 TW	> 1 kHz	yes	
Matter in extreme conditions	<1 fs	1 TW	10 Hz		10 ⁻³ -10 ⁻⁶
Macromolecular Single- Shot Coherent Imaging	10-30 fs	10 TW	1 kHz		10 ⁻¹ -10 ⁻²
Condensed Matter Research	1-10 fs	1 TW		yes	two color
state of the art	1-100 fs	≈10 GW	LCLS 120 Hz FLASH 10x800	only at FERMI	SASE 2•10 ⁻³ seeded 5•10 ⁻⁵

⇒ more power, shorter pulses, more pulses and more bandwidth control (chirp!)



X-ray FEL R&D Issues

Generic X-ray FEL*





achieved with different injector technologies

LCLS:	S-band 1 ¹ / ₂ cell RF gun with Cu photo-cathode
PITZ:	L-band 1 ¹ / ₂ cell RF gun with Cs ₂ Te photo-cathode
SACLA:	Pulsed diode with thermionic CeB ₆ cathode
SwissFEL:	S-band 2½ cell RF gun with Cu photo-cathode

open issues

...

- low emittance for cw operation ⇒*Fernando Sannibale, FRBBA01*
- operational robustness
- lower emittance always welcome



X-ray FEL R&D Issues cont.

Generic X-ray FEL*

^{*}number of BC and undulator lines may vary



Theory of CSR and microbunching instability made enormous progress during last decade

LCLS has set the standard for most projects: laser heater to cure microbunching instability harmonic cavity for flat current profile long, small angle bunch compressors to minimize CSR emittance growth

Arguably biggest issue today is phase and amplitude stability of upstream RF



X-ray FEL R&D Issues cont.

Generic X-ray FEL*

^{*}number of BC and undulator lines may vary



Not a feasibility issue, but major factor for investment and operation cost

R&D topics

- A & ϕ stability
- Reliable and cost effective RF sources
- Low power consumption for reasonable accelerating fields



Present XFELs linac technology & parameters driven by particle physics ancestors The LC Mantra of high gradient is not the key issues for X-ray FELs!

Make XFELs available for national labs and users ⇒ more economical solutions required without compromising scientific potential

Cost comparison linac technologies or Why doesn't everybody build s.c. & c.w.

Technology	Linac investment cost w/o building	Typical gradients (excl. fill factor)	Electric consumption		
Pulsed n.c. with SLED	≲ 10 M€/GeV	20 MV/m (S-band) 30 MV/m (C-band)	≈ 0.5 MW/GeV		
Pulsed superconducting	≈ 20 M€/GeV	24 MV/m	≈ 0.5 MW/GeV		
c.w. superconducting	≳ 30 M€/GeV	15 MV/m	≈ 5.0 MW/GeV		

Beware! This is not exact science !

cost optimization s.c. linac in c.w. mode from STFC-NLS design report



Relative total capital and 10-year operational linac costs



Effect of increased energy cost on Eacc optimisation over 10 years

cost optimization pulsed n.c. linac for SwissFEL



Advantage of C-band is in real-estate needs and electricity consumption

PAUL SCHERRER INSTITUT

SwissFEL C-Band Linac Module



X-ray FEL R&D Issues cont.

Generic X-ray FEL*





In vacuum, variable gap undulators: for X-ray FELs pioneered by SACLA with λ_{U} =18 mm

Apple II undulators: for X-ray FELs pioneered by FERMI@ELETTRA

Open questions

- What is the smallest acceptable gap 4,3,2mm? Decisive question for design of compact X-ray FELs Tests at SACLA foreseen
- Shorter period PM undulators
- Improved PM materials
- Feasibility of short period sc undulators

Excellent undulator technology reviews in proceedings of FEL'11

J. Bahrdt, "Pushing the Limits of Permanent Magnet Short Period Undulators"

S. Prestemon et al., "Development of Superconducting Undulators"



U15 in-vacuum undulator for SwissFEL



X-ray FEL R&D Issues cont.

Generic X-ray FEL*



SASE or Seeding ?

SASE principle



- + FEL amplifier without mirrors and without input signal \Rightarrow applicable for large wavelength range Baseline operation mode for LCLS, SACLA and FLASH
- Individual time slices in the bunch radiate independently \Rightarrow spiky time structure, spectrum,

intensity jitter by principle!

HHG seeding principle



HHG figure from http://www.newlightsource.org/events/presentations/open_meeting/Tisch.pdf

sFLASH seeding experiment at FLASH

Experiment for direct HHG seeding in XUV range

First seeding (38 nm) April-29, 2012



Courtesy

Katja Honkavara/DESY



HGHG principle



L.H. Yu, PRA 44, 5178 (1991)

slide courtesy Dong Wang SINAP





- **Two FEL lines** will cover different spectral regions.
- FEL-1, based on a single stage high gain harmonic generation scheme initialized by a UV laser, covers the spectral range from 80 nm down to 20 nm.



 FEL-2, in order to be able to cover the wavelength range from 20 nm down to 4 nm starting from a seed laser in the UV, is based on a double cascade of high gain harmonic generation. A magnetic electron delay line is used in order to improve the FEL performance by using the fresh bunch technique.

EEHG principle



G. Stupakov, PRL 102, 074801 (2009)

slide courtesy Dong Wang SINAP

Courtesy Dong Wang/SINAP



First lasing of EEHG FEL at Shanghai SDUV

(3rd harmonic, 350nm, April 2011)





Difficulties of seeding with external laser for short wavelength

- 1. External laser amplifier has $\lambda_{photon} \approx 1 \mu m$ Seeding FEL with $\lambda_{photon} = 1 nm$ requires generation of harmonic 1000!
- 2. Spectral power of e⁻ shot noise increases with $1/\lambda_{photon}$ Seed signal power has to exceed shot noise, otherwise SASE radiation will dominate



figure from http://www.newlightsource.org/events/presentations/open_meeting/Tisch.pdf

First Demonstration of Hard X-ray Self Seeding at LCLS Courtesy Paul Emma / LBNL, WEYB02



SASE and Seeded spectra recorded on single shots. The left panels are SASE with 150 pC, 3kA peak current, un-seeded. The FWHM of the SASE spectrum is 0.2 % Bandwidth. The right panels are the seeded beam with the same electron beam parameters. The FWHM of the seeded beam is 0.5 eV (5x10⁻⁵ bandwidth)

gonne



Single shot pulse energy from the gas detectors with 40pC charge

- Concept developed by Geloni, Kocharyan and Saldin, DESY 10-053 (2010).
- The mean seeded FEL power is 8 GW with a 2.5 GW SASE background at 8 keV for 40 pC bunch charge.
- Peak seeded power is in excess of 15 GW, comparable to SASE but with a spectral bandwidth reduction by the factor of 40.
- Next steps include system optimization of the LCLS undulator beamline including additional undulators which should increase seeded power and reduce intensity fluctuation.



SLAC-Argonne-TISNCM Collaboration

Seeding does not only improve bandwidth and longitudinal coherence but allows also for much better e⁻ to photon conversion with tapering of undulator strength



TOWARD TW-LEVEL, HARD X-RAY PULSES AT LCLS*

W.M. Fawley¹, J. Frisch¹, Z. Huang¹, Y. Jiao¹, H.-D. Nuhn¹, C. Pellegrini^{1,2}, S. Reiche³, J. Wu^{1†} ¹SLAC National Accelerator Laboratory, Menlo Park, CA 94025, USA ²Department of Physics and Astronomy, UCLA, Los Angeles, CA 90095-1547, USA ³Paul Scherrer Institute, Villigen PSI, 5232, Switzerland

First beam parties, reserve date in your agenda

- 2012 FERMI FEL II, 4nm HGHG seeding cascade, circular polarization with Apple II
- 2013 FLASH II, second FEL undulator line with variable gap and seeding upgrade
- 2013 SX FEL based on SCSS in SACLA undulator hall, *Hitoshi Tanaka WEYB01*
- 2015 European XFEL, HX FEL1, HX FEL2 and SX FEL3
- 2015 PAL XFEL HX and SX
- 2016 Shanghai SX FEL
- 2016 SwissFEL HX FEL
- 2017 LCLS II, HX and SX
- 2019 Shanghai HX FEL
- 2019 SwissFEL SX FEL
- ? NGLS
- ? MaRIE

FERMI FEL: 80 – 4 nm SEEDED HGHG

FEL I in operation (20-80 nm), FEL II commissioning starting this year aiming for 4nm with cascaded HGHG seeding! Talk Michele Svandrlik ,TUOBB03

~ 60 m Experimental Undulator Hall

@elettra

200 m Linac Building **FLASH II**

courtesy of Bart Faatz/DESY



Electron Beam Parameter	
Beam Energy	0.5 – 1.25 GeV
Normalized Emittance	1.4 mm mrad
Energy Spread	0.5 MeV
Peak Current	2.5 kA
Number of bunches /second	<8000***
Undulator Parameter	
Period	31.4 mm
Segment Length	2.5 m
Number of Segments	<=12
Focusing	F0D0
Radiation	
Wavelength SASE*	4-60 nm
Wavelength HHG*	10-40 nm





Photon beamlines

XFEL site Schenefeld

> XFEL site Oscorfer Born

Approx. 94% of tunnels finished

- 100% of 5.2m diameter
- Ca. 87% of 4.5m diameter
- 5429m of 5777m in total

Accelerator tunnel

Undulator tunnels Largest X-ray FEL facility

Curcopen X-ray laver project XFEL

Shead A gold FEET COLST Harvey and Brief COLST PROCEEDING TO COLST Harvey 17. 81 (2012) 17 GeV s.c. linac Min. λ_{photon} =0.5Å Commissioning starts 2015

XFEL Layout (as of April 2012)





PAL XFEL courtesy of In Soo Ko/PAL



PAL

	FEL wavelength [nm]	0.1
Electron	Beam energy [GeV]	10
Linac	Beam charge [nC]	> 0.2
	Beam emittance [mm-mrad]	< 0.5
	Injector Gun	Photocathode RF-gun
	Peak current at undulator [kA]	> 3
	Repetition rate	120 Hz
	Number of bunches	Single or Two
	Linac structure	S-band
Undulator	Undulator type	Out-vacuum
	Undulator period [cm]	2.46
	Undulator gap [mm]	6.8
	Undulator parameter, K	2.076
	Saturation length [m]	56
FEL	FEL radiation power [GW]	> 29
	Photon beam length [fs]	60
	FEL photons/pulse	> 1.0 E+12

Wavelength

Major Parameters

- Soft x-ray: 1 nm ~ 10 nm
- Hard X-ray: 0.7 ~ 0.1 nm
- · Extended to 0.06 nm Photon beam Length
- Nominal : 30 ~ 100 fs (200 pC)
- Short : < 5 fs (20 pC) Ultra short: < 0.5 fs by ESASE scheme
- **Undulator Beamline**

3 Hard X-ray / 2 Soft X-ray lines



- Untapered : 14 GW (4.7E+11 photons)
- Tapered : 55 GW (1.8E+12 photons)



Aramis

1-7 Å hard X-ray FEL for **SASE with reservations for self seeded** operation, In-vacuum, planar undulators with variable gap. User operation from 2017

Athos

7-70 Å soft X-ray FEL for **SASE & self seeded** operation . APPLE II undulators with variable gap and full polarization control. User operation from 2019

LCLS II

From https://slacportal.slac.stanford.edu/sites/lcls_public/lcls_ii/Published_Documents/CDR%20Index.pdf



FEL Roadmap at SINAP

courtesy of Dong Wang/SINAP



FEL Roadmap at SINAP

courtesy of Dong Wang/SINAP



FEL Roadmap at SINAP

courtesy of Dong Wang/SINAP

Parameters of Compact XFEL

Electron beam para	neters				
Energy/GeV	6.4				
Peak current/kA	3				
Bunch charge/pC	250				
Normalized slice emittance/mm-mrad	0.4				
RMS slice energy spread	0.01%				
Full bunch length/fs	100				
Undulator parameters					
Period/cm	1.6				
Segment length/m	4.8				
Full undulator length	70				
Peak undulator field/T	0.93				
К	1.4				
	6				
Gap/mm					

шШ

0.1

FEL parameters					
Radiation wavelength/nm	0.1				
ρ	3.41e-4				
Peak coherent power/GW	10				
Peak brightness/*	2e33				
Pulse repetition rate (Max.)/Hz	60				
3D gain length/m	2.156				
Saturation length/m	50				







Proposed MaRIE 1.0 (Matter - Radiation Interactions in Extremes) will be an Accelerator user facility, leveraging LANSCE 0.8 – GeV proton accelerator adding to it an XFEL - a unique source of very hard, coherent, brilliant photons





Pre-conceptual Design Parameters

Energy	12 GeV
Linac frequency	3 GHz
Linac type	RT Cu
Cavity gradient	25* MV/m
Maximum beamline θ	4 degrees
Bunch compressor 1	6 m
Bunch compressor 2	22 m

RF pulse duration	1.5 μs			
RF pulse rise time	0.8 μs			
RF peak power	80 MW			
RF Repetition rate	60 Hz			
Accelerator active length	480 m			
Accelerator length from	528 m			
injector to linac end				

FST 1943



X-ray FEL Megatable

blue: in operation, green: under construction or advanced planning, orange: design in progress

	LCLS	LCLS II	Eu-XFEL	SACLA	FLASH	FLASH II	FERMI	SwissFEL	PAL XFEL	Shanghai XFEL	NGLS	MaRIE
Shortest wavelength	1.5 Å	1 Å	0.5 Å	1 Å	40 Å	40 Å	40 Å	1 Å	1 (0.6) Å	1 Å	10 Å	0.3 Å
Undulator type hard X-ray.	Fixed gap	Variable gap	Variable gap	In- vacuum Var. gap	n.a.	n.a.	n.a.	In- vacuum var. gap	Variable gap	Variable gap	n.a.	?
Undulator type soft X-ray.	n.a.	Variable gap	Variable gap	n.a.	Fixed gap	Variable gap	Apple II	Apple II	Apple II	?	Var. gap & Apple	n.a.
Injector	S-band RF gun	S-band RF gun	L-band RF gun	Pulsed Diode	L-band RF gun	L-band RF gun	S-band RF gun	S-band RF gun	S-band RF gun	S-band RF gun	VHF c.w. RF Gun	?
Cathode	Cu	Cu	Cs ₂ Te	CeB ₆ (thermionic)	Cs ₂ Te	Cs ₂ Te	Cu	Cu	Cu	Cu	K ₂ CsSb	?
Main linac technology	n.c. Pulsed	n.c. pulsed	s.c. pulsed	n.c. pulsed	s.c. pulsed	s.c. pulsed	n.c. pulsed	n.c. pulsed	n.c. pulsed	n.c. pulsed	S.C. C.W.	n.c. pulsed
RF frequency	S-band	S-band	L-band	C-band	L-band	L-band	S-band	C-band	S-band	C-band	L-band	S -band
RF Rep. rate	120 Hz	120 Hz	10 Hz	60 Hz	10 Hz	10 Hz	10-50 Hz	100 Hz	60 Hz	60 Hz	n.a.	60 Hz
FEL pulses/RF pulse	1	1	2700	1	2700	2700	1	2	1	1	1 MHz c.w.	100
max. bunch charge	0.25 nC	0.25 nC	1 nC	0.2 nC	1 nC	1 nC	0.5 nC	0.2 nC	0.2 nC	0.2 nC	0.3nC	0.1 nC
max. electron energy	13.6 GeV	14 GeV	17.5 GeV	8 GeV	1.2 GeV	1.2 GeV	1.5 GeV	5.8 GeV	10 GeV	6.4 GeV	2.4 GeV	12 GeV
No. RF stations	81	81	29	69	5	5	15	34	49	?	?	?
Approx. facility length	1.7km	1.7km	3.4 km	0.8km	0.32 km	0.32 km	0.5 km	0.7km	1.1 km	0.6 km	0.8 km	1.0 km
Start operation	2009	2017	2015	2011	2005	2013	2010	2016	2015	2019	2023	?

No. X-ray FELs worldwide

(undulator lines with λ_{photon} <50Å)





courtesy Florian Grüner

X-ray FEL's Long Term Future



Some benefits from bandwidth and collimation properties of diffraction limited coherent X-ray beams

- Energy problem solved (see left)
- World communication network backbone based on 100 beam XFEL links with 10 Zettabyte/s
- Near light velocity space probe powered by XFEL beam from ground station on their way to Sirius
- Novel "Black&Decker iFEL" drill, cut, welding tool revolutionizes quality of do-it-yourself products

Thank you for your attention!

Thanks to the IPAC organizers for choosing this great place!