



### **Accelerators: The Final Frontier?**

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## Outline

- Future Accelerators for particle physics
  - What is needed & why
  - The Large Hadron Collider (LHC)
  - The Linear Collider (LC)
  - The Muon Collider (MC)
  - The Neutrino Factory (NF)
- Other scientific applications
  - Light sources, Spallation sources ...
- Other Future Accelerators
  - Laser-Plasma accelerators
- Other applications
  - Accelerators in Medicine
- Summary

No time



#### **Accelerators for particle physics**

#### What is needed, and why

2 routes to new knowledge about the fundamental structure of the matter

#### **High Energy Frontier**

New phenomena (new particles) created when the "usable" energy > mc<sup>2</sup> [×2]

#### **High Precision Frontier**

Known phenomena studied with high precision *may* show inconsistencies with theory



### **The Standard Model**

$$\begin{split} \mathcal{L} &= -\frac{1}{4} F^a_{\mu\nu} F^{a\mu\nu} + i\bar{\psi}D\psi \\ &+ \psi_i \lambda_{ij} \psi_j h + h.c. \\ &+ |D_\mu h|^2 - V(h) \\ &+ \frac{1}{M} L_i \lambda^{\nu}_{ij} L_j h^2 \text{ or } L_i \lambda^{\nu}_{ij} N_j \end{split}$$

The gauge sector (1)

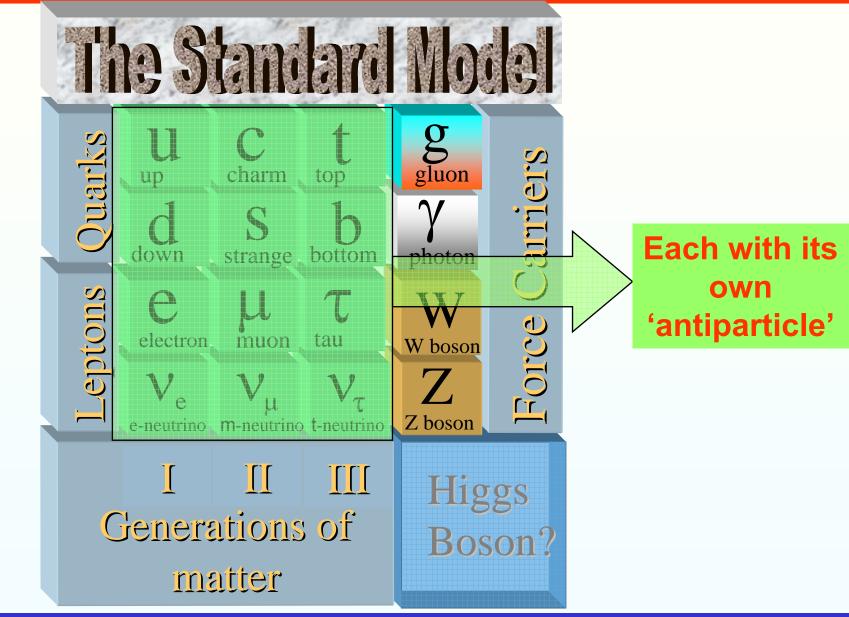
The flavor sector (2)

The EWSB sector (3)

The v-mass sector (4)



## **Particles** and **Forces**





## **The Standard Model**

#### The Standard Model Effective Lagrangean

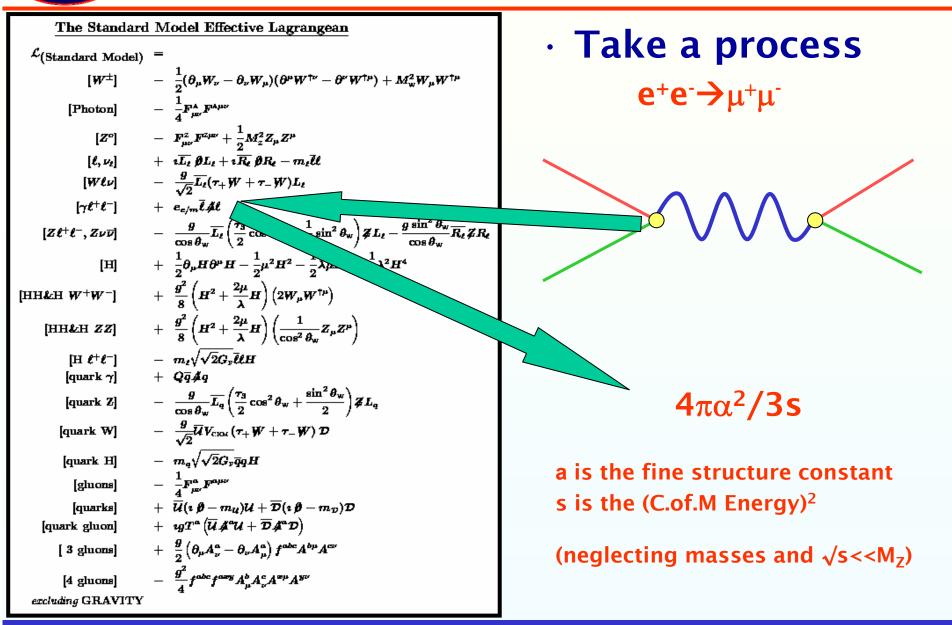
$\mathcal{L}_{(Standard Model)}$	=	
$[W^{\pm}]$	_	$rac{1}{2}(\partial_{\mu}W_{ u}-\partial_{ u}W_{\mu})(\partial^{\mu}W^{\dagger u}-\partial^{ u}W^{\dagger\mu})+M_{w}^{2}W_{\mu}W^{\dagger\mu}$
[Photon]	_	
$[Z^{o}]$	_	$F^x_{\mu u}F^{z\mu u}+rac{1}{2}M_x^2Z_\mu Z^\mu$
$[\ell,  u_{\ell}]$	+	$i\overline{L_{\ell}} \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ $
$[W\ell\nu]$	_	$rac{g}{\sqrt{2}}\overline{L_\ell}( au_+W+ auW)L_\ell$
[ <b>γ</b> ℓ <sup>+</sup> ℓ <sup>−</sup> ]		$e_{c/m}\bar{\ell}\mathcal{A}\ell$
$[Z\ell^+\ell^-, Z u\overline{ u}]$	_	$\frac{g}{\cos\theta_{w}}\overline{L_{\ell}}\left(\frac{\tau_{3}}{2}\cos^{2}\theta_{w}+\frac{1}{2}\sin^{2}\theta_{w}\right)\not\not\in L_{\ell}-\frac{g\sin^{2}\theta_{w}}{\cos\theta_{w}}\overline{R_{\ell}}\not\in R_{\ell}$
[H]	+	$\frac{1}{2}\partial_{\mu}H\partial^{\mu}H - \frac{1}{2}\mu^{2}H^{2} - \frac{1}{2}\lambda\mu H^{3} - \frac{1}{8}\lambda^{2}H^{4}$
[HH&H W <sup>+</sup> W <sup>-</sup> ]	+	$rac{g^2}{8} \left( H^2 + rac{2\mu}{\lambda} H  ight) \left( 2 W_\mu W^{\dagger \mu}  ight)$
[HH&H ZZ]	+	$rac{g^2}{8}igg(H^2+rac{2\mu}{\lambda}Higg)igg(rac{1}{\cos^2 heta_w}Z_\mu Z^\muigg)$
[H l+l-]		$m_t \sqrt{\sqrt{2}G_y} \tilde{\ell} \ell H$ The Higgs Sector
[quark 7]	+	Qq#q
[quark Z]	_	$\frac{\boldsymbol{g}}{\cos\boldsymbol{\theta}_{\mathbf{w}}}\overline{L_{q}}\left(\frac{\tau_{3}}{2}\cos^{2}\boldsymbol{\theta}_{\mathbf{w}}+\frac{\sin^{2}\boldsymbol{\theta}_{\mathbf{w}}}{2}\right)\boldsymbol{\not} \!$
[quark W]	—	$rac{g}{\sqrt{2}}\overline{\mathcal{U}}V_{ ext{CEOA}}\left( au_{+}oldsymbol{W}+ au_{-}oldsymbol{W} ight)\mathcal{D}$
[quark H]		$m_q \sqrt{\sqrt{2}G_y} \overline{q} q H$
[gluons]	_	$\frac{1}{-F^{\alpha}}F^{\alpha}F^{\alpha}F^{\alpha}$
[quarks]	+	$\overline{\overline{\mathcal{U}}}(\imath \ \boldsymbol{\theta} - m_{\mathcal{U}})\mathcal{U} + \overline{\mathcal{D}}(\imath \ \boldsymbol{\theta} - m_{\mathcal{D}})\mathcal{D}$
[quark gluon]	+	$\iota g T^{lpha} \left( \overline{\mathcal{U}}  {{}\!$
[ 3 gluons]	+	$rac{g}{2} \left(  heta_{\mu} A^a_{ u} -  heta_{ u} A^a_{\mu}  ight) f^{abc} A^{b\mu} A^{c u}$
[4 gluons]	_	$rac{g^2}{4} f^{abc} f^{aary} A^b_\mu A^c_ u A^{\mu\mu} A^{\mu u}$
excluding GRAVITY		-

#### **The Parameters**

- · 6 quark masses
  - m<sub>u</sub>, m<sub>c</sub>, m<sub>t</sub>
  - $m_{d_1} m_{s_1} m_{b_2}$
  - 3 lepton masses
    - $\mathbf{m}_{\mathbf{e},} \mathbf{m}_{\mu,} \mathbf{m}_{\tau}$
- 2 vector boson masses

- 1 Higgs mass
  - M<sub>h</sub>
- · 3 coupling constants
  - $\mathbf{G}_{\mathbf{F},\alpha,\alpha_{s}}$
- · 3 quark mixing angles
  - $\theta_{12,}\theta_{23,}\theta_{13}$
- 1 quark phase
  - δ

# The Standard Model in action



JAI



#### How good is the Standard Model?

The Standard Model Effective Lagran

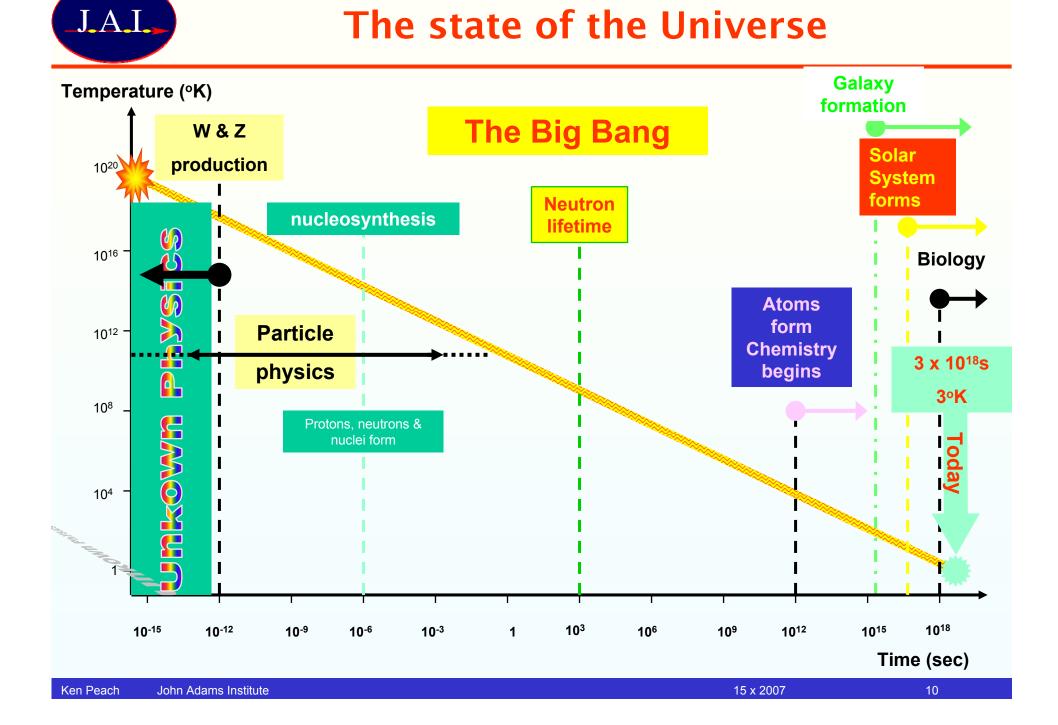
		¥				1410
$\mathcal{L}_{(\text{Standard Model})}$	=				(5)	
[ <i>W</i> <sup>±</sup> ]	_	$-rac{1}{2}(\partial_{\mu}W_{ u}-\partial_{ u}W_{\mu})(\partial^{\mu}W^{\dagger u}-\partial^{ u}W^{\dagger})$	$ ^{\mu} angle + M_{w}^{2}W_{\mu}W^{\dagger\mu}$		$\Delta \alpha^{(5)}_{had}(m_Z)$	0.02
[Photon]	_	1 - F <sup>A</sup> , F <sup></sup>			m <sub>z</sub> [GeV]	91.1
[ <b>Z</b> °]	_	$F^2_{\mu u}F^{Z\mu u}+rac{1}{2}M_z^2Z_\mu Z^\mu$			Γ <sub>z</sub> [GeV]	2.4
$[\ell,  u_t]$	+	$i\overline{L_{\ell}} \ \theta L_{\ell} + i\overline{R_{\ell}} \ \theta R_{\ell} - m_{\ell}\overline{\ell}\ell$			$\sigma_{\sf had}^0$ [nb]	41
$[W\ell  u]$	_	$rac{g}{\sqrt{2}}\overline{L_\ell}( au_+W+ auW)L_\ell$			R	20
$[\gamma \ell^+ \ell^-]$		$e_{e/m}\bar{l}\mathcal{A}l$	_		A <sup>0,I</sup> <sub>fb</sub>	0.01
$[Z\ell^+\ell^-, Z u\overline{ u}]$	_	$\frac{g}{\cos\theta_{\rm w}}\overline{L_{\ell}}\left(\frac{\tau_{\rm s}}{2}\cos^2\theta_{\rm w}+\frac{1}{2}\sin^2\theta_{\rm w}\right)\ddot{q}$	$dL_{\ell} - \frac{g\sin^2\theta_w}{\cos\theta_w}\overline{R_{\ell}}ZR_{\ell}$		$A_{I}(P_{\tau})$	0.1
[H]	+	$\frac{1}{2}\partial_{\mu}H\partial^{\mu}H-\frac{1}{2}\mu^{2}H^{2}-\frac{1}{2}\lambda\mu H^{3}-$	$\frac{1}{8}\lambda^2 H^4$		R <sub>b</sub>	0.21
[HH&H W <sup>+</sup> W <sup>-</sup> ]	+	$rac{m{g}^2}{8}igg(m{H}^2+rac{2\mu}{\lambda}m{H}igg)igg(2W_{\mu}W^{\dagger\mu}igg)$			R	0.1
[HH&H ZZ]	+	$rac{g^2}{8}\left(H^2+rac{2\mu}{\lambda}H ight)\left(rac{1}{\cos^2 heta_{ m w}}Z_{\mu}Z^{\mu} ight)$			A <sup>0,b</sup>	0.0
[H ℓ <sup>+</sup> ℓ <sup>−</sup> ]		$m_t \sqrt{\sqrt{2}G_F} \bar{\ell} \ell H$			A <sup>ID</sup> <sub>fb</sub>	0.0
$[{\rm quark}\; \boldsymbol{\gamma}]$		Qq4q			A <sub>b</sub>	0
[quark Z]	-	$-\frac{g}{\cos\theta_{\rm w}}\overline{L_q}\left(\frac{\tau_{\rm 3}}{2}\cos^2\theta_{\rm w}+\frac{\sin^2\theta_{\rm w}}{2}\right)\not\not\in I$	Lq		A <sub>c</sub>	0
[quark W]	_	$rac{g}{\sqrt{2}}\overline{\mathcal{U}}V_{ ext{ ext{cross}}}\left( au_+ oldsymbol{W} +  auoldsymbol{W} ight) \mathcal{D}$	19 magaurar	nonto	A <sub>I</sub> (SLD)	0.1
[quark H]	_	$m_q \sqrt{\sqrt{2}G_r} \overline{q} q H$	18 measurer			
[gluons]	_	1 - F <sup>a</sup> , F <sup>aµ</sup>	5 free param	eters	$sin^2 \theta_{eff}^{lept}(Q_{fb})$	0.2
[quarks]		$\frac{4}{\overline{\mathcal{U}}}(\imath  \boldsymbol{\theta} - m_{\mathcal{U}})\mathcal{U} + \overline{\mathcal{D}}(\imath  \boldsymbol{\theta} - m_{\mathcal{D}})\mathcal{D}$	<mark>χ²= 18.1/13 d</mark>	l.o.f.	m <sub>w</sub> [GeV]	80
[quark gluon]		${m u} {m T}^{a} \left( \overline{{m u}}  {m A}^{a} {m U} + \overline{{m D}}  {m A}^{a} {m D}  ight)$	<b>3 &gt; 1</b> σ		Г <sub>w</sub> [GeV]	2
[ 3 gluons]		${g\over 2} \Big(  heta_\mu A^a_ u -  heta_ u A^a_\mu \Big) f^{abc} A^{b\mu} A^{c  u}$	1 > 2σ		m <sub>t</sub> [GeV]	1
[4 gluons]	_	$rac{g^2}{4} f^{abc} f^{aay} A^b_\mu A^c_ u A^{a\mu} A^{\mu u}$				
excluding GRAVITY			Almost too	goou		



Ken Peach

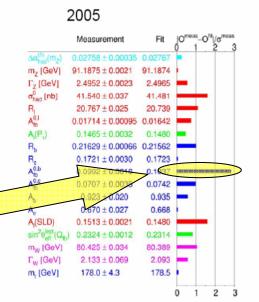


- The Standard Model is a very good *description* of the Universe at the particle scale (~2M<sub>W</sub>)
  - But does not explain many things
    - $\cdot$  Why so many particles?
    - $\cdot$  Why so many forces?
    - $\cdot$  What is mass?
      - Why do particles have the masses they have?
    - $\cdot$  How do neutrinos get mass?
      - Are neutrinos different? How do they fit in?
    - · What is Dark Matter? Dark Energy?
    - · Why is matter different from antimatter?
      - (Where did all the antimatter go?)
    - $\cdot$  Where does gravity fit in?



What do we need to make progress?

- · To reach higher energy
  - To take us beyond the LEP/Tevatron energy scale 2005
    - ·~100-200GeV
- To reach higher precision
  - $10 \times$  statistics would make this effect (if real)  $8\sigma$
- New types of accelerator
  - Neutrino factories
  - Beta beams
  - Muon colliders ...



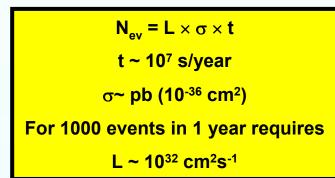
- We can accelerate *stable* particles
  - *"Stable"* means "with a lifetime long enough to capture and accelerate them
     in practice, > ~μ-second
- Hadrons
  - p, d, t, α, ... nuclei (up to Pb) & antiprotons
     · Hadrons contain "partons" (quarks, gluons...)
- · Leptons
  - $e^{\pm}$ ,  $\mu^{\pm}$ 
    - · Leptons are "point-like"
      - (at our present energy scales)



- The *Energy* must be sufficiently high that the process of interest can occur
- The Luminosity must be sufficiently high that a sufficient number of events are obtained in a "reasonable" time
  - (a few years)

For fixed target (esp. neutrino experiments) the equivalent parameter is

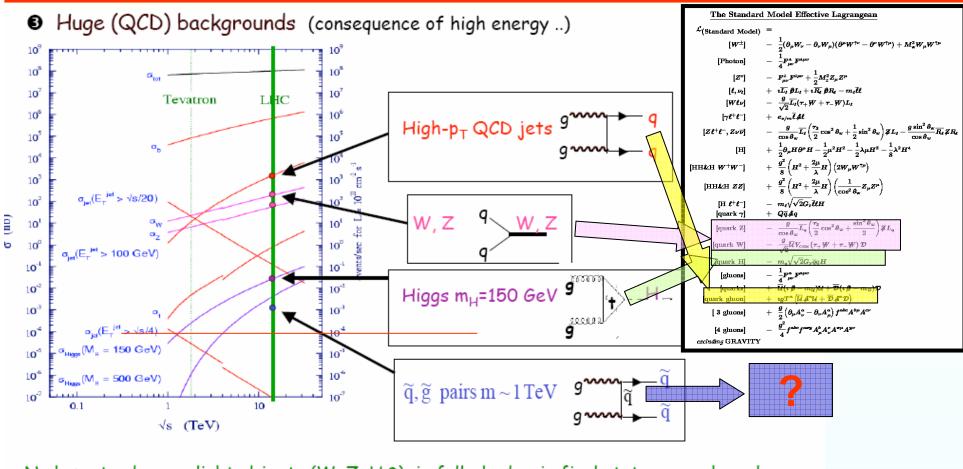
**Beam Power or Protons on Target (POT)** 



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#### An example – the LHC

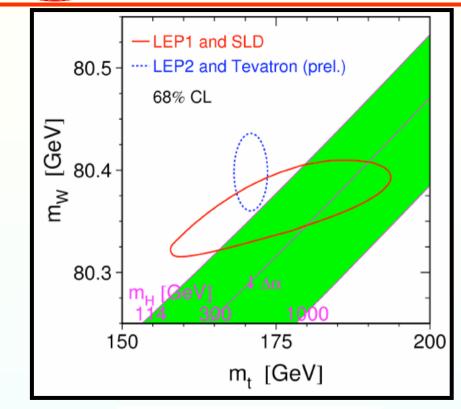


- No hope to observe light objects (W, Z, H ?) in fully-hadronic final states  $\rightarrow$  rely on I,  $\gamma$
- Fully-hadronic final states (e.g. q<sup>\*</sup> → qg) can be extracted from backgrounds only with hard O(100 GeV) p<sub>T</sub> cuts → works only for heavy objects
- Mass resolutions of ~ 1% (10%) needed for I,  $\gamma$  (jets) to extract tiny signals from backgrounds
- · Excellent particle identification: e.g. e/jet separation

Gianotti, LP05

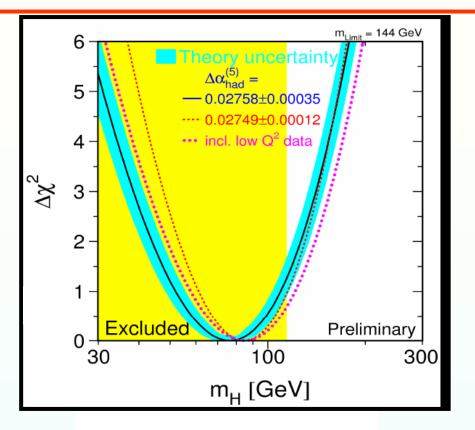


#### What are the big issues?



$$\begin{split} \mathsf{M}_{H} &= 76^{+33} \text{-}_{24} \text{ GeV} \\ \text{Incl. theory uncertainty:} \\ \mathsf{M}_{H} &< 144 \text{ GeV} (95\%\text{CL}) \end{split}$$

Direct search limit (LEP-2):  $M_H > 114 \text{ GeV} (95\% \text{CL})$ 



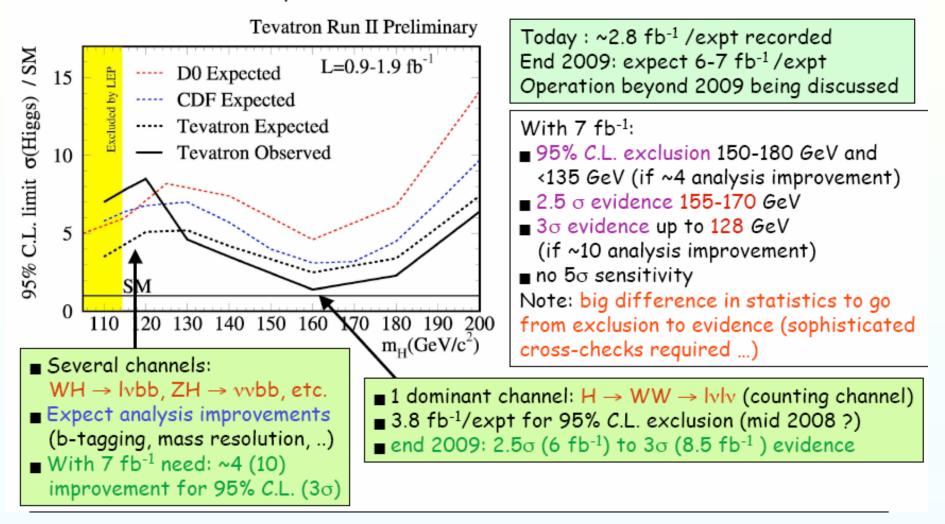
Probability M<sub>H</sub>>114 GeV: 15%

Renormalise probability for M<sub>H</sub>>114 GeV to 100%: M<sub>H</sub> < 182 GeV (95%CL)



#### **The Tevatron Search**

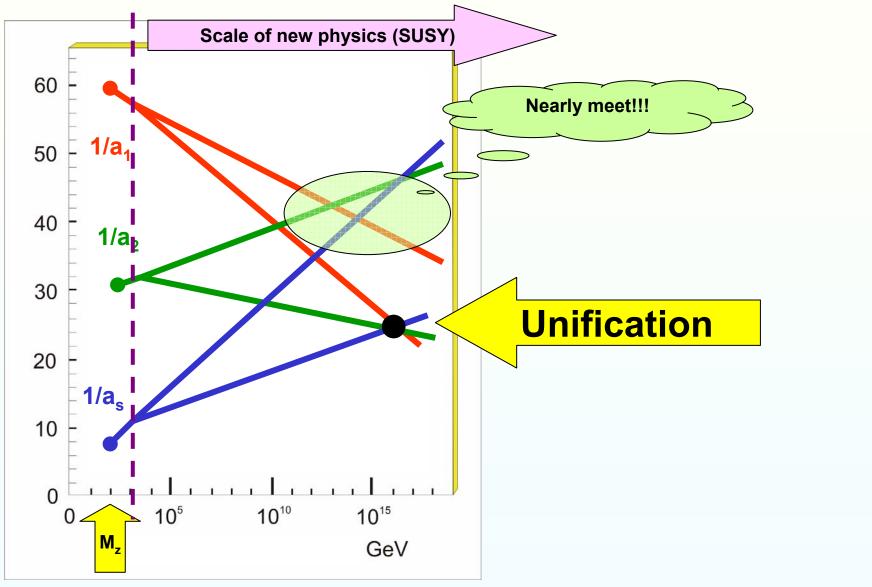
What about the "competition" with Tevatron?



#### After Gianotti, 07; Plot from Kim LP07

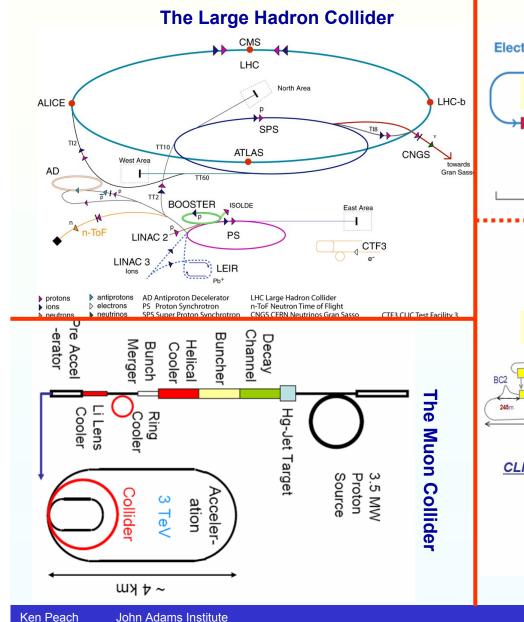


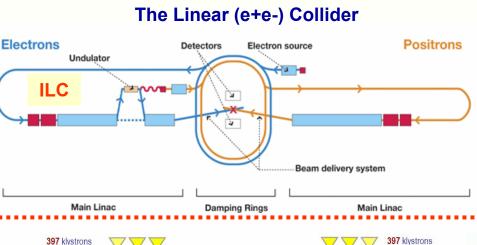
#### **Unification of the forces?**

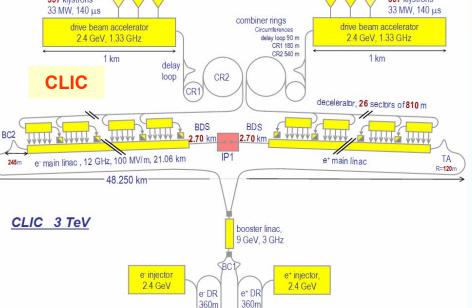




#### **Future Accelerators for particle physics**







15 x 2007

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#### **The Large Hadron Collider**

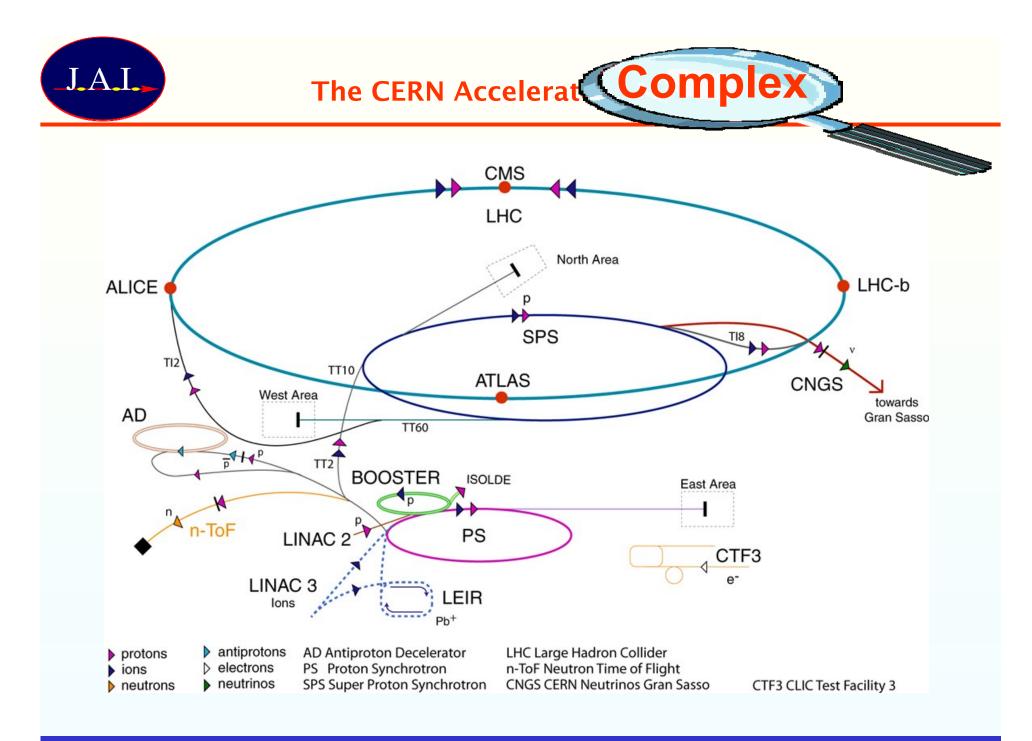


#### **The Large Hadron Collider**

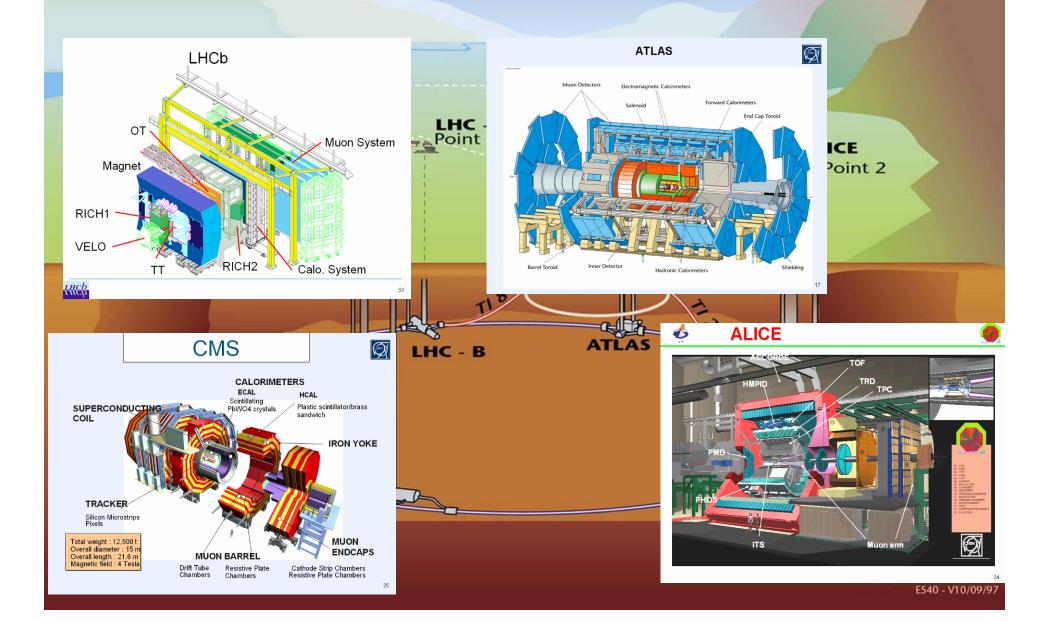
#### • The two main goals are:

- Find the Higgs
  - If it exists!!!
- Find the new physics
  - If it exists!!!
- $\cdot$  We know ~ the energy scales
  - $M_H < 250 \text{GeV}$ ;  $E_{NP} < 1 \text{TeV}$
- $\cdot$  pp collisions at high energy
  - Collision energy ~10% of total energy
    - Need a total collision energy >10TeV
  - Can calculate the cross-sections
     Need a luminosity > 10<sup>33</sup>cm<sup>2</sup>/s
- The Large Hadron Collider (LHC) @ CERN
  - $E \sim 14TeV$ ;  $L \sim 10^{34} cm^2/s$

The Standard	I M	odel Effective Lagrangean
$\mathcal{L}_{(\text{Standard Model})}$	=	
[ <i>W</i> <sup>±</sup> ]	_	$rac{1}{2}(  heta_{\mu}W_{ u} -  heta_{ u}W_{\mu})(  heta^{\mu}W^{\dagger  u} -  heta^{ u}W^{\dagger \mu}) + M^2_w W_{\mu}W^{\dagger \mu}$
[Photon]	-	$\frac{1}{4} \overline{F}^{A}_{\mu\nu} \overline{F}^{A\mu\nu'}$
[ <b>Z</b> °]	_	$F^z_{\mu u}F^{z\mu u}+rac{1}{2}M^2_zZ_\mu Z^\mu$
$[\ell, \nu_{\ell}]$	+	$i\overline{L_{\ell}} \not \! B L_{\ell} + i\overline{R_{\ell}} \not \! B R_{\ell} - m_{\ell} \vec{\ell} \vec{\ell}$
[Wev]	-	$\frac{g}{\sqrt{2}}\overline{L_{\ell}}(\tau_+W+\tauW)L_{\ell}$
[ <b>γ</b> ℓ <sup>+</sup> ℓ <sup>−</sup> ]		e <sub>c/m</sub> t#l
$[Z\ell^+\ell^-, Z u\overline{ u}]$	_	$\frac{g}{\cos\theta_{\rm w}}\overline{L_\ell}\left(\frac{\tau_3}{2}\cos^2\theta_{\rm w}+\frac{1}{2}\sin^2\theta_{\rm w}\right)\not\not\in L_\ell-\frac{g\sin^2\theta_{\rm w}}{\cos\theta_{\rm w}}\overline{R_\ell}\not\subset R_\ell$
[H]	+	$\frac{1}{2} \partial_{\mu} H \partial^{\mu} H - \frac{1}{2} \mu^2 H^2 - \frac{1}{2} \lambda \mu H^3 - \frac{1}{8} \lambda^2 H^4$
[HH&H W <sup>+</sup> W <sup>-</sup> ]	+	$rac{g^2}{8}\left(H^2+rac{2\mu}{\lambda}H ight)\left(2W_\mu W^{\dagger\mu} ight)$
[HH&H ZZ]	+	$\frac{g^2}{8}\left(\boldsymbol{H}^2+\frac{2\mu}{\lambda}\boldsymbol{H}\right)\left(\frac{1}{\cos^2\theta_{w}}\boldsymbol{Z}_{\mu}\boldsymbol{Z}^{\mu}\right)$
[H ℓ <sup>+</sup> ℓ <sup>-</sup> ]	_	$m_t \sqrt{\sqrt{2}G_F} \ell \ell H$
$[quark \gamma]$		Qq#q
[quark Z]	-	$\frac{\boldsymbol{g}}{\cos\boldsymbol{\theta}_{w}}\overline{L_{q}}\left(\frac{\boldsymbol{\tau}_{3}}{2}\cos^{2}\boldsymbol{\theta}_{w}+\frac{\sin^{2}\boldsymbol{\theta}_{w}}{2}\right)\boldsymbol{\not\!\!Z}L_{q}$
[quark W]	-	$rac{g}{\sqrt{2}}\overline{\mathcal{U}}V_{ ext{CXXX}}( au_+ oldsymbol{W} +  au oldsymbol{W})  \mathcal{D}$
[quark H]		$m_q \sqrt{\sqrt{2}G_y} \overline{q} q H$
[gluons]	_	$\frac{1}{4} F^{\alpha}_{\mu\nu} F^{\alpha\mu\nu}$
[quarks]	+	$\frac{d}{dl}(\imath \not p - m_{\mathcal{U}})\mathcal{U} + \overline{\mathcal{D}}(\imath \not p - m_{\mathcal{D}})\mathcal{D}$
[quark gluon]	+	$ugT^{lpha}\left(\overline{\mathcal{U}}\not\hspace{-0.15cm}\not\hspace{-0.15cm}\mathcal{A}^{lpha}\mathcal{U}+\overline{\mathcal{D}}\not\hspace{-0.15cm}\not\hspace{-0.15cm}\mathcal{A}^{lpha}\mathcal{D} ight)$
[ 3 gluons]	+	$rac{g}{2}\left(  heta_{\mu}A^{a}_{ u} -  heta_{ u}A^{a}_{\mu} ight) f^{abc}A^{b\mu}A^{c u}$
[4 gluons]	-	$rac{g^2}{4} f^{abc} f^{awy} A^b_\mu A^c_ u A^{w\mu} A^{y u}$
excluding GRAVITY		



#### **The Large Hadron Collider**

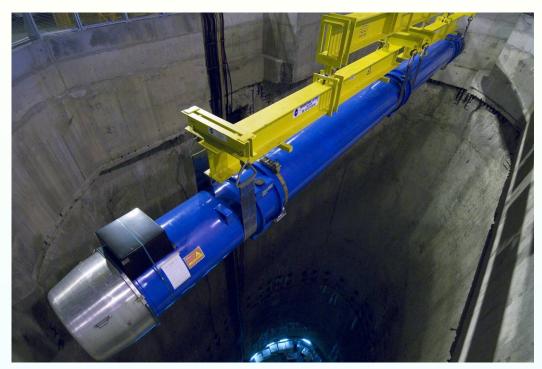




#### **The LHC installation**



## J.A.I., Descent of the last magnet, 26 April 2007

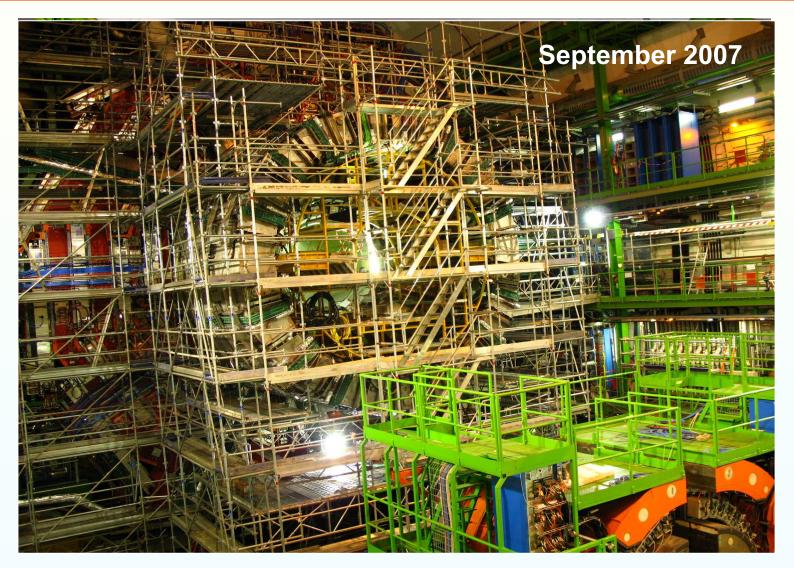


#### 30'000 km underground at 2 km/h!



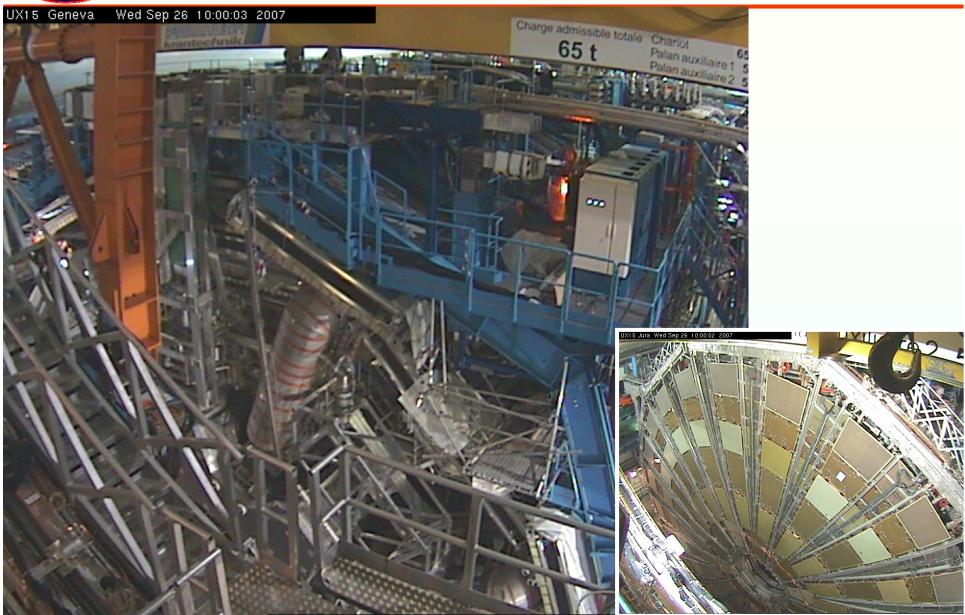








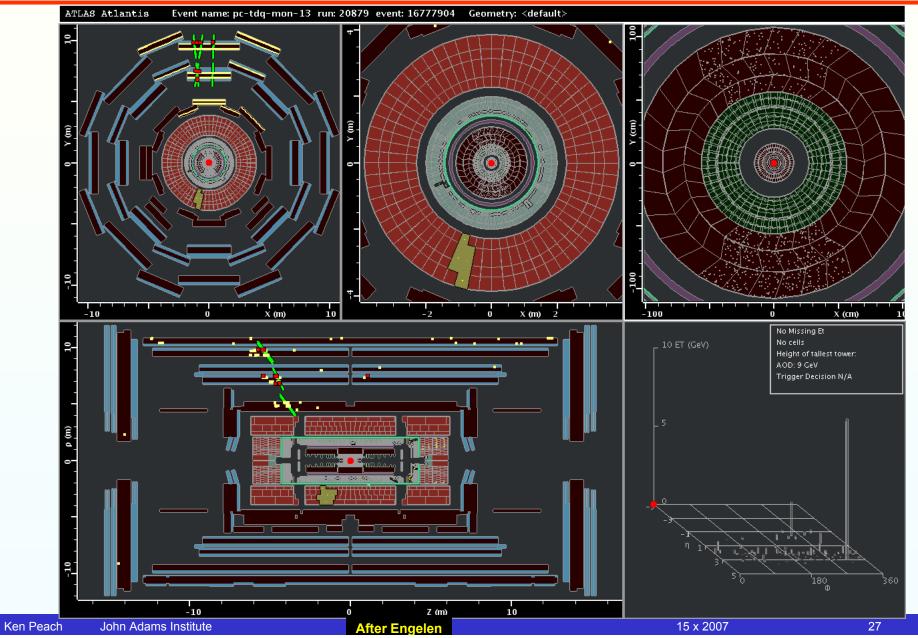




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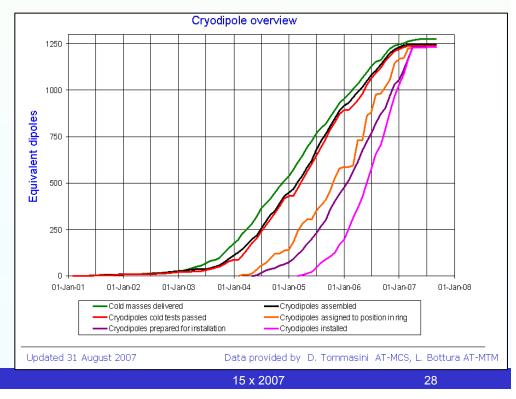
#### Already taking "data" (cosmics)



# J.A.I.

### LHC status

- · Machine installed, commissioning
- Experiments nearly installed, commissioning
- Due for completion in Spring 2008
- First collisions end summer 2008
- First results 2009
  - Higgs, SUSY
    - $\cdot$  or something else

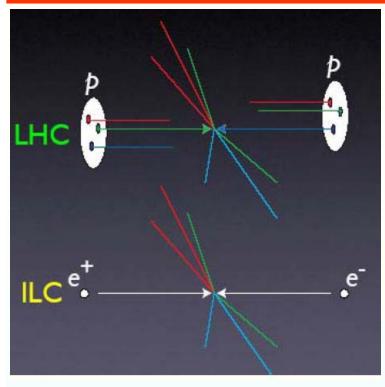




#### The Linear (e+e-) Collider

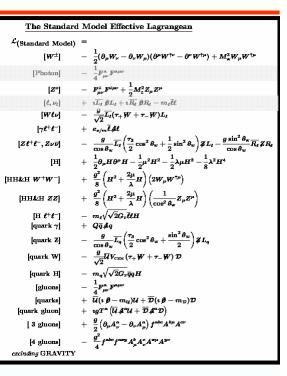


#### Why an e+e- collider?



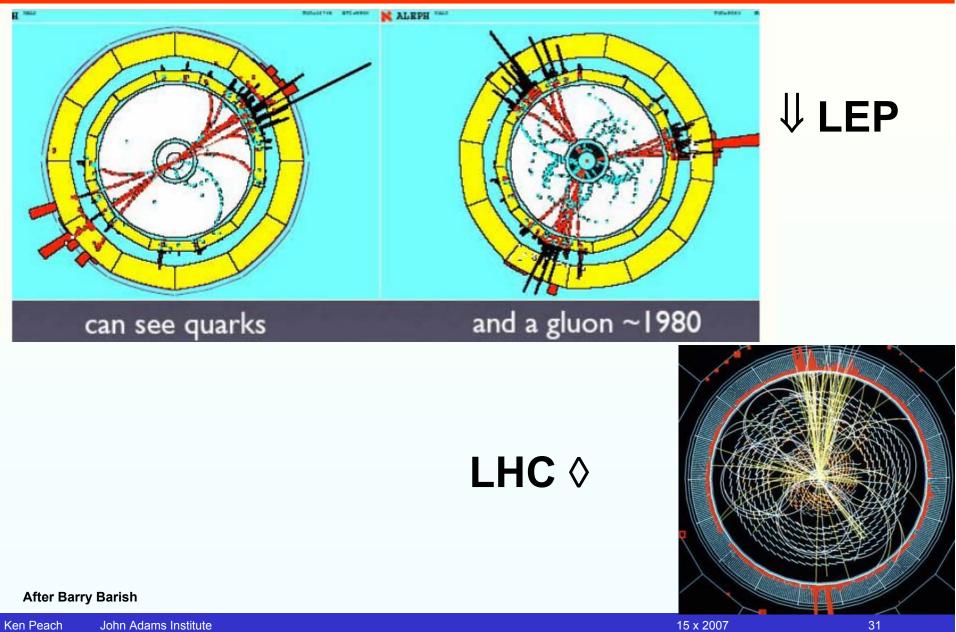
#### • elementary particles

- well-defined
  - energy,
  - angular momentum
- uses full COM energy
- produces particles democratically
- can mostly fully reconstruct events





#### Why an e+e- collider?





Why a *linear* e+e- collider?

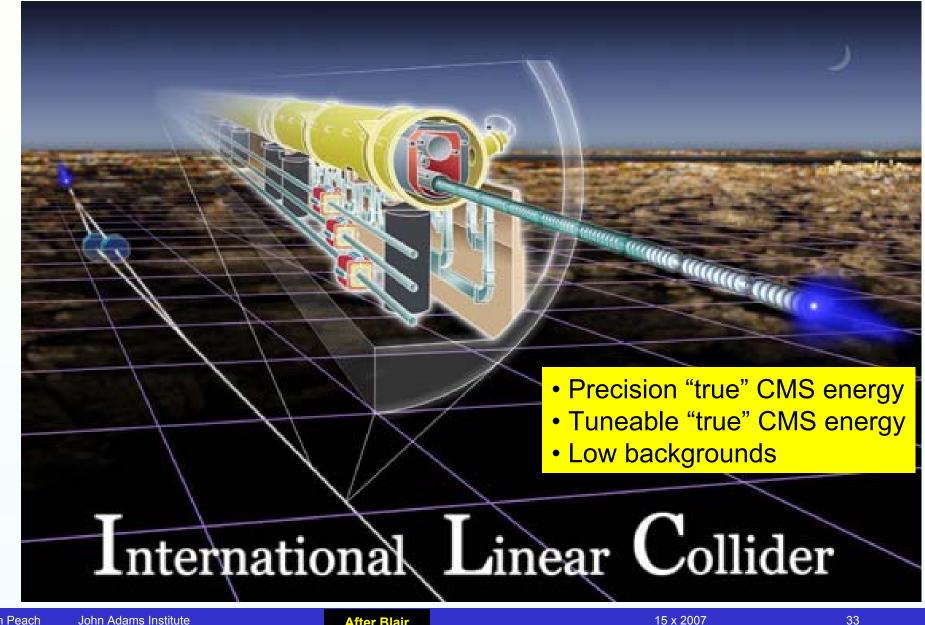
# Synchrotron Radiation!

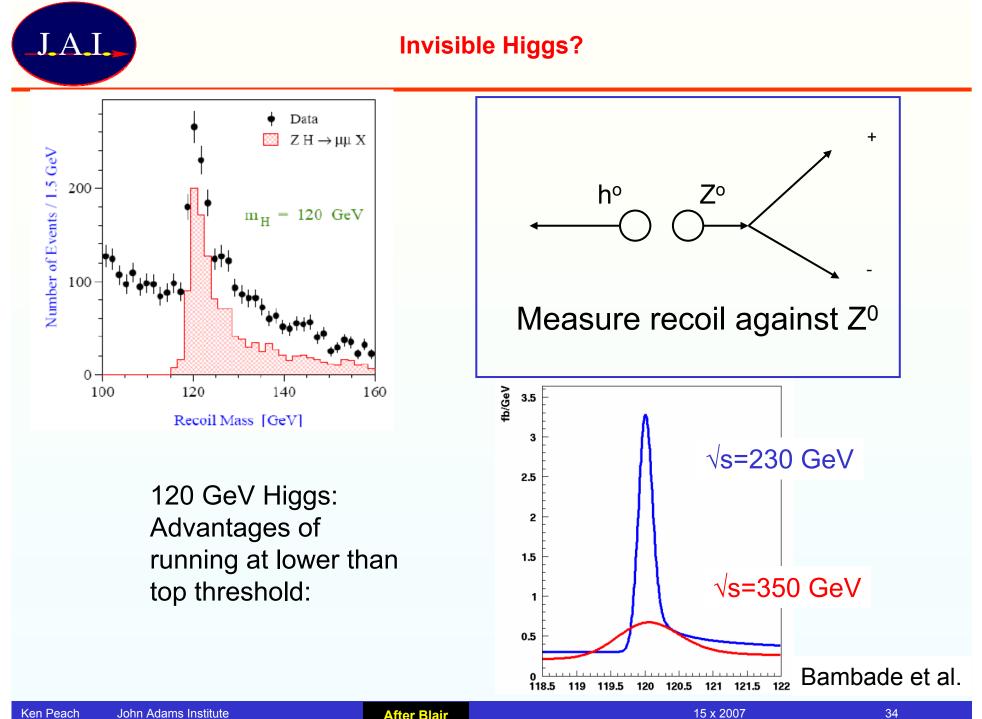
### or rather

# the lack of it in a linear machine



#### **Key ILC Properties**





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Ken Peac

#### **RDR vs ILC Physics Goals**

- E<sub>cm</sub> adjustable from 200 500 GeV
- Luminosity  $\rightarrow \int Ldt = 500 \text{ fb}^{-1}$  in 4 years
- Ability to scan between 200 and 500 GeV
- Energy stability and precision below 0.1%
- Electron polarization of at least 80%
- The machine must be upgradeable to 1 TeV

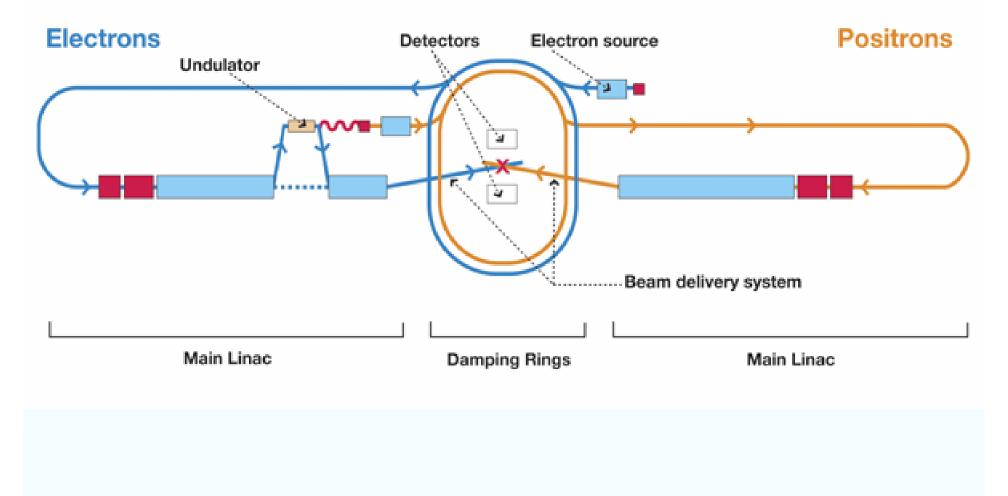
The RDR Design meets these "requirements," including the recent update and clarifications of the reconvened ILCSC Parameters group!

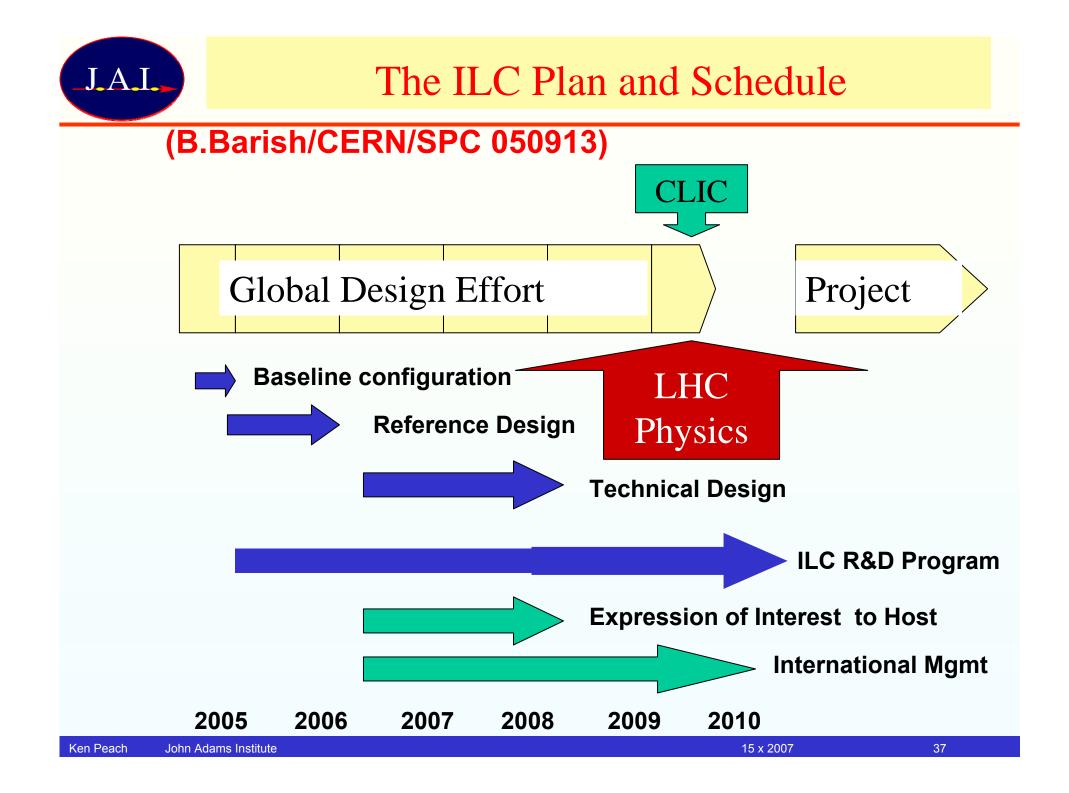
7-Feb-07 GDE/ACFA Closing Beijing	Global Design Ef	ffort		6	
B. Barish, Beijing 2007	,				
John Adame Instituto	After Diete		15 v 2007		



## **ILC Layout**

#### 500 GeV machine



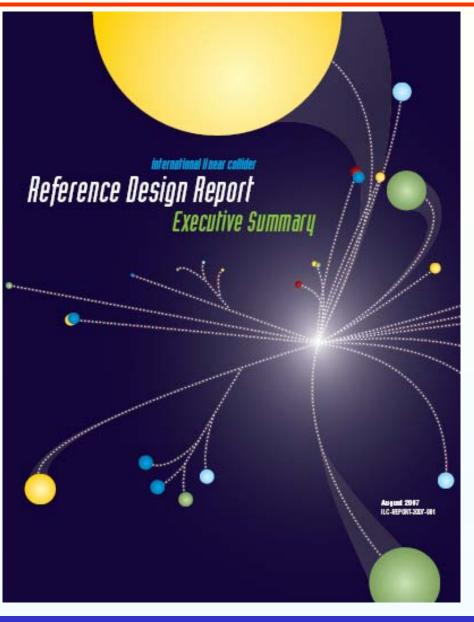






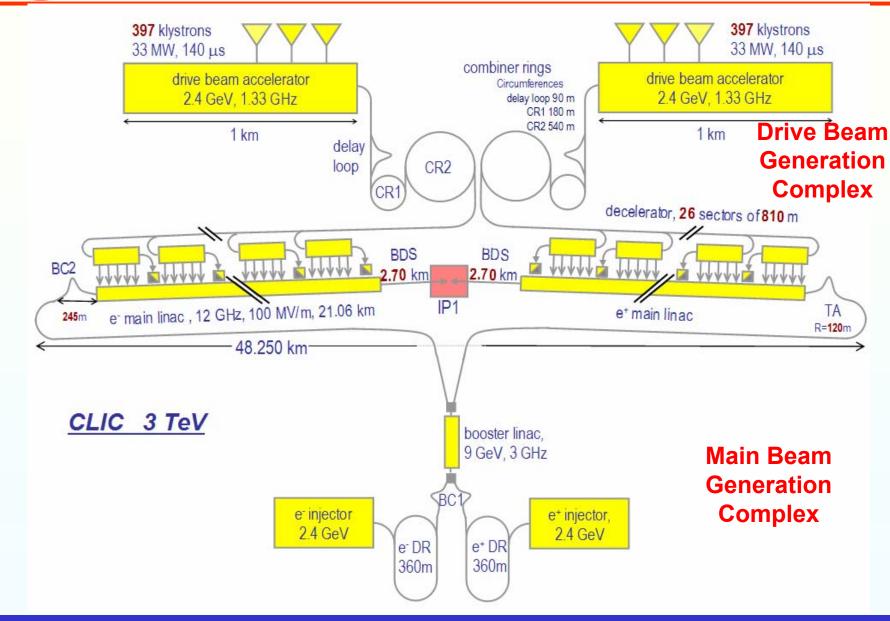
# Compact Linea

# Collider



Ken Peach John Adams Institute

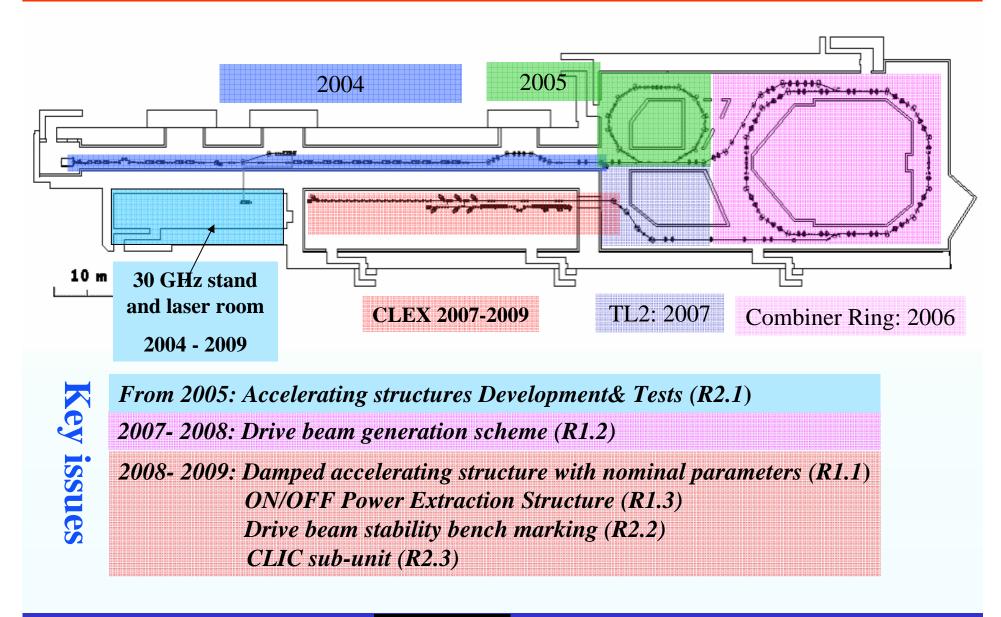
**CLIC – overall layout** 



JAI.

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### **CLIC Test Facility (CTF3)**



JAI



Main Linac RF frequency	30 GHz → 12 GHz
Accelerating field	150 MV/m → 100 MV/m
Overall length @ E <sub>CMS</sub> = 3 TeV	33.6 km → 48.2 km

- Substantial cost savings and performance improvements for 12 GHz / 100 MV/m indicated by parametric model (flat optimum in parameter range)
- Promising results already achieved with structures in test conditions close to LC requirements (low breakdown rate) but still to be demonstrated with long RF pulses and fully equipped structures with HOM damping.
- Realistic feasibility demonstration by 2010



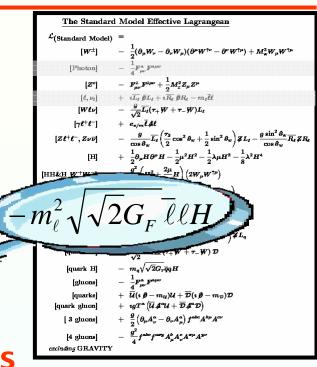
### **The Muon Collider**



#### A Muon Collider???



- 1. For some processes, muons are better than electrons
  - $\sigma$ (μ<sup>+</sup>μ<sup>-</sup>→H) is (m<sub>µ</sub>/m<sub>e</sub>)<sup>2</sup> ×  $\sigma$ (e<sup>+</sup>e<sup>-</sup>→H)
    - (40000 times larger)
- 2. If CLIC does *not* work, thin may be the only route to multi-TeV collisions under clean conditions
- Why not?
  - **Muon lifetime is only 2μsec!**
  - Need to produce, collect, cool and accelerate large numbers (>>10<sup>13</sup>) muons per second

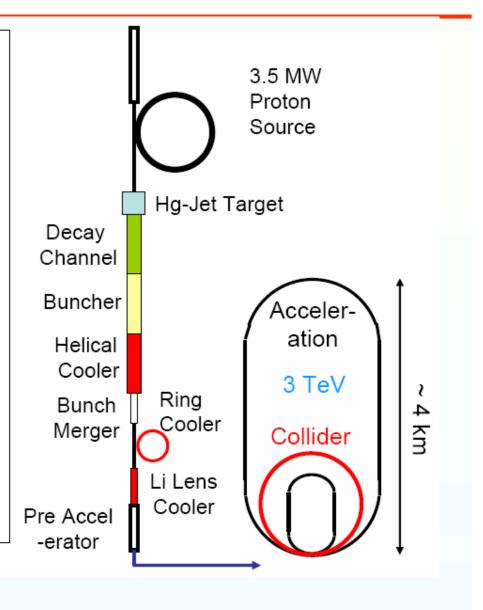


### Main Components of a Muon Collider

Proton Driver

JAL

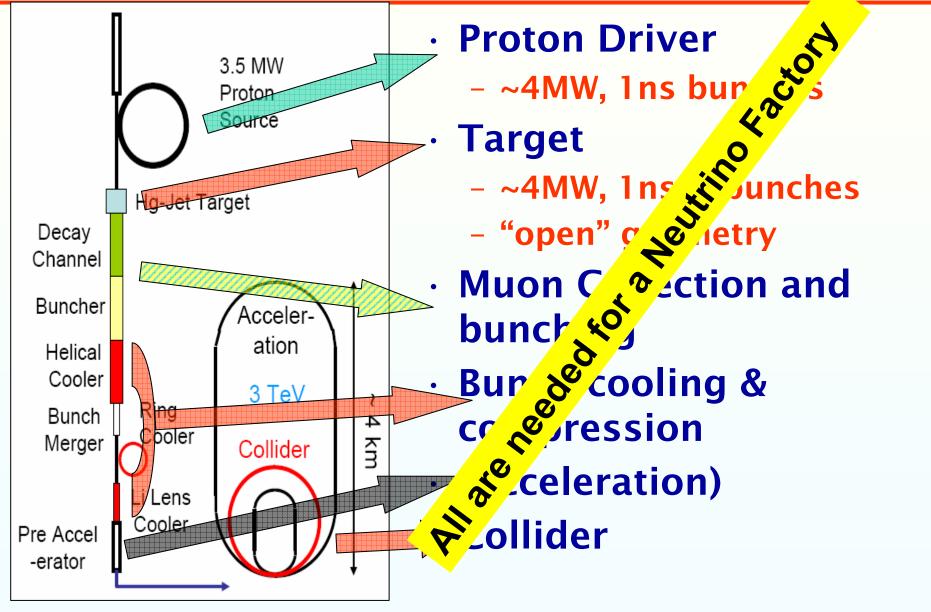
- primary beam on production target
- Target, Capture, and Decay
  - create π; decay into μ
- Bunching & Phase Rotation
  - reduce  $\Delta E$  of bunch
- Cooling
  - · reduce 6D emittance
- Acceleration
  - 130 MeV → up to 1.5 TeV
- Storage Ring
  - store for ~1000 turns
  - One IP



#### **Steve Geer**



#### **Challenges for the Muon Collider**





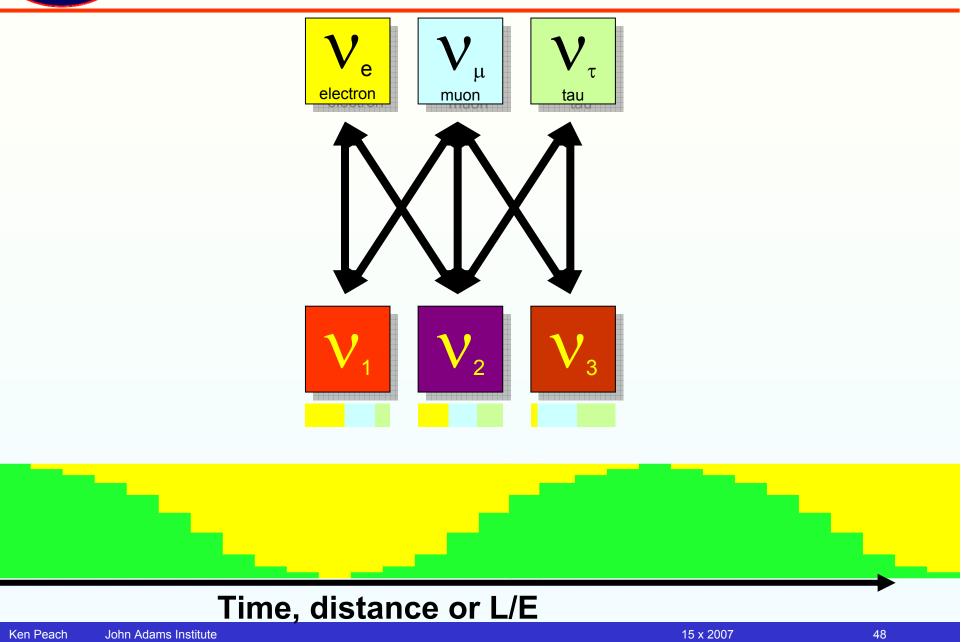
#### **Neutrino Factory**



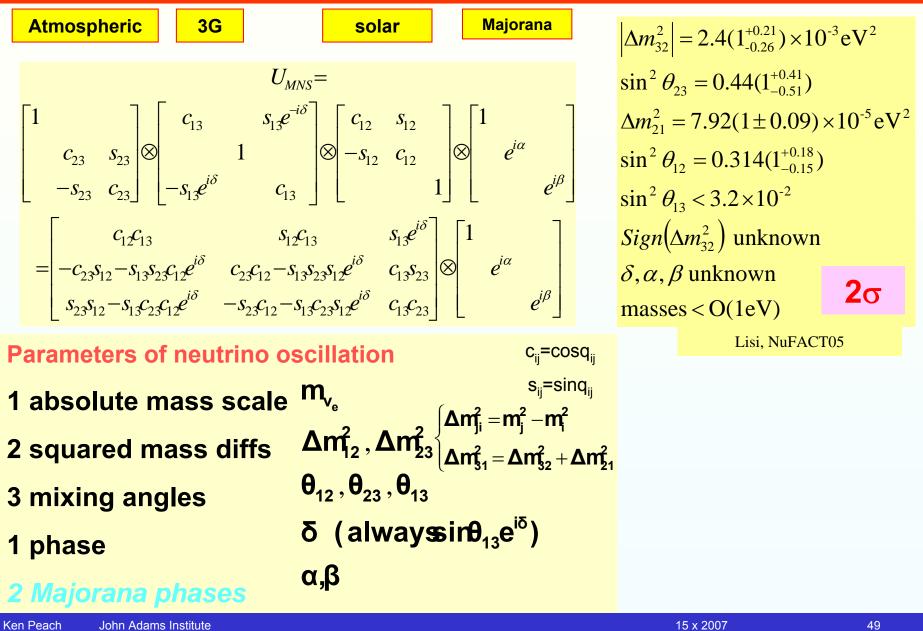
- 1950's and early 60's - Nature (and existence) of the neutrino · (Reines & Cowan, Lederman, Schwartz and Steinberger) • Late 1960s, 1970s, 1980s Structure of the nucleon •  $F_2$ ,  $xF_3$  etc New facilities Structure of the weak current allow old · Neutral currents,  $sin_2\theta_w$  etc physics to be Now, and future done much better - Nature of the neutrino Neutrino Mass and Neutrino Oscillations Standard Model assumption of massless neutrinos is wrong!
  - Note: difficult to add neutrino mass to SM a la Higgs
  - Lack of Charge  $\rightarrow$  additional mass-like (Majorana) terms



## **Neutrino Mixing**



### J.A.I. Neutrino Mixing





$$\begin{split} P(\nu_{\mu} \Rightarrow \nu_{e}) &= \\ & 4c_{13}^{2}s_{12}^{2} \left(c_{12}^{2}c_{23}^{2} - s_{12}^{2}s_{13}^{2}s_{23}^{2} - 2c_{12}c_{23}s_{12}s_{23}s_{13}\cos\delta\right) \sin^{2}\left(\frac{\Delta m_{21}^{2}L}{4E}\right) \\ & + 8c_{13}^{2}s_{12}s_{13}s_{23}\left(c_{12}c_{23}\cos\delta\right) - s_{12}s_{13}s_{23}\right) \cos\left(\frac{\Delta m_{32}^{2}L}{4E}\right) \sin\left(\frac{\Delta m_{21}^{2}L}{4E}\right) \\ & + 4c_{13}^{2}s_{13}^{2}s_{23}^{2}\sin^{2}\left(\frac{\Delta m_{13}^{2}L}{4E}\right)\left(1 + \left(1 - 2s_{13}^{2}\right)\frac{2a}{\Delta m_{31}^{2}}\right) - \nu_{\mu} \Rightarrow \overline{\nu}_{\mu} \Leftrightarrow a \Rightarrow -a \\ & - 8c_{13}^{2}c_{12}c_{23}s_{12}s_{13}s_{23}\sin\delta\sin\left(\frac{\Delta m_{32}^{2}L}{4E}\right)\sin\left(\frac{\Delta m_{31}^{2}L}{4E}\right)\sin\left(\frac{\Delta m_{31}^{2}L}{4E}\right)\sin\left(\frac{\Delta m_{31}^{2}L}{4E}\right) \\ & - 8c_{13}^{2}s_{13}^{2}s_{23}^{2}\cos\left(\frac{\Delta m_{32}^{2}L}{4E}\right)\sin\left(\frac{\Delta m_{31}^{2}L}{4E}\right)\sin\left(\frac{\Delta m_{21}^{2}L}{4E}\right)\left(1 - 2s_{13}^{2}\right)\frac{aL}{4E} \end{split}$$

Where is the electron density ; r is the density (g/cm<sup>3</sup>) ; E is the neutrino energy (GeV)

(Richter: hep-ph/0008222)

c<sub>ij</sub>=cosq<sub>ij</sub>, s<sub>ij</sub>=sinq<sub>ij</sub>



#### Neutrinos

- v<sub>e</sub> disappearance
- $v_e \rightarrow v_\mu$  appearance
- $v_e \rightarrow v_{\tau}$  appearance

 $v_{\mu}$ disappearance $v_{\mu} \rightarrow v_{e}$ appearance $v_{\mu} \rightarrow v_{\tau}$ appearance

... and the corresponding antineutrino interactions

Note: the beam requirements for these experiments are:

high intensity

known flux

known spectrum

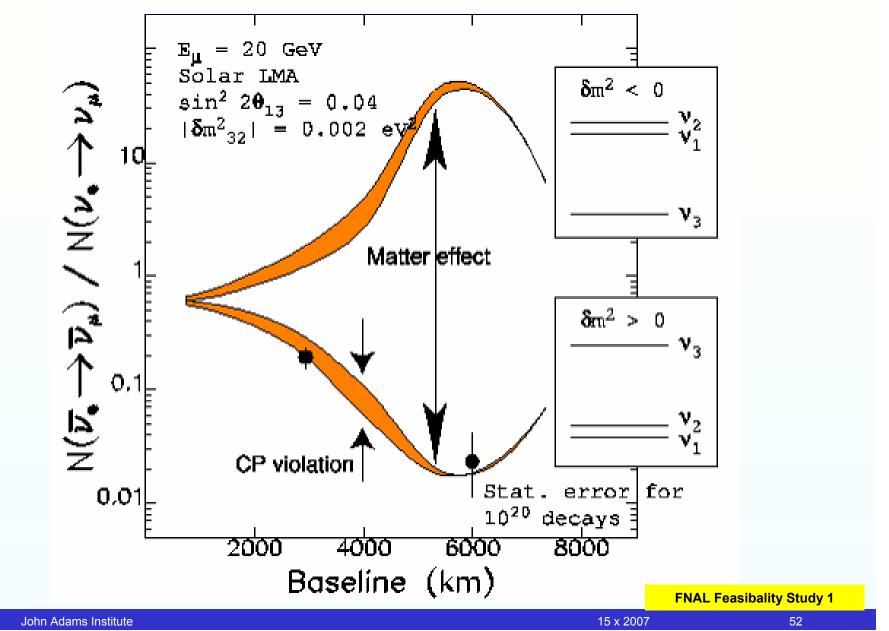
known composition

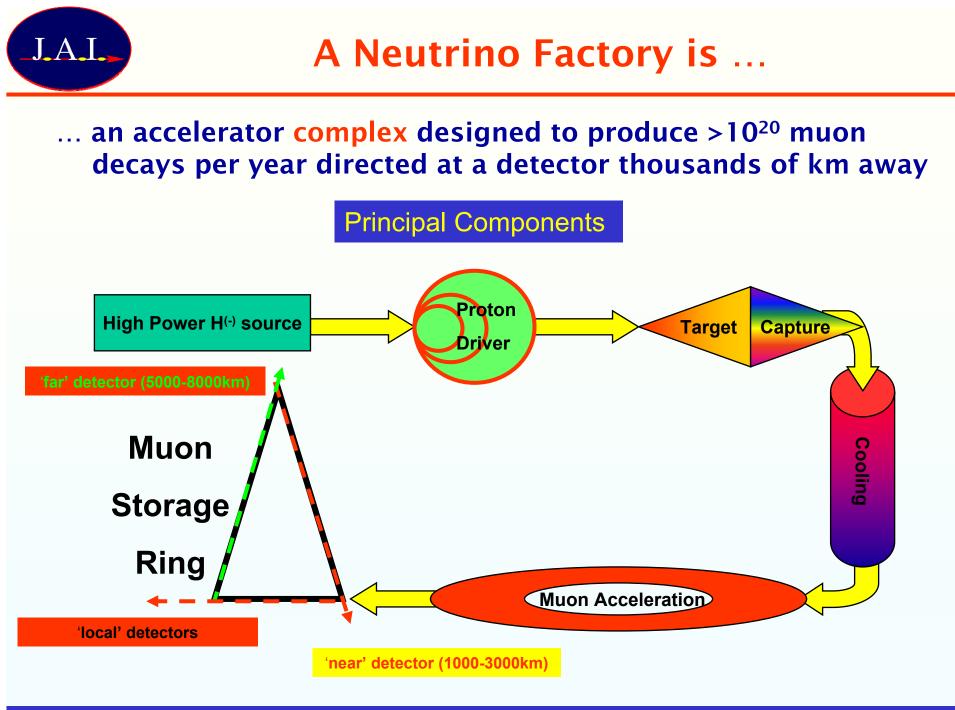
(preferably no background)



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#### **CP-violation**





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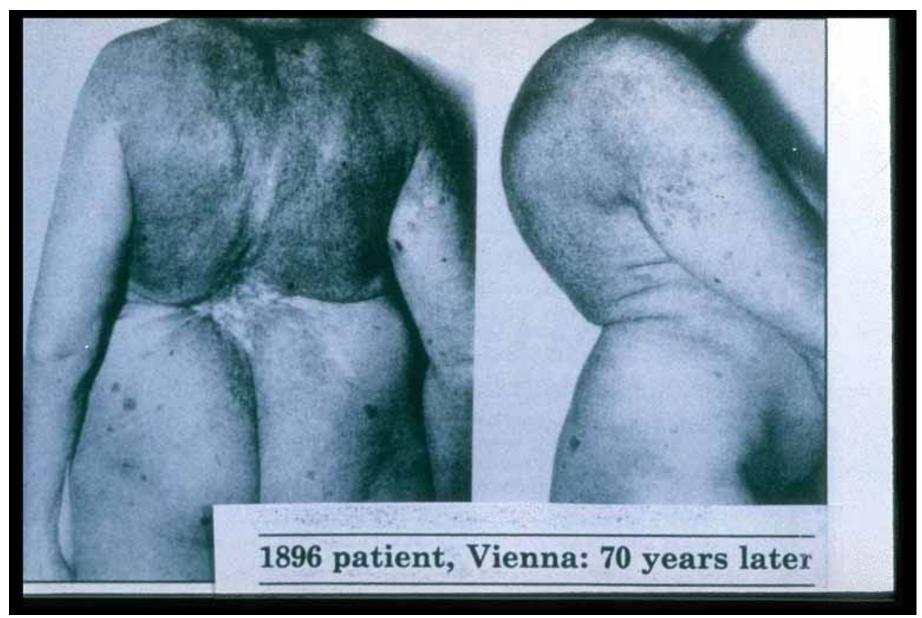
- Parameters
  - Need to know that  $\theta_{13}$  is not zero
    - · Other parameters well known to fix ( $E_{\mu}$ ,L)
- Technology
  - Proton driver
    - RCS or LINAC?
  - Proton energy?
     HARP, E910, MIPP
  - Target
    - · MW beam power
      - Mercury, solid, liquid-cooled, pellet, ...
  - Pion/muon collection and/or cooling
    - · Magnetic Horns or Solenoids?
    - · Phase Rotators, FFAG's, cooling?
  - RF and acceleration
    - · RLA's or FFAG's?
  - Muon Storage Ring
    - · Racetrack, triangular or bow-tie
    - · Conventional or FFAG?
- Other uses of high power protons & muons?



- Other scientific applications
  - Light sources, Spallation sources, FELs, ERLs ...
- Other Future Accelerators
  - Laser-Plasma accelerators
- Other applications
  - Accelerators in Medicine



X-ray therapy began within months of Roentgen's discovery



Slide from Gillies McKenna, Oxford

Conclusion



- Particle Accelerators have an exciting future
  - In particle physics
    - · LHC, LC, CLIC, NF, factories ...
  - In other sciences
    - $\cdot$  Light sources, FELs, spallation sources
  - In society
    - Medical accelerators (isotopes, hadrontherapy...)
- And they are *fun too*!