



12th International Workshop COOL'19

Budker Institute of Nuclear Physics SB RAS

September 23 - 27, 2019

Electron Cooling in the NICA Project: Status and Problems

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¹JINR Dubna, ²Budker INP SB RAS

Outline

Introduction: NICA project goals and three stages of the NICA project

1. Stage I: Experiment “The Baryonic Matter at Nuclotron”

**2. Stage II : Search for The Mixed Phase and New Physics In Heavy Ions' Collisions
at NICA Collider**

3. Stage II-a: The basic configuration of the NICA complex

3.1. Status of The NICA civil construction (11.05.2017)

4. Stage II-b: The project (full) configuration of the NICA complex

5. Stage III: Polarized Beams' Mode of The Collider



Introduction: Goals and Three Stages of the NICA Project

- We intend to study the Universe
 - as it was **13.799 ± 0.021 billion years ago** (Planck Mission data, 2015)
 - and $\sim 10\text{-}100 \mu\text{s}$ after Big Bang;
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Stage I: Fixed target experiment “The Baryonic Matter at Nuclotron” (BM@N)

2019 => ...

Stage II : Search for The Mixed Phase and New Physics In Heavy Ion Collisions

at NICA Collider

2022 => ...

Stage III: Polarized Beams' Mode of The Collider 2025 => ...



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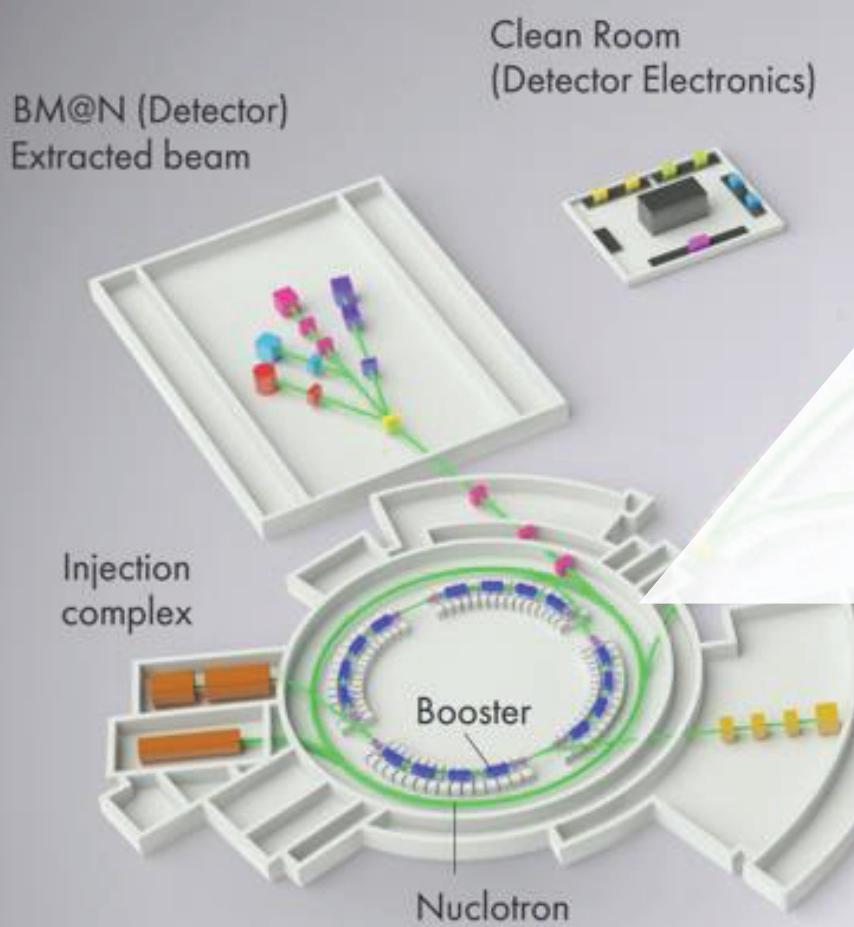
2022 => ...

Stage III: Polarized Beams' Mode of The Collider 2025 => ...

**Stage I & II – NICA White Paper: EPJ A. V 52, N8, 2016. 267 p. and
http://theor0.jinr.ru/twiki/pub/NICA/WebHome/WhitePaper_10.01.pdf**

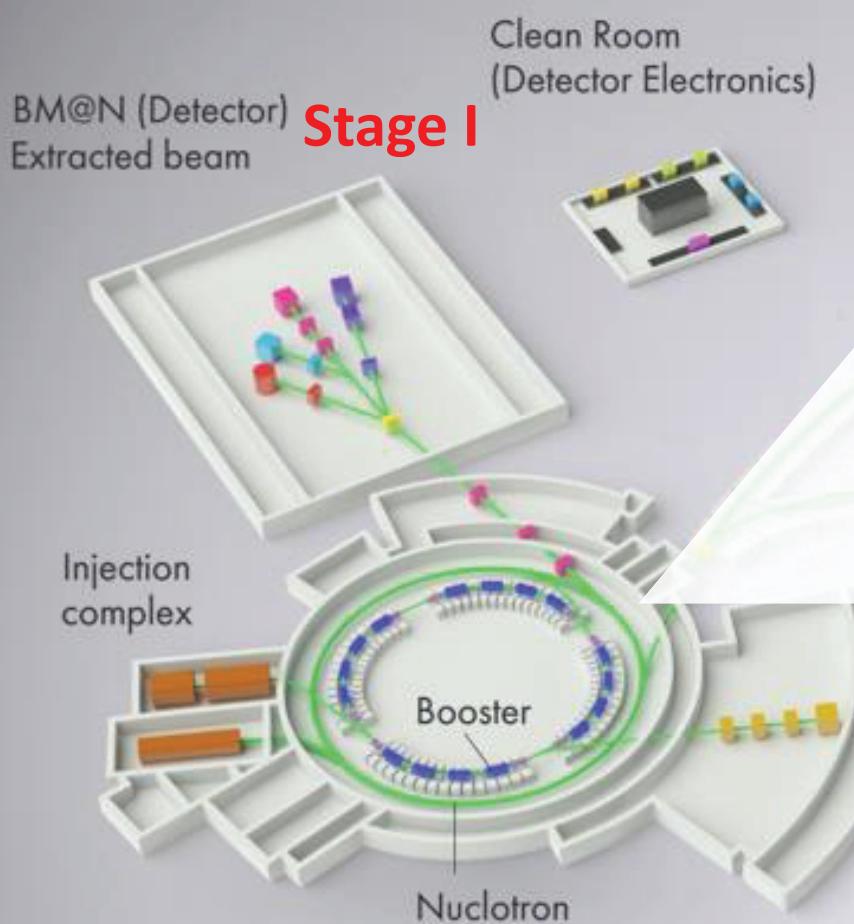


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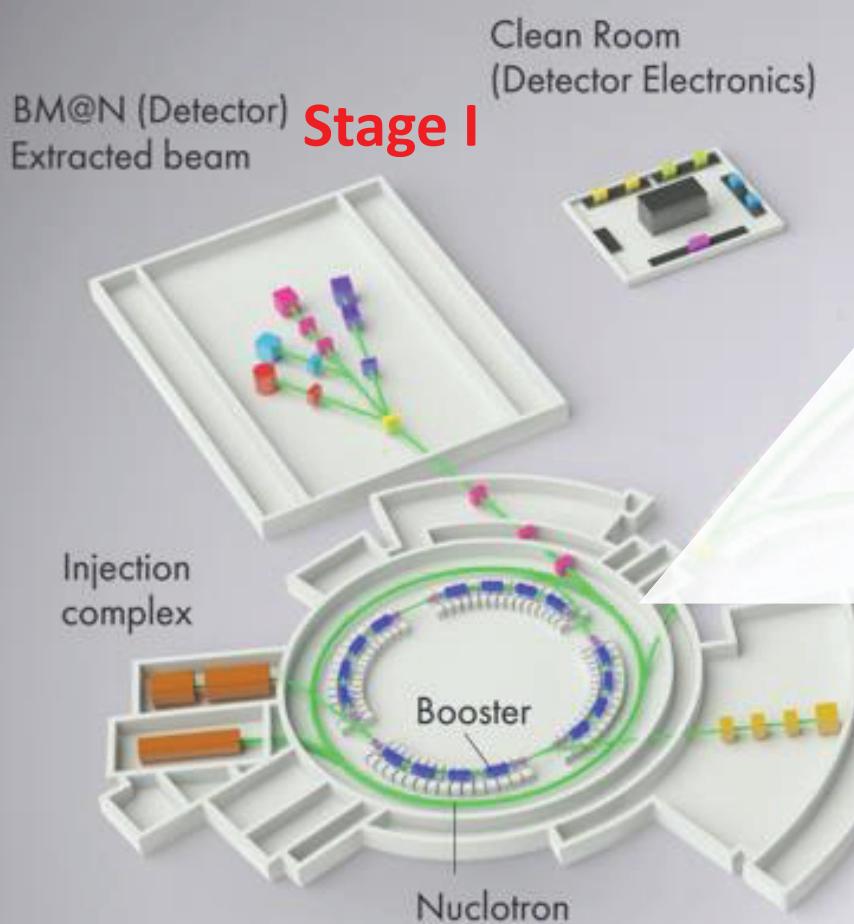
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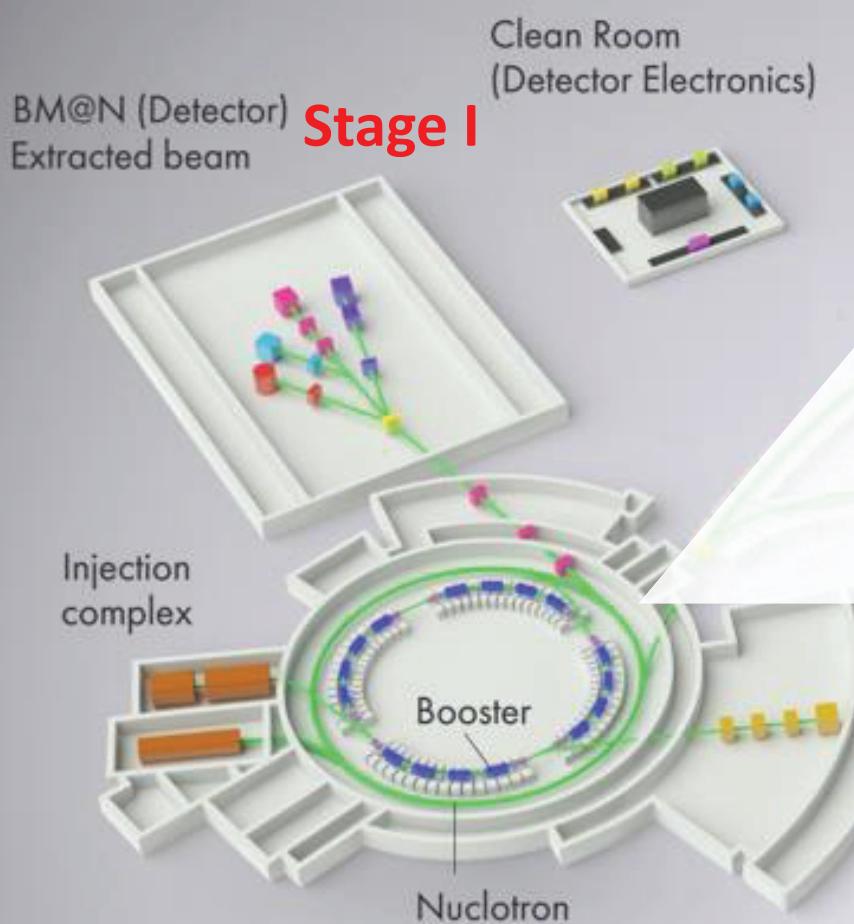
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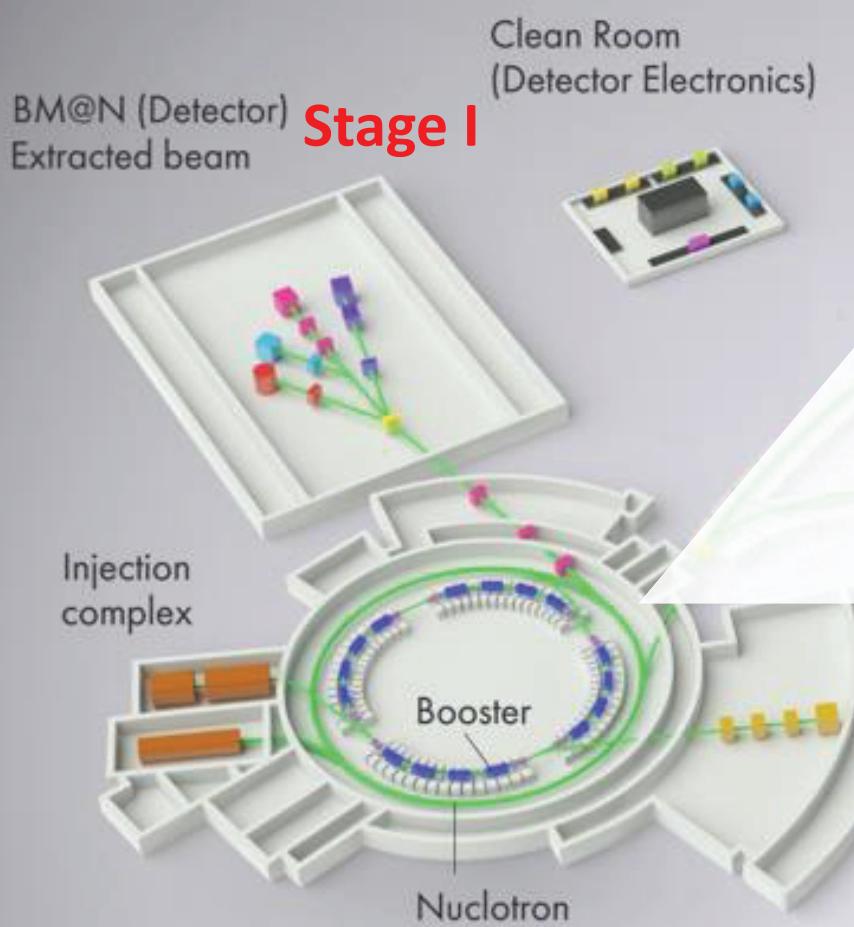
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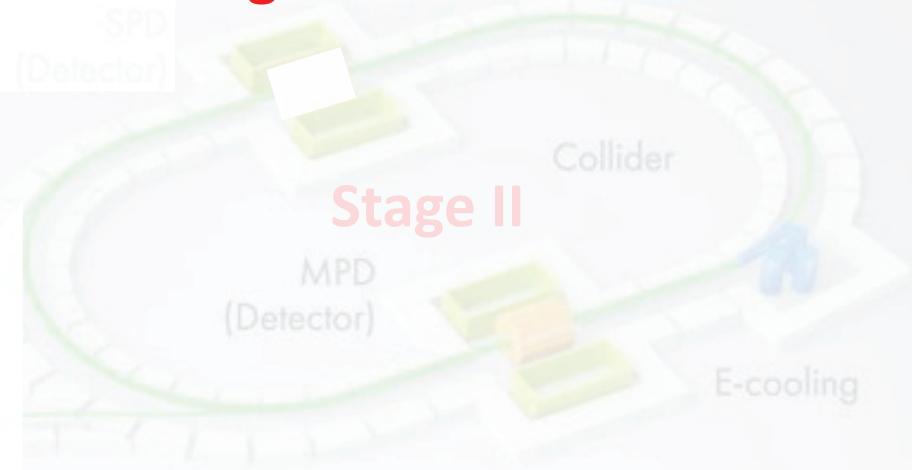
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Stage I

NICA – How will it look like

Stage III



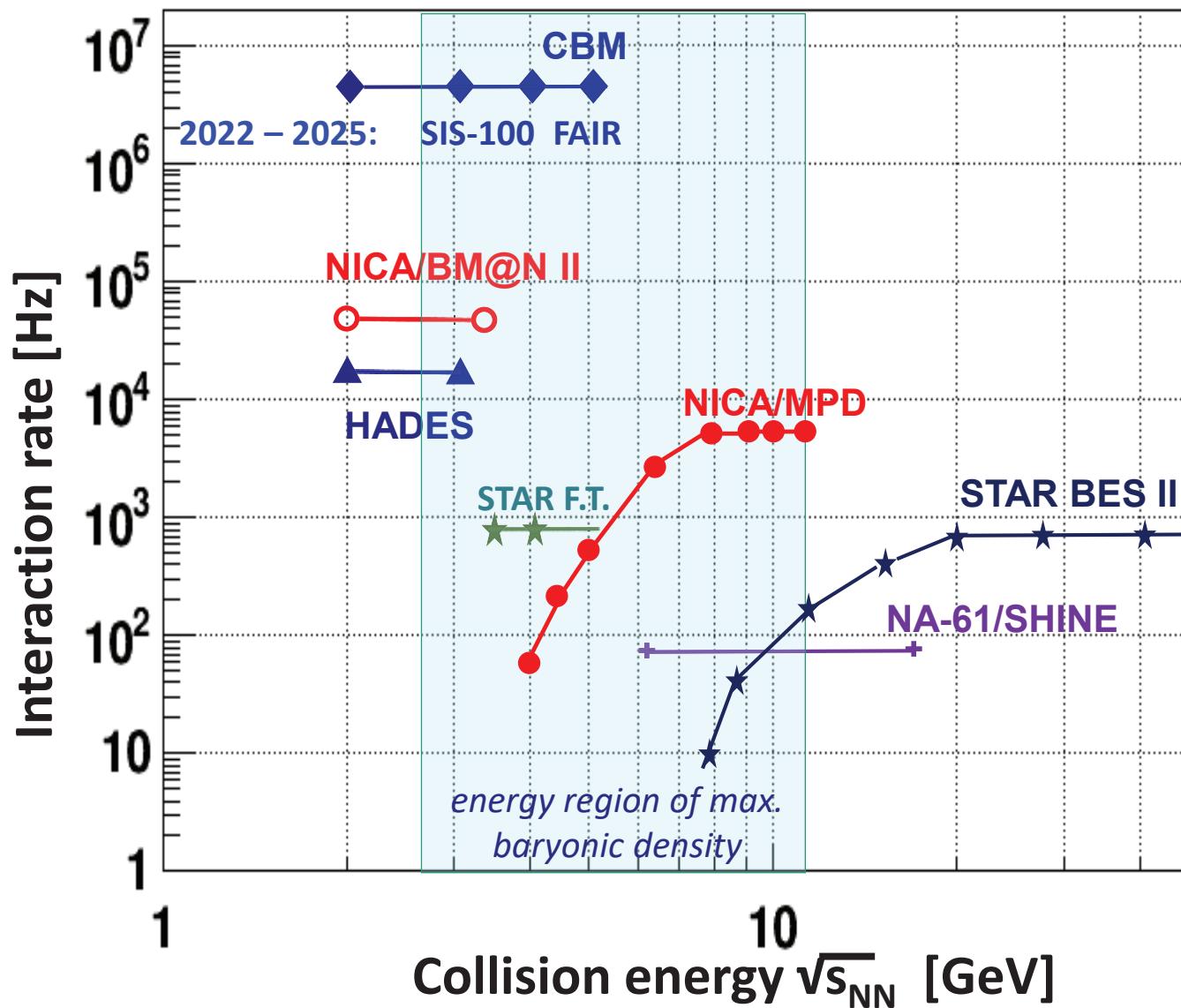
Stage II



Center NICA

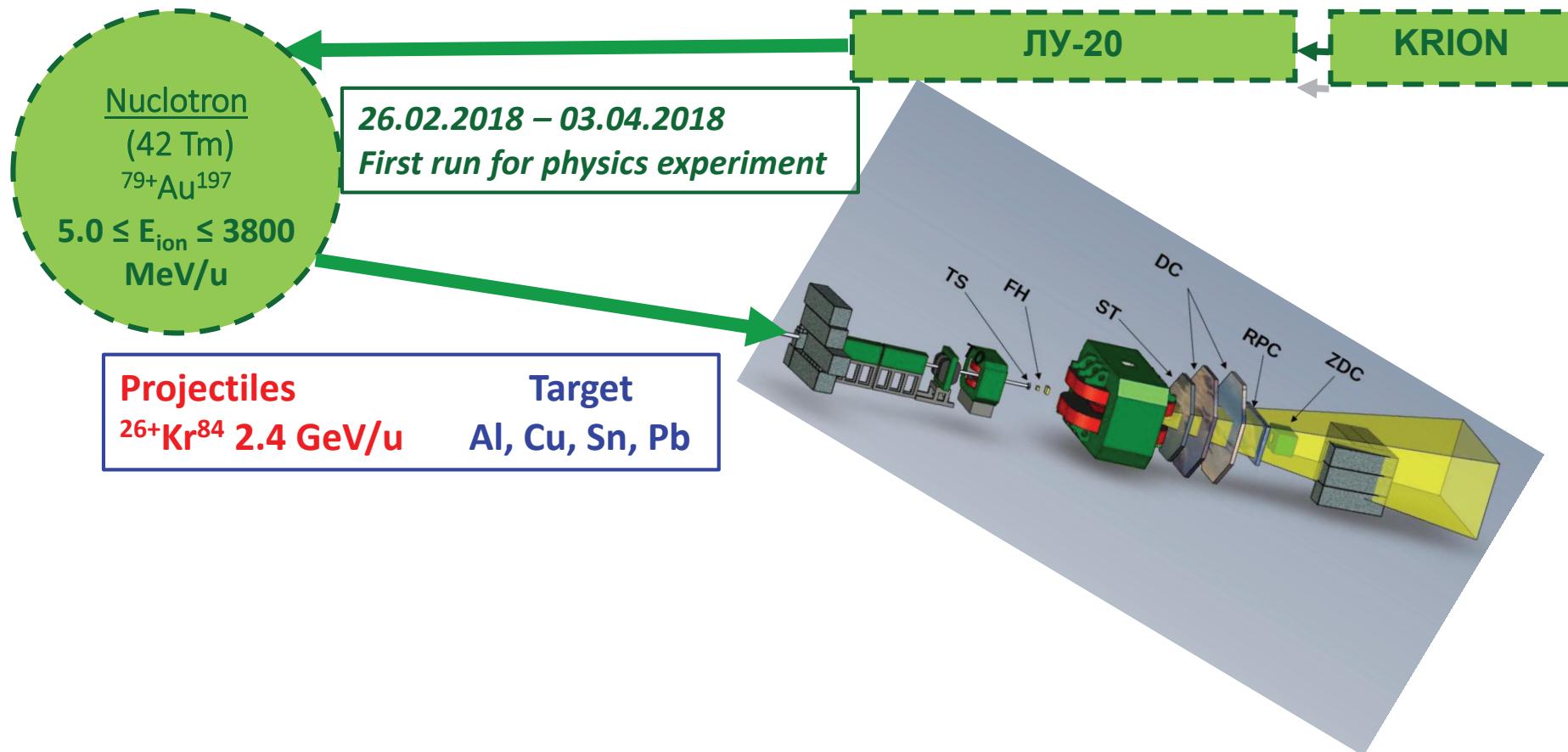
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Present and Future Heavy Ion Experiments



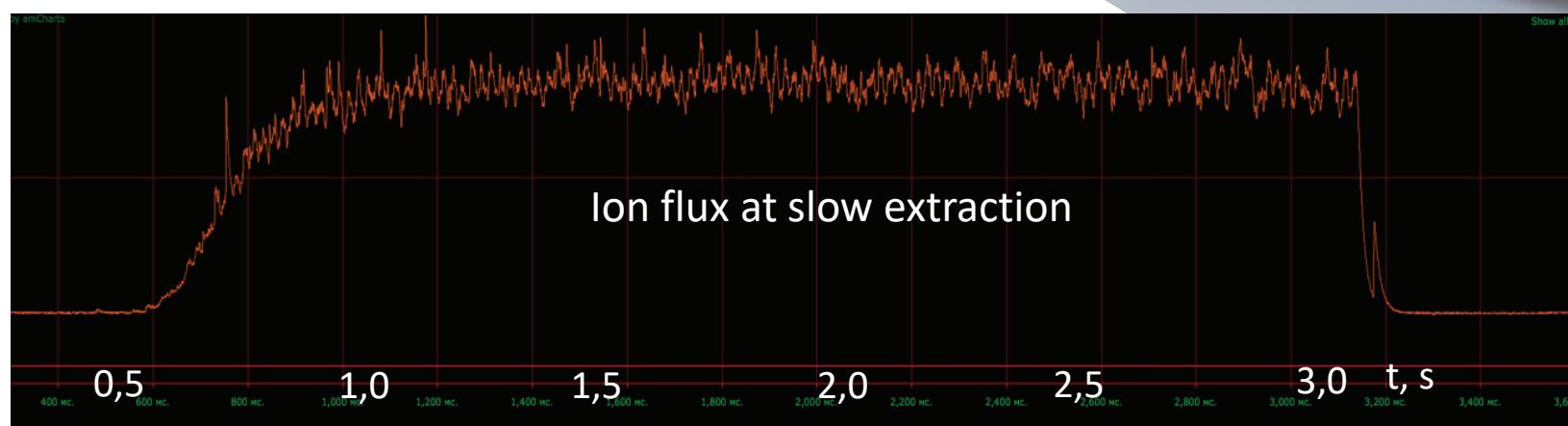
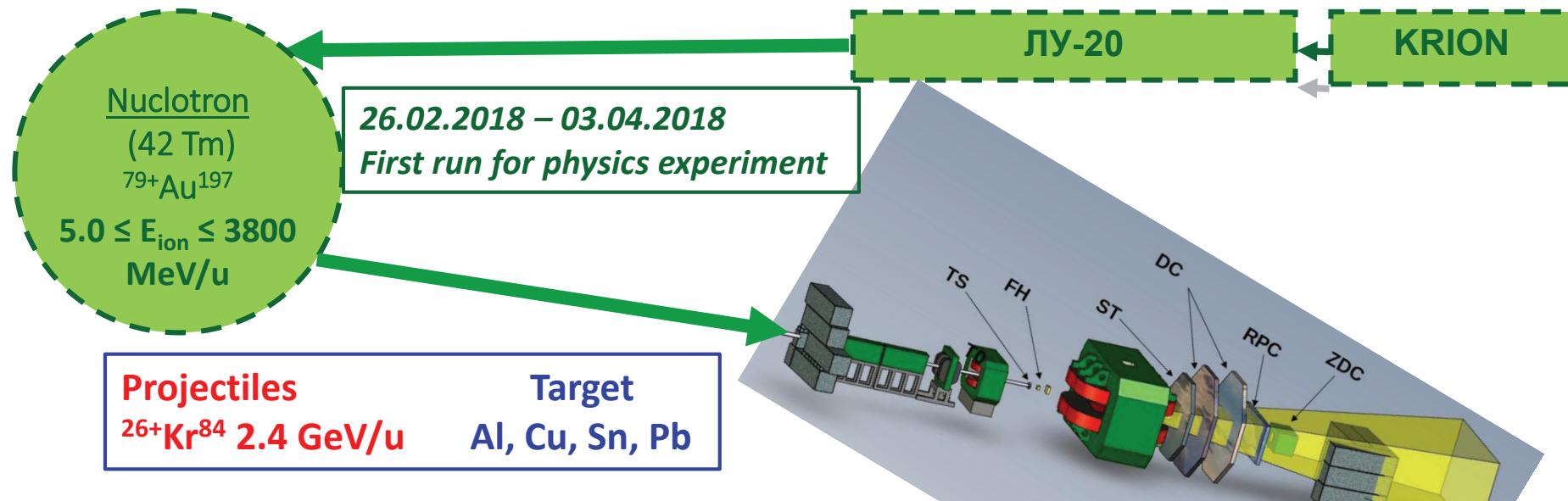
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BM@N will study the same physics in heavy ion collisions as NICA/MPD and CBM/FAIR, but in fixed target experiment at lower energy $\sqrt{s} = 2.3 \div 3.47 \text{ GeV/u}$.

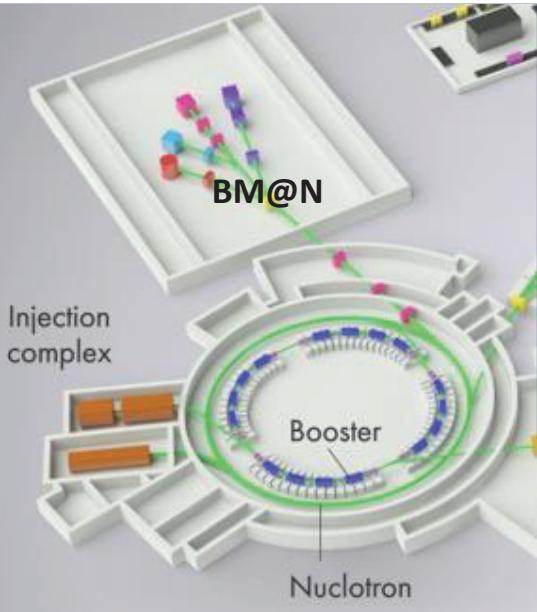


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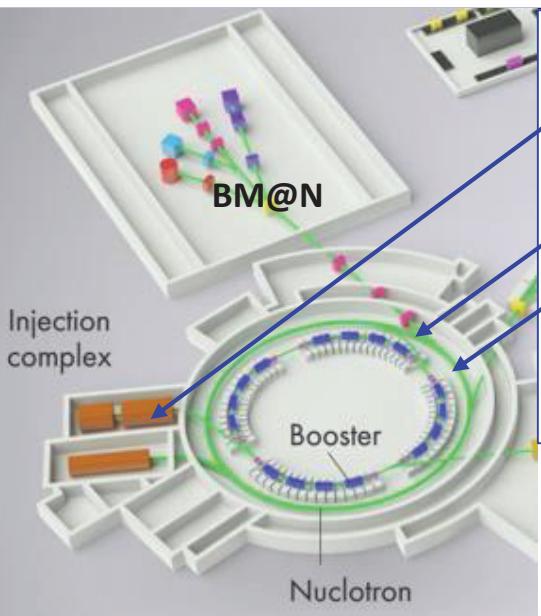
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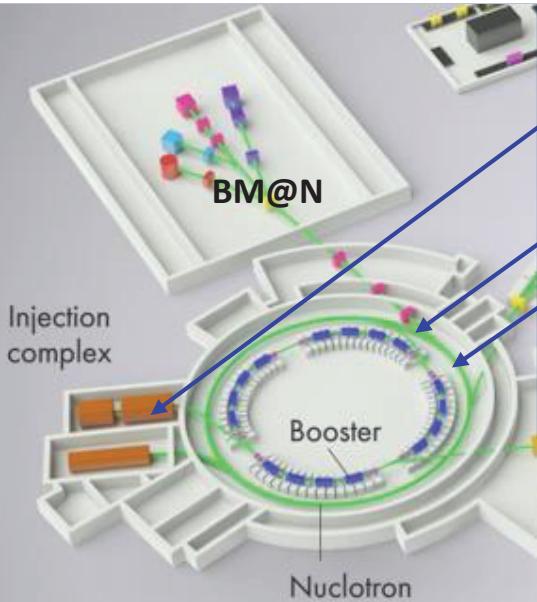
- **Injector – KRION + HILAc + Beam transfer line (BTL) from injector to the Booster**
- **Booster (under mounting and commissioning)**
- **BTL Booster - Nuclotron (under manufacturing at BINP, to be delivered to JINR in March 2020, commissioning – May 2020)**
- **Upgrade of BTL Nuclotron – BM@N (slow extraction)**

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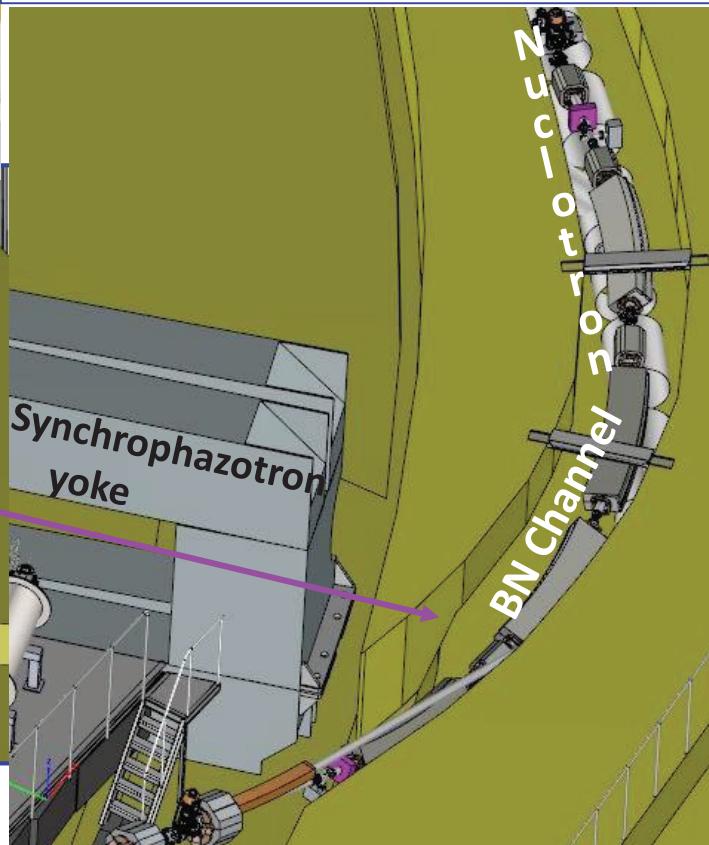
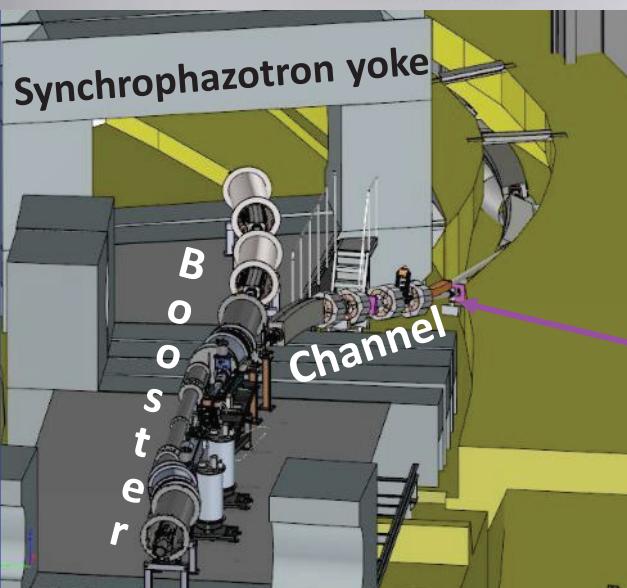


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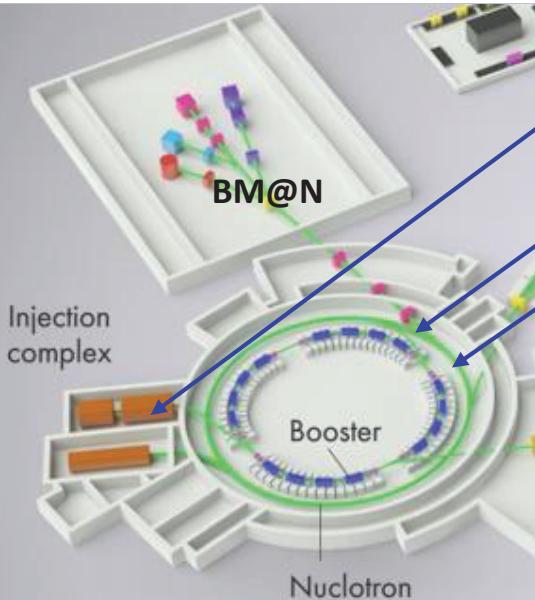
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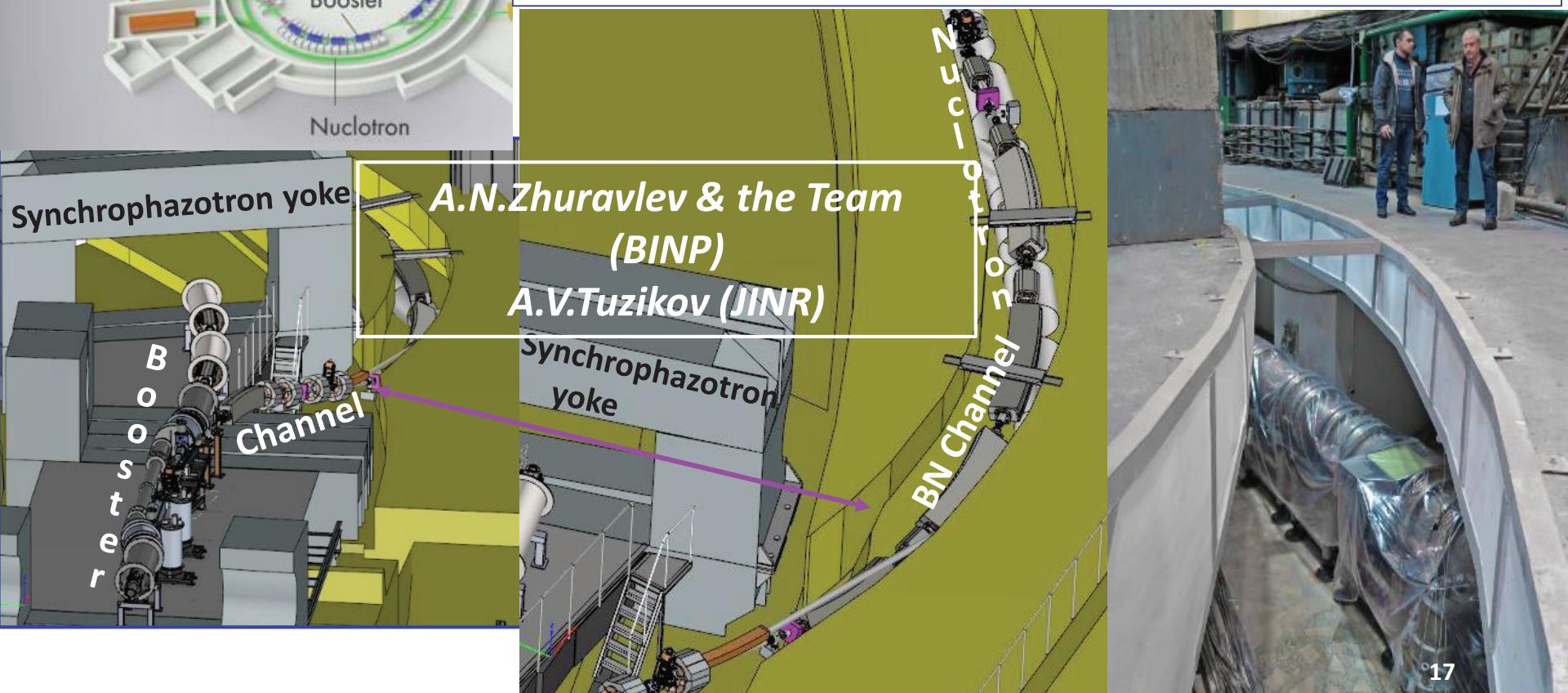
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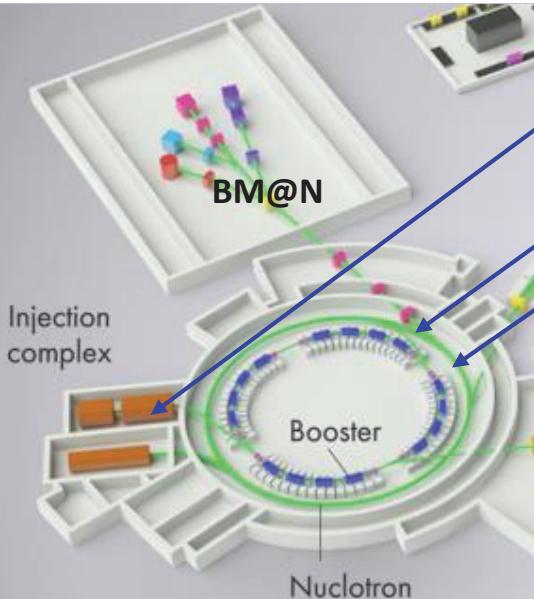
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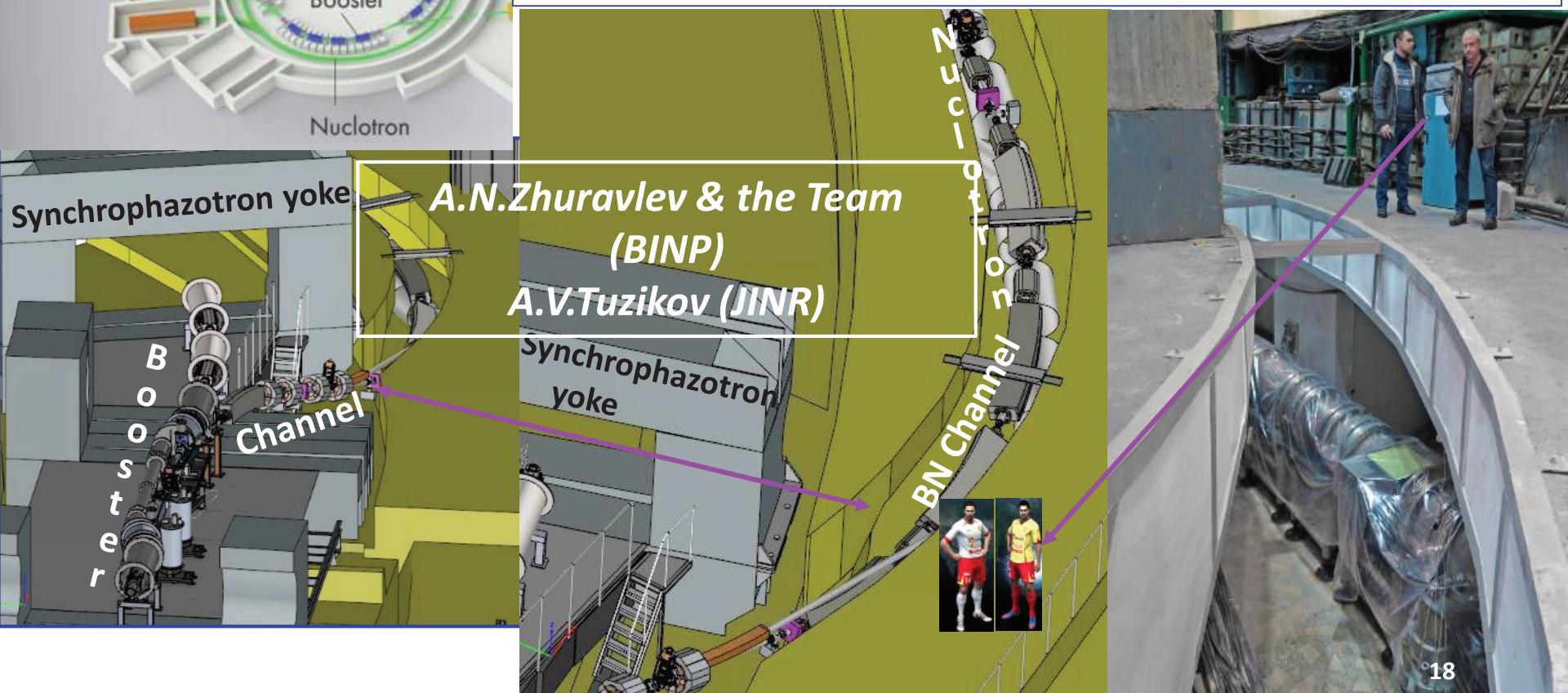
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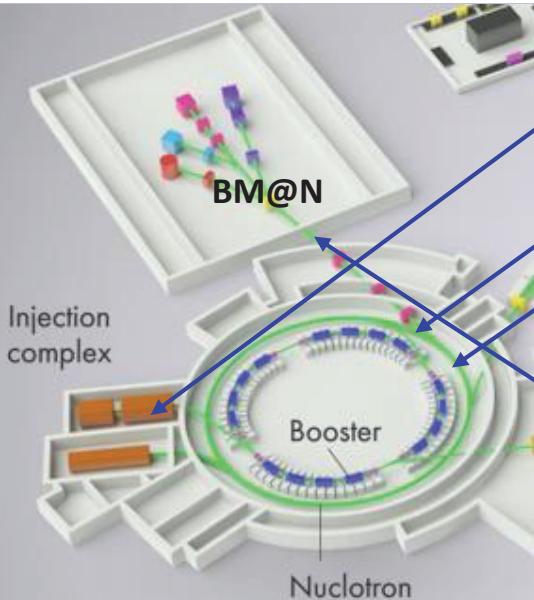
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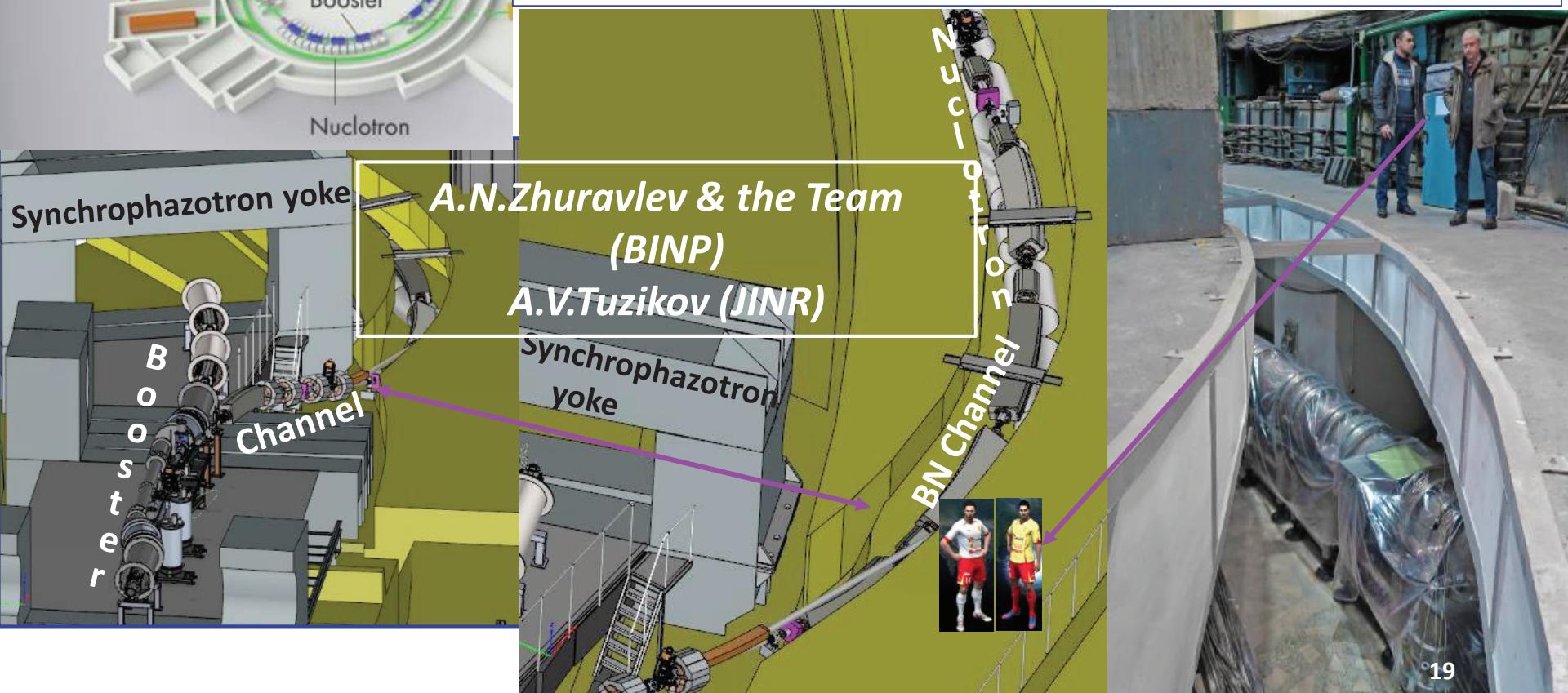
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1. Stage I: Experiment “The Baryonic Matter at Nuclotron”

Booster Mounting and Nearest Plans

100% of dipoles have been mounted
Quads tests and mounting is in progress



Superconducting connections and cables
yoke



Booster PSs installed on the top of Synchrophasotron

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Booster PSs installed on the top of Synchrophasotron yoke

Nearest Plans

1. November 2019 – mounting completion, technological run;
2. December 2019 - January 2020 – ion injection and acceleration in Booster;
3. May - June 2020 – ion acceleration and extraction, injection into BTL and transportation to Nuclotron

2020 – first run with heavy ions at maximum energy of Nuclotron.

Electron Cooling Application to The NICA Booster

The goals of e-cooling application:

- 1) storage of ions at injection,
- 2) formation of a dense bunched ion beam.

1. Injection – three schemes of injection into the Booster:

$^{197}\text{Au}^{31+}$ 1-1. Single turn injection

1-2. Multiturn injection

1-3. Multicycle injection

Ion energy 3.24 MeV/u, $E_e = 1.765 \text{ keV}$

2. Electron cooling on energy plateau at acceleration

$$E_{\text{ion}} = 50 - 80 \text{ MeV/u}, E_e = 27.233 - 43.573 \text{ keV}$$

Characteristic cooling time $\leq 1 \text{ sec}$

Why such an energy choice: to avoid space charge limitation at the cooled beam shrinking.



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Electron cooling will be used
in all three schemes.

Characteristic cooling time $\sim 0.1 \text{ sec.}$

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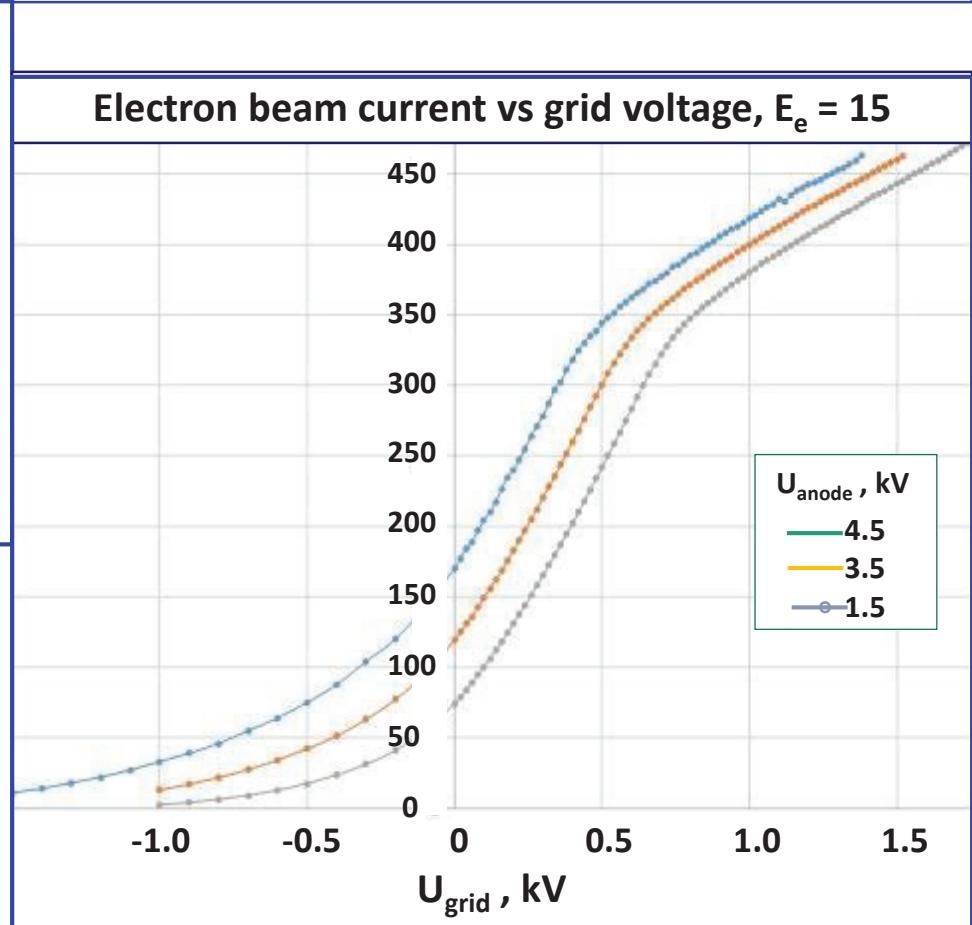
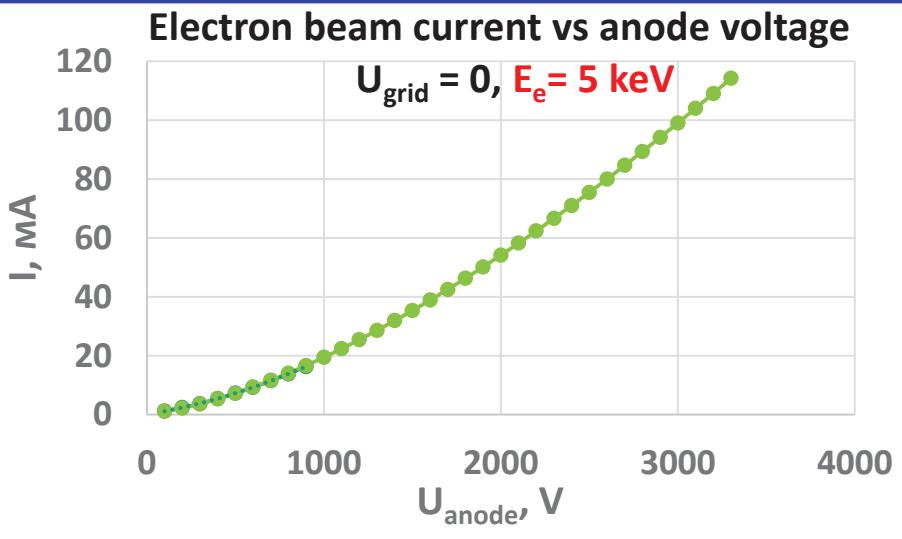
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Electron Cooling Application to The NICA Booster

Booster Electron Cooler Commissioning

A.Kobets, S.Melnikov, O.Orlov, S.Semenov, A.Sergeev, Al. Sidorin

July 09-24, 2019



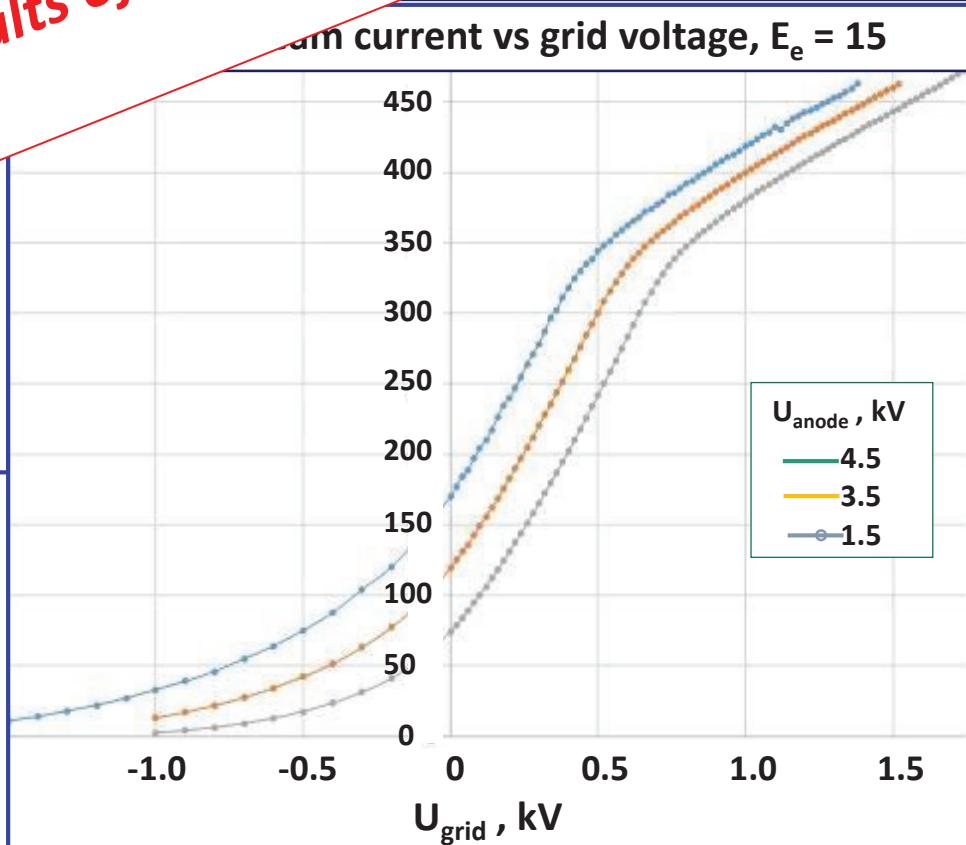
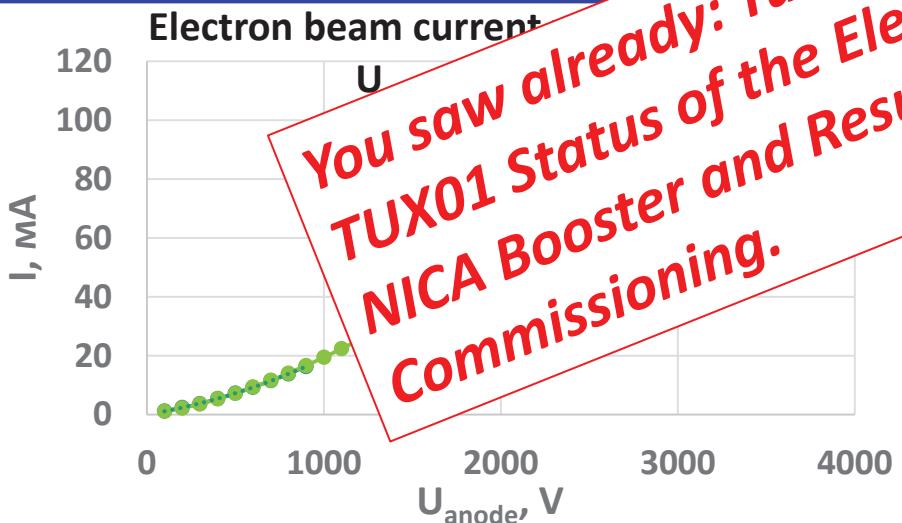
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Electron Cooling Application to The NICA Booster

Booster Electron Cooler

A.Kobets, S.Melnikov, O.Orlov

You saw already: Tue 9:00 M.Bryzgunov
TUX01 Status of the Electron Cooler for
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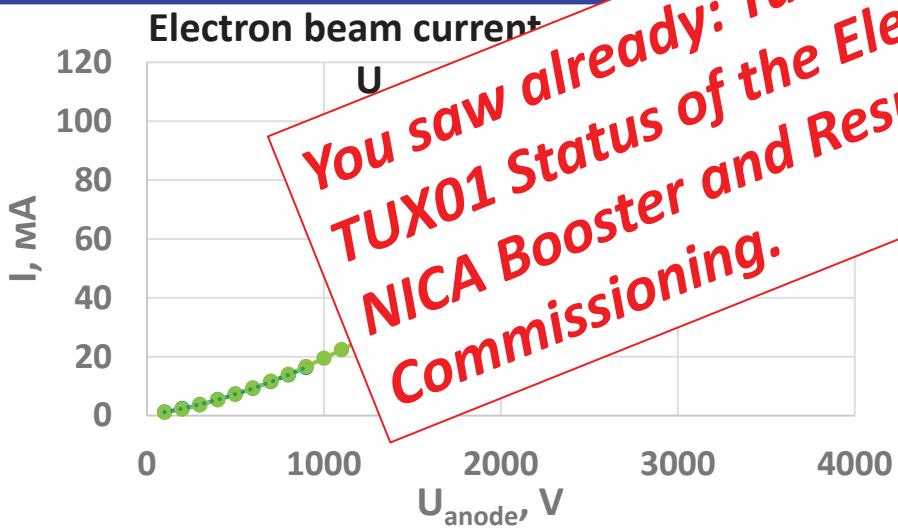
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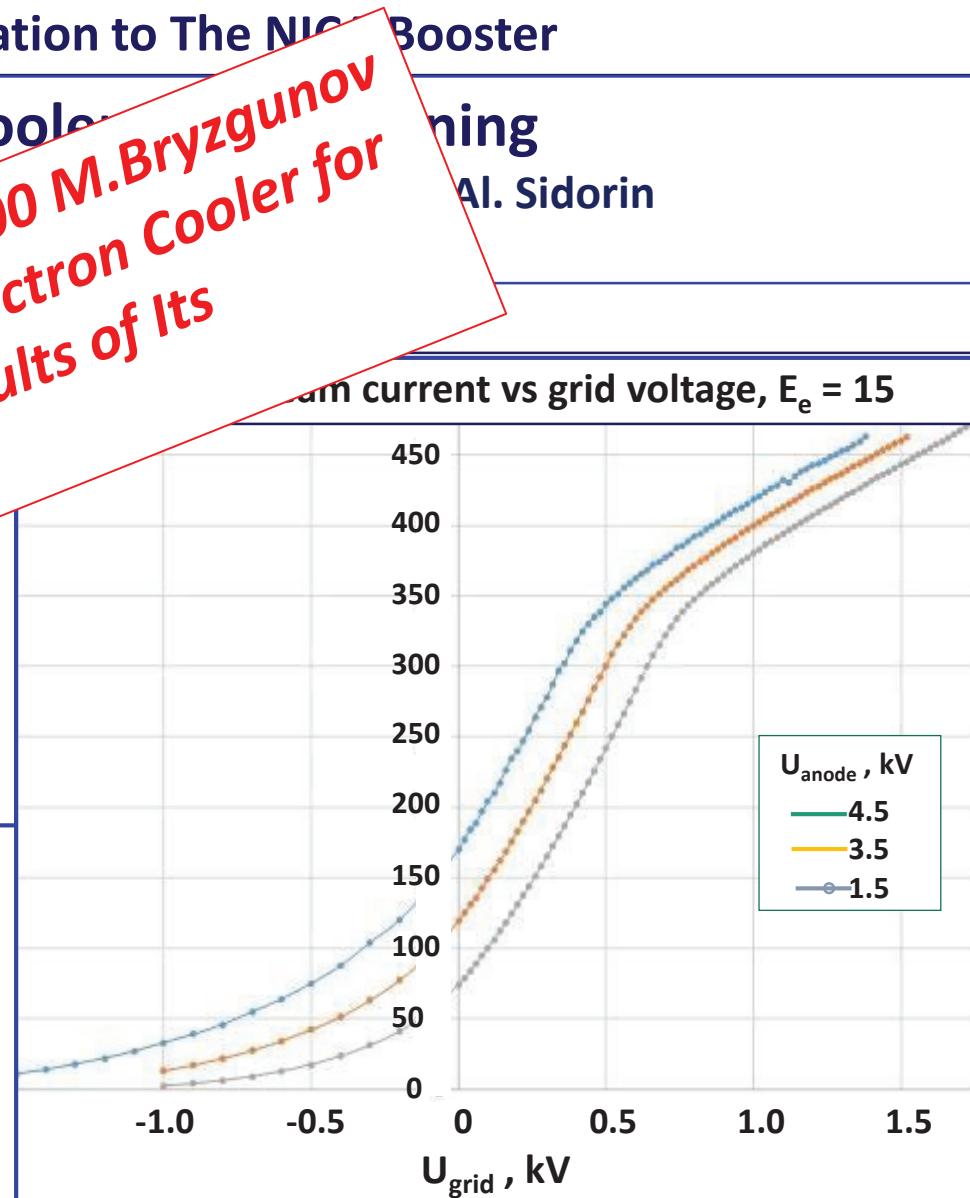
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NICA team is learning e-cooling
technique (contnd. next week!)

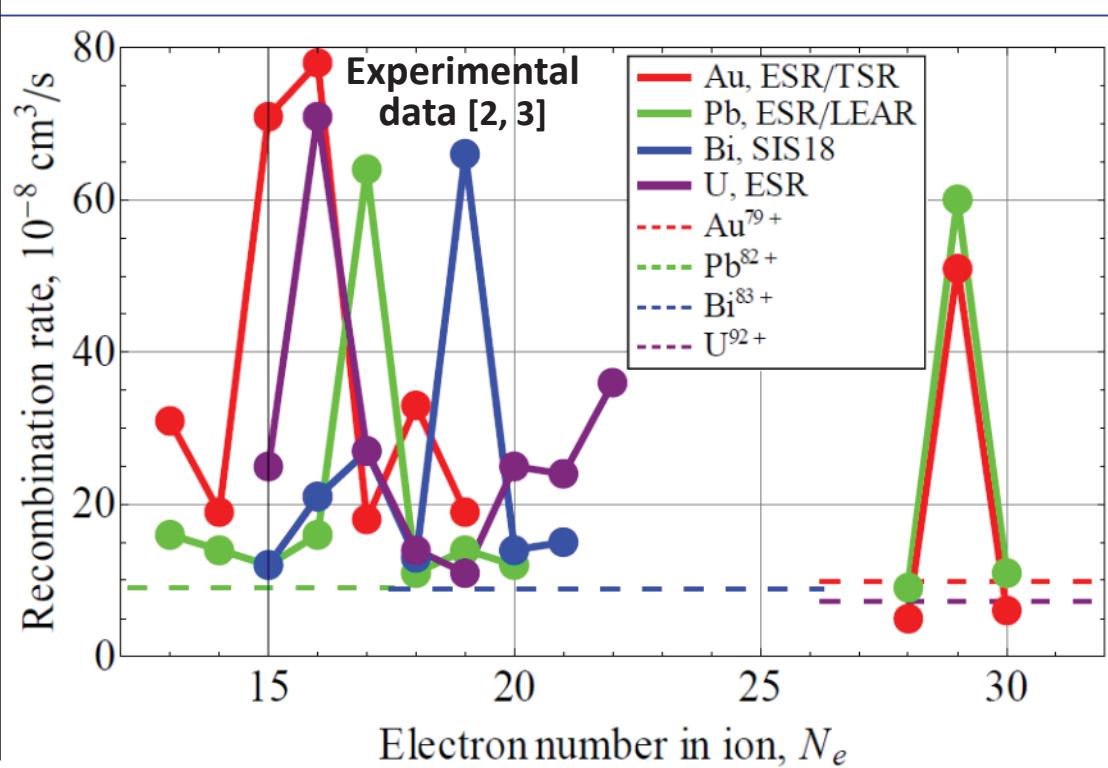
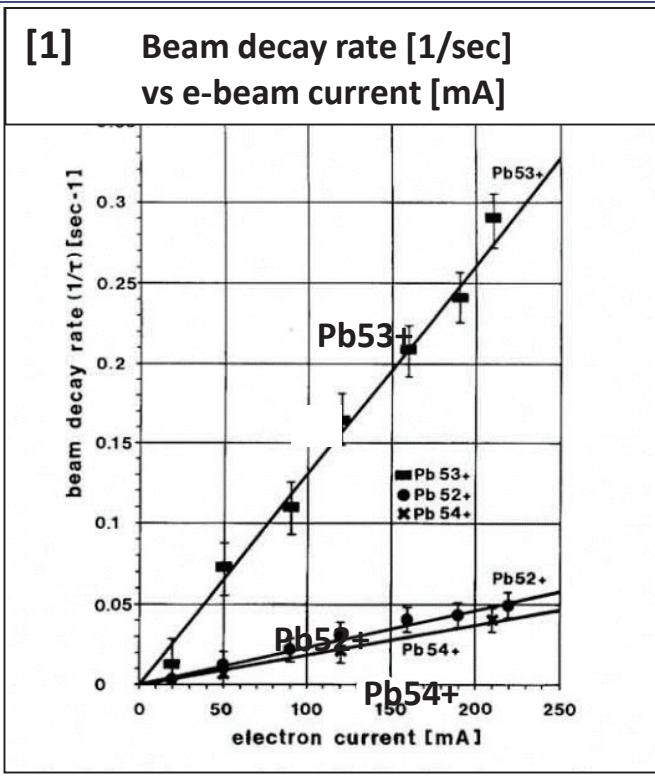


Stage I: Experiment “The Baryonic Matter at Nuclotron”

Electron Cooling Application to The NICA Booster

Why do we need to use $^{197}\text{Au}^{31+}$:

- because of limited capability of KRION ion source (ionization time, intensity, etc.),
- to avoid ion loss due to electron-ion recombination in electron cooler.



1. S.Baird, I.Meshkov, D.Moehl, et al., Phys. Lett. B, 361, 1995, pp.184–186.
2. I. Meshkov, Report at working seminar at CERN 14.03.2011
3. O. Uwira, A. Mueller, A. Wolf, et al., Hyperfine Interactions 108 (1997) 149–154 (Au49+, 50+, 51+)
4. A.V. Philippov, A.B. Kuznetsov, I.N. Meshkov, in Proc. COOL'11, <https://www.jacow.org>, THCOB01

Stage I: Experiment “The Baryonic Matter at Nuclotron”

Electron Cooling Application to The NICA Booster

Dependence of recombination rate $1/\tau_{\text{rec}}$ on electron shell structure has resonance character: $1/\tau_{\text{rec}}$ is

- minimal if outer (sub)shells are completely filled, either contain even number of electrons;
- maximal if outer (sub)shells have one vacancy or odd number of electrons.

$^{197}\text{Au}^{31+}$ ion contains $79 - 31 = 48$ electrons in the outer subshell of this ion.

Ion	Electron number in the ion	Atomic shell structure											Measured rate [$10^{-8} \text{ cm}^3 \text{s}^{-1}$]
Pb^{52+}	30	1s ²	2s ²	2p ⁶	3s ²	3p ⁶	4s ²	3d ¹⁰					11 (full subshell)
Pb^{53+}	29	1s ²	2s ²	2p ⁶	3s ²	3p ⁶	4s ²	3d ¹⁰	4p ¹				60 (1 in subshell)
Pb^{54+}	28	1s ²	2s ²	2p ⁶	3s ²	3p ⁶	4s ²	3d ⁸					9 (2 vacancies)
Au^{29+}	50	1s ²	2s ²	2p ⁶	3s ²	3p ⁶	4s ²	3d ¹⁰	4p ⁶	5s ²	4d ¹⁰	5p ²	
Au^{30+}	49	1s ²	2s ²	2p ⁶	3s ²	3p ⁶	4s ²	3d ¹⁰	4p ⁶	5s ²	4d ¹⁰	5p ¹	
Au^{31+}	48	1s ²	2s ²	2p ⁶	3s ²	3p ⁶	4s ²	3d ¹⁰	4p ₆	5s ²	4d ¹⁰		
Au^{49+}	30	1s ²	2s ²	2p ⁶	3s ²	3p ⁶	4s ²	3d ¹⁰					6 (full subshell)
Au^{50+}	29	1s ²	2s ²	2p ⁶	3s ²	3p ⁶	4s ²	3d ⁹					50 (1 vacancy)
Au^{51+}	28	1s ²	2s ²	2p ⁶	3s ²	3p ⁶	4s ²	3d ⁸					5 (2 vacancies)
Max. number of electrons in the shell of Au atom	2	4	10	12	18	20	30	36	38	48	54		



Electron Cooling Application to The NICA Booster

Estimates of recombination rate in Electron cooler of NICA Booster

$$\tau_{rec}^{-1} = \frac{\alpha}{\beta\gamma^2} \cdot \frac{J_e}{ce} \cdot \eta_{cool},$$

α is recombination coefficient, β, γ are relativistic factors,
 J_e is electron current density, c is the speed of light, e is electron charge,
 η is ratio of the cooling section length l_{cool} to the ring circumference C_{ring} .

Choosing the ion Au³¹⁺ and the parameter values

$$\alpha = 10 \times 10^{-8} \text{ cm}^3 \cdot \text{s}^{-1}, l_{cool} = 1.5 \text{ m}, C_{ring} = 210.96 \text{ m}, \phi_{cathode} = 30 \text{ mm}$$

we find for the NICA Booster:

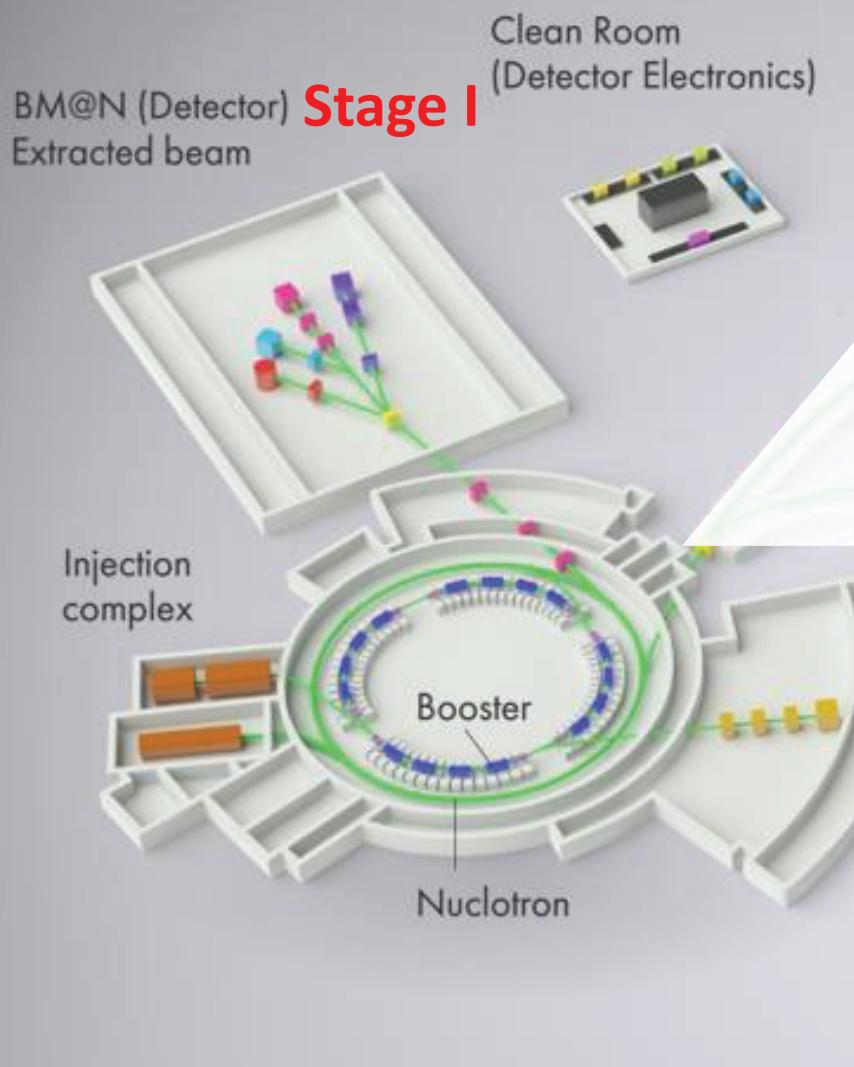
at injection: $E_{ion} = 3.2 \text{ MeV/u}$ $E_e = 1.76 \text{ keV}$, $I_e = 0.2 \text{ A} \Rightarrow \tau_{rec} = 13.3 \text{ sec}$;
 lifetime on vacuum 4.4 sec.

on “plateau”: $E_{ion} = 80 \text{ MeV/u}$: $E_e = 27.2 \text{ keV}$, $I_e = 1.0 \text{ A} \Rightarrow \tau_{rec} = 14.6 \text{ sec}$.

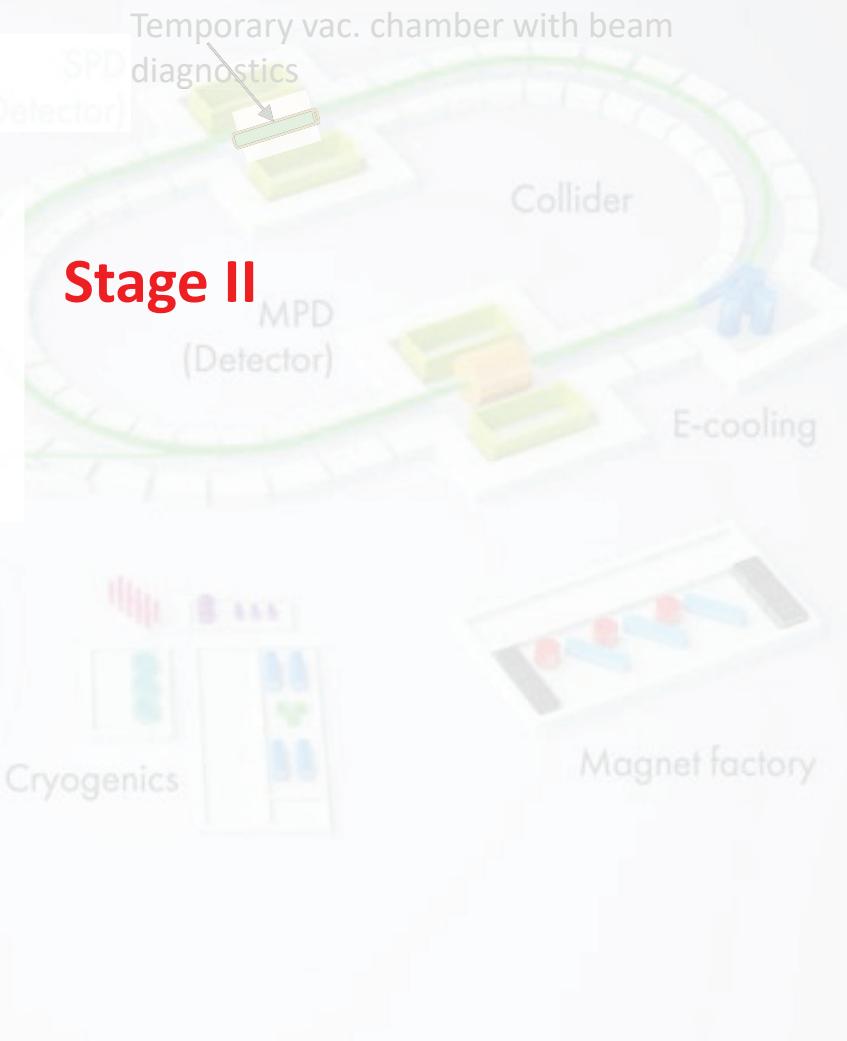
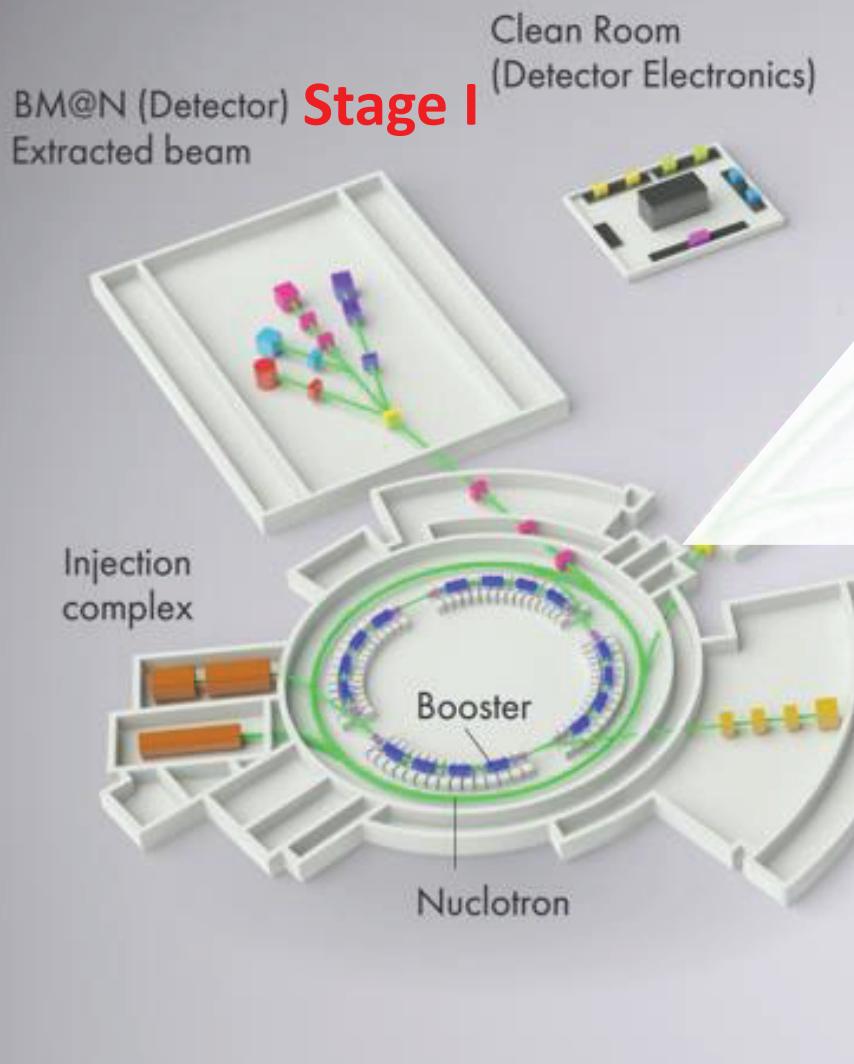
These values of the ion beam life time limited by ion recombination in electron cooler
 are sufficient for NICA operation.

Note that application of Au³⁰⁺ will reduce recombination time by 5 – 10 times.

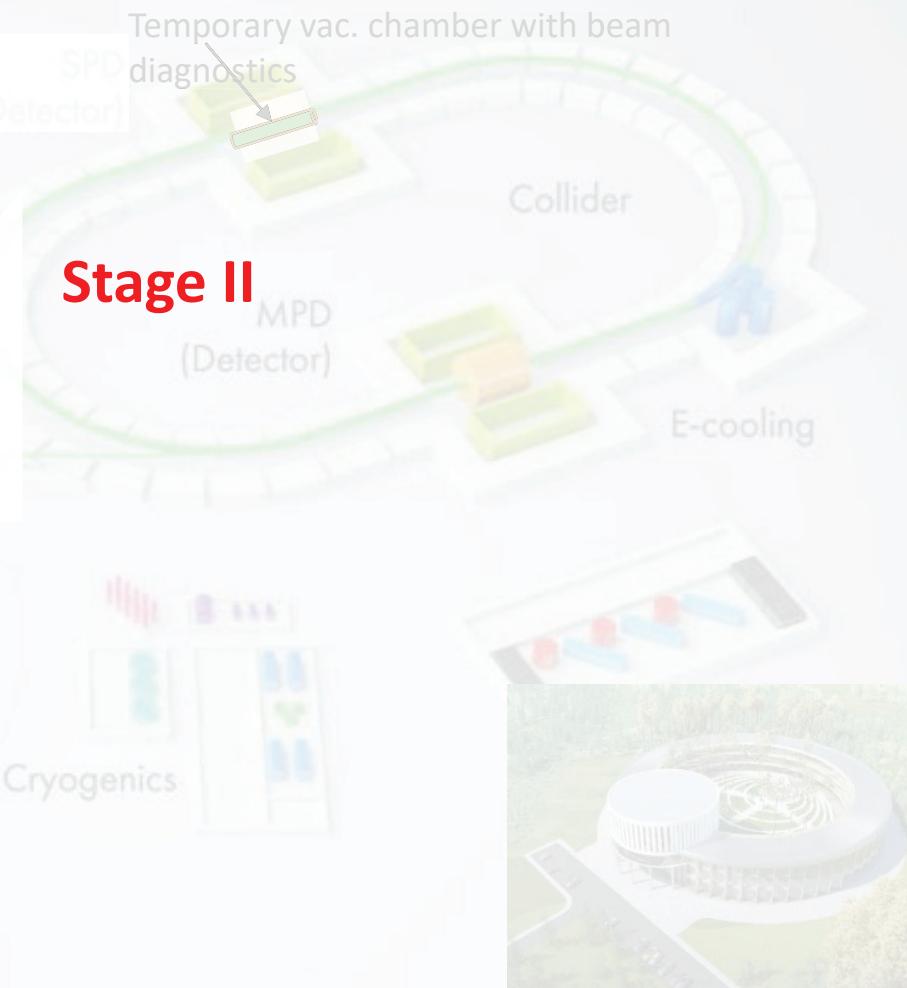
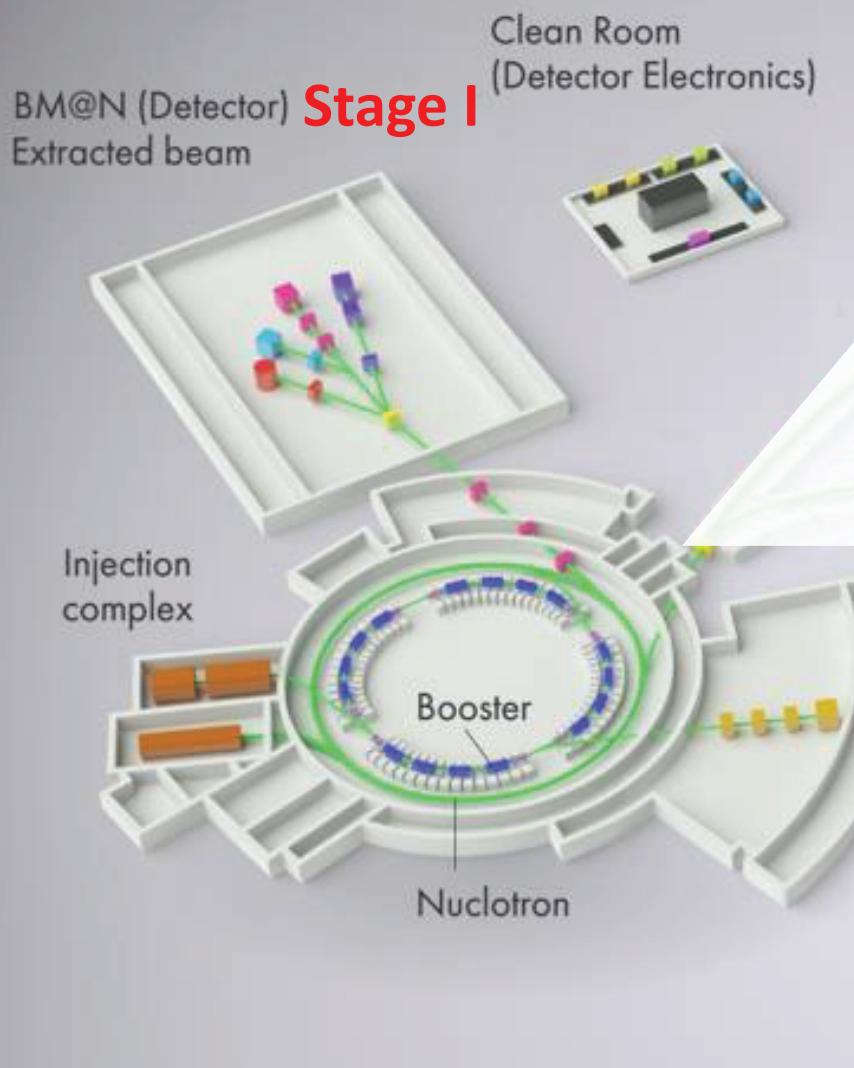
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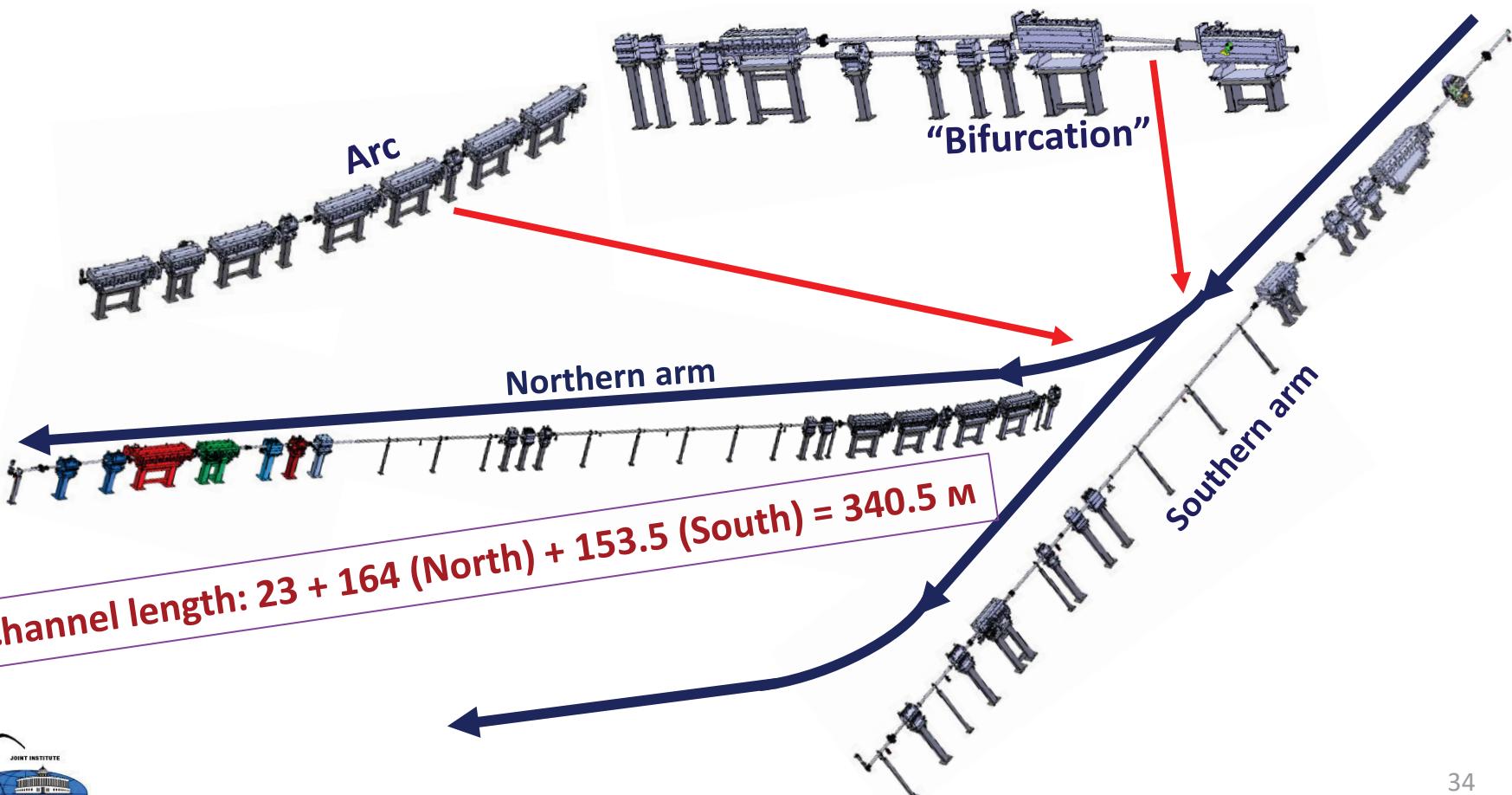
2. Stage II : Search for The Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

Beam Transfer Channel Nuclotron - Collider

Physics project and coordination – A.Tuzikov (JINR)

Working project, fabrication – SigmaPhi C° (France)

Mounting and commissioning – common work



2. Stage II : Search for The Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

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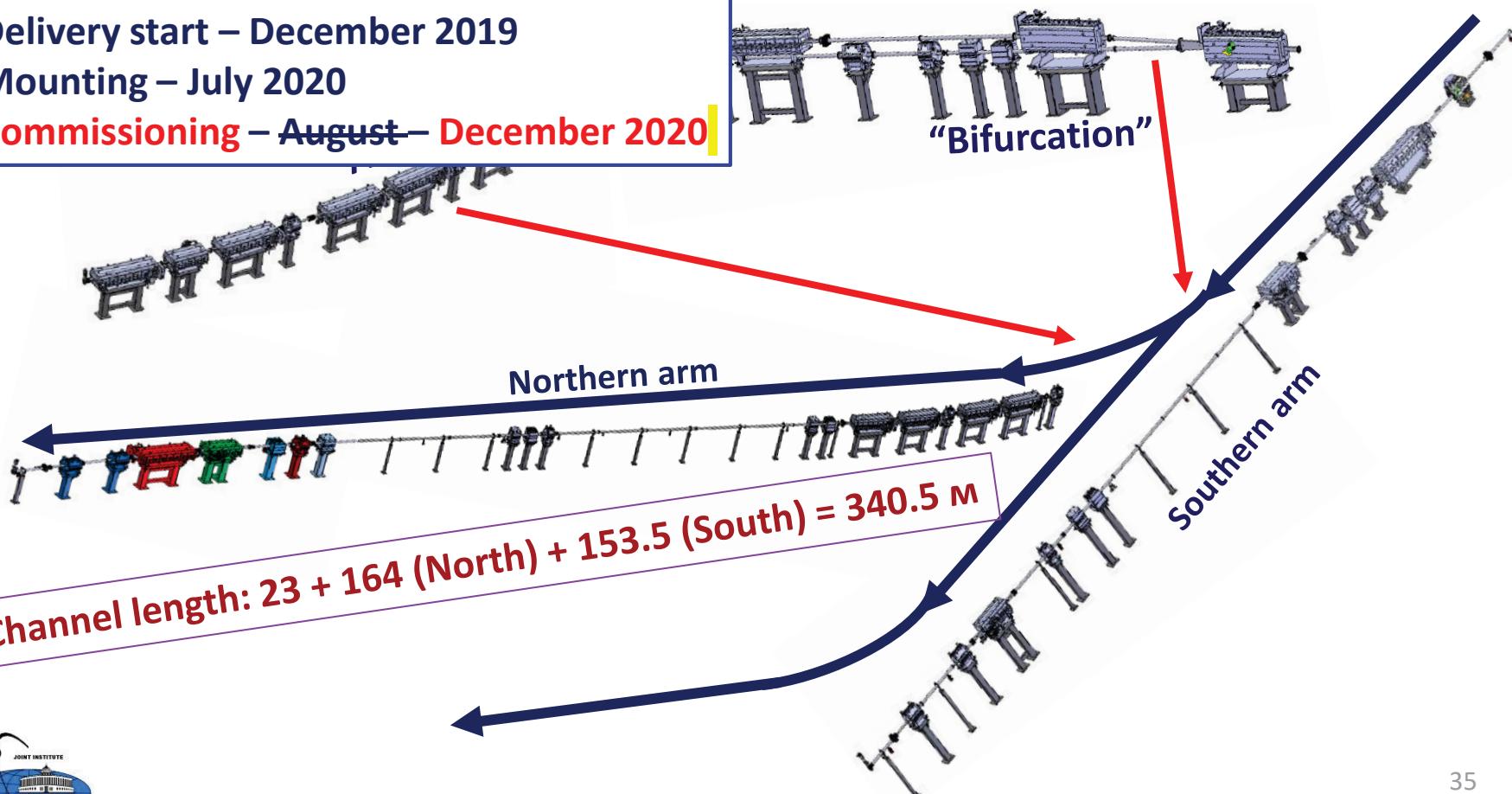
Working project, fabrication – SigmaPhi C° (France)

Mounting and commissioning – common work

Delivery start – December 2019

Mounting – July 2020

Commissioning – August – December 2020



2. Stage II : Search for The Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: Focusing System

September 2019

SC magnets	Booster	Collider
Dipoles	40	88 (18)
Quads	24 (8)	82 (1)

2. Stage II : Search for The Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: Focusing System

September 2019

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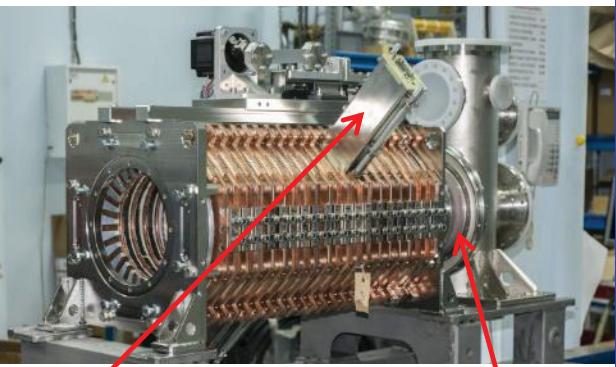
Collider:
Mounting start – September 2021
Commissioning – February 2022



2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: RF System (BINP, A.Tribendis and the team)

Barrier bucket system RF1
(Rectangular periodic pulses)



Pulse generator Ceramic gap
Test assembly at BINP
(Vac. chamber is not installed yet)

Harmonic systems

RF2 cavity prototype RF2 and RF3



$f_0 = 11.1 \div 13.6 \text{ MHz}$
 $Q_0 = 3\,600 (@ 12.2 \text{ MHz})$
Ready for high power tests.

Photos of May 2019

RF3 cavity prototype



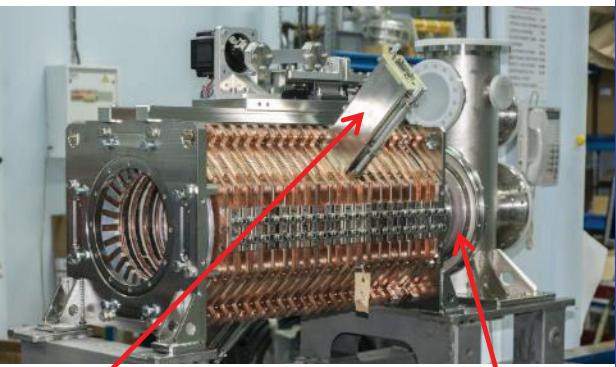
$f_0 = 34.1 \div 39.4 \text{ MHz}$
 $Q_0 = 6\,900 (@ 39.0 \text{ MHz})$

Test assembly at the BINP Workshop

2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: RF System (BINP, A.Tribendis and the team)

Barrier bucket system RF1
(Rectangular periodic pulses)



Pulse generator Ceramic gap
Test assembly at BINP
(Vac. chamber is not installed yet)



Friendly inspection
21.03.2019

Harmonic systems

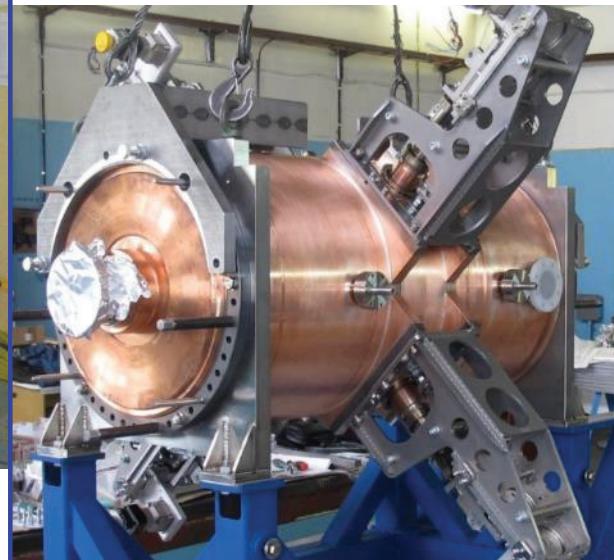
RF2 cavity prototype RF2 and RF3



$f_0 = 11.1 \div 13.6 \text{ MHz}$
 $Q_0 = 3\,600 (@ 12.2 \text{ MHz})$
Ready for high power tests.

Photos of May 2019

RF3 cavity prototype



$f_0 = 34.1 \div 39.4 \text{ MHz}$
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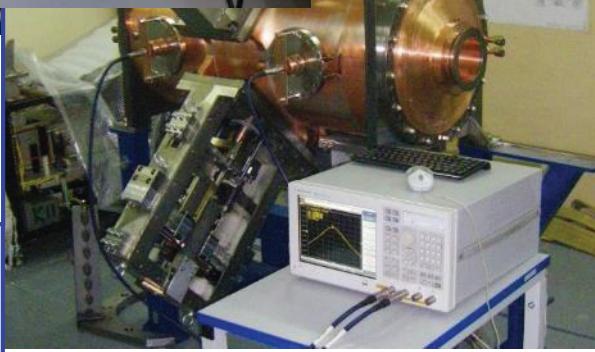
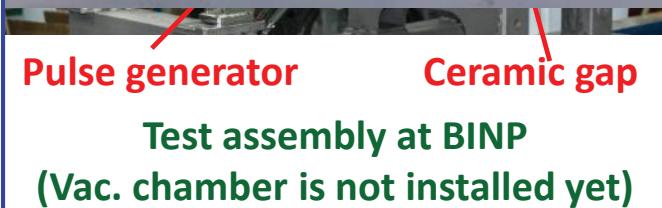
Mixed Phase and New Physics Experiments at NICA Collider

BINP, A.Tribendis and the team)

Photos of May 2019



23.09.2019 RF-2 and All The Team



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Harmonic systems
Prototype RF2 and RF3



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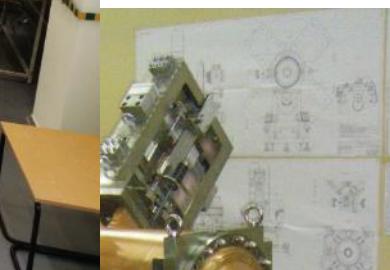
23.09.2019 RF-2 and All The Team



24.09.2019 RF-2 Visit of the NICA MAC

harmonic systems

prototype RF2 and RF3



RF3 cavity prototype



$f_0 = 13.6\text{MHz}$
 $\varnothing 12.2\text{MHz}$)
high power tests.

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Test assembly at the BINP Workshop



23.09.2019 RF-2 and All The Team



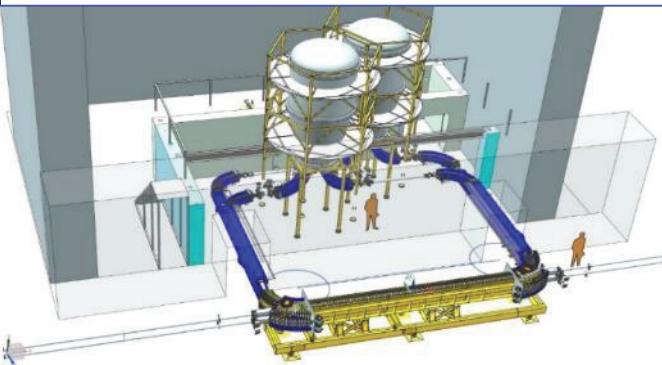
24.09.2019 The same and RF-3



24.09.2019 RF-2 Visit of the NICA MAC

2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: HV Electron Cooler (BINP, V.Parkhomchuk and the team)



Suppression of
Ion Beam Space Charge
and IBS.
Luminosity Optimization

Peculiarity of Application of the HV Electron Cooler at NICA Collider

1. Energy range $1 - 4.5 \text{ GeV/u}$ $^{197}\text{Au}^{79+}$ ions
 $2 - 4.5 \text{ GeV}$ protons
 $0.55 - 2.45^*) \text{ MeV}$ electrons
2. Electron beam current $0.1 - 1.0^*) \text{ A}$
3. Solenoid magnetic field $0.05 - 0.2^*) \text{ T}$
4. Two electron beams operated simultaneously $^*)$
5. Suppression of ion beam space charge and IBS
6. Suppression of ion recombination with cooling electrons $^*)$

$^*)$ Never reached before in electron coolers

Optimum luminosity of an ion collider can be obtained if sum of the betatron tune shifts caused by both bunch space charge (Δq) and beam-beam effect (ξ) for both colliding beams does not exceed certain critical value (a shift to nearest betatron resonances):

$$\Delta q + \xi \leq \Delta Q_{max}$$

This criterion allows one to find optimum value of the intensity of both colliding beams (see next slide) $^*)$.

$^*)$ I.Meshkov, Luminosity of an Ion Collider, Phys. Part. Nucl. (2019) v. 6

2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collision
See details in
V.Parkhomchuk
report, THX02



Suppression of
Ion Beam Space Charge
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Electron Cooler (BINP, V.Parkhomchuk and the team)

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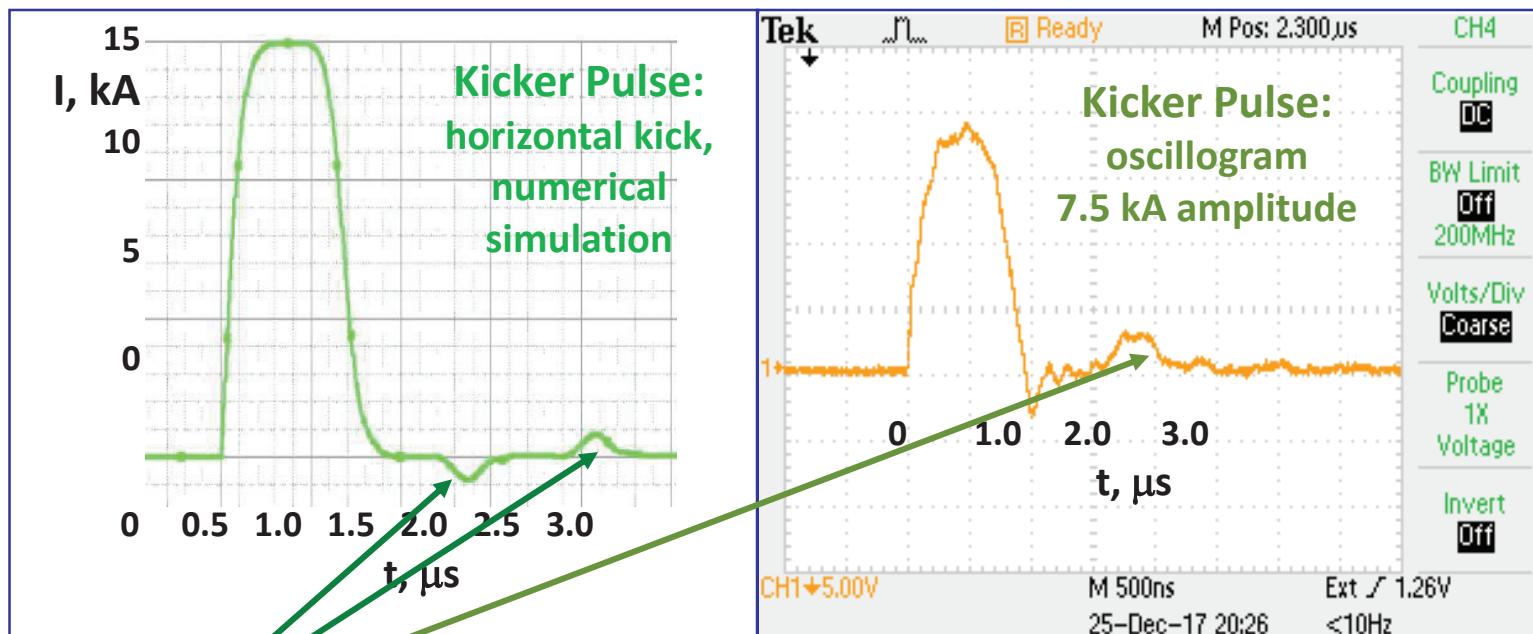
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2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: Injection and Stacking with RF Barrier Voltage at Electron Cooling Application

Kicker Pulse



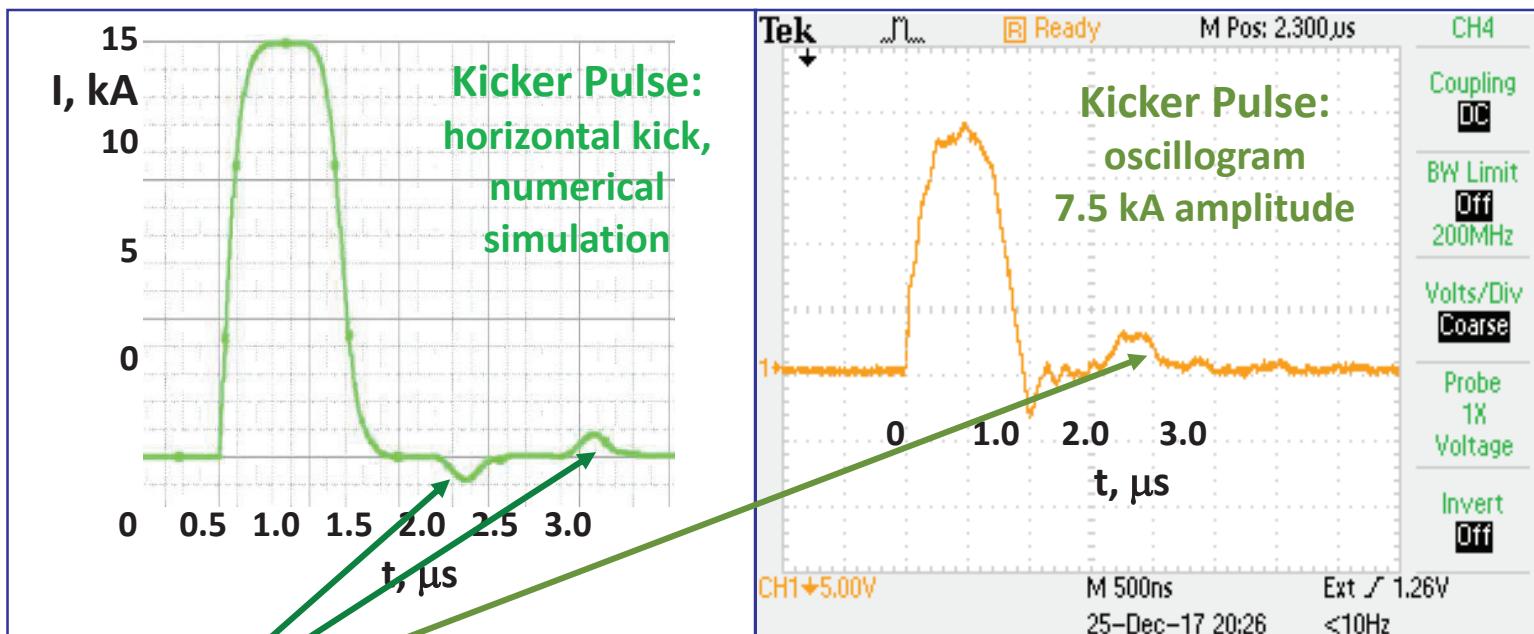
Ripples of the kicker pulse, which “heat” the ion stack

2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: Injection and Stacking with RF Barrier Voltage at Electron Cooling Application

Why do we need
HV e-cooling...

Kicker Pulse



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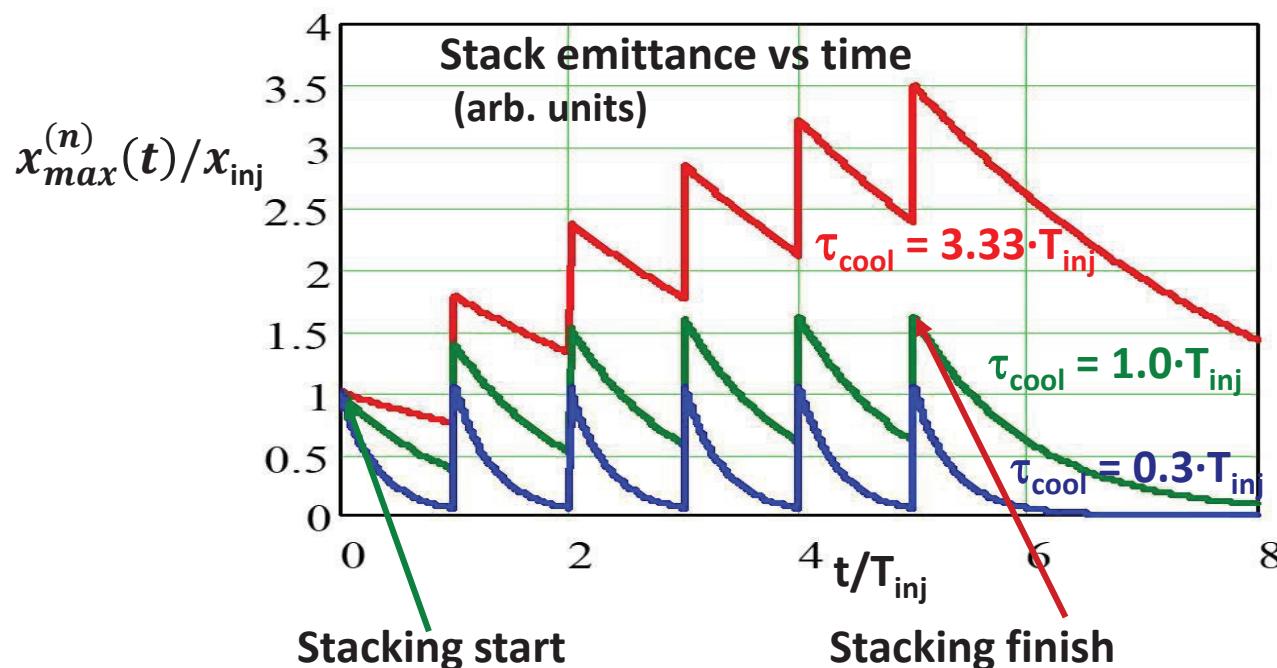
2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: Injection and Stacking with RF Barrier Voltage at Electron Cooling Application

Stack Transverse size vs Time (recurrent formula)

$$x_{max}^{(n)}(t) = x_{inj} \cdot \left(\sum_{m=1}^n e^{-(m-1)\alpha} (1+r)^{m-1} \right) \cdot e^{-\alpha(t-n)}, \quad (n-1) \leq t \leq n$$

$\alpha = T_{inj}/\tau_{cool}$ is cooling decrement, T_{inj} is injection period, t is time in T_{inj} units,
 $r = \Delta x/x_{max}$ is increase of the stack x-size under action of the kicker pulse ripples.



2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

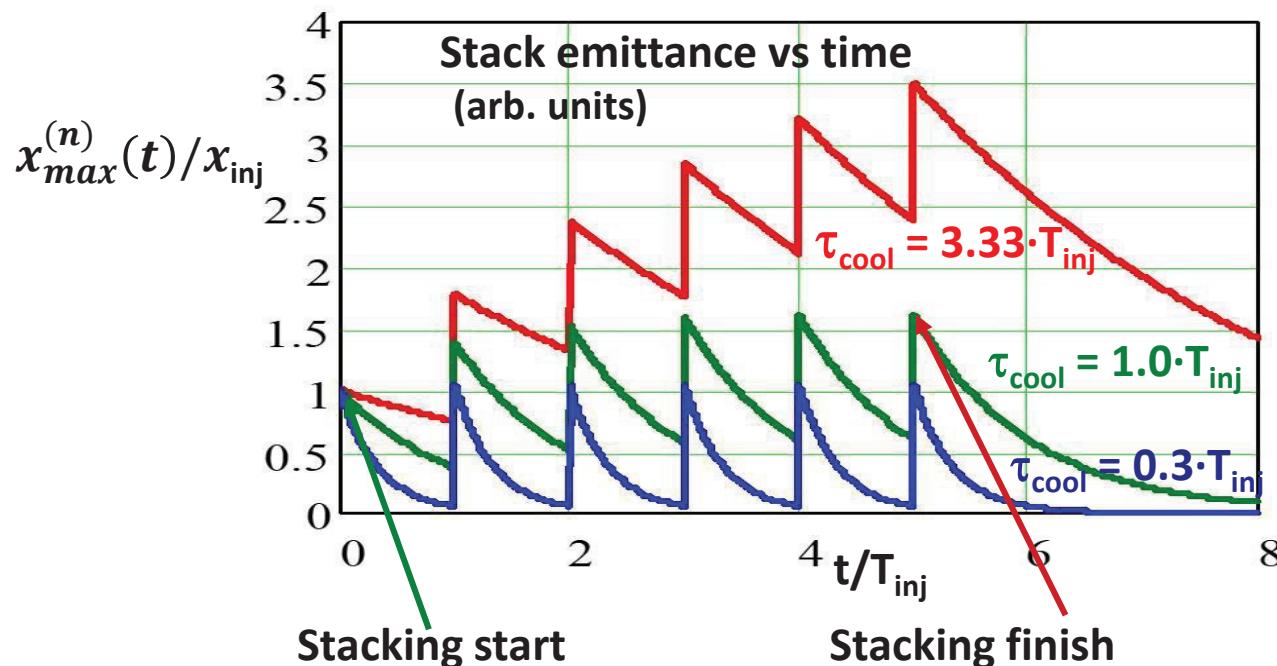
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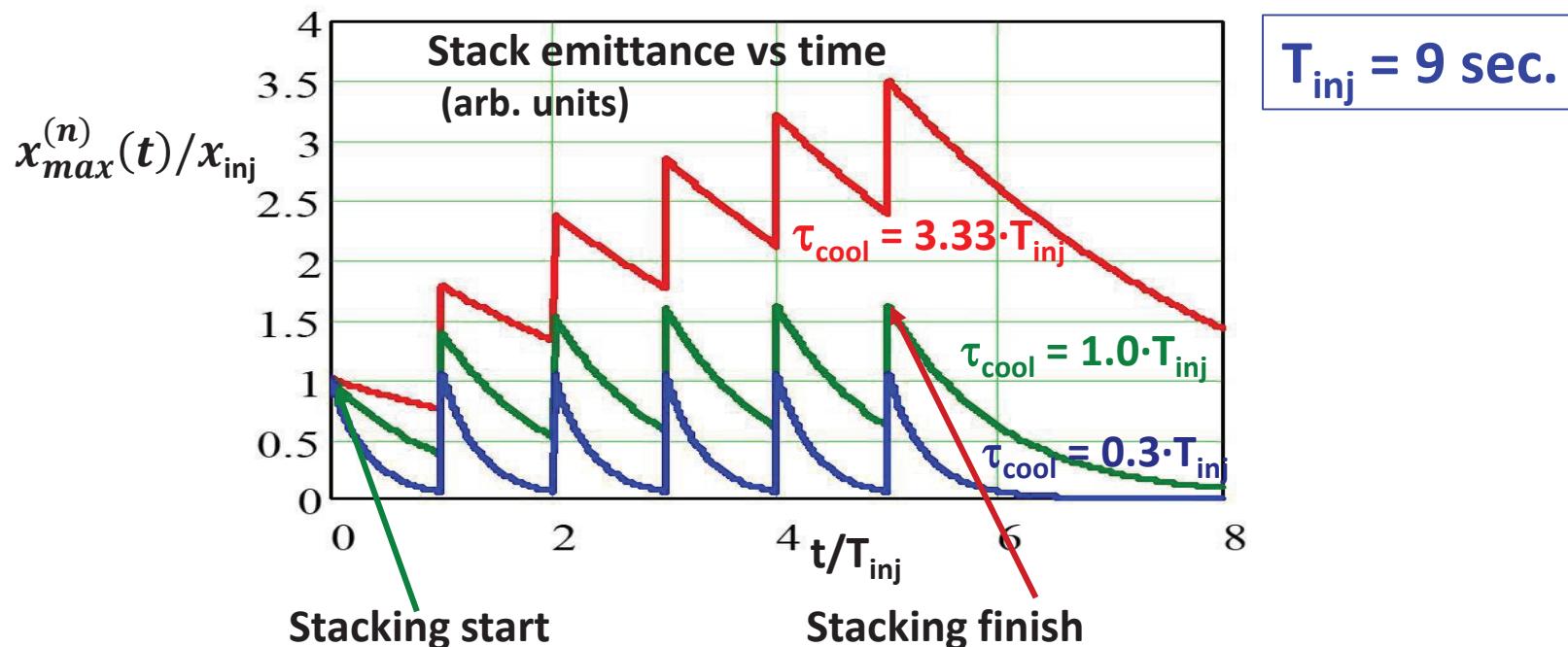
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The Collider: Injection and Stacking with RF Barrier Voltage at Electron Cooling Application

Equilibrium of electron cooling and heating by the kicker pulses

Just after $(n + 1)$ -st injection we have the beam x-size value $x^{(n+1)}$:

$$x_{max}^{(n+1)} = x_{max}^{(n)} \cdot e^{-\alpha} \cdot (1 + r) + x_{inj} \quad (1)$$

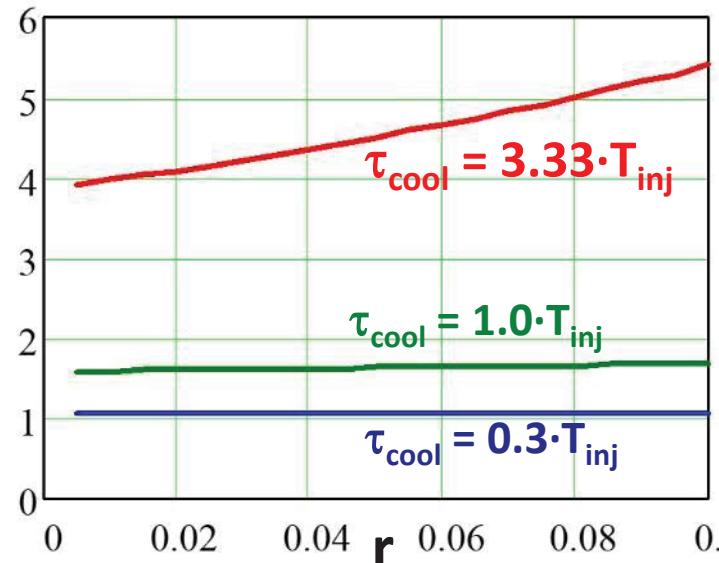
The cooled n-th stack interrupted by ripples new injected bunch

$$\text{Equilibrium condition: } x_{equi} = x_{max}^{(n+1)} = x_{max}^{(n)} \quad (2)$$

$$\text{Inserting (2) in (1) we find: } x_{equi} = \frac{x_{inj}}{1 - (1 + r) \cdot e^{-\alpha}} \quad (3)$$

It follows from (3): $r < e^{-\alpha} - 1$

$$x_{equi}/x_{inj}$$



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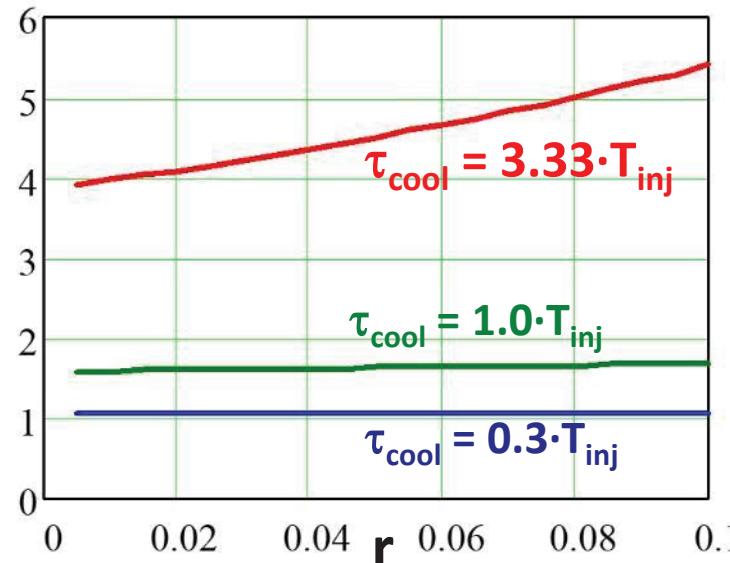
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2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: Injection and Stacking with RF Barrier Voltage at Electron Cooling Application

Electron Cooling, IBS and Other Limitations

N.V.Mityanina

Ions $^{197}\text{Au}^{79+}$
at the energy
of 3 GeV/u

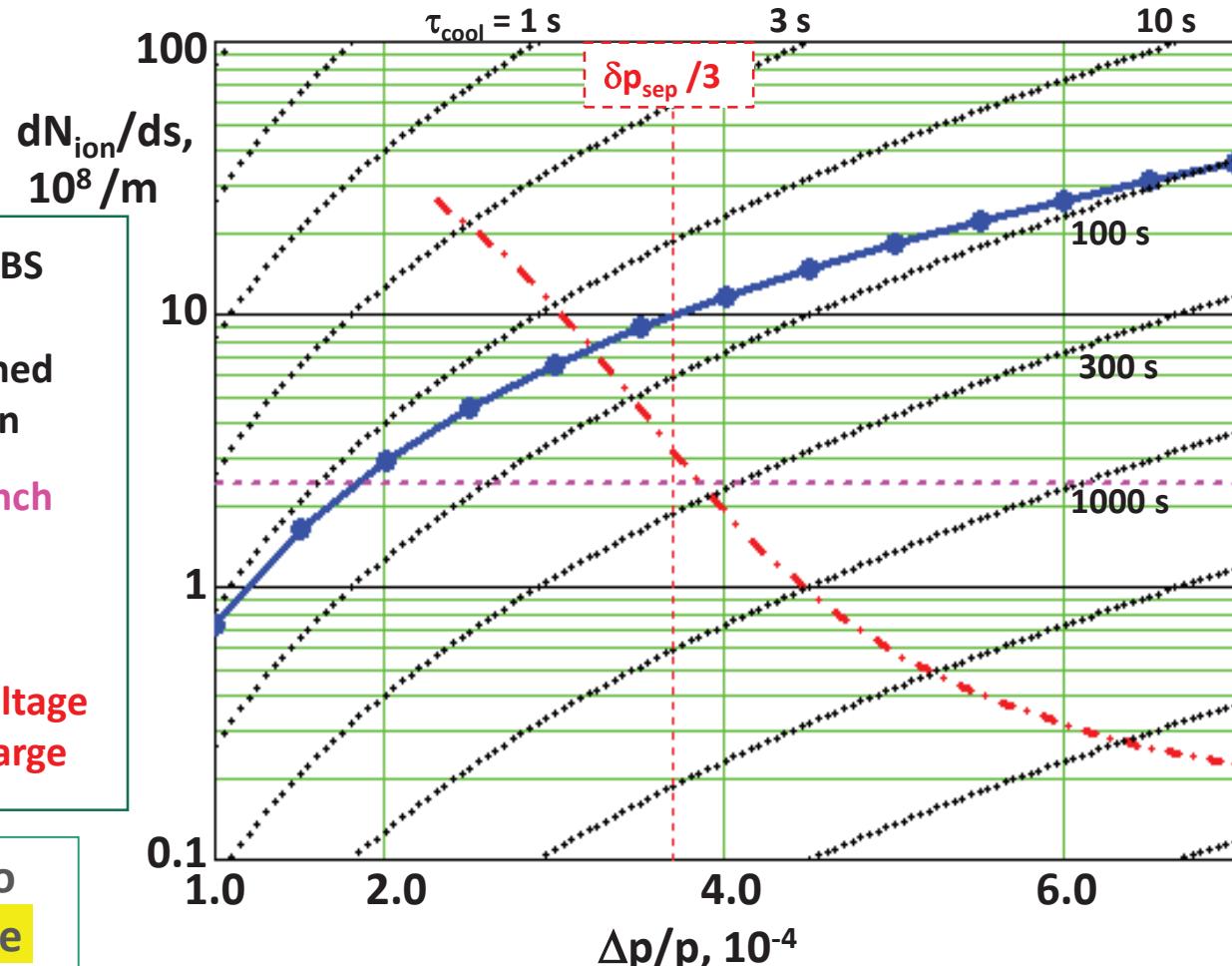
Equilibrium between IBS
and electron cooling;
beam emittance is defined
by equilibrium condition

Project ion number/bunch

Keil-Schnell Criterion
at $(Z/n)_{\max} = 70$ Ohm

Limitation by barrier voltage
and the beam space charge

Electron cooling is supposed to
be linear with $\Delta p/p$, or average
decrement over the beam.



2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: Injection and Stacking with RF Barrier Voltage
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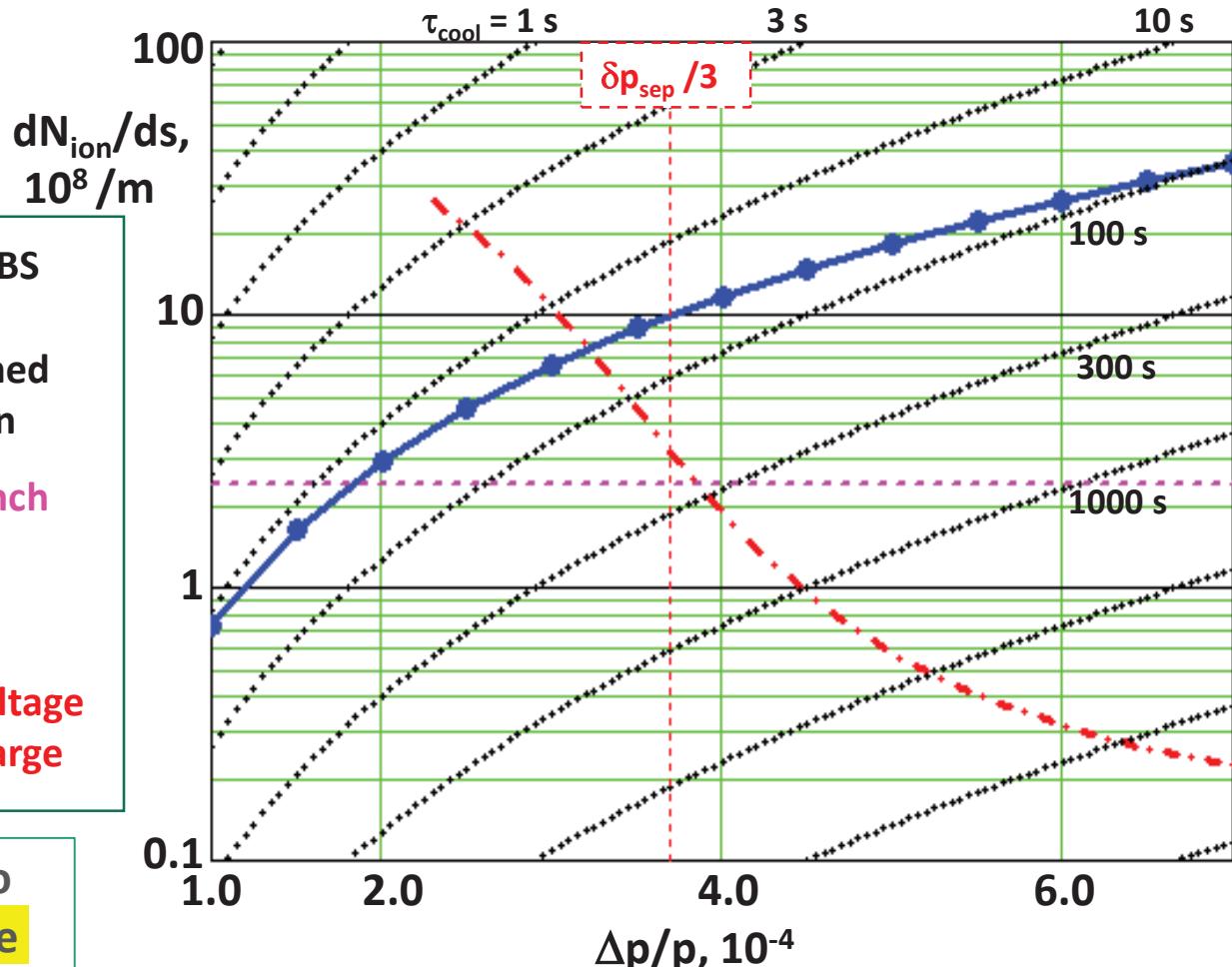
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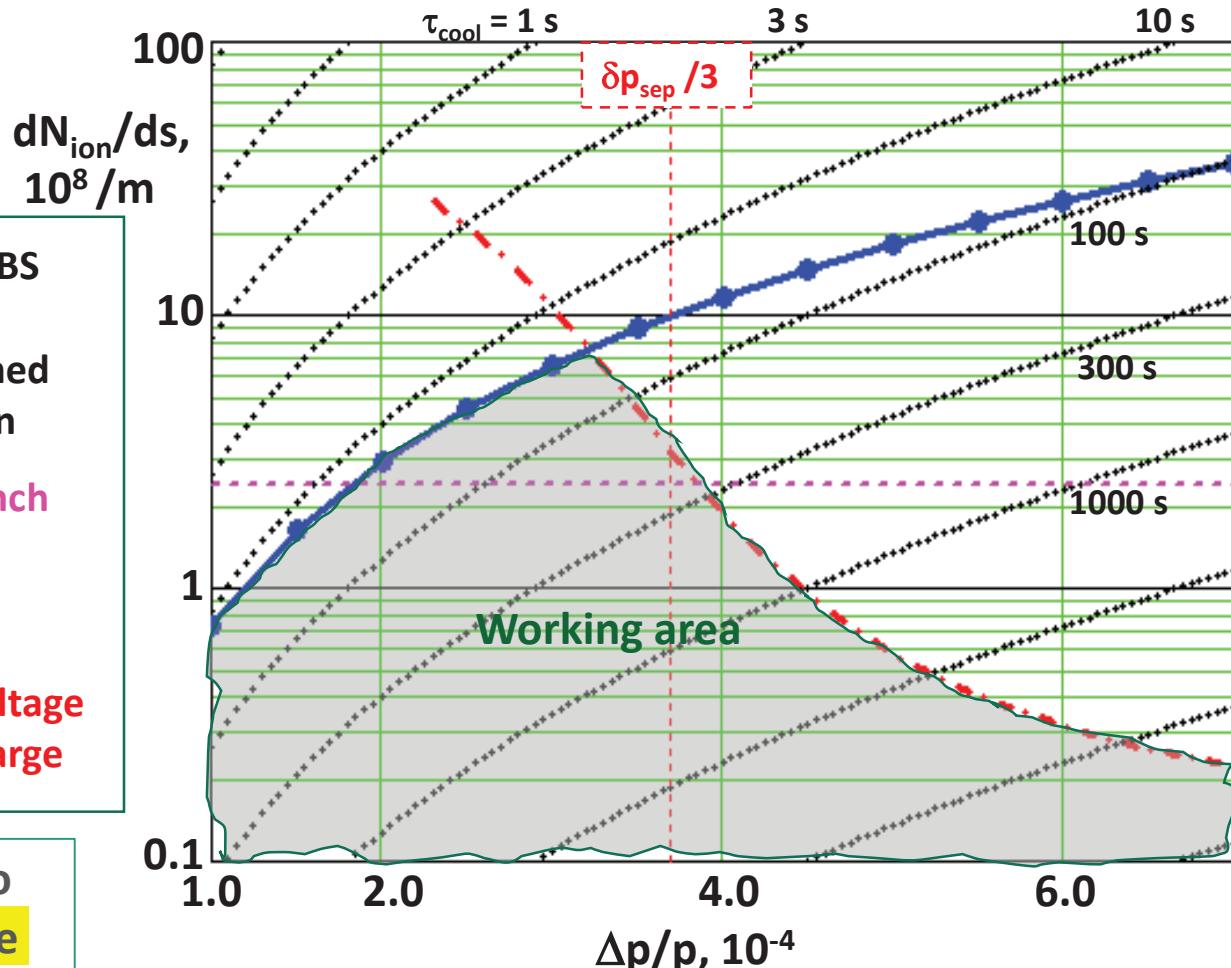
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Project ion number/bunch

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The Collider: Injection and Stacking with RF Barrier V at Electron Cooling Application

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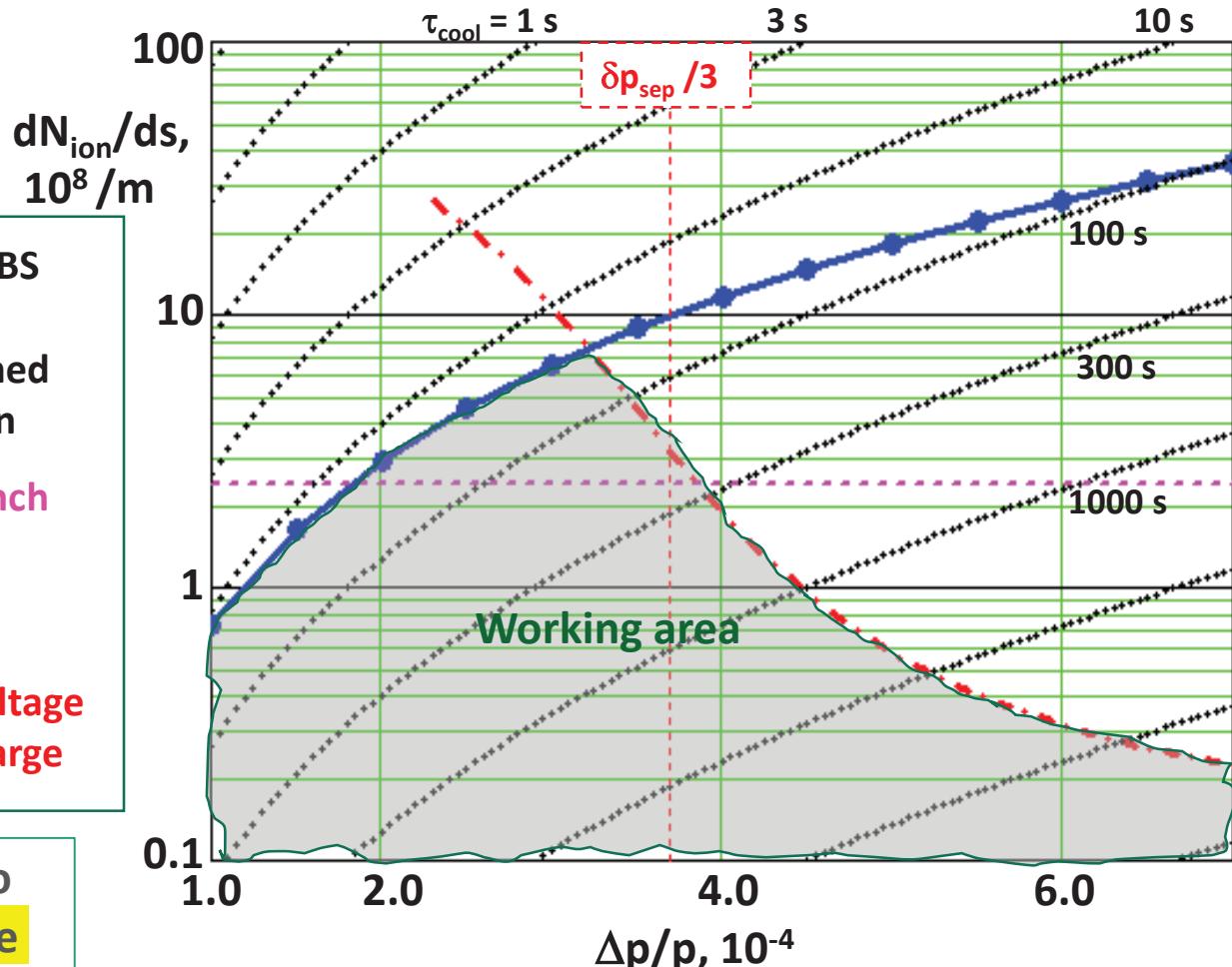
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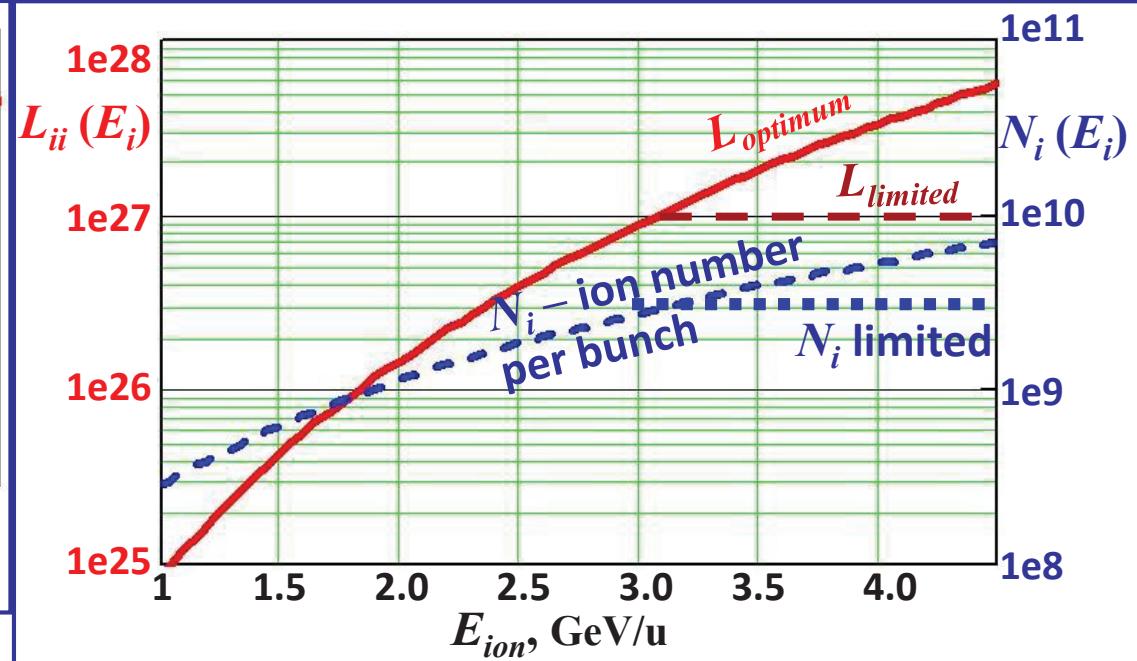
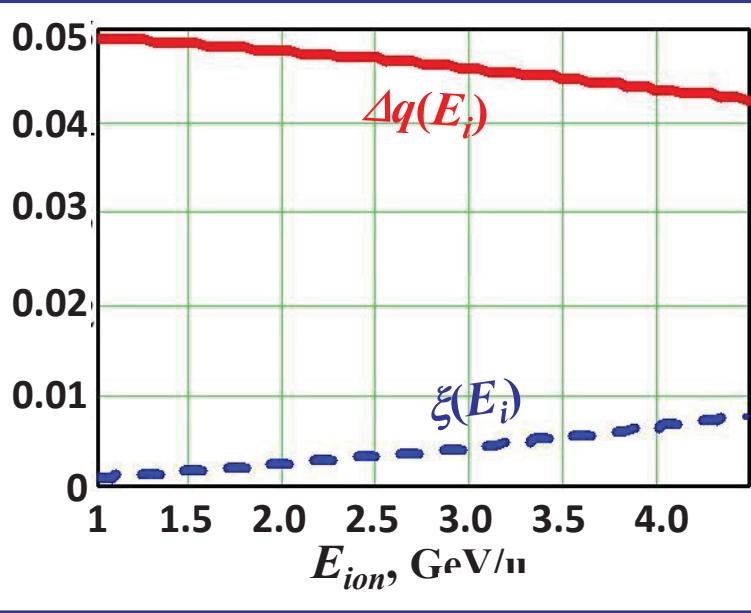


2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: HV Electron Cooler (BINP, V.Parkhomchuk and the team)

Suppression of Ion Beam Space Charge and IBS. Luminosity Optimization

An example: Optimum luminosity for the NICA Collider: $\Delta Q_{\max} = \Delta q(E_i) + \xi(E_i) = 0.05$



Using an effective electron cooling one can hope to obtain $\Delta Q_{\max} = 0.1!$

However, luminosity of NICA Collider will be limited, first of all, by intensity of injection chain...
...and MPD capability.

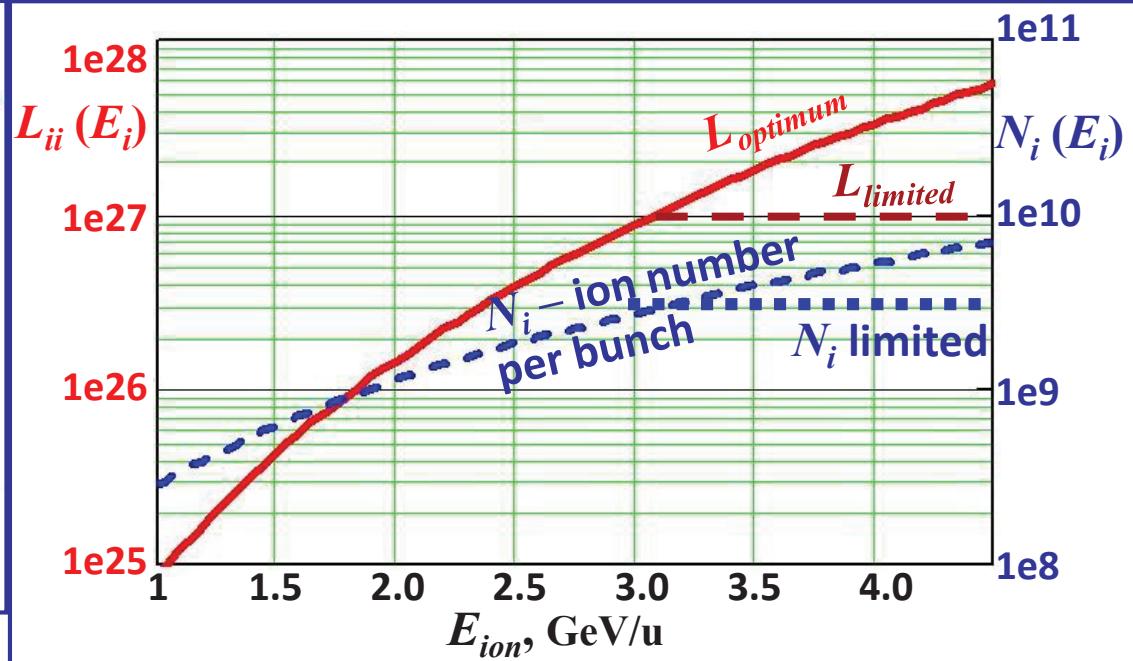
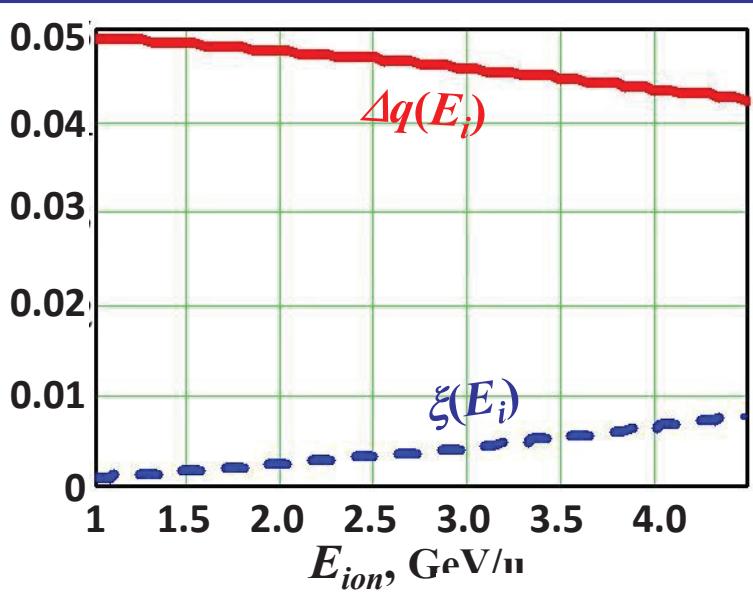
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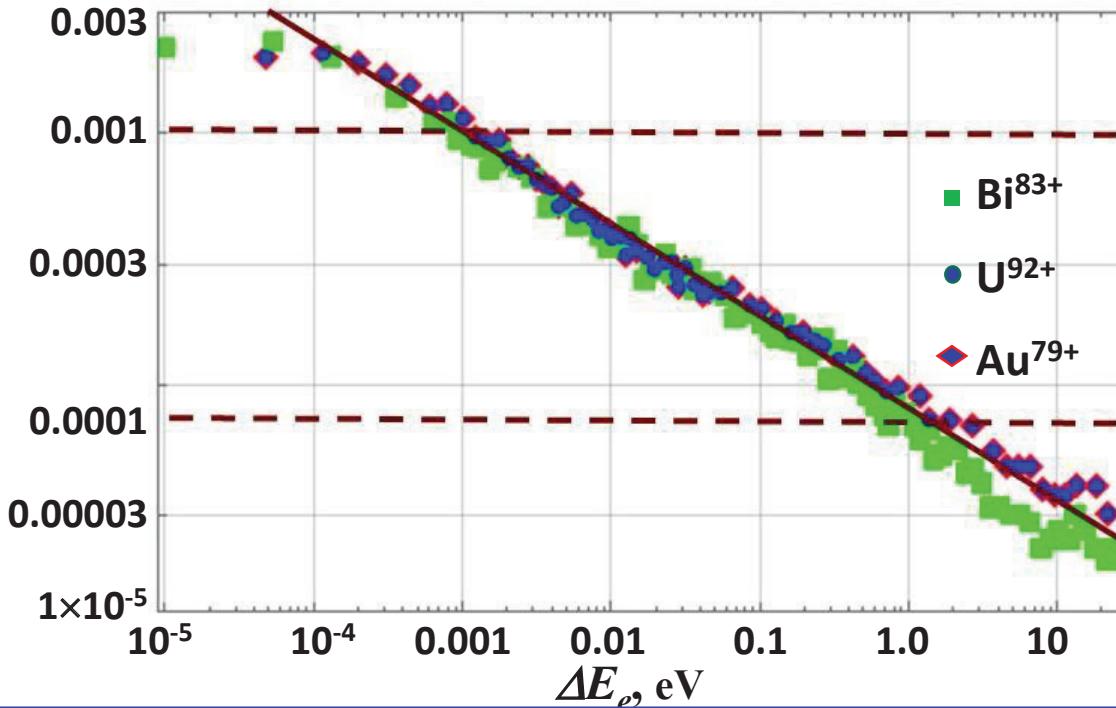
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The Collider:

Recombination in E-Cooler of Collider and its suppression



Bi⁸³⁺ A.Hoffknecht et al. Phys. Rev. A.mV.63 (2000) 012702

U⁹²⁺ W.Shi et al. Eur.Phys. J. D. 15 (2001) P. 145 – 154

Au⁷⁹⁺ Recalculated from U⁹²⁺ data

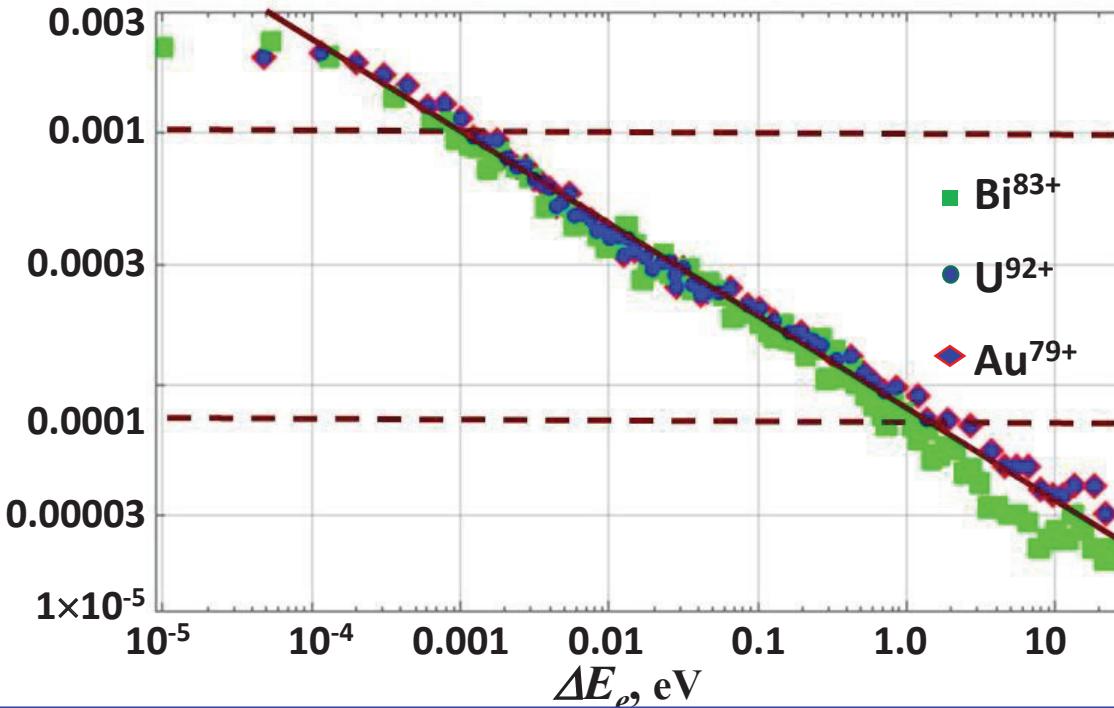
Fit equation:

$$\alpha_z(\Delta E_e) = 7 \cdot 10^{-13} \cdot Z^2 T^k, k = -0.3849673 \text{ cm}^3/\text{s}$$

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Au^{79+} Recalculated from U^{92+} data

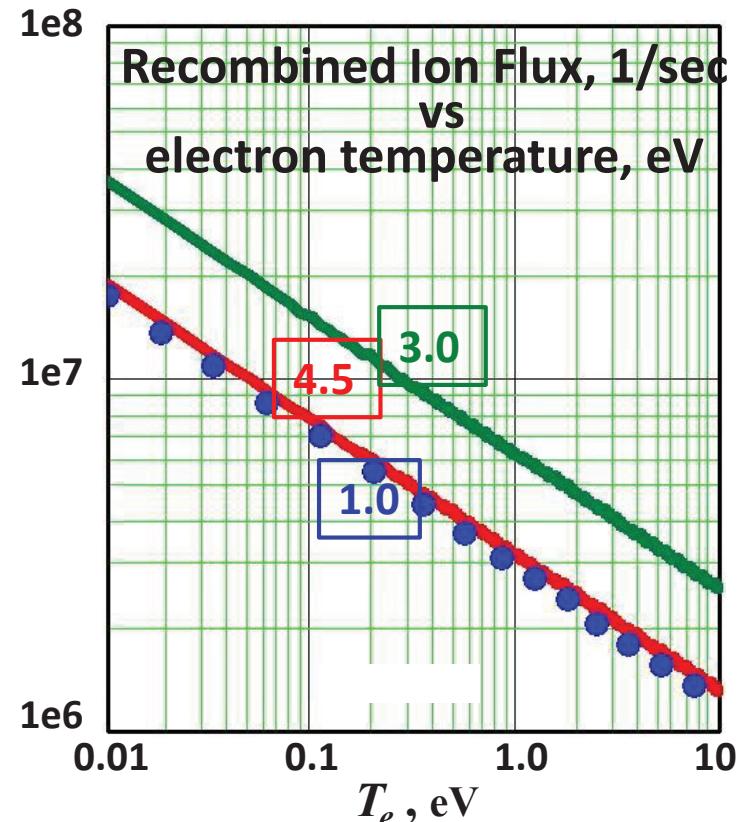
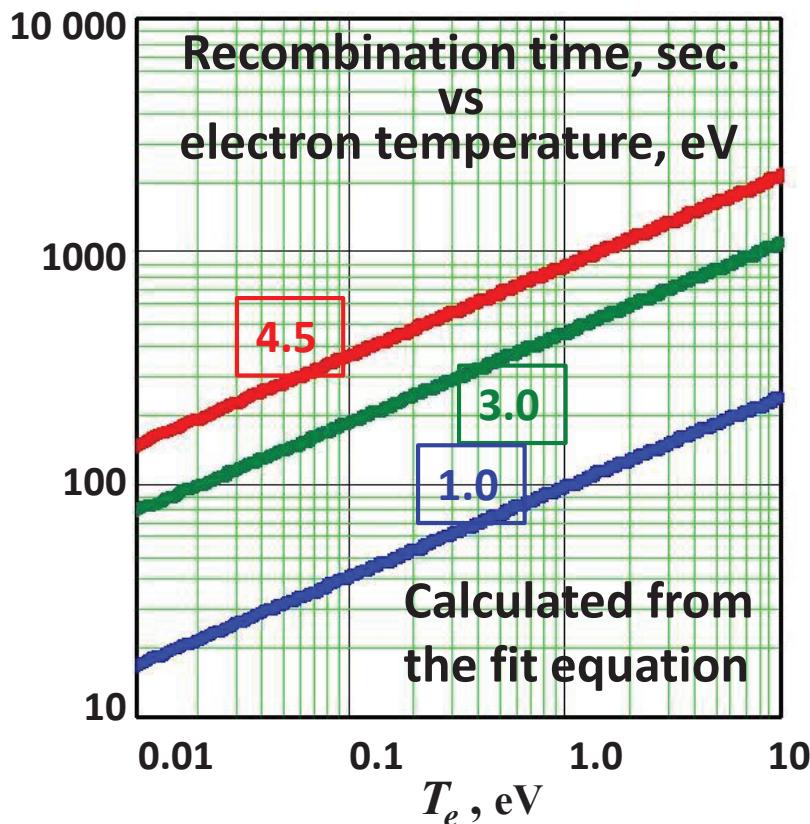
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The Collider: HV Electron Cooler (BINP, V.Parkhomchuk and the team)

Flux of The Recombined Ions and Its Suppression



$^{197}\text{Au}^{79+}$, $E_{\text{ion}} = 1 - 4.5 \text{ GeV/u}$ (shown in squares in diagrams)

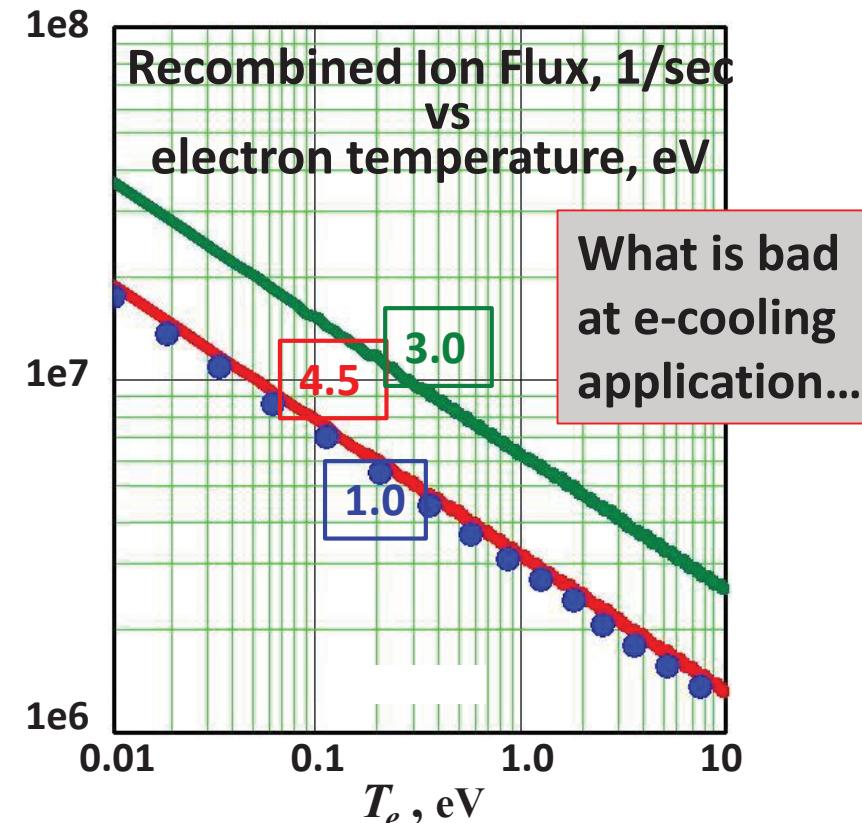
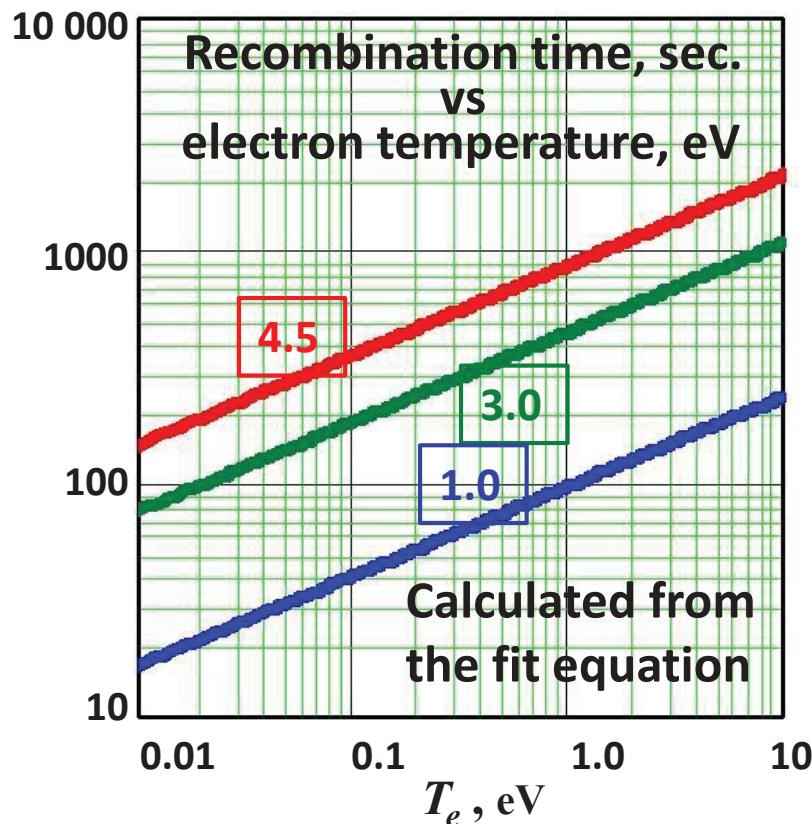
Electron cooler: $I_e = 1 \text{ A}$, $r_e = 3 \text{ mm}$, $n_e = 1.41 \times 10^8 \text{ cm}^{-3}$

cooling section: length = 6 m, solenoid field = 0.2 T

2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

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Flux of The Recombined Ions and Its Suppression

The flux of recombined ions

- 1) shortens beam lifetime
- 2) generates parasitic signal in the detector – ***collimation*** (!),
- 3) increases radiation level in vicinity of the NICA facility.



2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

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What is bad
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The flux of recombined ions can be suppressed by excitation of transverse temperature of electrons:

*P.Beller, K.Beckert, B.Franzke, C.Kozhukharov, F.Nolden, M.Steck,
NIM A V. 532 (2004) P. 427 - 432*



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NIM A V. 532 (2004) P. 427 - 432*

Strong magnet field at cooling section is necessary for reduction of recombination without decreasing cooling rate.

In the electron cooler of the NICA Collider solenoid magnetic field is

$$B_{\text{sol}} \leq 0.2 \text{ T.}$$



2. Stage II : Search for the Mixed Phase and New Physics in Heavy Ions' Collisions at NICA Collider

The Collider: HV Electron Cooler (BINP, V.Parkhomchuk and the team)

Flux of The Recombined Ions and Its Suppression in NICA Collider

$$^{197}\text{Au}^{79+}, E_{\text{ion}} = 4 \text{ GeV/u},$$

Electron cooler: $I_e = 1 \text{ A}$, $r_e = 3 \text{ mm}$, $n_e = 1.41 \times 10^8 \text{ cm}^{-3}$

$$C_{\text{ring}} = 503,04 \text{ m}$$

(not 580 m!)

cooling section: length = 6 m, solenoid field = 0.2 T

$T_e, \text{ eV}$	$\tau_{\text{cool}}, \text{ sec}$	$\tau_{\text{rec1}}, \text{ sec}$	$\tau_{\text{rec2}}, \text{ sec}$
0.1	18.45	87.6	295
1.0	21.91	209.9	717
10	28.8	1153	1740
100	40.35	4843	2800

V.V.P. Formula

Fit of experimental data (slide 24)

M. Bell and J.S. Bell. Particle
Accelerators. 1982. V. 12. P. 49-52

Flux of The Recombined Ions and Its Suppression in NICA Collider

Electron cooler:

$^{197}\text{Au}^{79+}$, $E_{\text{ion}} = 4 \text{ GeV/u}$,

$I_e = 1 \text{ A}$, $r_e = 3 \text{ mm}$, $n_e = 1.41 \times 10^8 \text{ cm}^{-3}$

cooling section: length = 6 m, solenoid field = 0.2 T

$C_{\text{ring}} = 503,04 \text{ m}$
(not 580 m!)

T_e , eV	τ_{cool} , sec	τ_{rec1} , sec	τ_{rec2} , sec
0.1	18.45	87.6	295
1.0	21.91	209.9	717
10	28.8	1153	1740
100	40.35	4843	2800

V.V.P. Formula

Fit of experimental data (slide 24)

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What about stochastic cooling?

Both methods are complimentary and will be applied to *ion beams* of the NICA Collider (Poster reports of Konstantin Osipov TUPS17 and Ivan Gorelyshev TUPS18, both JINR, dedicated to stochastic cooling application to NICA have been presented on Tuesday).

Concerning **the proton mode** it is much more complicated problem. We loose stochastic signal level by factor $Z = 79$ and get increasing of cooling time by factor $N_{\text{proton}}/N_{\text{ion}}$. This problem is under investigation by NICA group.



Conclusion

The presented talk describes the present status of the NICA project with focus on the problems related to application of electron cooling at the steps of ion injection and acceleration in the NICA Booster and ion storage in the NICA Collider.

More detailed analysis of ion storage with barrier RF voltage application (including “moving” barriers) and bunched beam formation in the Collider is described in the next talk of Natalya Mityanina.

Both talks demonstrate that electron cooling is a key issue of the NICA project.











Thank you for your attention!
Спасибо за внимание!