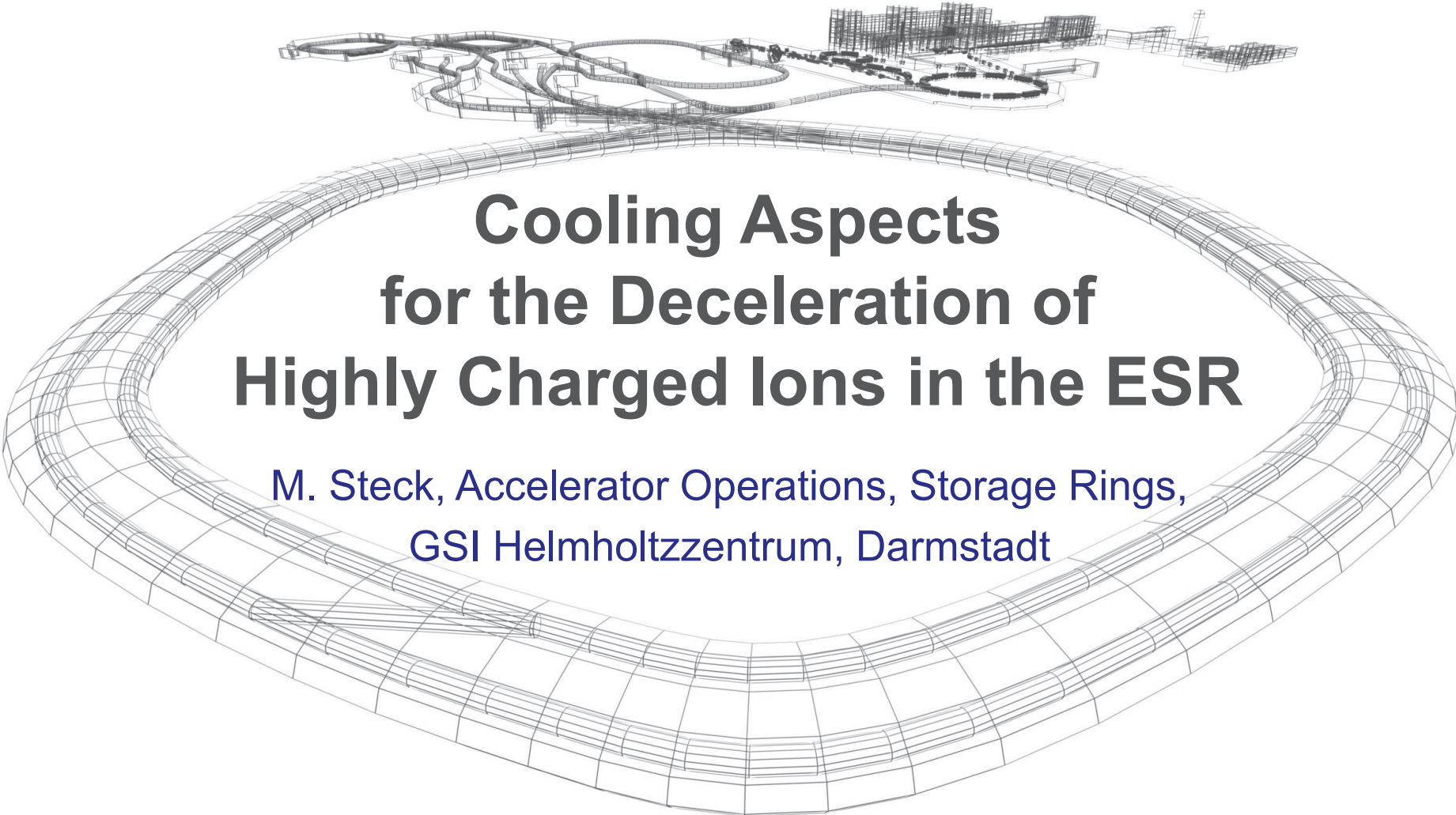


A detailed wireframe model of the ESR (Extensible Storage Ring) at GSI. The model shows a large, circular storage ring with a complex internal structure, including various bending magnets and straight sections. The ring is surrounded by a network of pipes and support structures. In the background, there are several buildings and other facilities of the GSI complex.

Cooling Aspects for the Deceleration of Highly Charged Ions in the ESR

M. Steck, Accelerator Operations, Storage Rings,
GSI Helmholtzzentrum, Darmstadt

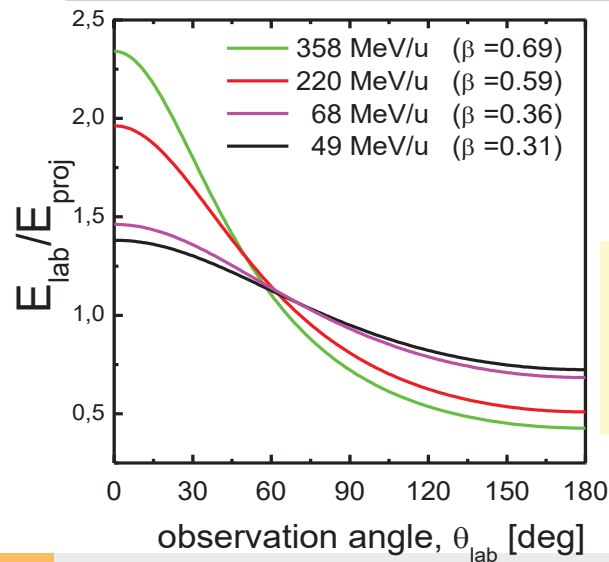
Decelerated Beams for Atomic Physics

relativistic Doppler transformation

$$E_{\text{lab}} = \frac{E_{\text{proj}}}{\gamma \cdot (1 - \beta \cdot \cos \theta_{\text{lab}})}$$

relativistic transformation of solid angle:

$$\Delta\Omega_{\text{lab}} = \frac{\Delta\Omega_{\text{proj}}}{\gamma^2 \cdot (1 - \beta \cdot \cos \theta_{\text{lab}})^2}$$

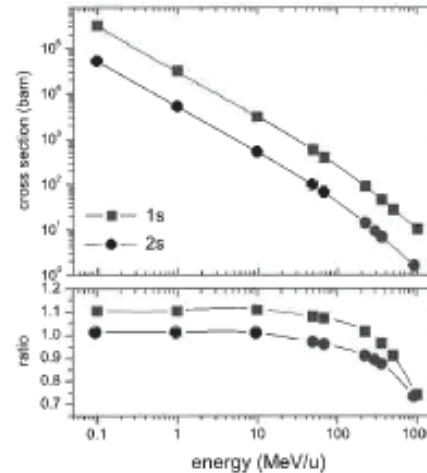


E_{lab} : photon energy in lab system

E_{proj} : photon energy in emitter system

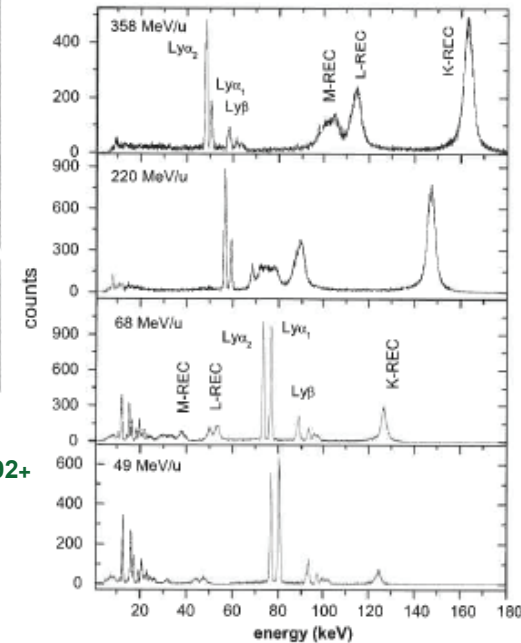
deceleration reduces:

- dependence on velocity and observation angle θ_{lab}



radiative recombination U^{92+}

J. Eichler, T. Stöhrer / Physics Reports 439 (2007) 1–99



X-rays of REC $\text{U}^{92+} \rightarrow \text{N}_2$

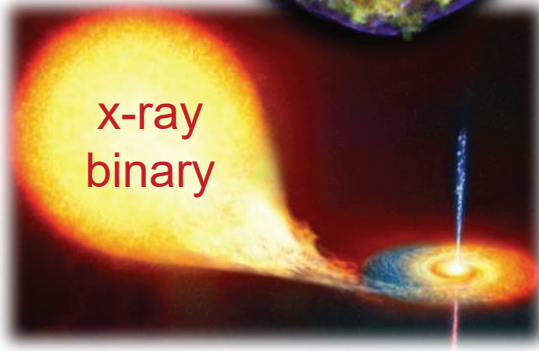
deceleration:

- provides increased reaction cross sections
- opens new reaction channels (transitions, polarization)

Decelerated Beams for Nuclear Astrophysics



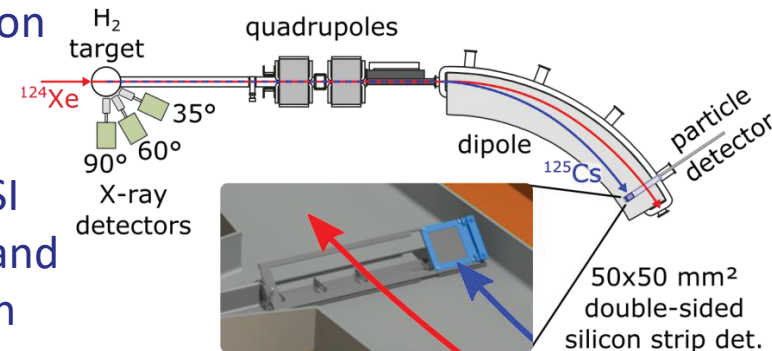
supernova
remnant



x-ray
binary

To investigate the origin of elements in stars, nuclear astrophysics aims for challenging reaction studies on rare ion beams.

Heavy ion storage rings at GSI provide unique possibilities and unrivalled conditions for such experiments.

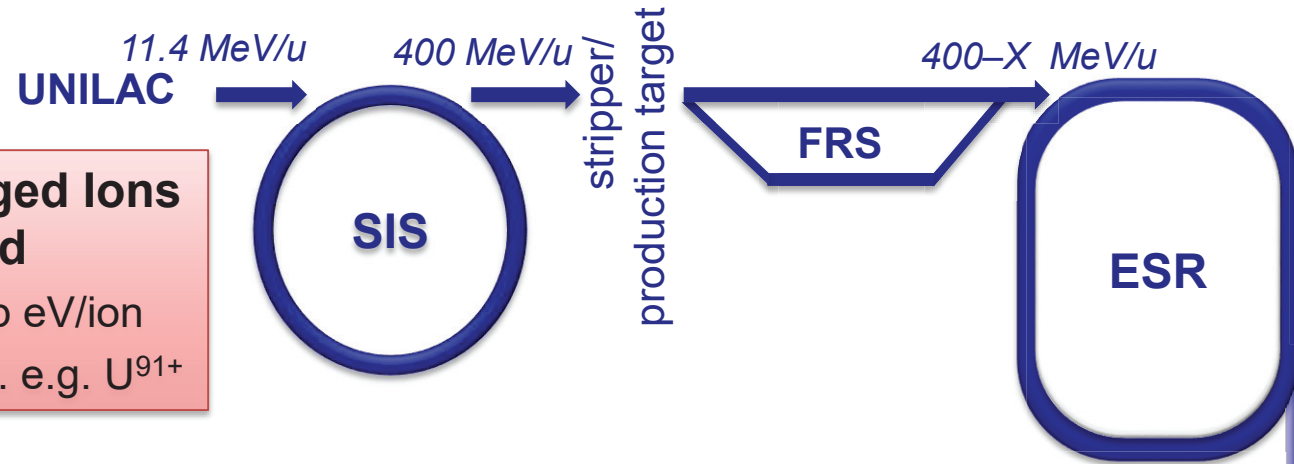


Experimental setup of the reaction study $^{124}\text{Xe}(p,\gamma)$ in the ESR at beam energies as low as 5.5 MeV/u.

- high energy production & separation of radioactive beams in FRS
 - most efficient and versatile technique available
- deceleration to energies below 10 MeV/u
 - access to the famous Gamow window relevant for nuclear physics of stars
- cooled beam in combination with a thin gas jet target
 - inverse kinematics studies at unmatched energy resolution
- storage and recycling of the rare ion beam
 - extremely efficient technique for studies on beams of limited intensity

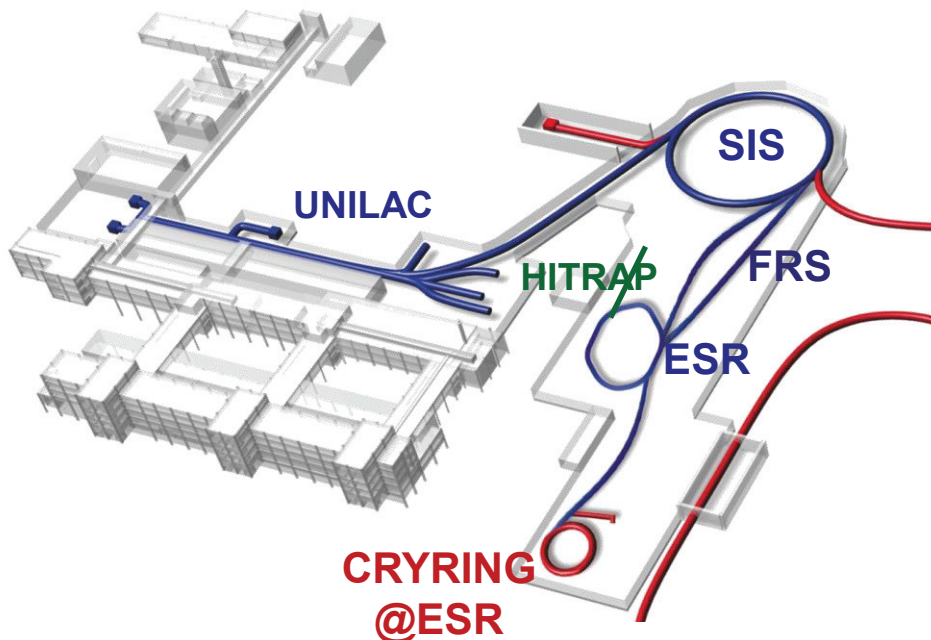
Jan Glorius

HITRAP and CRYRING@ESR



Slow, heavy, highly charged ions stored and well controlled

Energy range 10 MeV/u to sub eV/ion
 10^5 to 10^7 highly charged ions. e.g. U^{91+}



4 K

HITRAP

linear decelerator

4 MeV/u

4 MeV/u ... 100 keV/u

CRYRING@ESR

ring decelerator

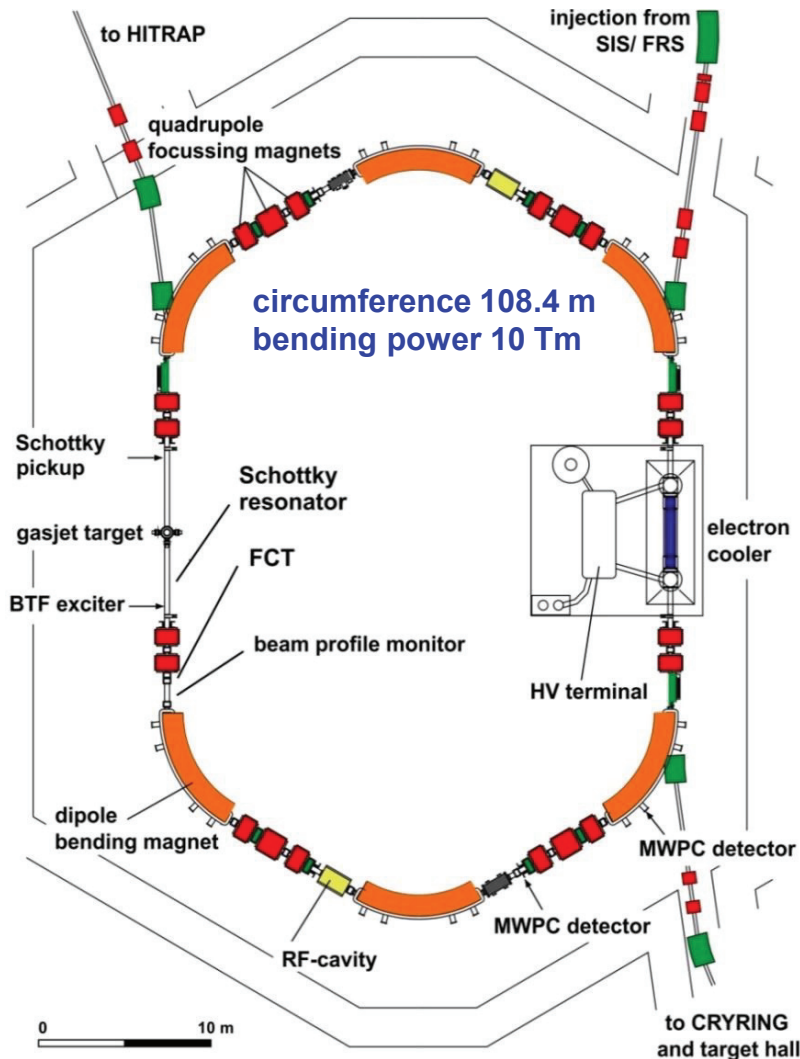
4–10 MeV/u

Goals for Deceleration in the ESR



- **good beam quality of stored beam $\Rightarrow$ high precision experiments**
- **high luminosity with internal target for internal experiment**
- **long lifetime of stored beam $\Rightarrow$ vacuum pressure 10^{-11} mbar or below**
- **fast extraction at 4 MeV/u for HITRAP, defined by the HITRAP linac**
- **fast extraction below ~ 10 MeV/u for CRYRING, defined by bending power**
- **slow extraction employing charge changing processes (cooler, target)**
- **maximum intensity (particles per time interval) to HITRAP/CRYRING**
i.e. large particle number, short cycle time

The Heavy Ion Storage Ring ESR



Fast injection (stable ions / RIBs)

Stochastic cooling (≥ 400 MeV/u)

Electron cooling (3 - 430 MeV/u)

Laser cooling (C^{3+} 120 MeV/u)

Internal gas jet target

Deceleration (down to 3 MeV/u)

Fast extraction (HITRAP/CRYRING@ESR)

Slow (resonant) extraction

Ultralow extraction (charge change)

Beam accumulation

Multi charge state operation

Schottky mass spectrometry of RIBs

Isochronous mode (TOF detector)

Beam Cooling at the ESR

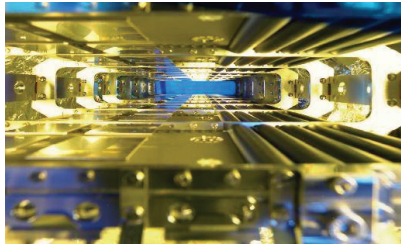
stochastic pre-cooling on the injection orbit

energy 400 (-550) MeV/u

bandwidth 0.8 GHz (range 0.9-1.7 GHz)

$\delta p/p = \pm 0.35\%$ $\rightarrow$ $\delta p/p = \pm 0.01\%$

$\epsilon = 10\ \mu\text{m}$ $\rightarrow$ $\epsilon = 2\ \mu\text{m}$



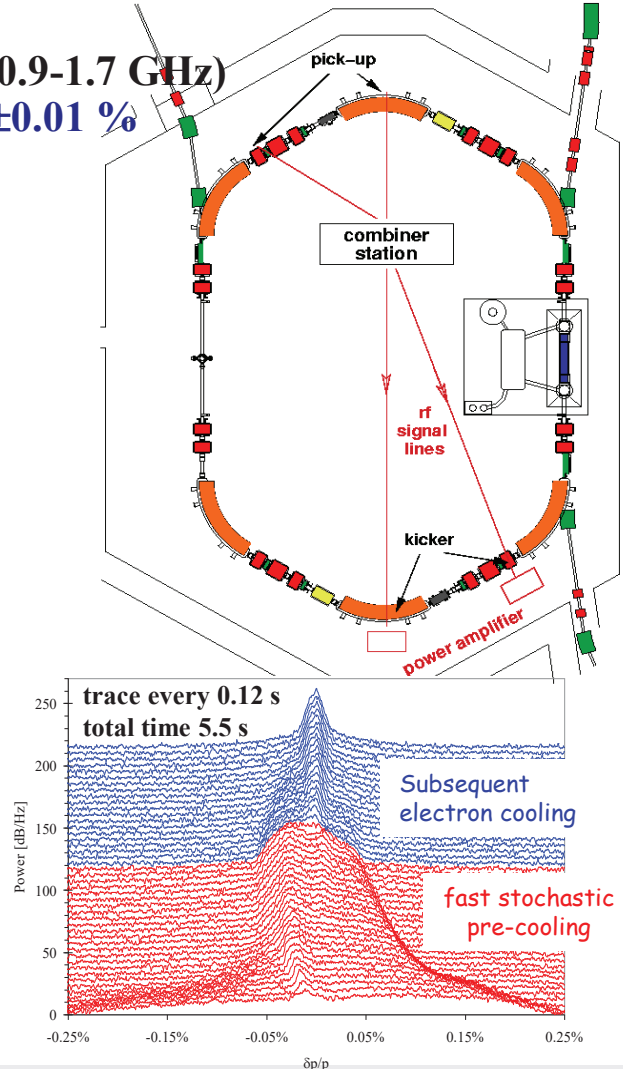
electrodes



combiner station



power amplifiers



electron cooling

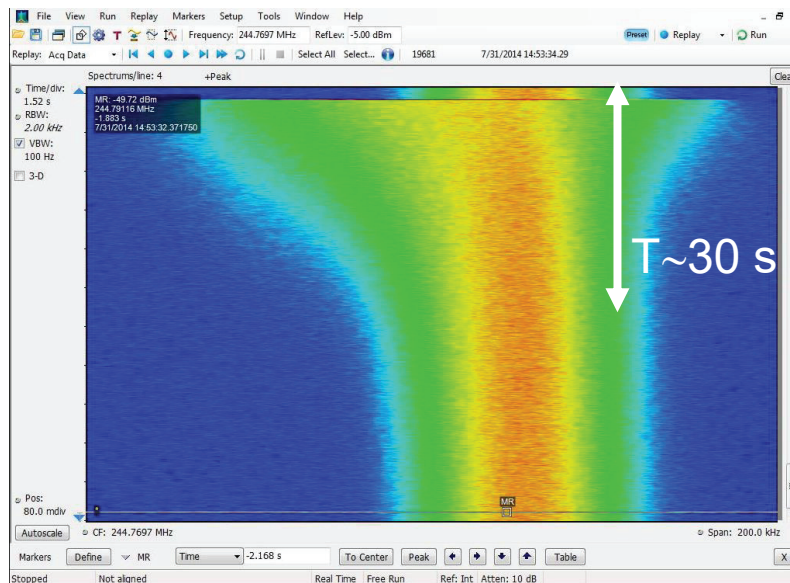


electron energy 1.6 – 250 keV
electron current 0.001 – 1 A
beam diameter 50.8 mm
magnetic field 0.01 - 0.2 T
collection efficiency 0.9998
transverse temperature 0.1 eV
longitudinal temperature $\sim 0.1\ \text{meV}$
vacuum $1 \times 10^{-11}\ \text{mbar}$

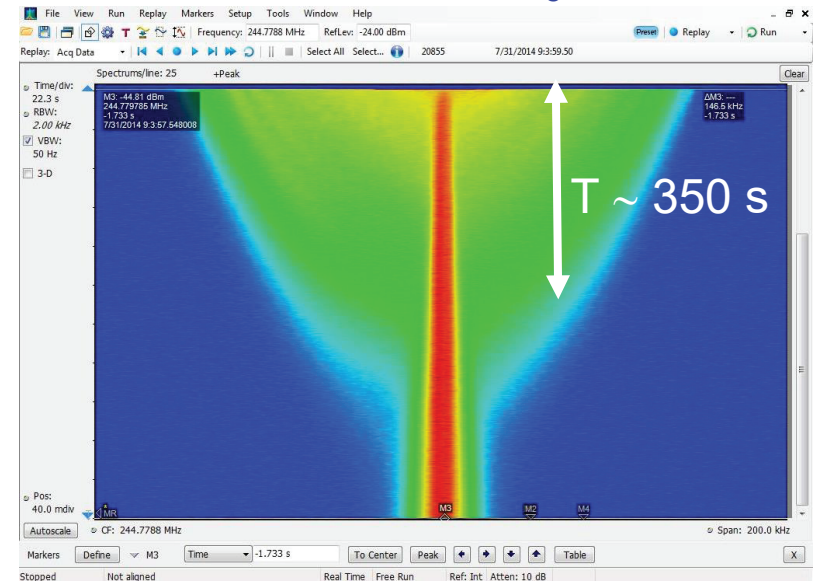
Cooling after Injection

- The injection energy of the ESR is flexible (30- 500 MeV/u)
- the production rate of highly charged heavy ions increases with energy, but the larger energy reduction conflicts $\Rightarrow$ the production rate needs to be traded off
- electron cooling can be applied for any energy ≤ 430 MeV/u
- stochastic cooling is presently available at 400 MeV/u but could be extended to higher energies (up to 550 MeV/u)
- at 400 MeV/u stochastic cooling is initially faster than electron cooling

protons 400 MeV stochastic cooling



electron cooling ($I_e=0.25$ A)

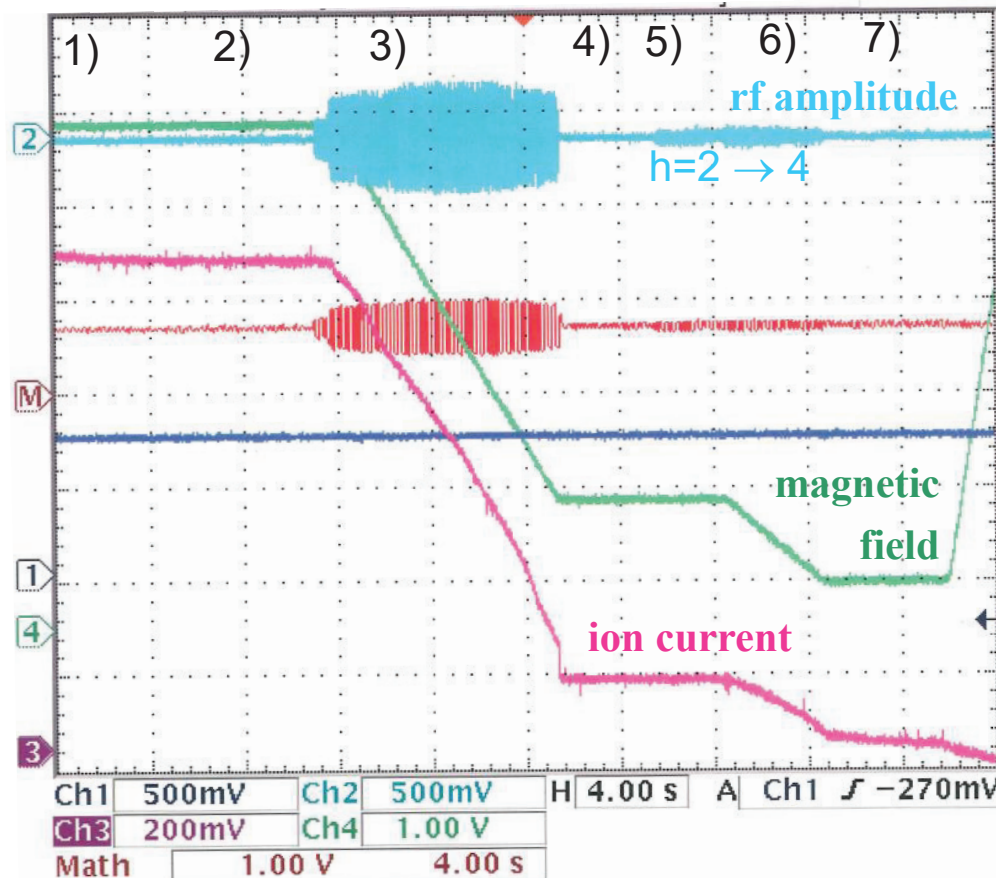


Typical ESR Deceleration Cycle

Ni^{28+} 400 $\rightarrow$ 30 $\rightarrow$ 4 MeV/u

1100 μA $\rightarrow$ 180 μA $\rightarrow$ 25 μA
45% 37 %

cycle time 45 s

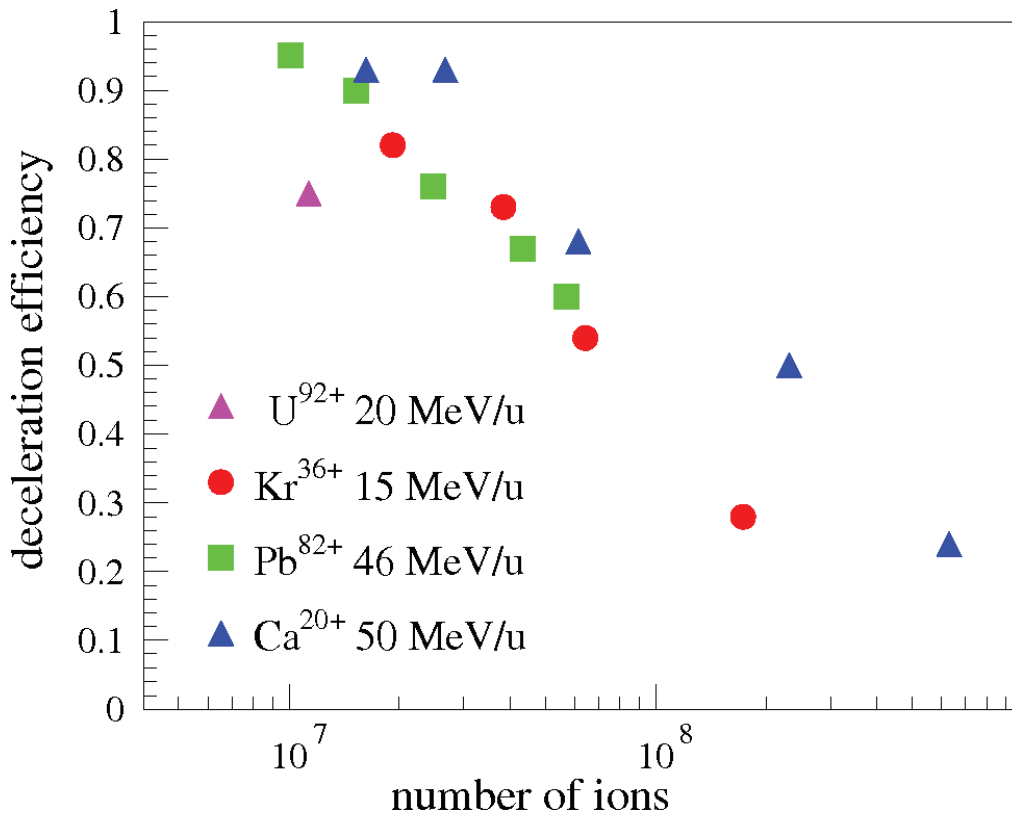


- 1) injection (at 400 MeV/u)
- 2) stochastic/electron cooling
(controlled ramp of cooler HV is difficult)
- 3) deceleration to 30 MeV/u
- 4) electron cooling
- 5) changing rf harmonic 2 $\rightarrow$ 4
- 6) deceleration to final energy
- 7) electron cooling, bunching, extraction

maximum change of p ($B\rho$): 13

Efficiency of Deceleration

Relative losses increase with increase of the initial stored particle number most likely due to Intrabeam Scattering



lower energies than 15 MeV/u require a change to a higher rf harmonic

in a bunched beam during deceleration transverse emittance is growing due to intrabeam scattering

- no cooling,
- larger emittance of bunched beam

and

due to adiabatic emittance growth ($\propto 1/\beta\gamma$)

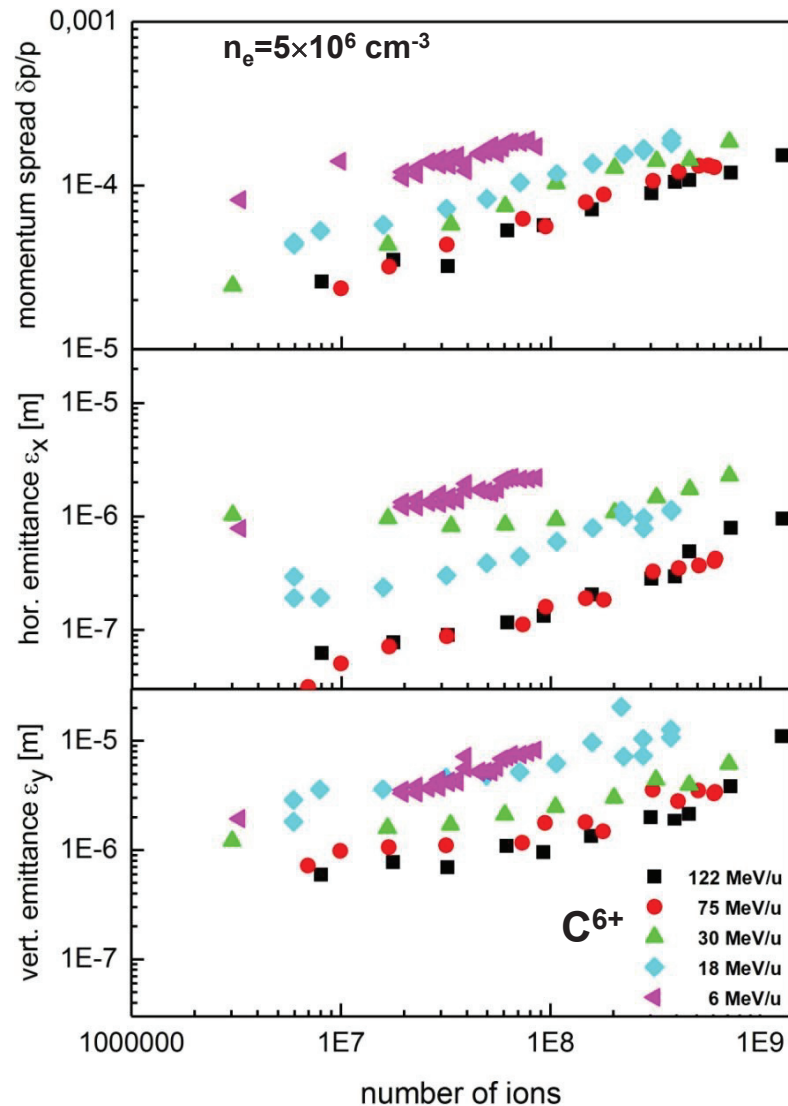
during deceleration the beam is also approaching the space charge dominated regime

$$\Delta Q_x = \frac{r_p Z^2 N g}{\pi A \beta^2 \gamma^3 B (\epsilon_x + \sqrt{\epsilon_x \epsilon_y Q_x / Q_y})}$$

$$\Delta Q_x \simeq \frac{r_p Z^2 N}{2\pi A \beta^2 \gamma^3 B \epsilon_x},$$

$$g = 1, \epsilon_x = \epsilon_y, Q_x = Q_y$$

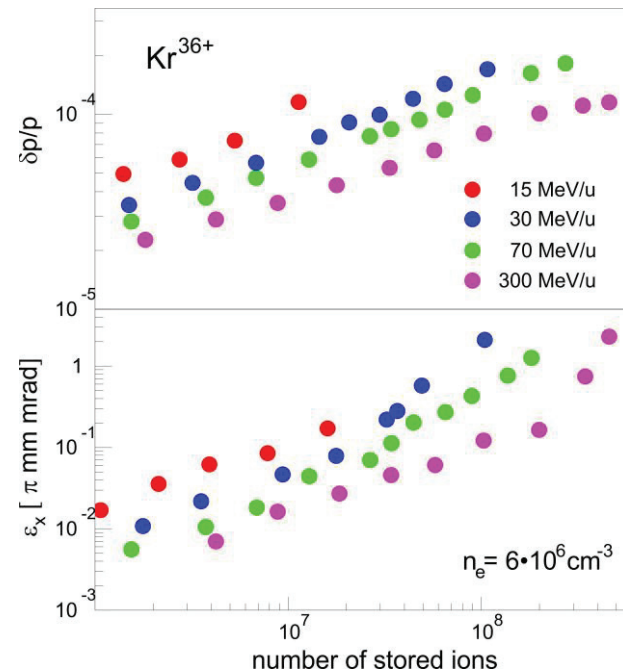
Beam Parameters of Decelerated Ions



Intrabeam Scattering rate

$$\tau_{IBS}^{-1} = \frac{Q^4 e^4}{(A m_i)^2} \cdot \frac{N}{C \varepsilon_h \varepsilon_v \delta p / p} \cdot \frac{1}{(\gamma^4 \beta^3 c^3)} \cdot 4 \pi L_C^{IBS}$$

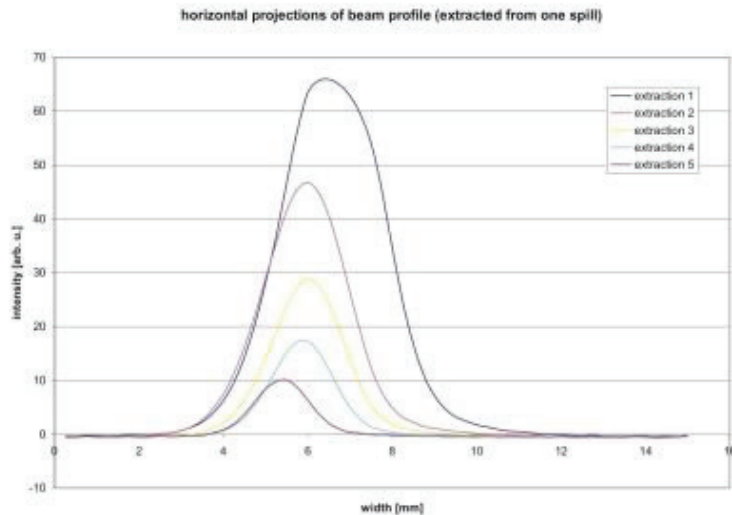
in case of equipartitioning:
emittance increases $\approx$ proportional $\beta \gamma$



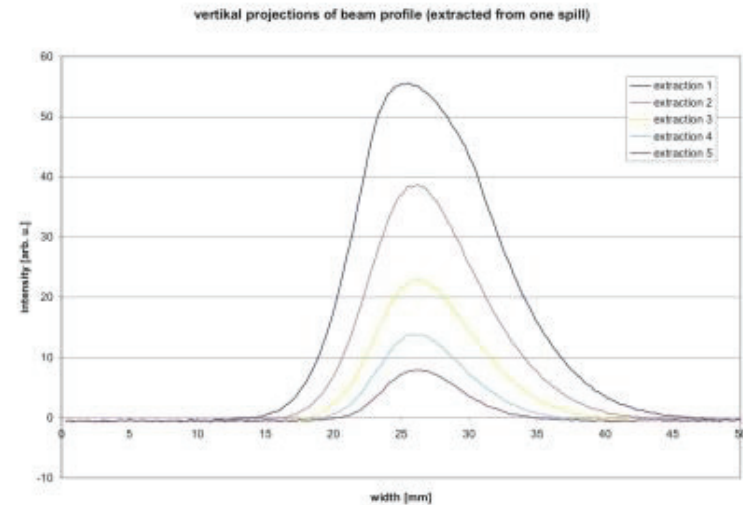
Emittance of Extracted Decelerated Beam

fast extracted beam
 Ca^{20+} 4 MeV/u

horizontal profile



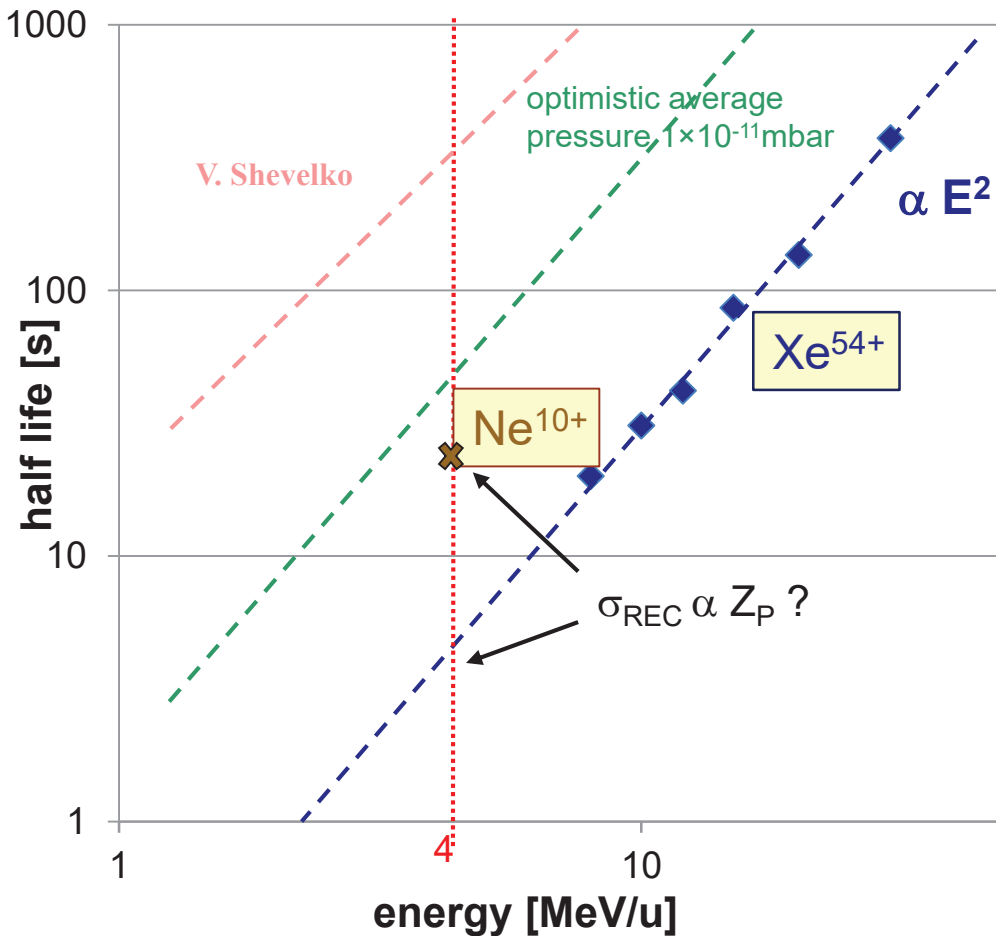
vertical profile



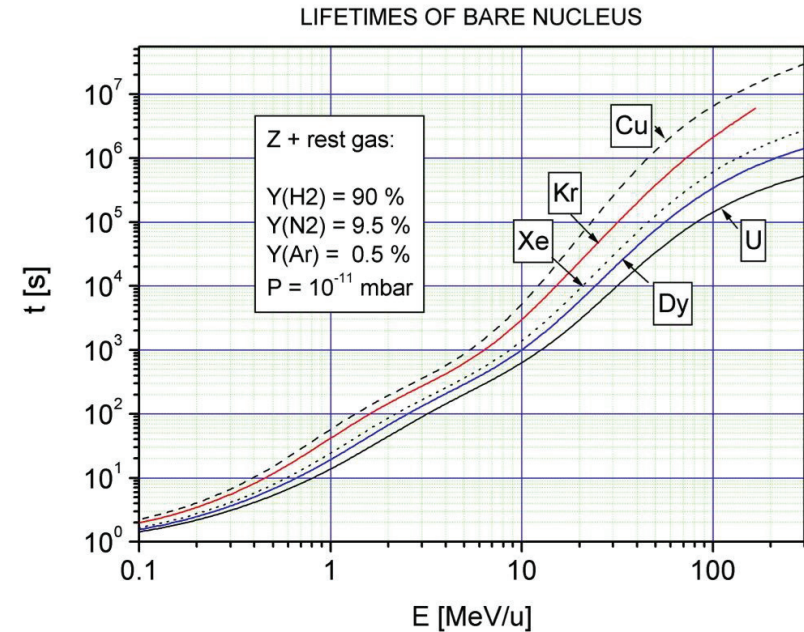
the increase of the beam emittance with intensity due to IBS
reflects in the beam size of the extracted cooled beam

Low Energy Beams – Vacuum

measurement of beam half life with
an average pressure of $p \approx 1 \times 10^{-10}$ mbar



simulation for CRYRING V. Shevelko



lifetime in the residual gas and
cooling time can be of similar value
 $\Rightarrow$ excellent vacuum and fast cooling
are crucial for low energy
highly charged ions

Deceleration to 4 MeV/u

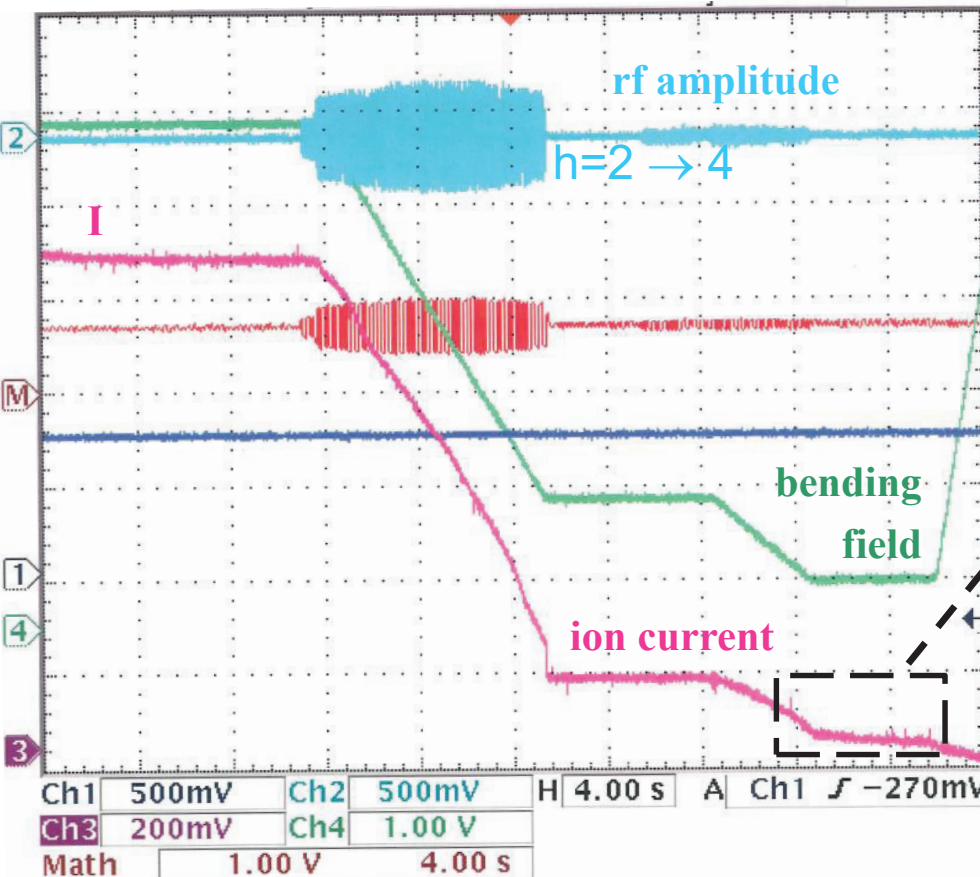
Ni^{28+} 400 $\rightarrow$ 30 $\rightarrow$ 4 MeV/u

1100 μA $\rightarrow$ 180 μA $\rightarrow$ 25 μA

45%

37 %

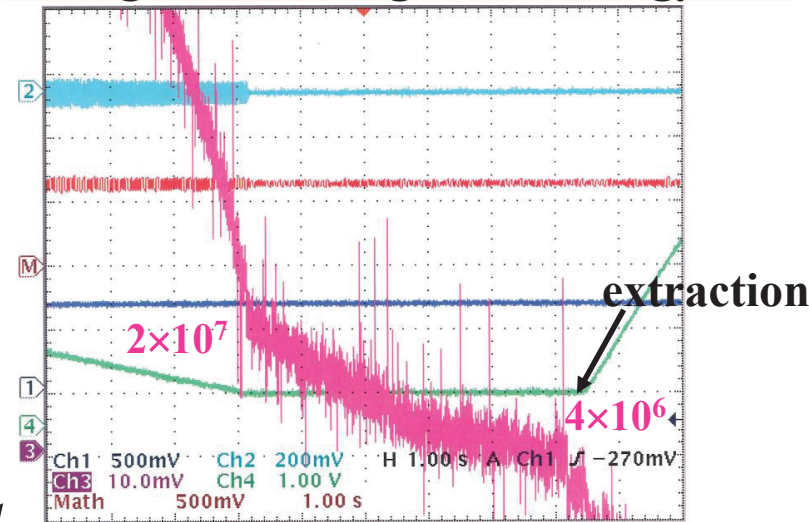
cycle time 45 s



Main losses:

End of ramp

Storage and cooling at low energy

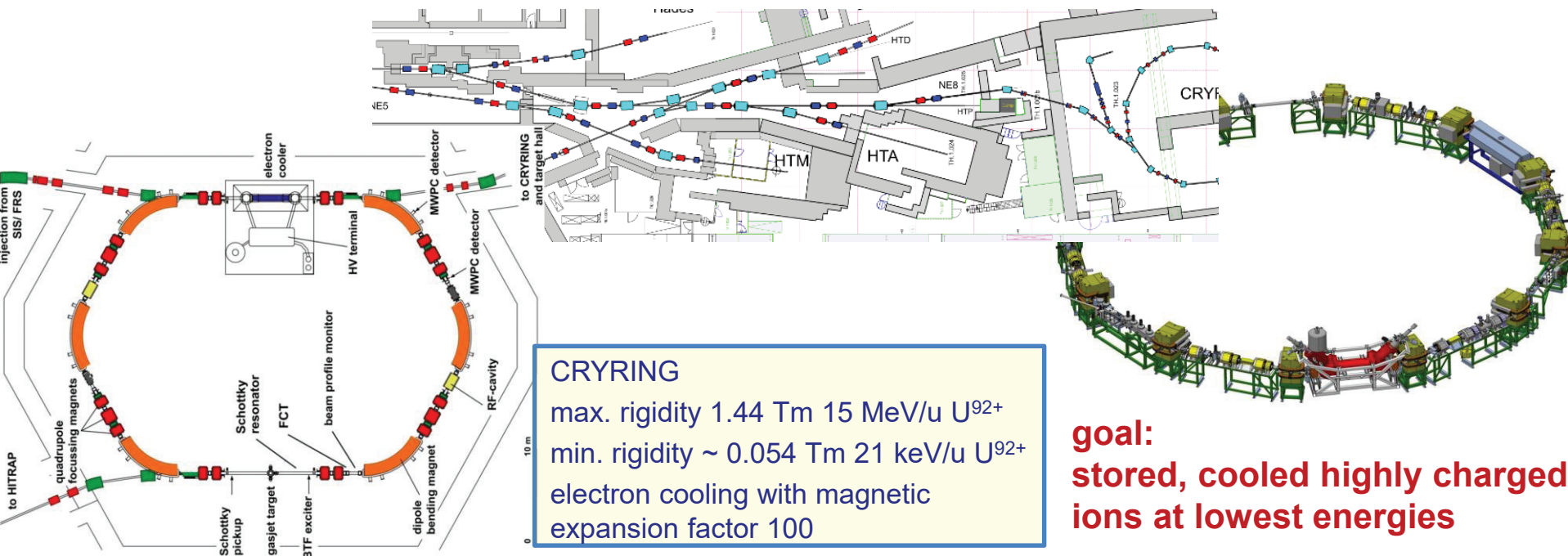


beam half life (vacuum dominated):

30 MeV/u: $T_{1/2} \approx 480$ s

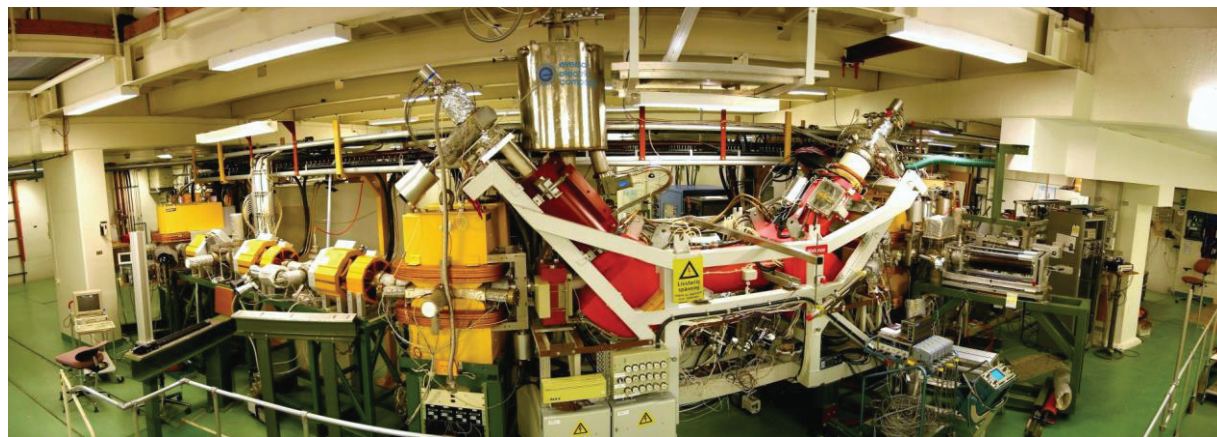
4 MeV/u: $T_{1/2} \approx 2$ s

Connecting ESR to CRYRING@ESR



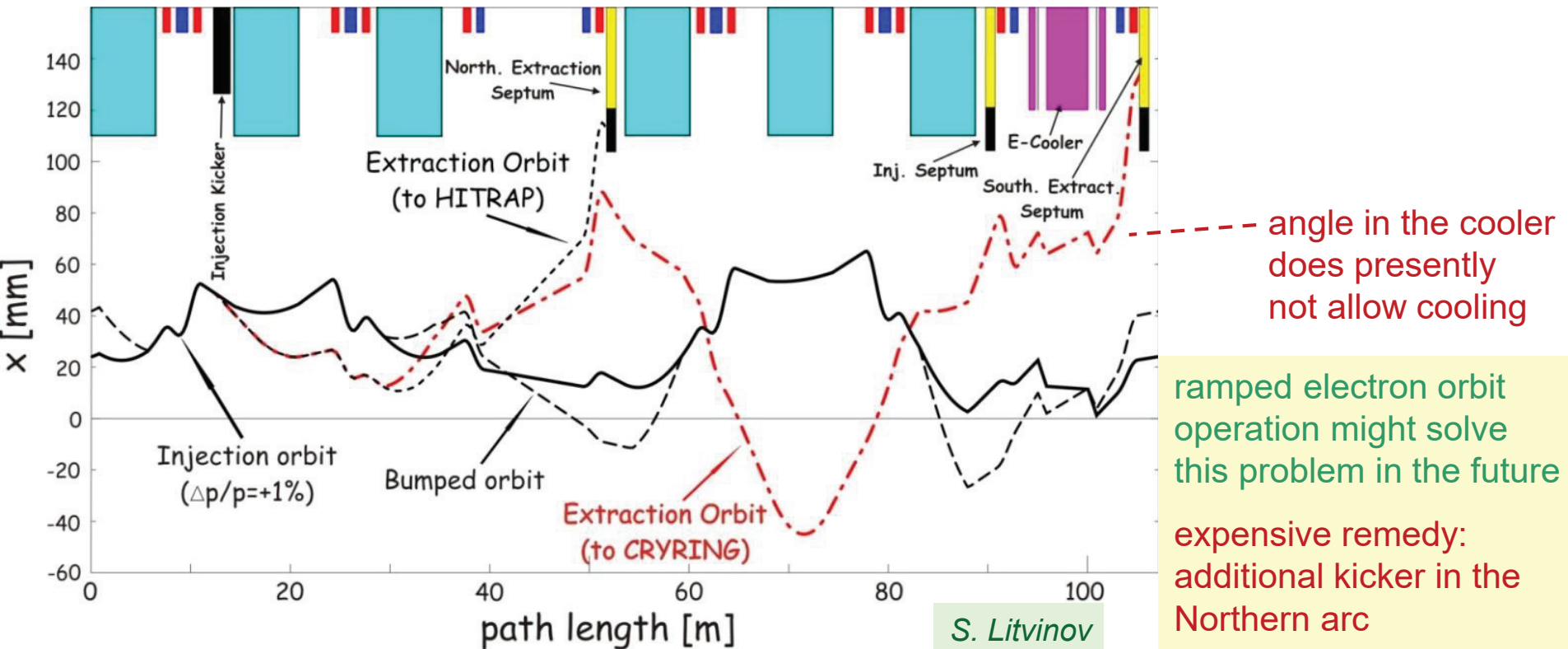
ESR decelerates cooled highly charged ions or rare isotope beams from 400 MeV/u to 4-15 MeV/u

goal:
stored, cooled highly charged
ions at lowest energies

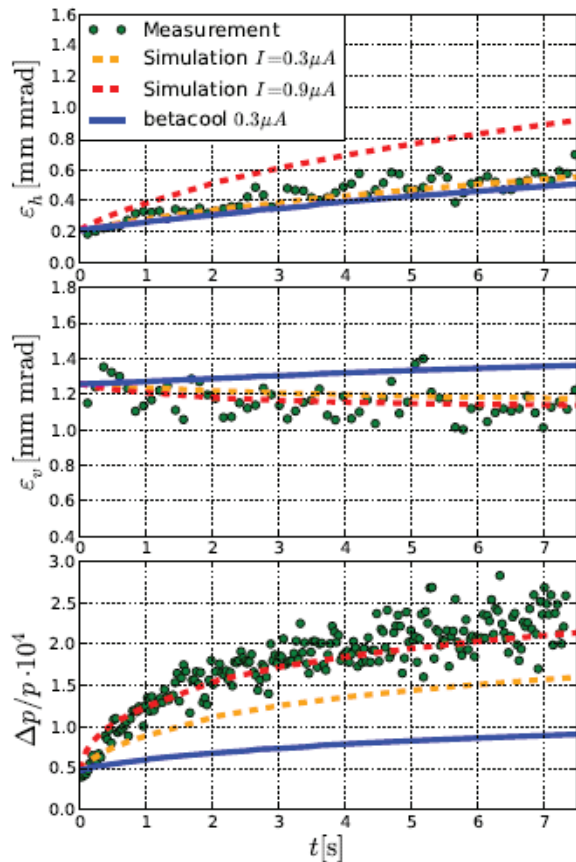


Fast Beam Extraction from ESR to CRYRING

fast beam extraction from ESR to CRYRING is employing a sophisticated scheme with a distorted closed orbit for the circulating beam before extraction to CRYRING
consequence: **no cooling** during extraction, risk of unacceptable heating and emittance growth, particularly for highly charged ions

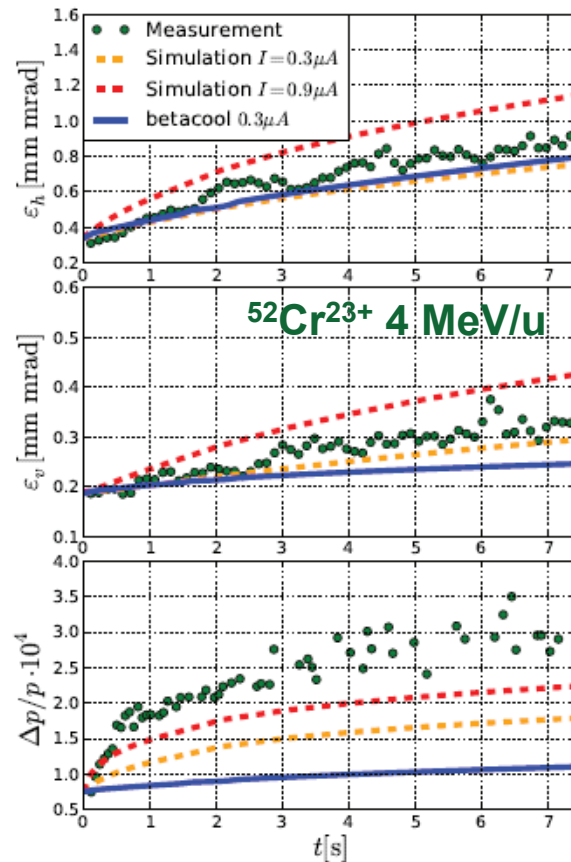


Growth of Beam Emittance due to Intrabeam Scattering



(a) Start from 30mA cooler current

P. Görden, Master Thesis, 2011



(b) Start from 15mA cooler current

stronger (faster) cooling
results in faster IBS heating

for highly charged ions IBS
is expected to scale with Q^4/A^2
⇒ highly charged uranium heats
up 10 times faster than Cr^{23+}

It is an open issue whether an
efficient beam transport from
ESR to CRYRING can be
achieved with the present
extraction mode.

**Similar observations at LEIR
with Pb^{53+} at 4.2 MeV/u:
heating time of order of 10 ms**

Space Charge Limit

Space charge limit due to incoherent tune shift

$$\Delta Q_x = \frac{r_p Z^2 N g}{\pi A \beta^2 \gamma^3 B (\epsilon_x + \sqrt{\epsilon_x \epsilon_y Q_x / Q_y})}$$
$$\Delta Q_x \simeq \frac{r_p Z^2 N}{2\pi A \beta^2 \gamma^3 B \epsilon_x}, \quad g = 1, \epsilon_x = \epsilon_y, Q_x = Q_y$$

at 4 MeV/u, for $\Delta Q_x = 0.1$ and $\epsilon_x = 1$ mm mrad:

$$N \leq 8.7 \times 10^8 \frac{A}{Z^2} \cdot B$$

Coasting (B=1): Ne¹⁰⁺: 1.8×10^8
 Kr³⁶⁺: 5.8×10^7
 U⁹²⁺: 2.5×10^7

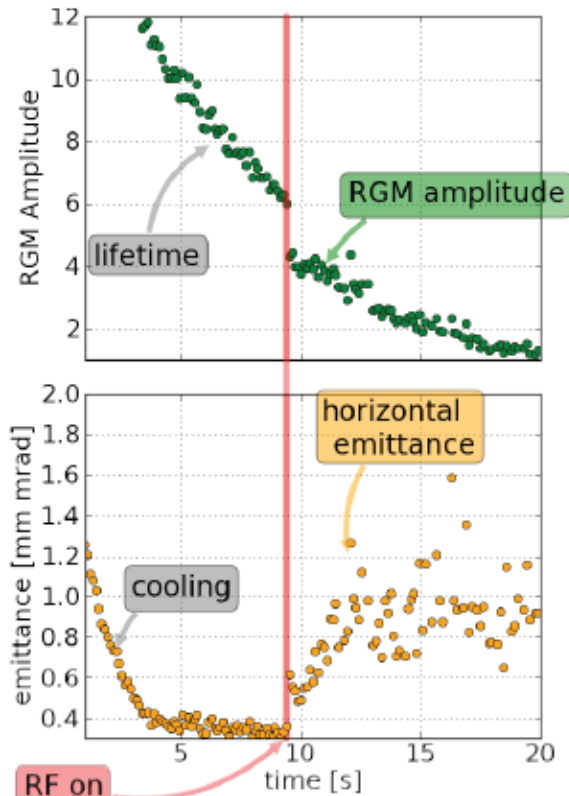
Bunched ($\leq 1 \mu\text{s}$, B=1/5): Ne¹⁰⁺: 4×10^7
 Kr³⁶⁺: 1×10^7
 U⁹²⁺: 5×10^6

For bunched beam the intensity limit is reduced by the bunching factor.

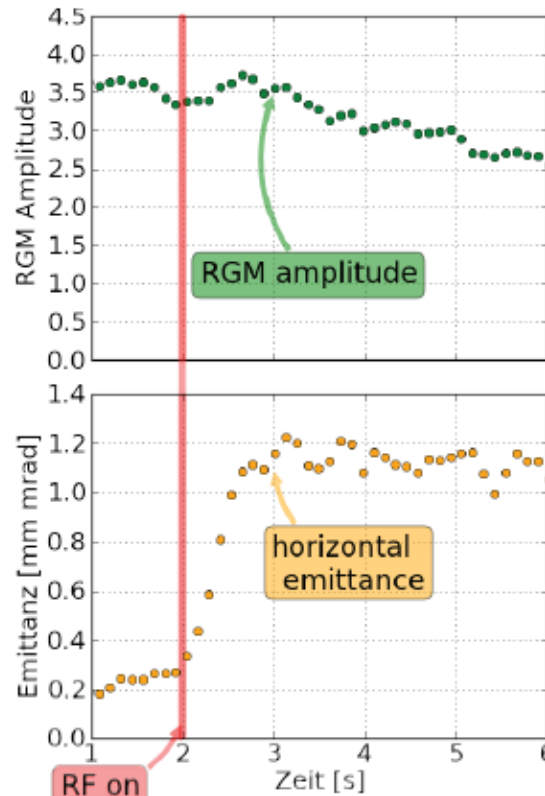
Experimental observation: with rf amplitude 100 V (h=1) the bunching factor is $B \approx 1/20$.

Loss during Bunching a Low Energy Beam

Growth after Bunching the Cooled Coasting Beam



(a) Beam loss: bunching of $^{40}_{18}\text{Ar}^{18+}$

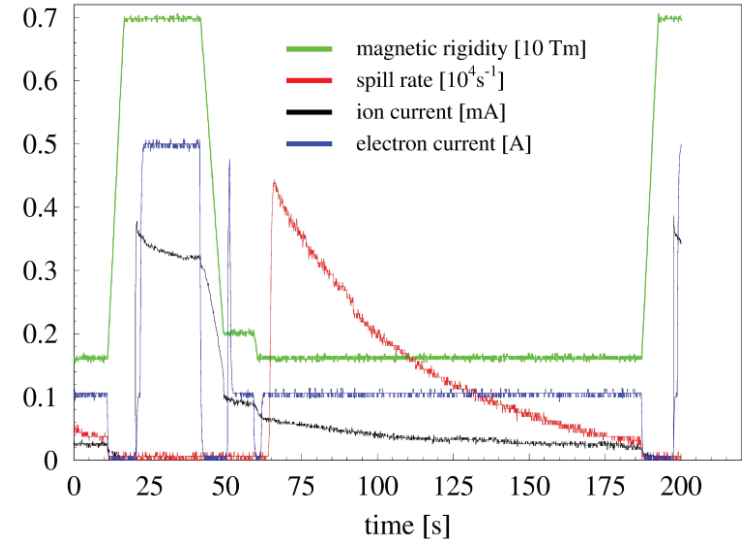
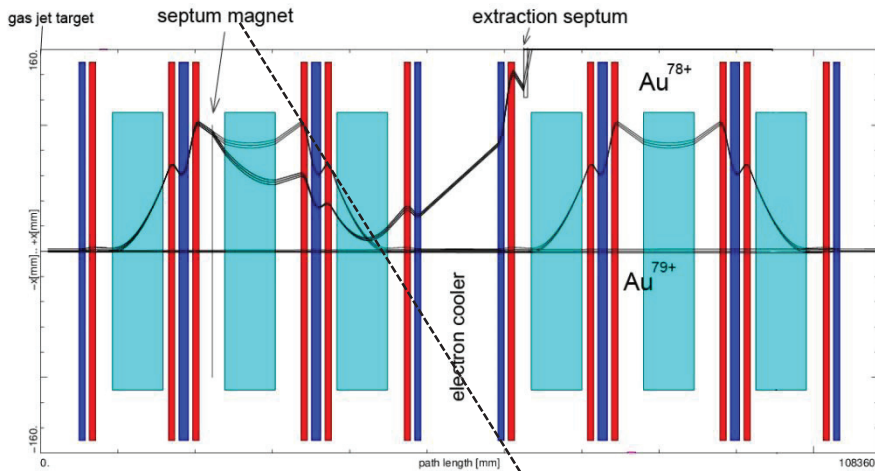


(b) No beam loss: bunching of $^{52}_{24}\text{Cr}^{23+}$

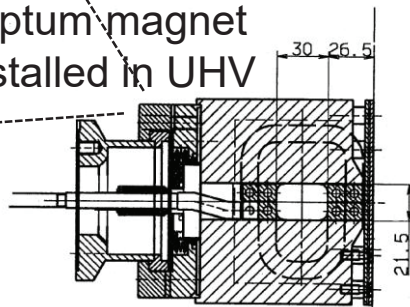
*P. Görgen
Master Thesis, 2011*

⇒ Precise control of cooling, tune and rf amplitude down to small values

Slow Extraction by Charge Change



septum magnet
installed in UHV



diagnostics in the beamline



extracted cooled beam
after deceleration (to ≥ 10 MeV/u)
small emittance
good control of extraction rate
extraction time minutes to hours

Additional Options

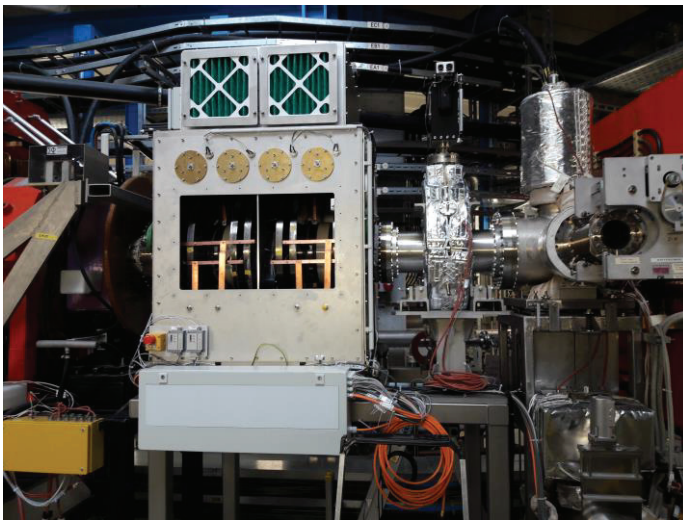
cooling in the storage ring allows accumulation of higher beam intensity

important for low intensity primary beams and secondary (rare isotope) beams

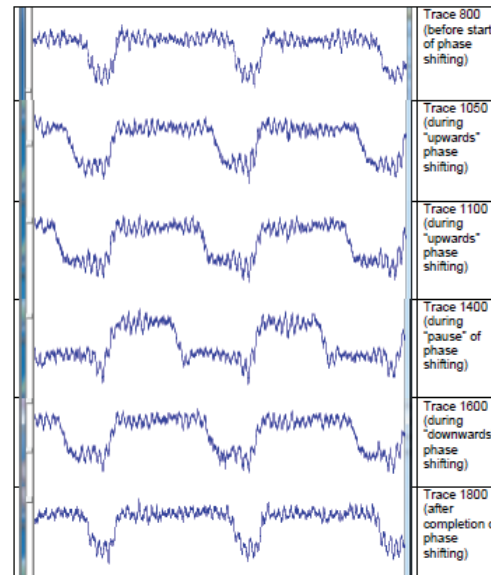
various methods are available, all based on cooling

rf stacking (stochastic/electron cooling on injection orbit, electron cooling on stack orbit)

barrier bucket stacking (fixed barriers, moving barriers, $h=1$ unstable fixed point)



new ESR barrier bucket cavity



bunch compression
and decompression
of an uncooled beam

also suitable for flexible creation of bunches at very low energy

Optimization of the deceleration cycle of the ESR mainly means optimization of the parameters of the cooling systems (reduction of cooling time)

in addition: faster ramping is supported by the hardware

new controls should be seminal to employ this feature

Main goal of the coming re-commissioning of the ESR is short cycle times

optimization of cooling

flexible ring operation.

The main tool is the new FAIR-type control system LSA (LHC software architecture).

It should provide increased flexibility in tuning the subsystems.

Thanks to you

and to my collaborators

C. Dimopoulou, R. Heß, R. Joseph, C. Kleffner,

S. Litvinov, F. Nolden, D. Winters