X-Ray Investigation on the Superconducting Source for lons (SuSI)

D. Neben*, J. Fogleman, B. Isherwood, D. Leitner, G. Machicoane, S. Renteria, J. Stetson and, L. Tobos

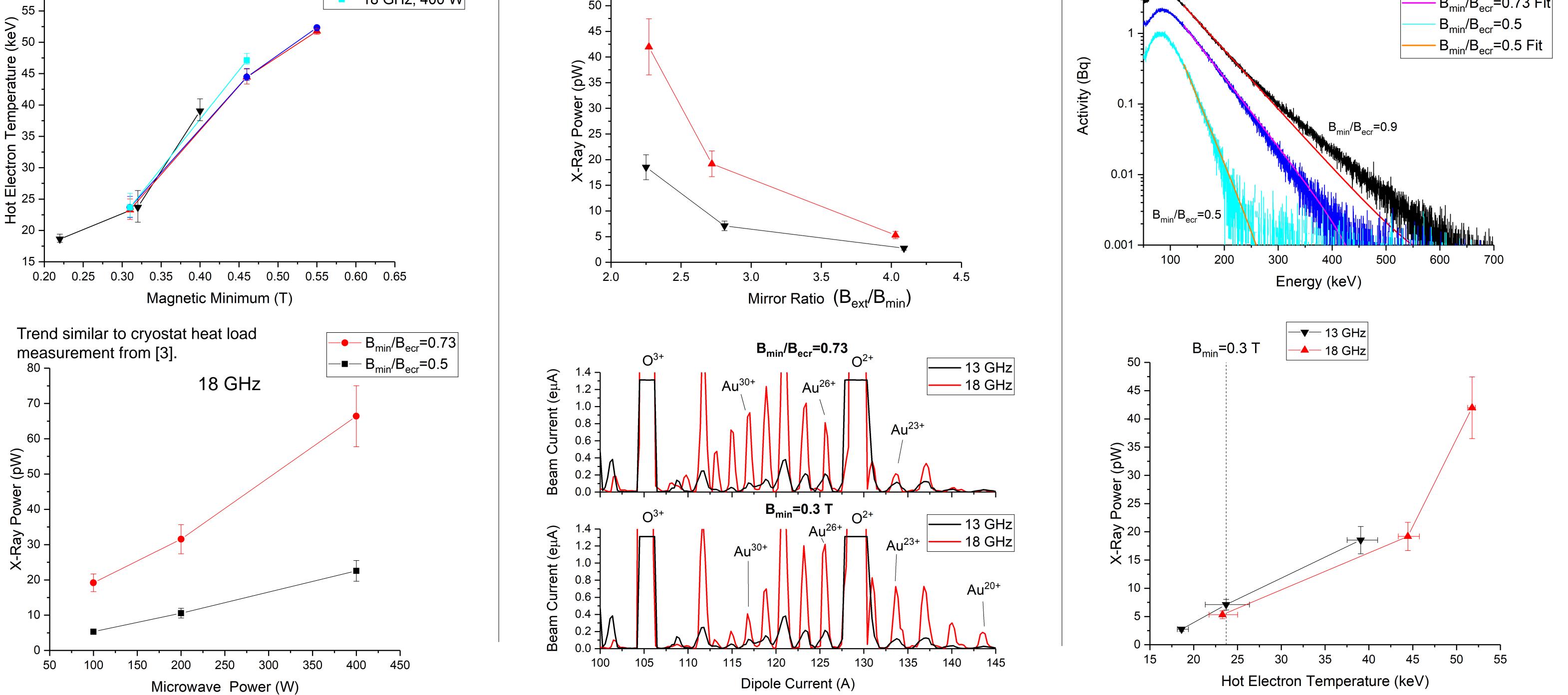
^{*}neben@nscl.msu.edu

Abstract



Heavy ion facilities such as the National Superconducting Cyclotron Laboratory (NSCL) often use ECR Ion Sources (ECRIS) for the production of highly charged ions to increase the efficiency of accelerating structures. Axial bremsstrahlung emission was studied on the Superconducting Source for Ions (SuSI) at the NSCL for 18 GHz and 13 GHz operation with oxygen. The hot electron temperatures were estimated from the bremsstrahlung high energy tail and seem to depend only on magnetic minimum in the same way as was found on VENUS [1], even in the case where 18 GHz and 13 GHz frequencies were compared for similarly sized ECR zones. Additionally, the time independent x-ray power increased at a significantly larger rate when operating the source in known regions of instability such as where the magnetic minimum approaches the ECR zone [2]. The results are discussed in the context of electron losses due to magnetic confinement.

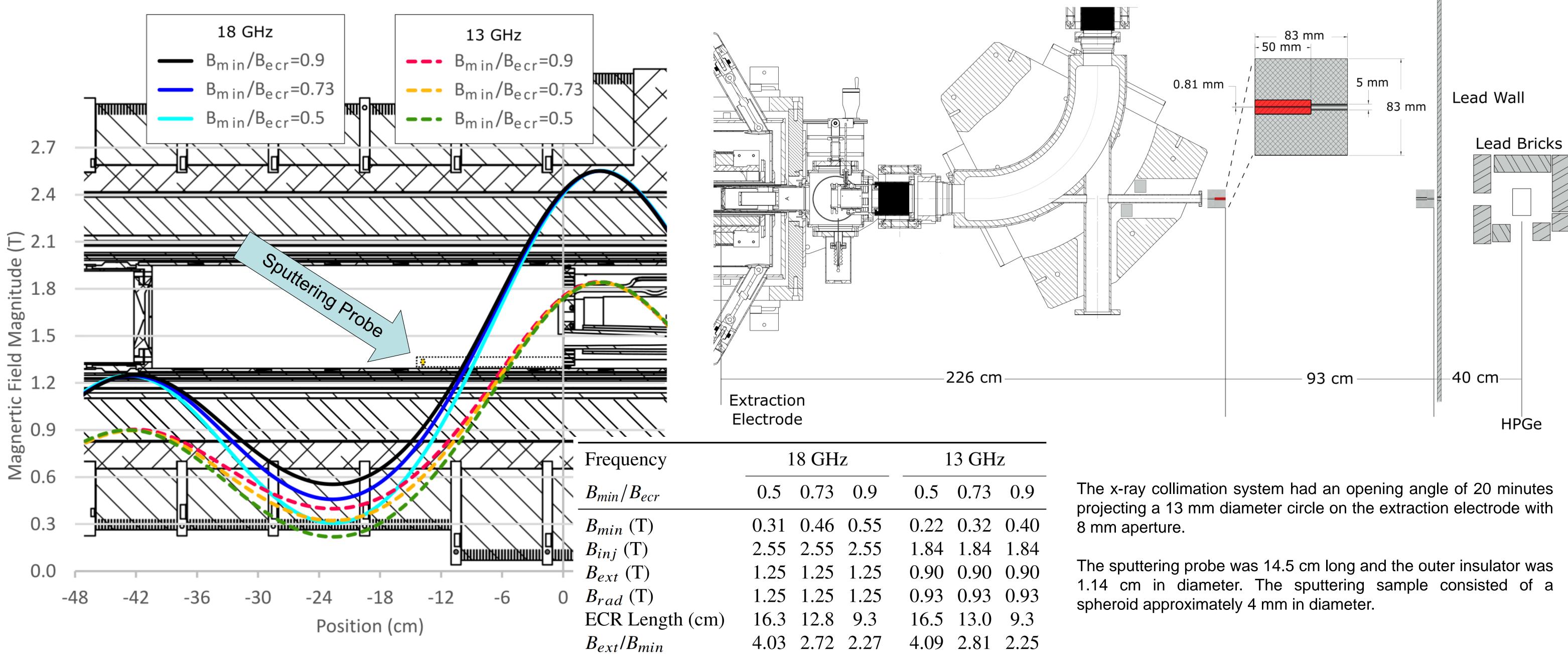
Trend similar to the findings from [1].		18 GHz operation produced more x-ray	100 W		Bremsstrahlung fitting performed between 125 keV	18 GHz	
	— ▼ — 13 GHz, 100 W	power than 13 GHz.			and 1440 keV.		B _{min} /B _{ecr} =0.9
	18 GHz, 100 W			— ▼ —13 GHz	10 ₃	100 W	—— B _{min} /B _{ecr} =0.9 Fit
⁶⁰	● 18 GHz, 200 W	55 J		▲ 18 GHz			—— B _{min} /B _{ecr} =0.73
1]					



Hot electron temperature increased linearly with B_{min} similarly to what was reported in [1]. X-ray power increased linearly with microwave power in a manner similar to how the VENUS cryostat heat load increased with microwave power [3].

The plasma density was lower for 13 GHz than at 18 GHz for constant microwave power, magnetic minimum, and ECR zone.

X-ray power doubled for a 17% increase in hot electron temperature, and could be the result of instabilities that appear for $B_{min}/B_{ecr} > 0.7$ [2].



REFERENCES:

[1] J. Benitez et. al., in Proc. ECRIS'16, 2016, paper MOCO04, pp. 23-29. [2] O. Tarvainen et al., Plas. Sourc. Sci. Technol., vol. 23, p. 025020, 2014. [3] D. Leitner et. al., Rev. Sci. Instrum., vol. 79, p. 033302, 2008.

ACKNOWLEDGEMENTS:

This work was supported by Michigan State University and the National Science Foundation: NSF Award Number PHY-1415462 The NSCL is funded in part by Michigan State University and the National Science Foundation.



