Measurement of phase-space density evolution in MICE Step IV

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Experimental apparatus at present (Step IV)

All the detectors are installed and working

- Three time-of-flight (TOF) detector stations
- Two Cherenkov counters and a downstream calorimetry module
- Two scintillating-fibre trackers
- Part of the cooling channel (no RF yet)
 - Two Spectrometer Solenoids (SS), each composed of 5 coils
 - The two downstream match coils are currently not turned on
 - An Absorber Focus Coil (AFC) module, made of 2 coils + absorber



Beam cooling setting optimization

Suitable optics have been found using two approaches to conjointly optimize transmission and cooling performance:

- Linear optics, scan in the parameter space of magnet currents;
- Genetic algorithm, best sets of optics bear the next generation, penalize transmission loss and encourage emittance reduction.

The **bottom lattice** setting is expected to have one of the best cooling performance – transmission trade-off and is **presented in the following**



Reproduction of optical functions in the simulation

Reliable reproduction of the data in the simulation for

- $\circ \sim 6 \, \text{mm}$ input beam
- $140 \,\mathrm{MeV}/c$ central momentum 0
- \circ -0.68 central α_{\perp}

[mm]Tg

1000

800

600

400

200

- 787 mm central β_{\perp}
- \rightarrow Excellent tool to systematically test novel density methods



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Particle selection

Series of cuts applied to both data and simulation:

- Muon tagging using TOF01 (Particle ID)
- Upstream reference plane hit (Reference)
- Good track reconstruction quality (Quality)
- Track within the tracker fiducial (Aperture)
- Long. momentum $\in [135, 145] \text{ MeV}/c$ (Momentum)



All

Transverse normalised RMS emittance

4D normalised RMS emittance:

$$\epsilon_n = \frac{1}{m} |\Sigma|^{\frac{1}{4}}, \qquad (1)$$

with D the determinant of the covariance matrix defined as

$$\Sigma = \begin{pmatrix} \sigma_{xx} & \sigma_{xp_x} & \sigma_{xy} & \sigma_{xp_y} \\ \sigma_{p_xx} & \sigma_{p_xp_x} & \sigma_{p_xy} & \sigma_{p_xp_y} \\ \sigma_{yx} & \sigma_{yp_x} & \sigma_{yy} & \sigma_{yp_y} \\ \sigma_{p_yx} & \sigma_{p_yp_x} & \sigma_{p_yy} & \sigma_{p_yp_y} \end{pmatrix}.$$
 (2)



The RMS emittance is directly related to the **volume of the RMS** ellipsoid through $\epsilon_n = \sqrt{2V_{\text{RMS}}}/(m\pi)$ and as such is the most common probe of average phase-space density:

$$\rho_{\rm RMS} = \frac{N}{V_{\rm RMS}} = \frac{N}{\frac{1}{2}m^2\pi^2\epsilon_n^2} = \frac{N}{\frac{1}{2}\pi^2|\Sigma|^{\frac{1}{2}}} \quad [{\rm mm}^{-2}({\rm MeV}/c)^{-2}].$$
(3)

 \rightarrow It follows from Liouville's theorem that the phase-space volume should be conserved

Emittance evolution



Power of density estimation

The emittance plot exhibits two obvious challenges:

- o transmission losses yield apparent emittance reduction;
- filamentation in SSD yield apparent emittance growth.
- The key to solving both problems lies in density estimation:
 - Estimate density in the transverse 4D phase-space (center);
 - Select an **identical fraction** of the beam upstream and downstream from within the **densest area of the space** (*right*);
 - · Define cooling figure of merits on these subsamples;
 - \rightarrow The core, unlike the tails, is **<u>transmitted</u>** and <u>linear</u>.



Transverse single-particle amplitude

Single particle amplitude is defined as

$$A_{\perp} = \epsilon_n \mathbf{u}^T \Sigma^{-1} \mathbf{u} \tag{4}$$

with $\mathbf{u} = \mathbf{v} - \boldsymbol{\mu}$, the centered phase-space vector (\mathbf{v}) of the particle.

Amplitude follows a χ^2 distribution with d degrees of freedom with $\langle A_{\perp} \rangle = 4\epsilon_n$.

It is related to the **volume** of an ellipse similar to the RMS ellipse, going through \mathbf{v} .



$$\rho(\boldsymbol{v}_i) = \frac{1}{(2\pi)^2 |\Sigma|^{\frac{1}{2}}} \exp\left[-\mathbf{u}^T \Sigma^{-1} \mathbf{u}/2\right] = \left[\frac{1}{4\pi^2 m^2 \epsilon_n^2} \exp\left[-\frac{A_\perp}{2\epsilon_n}\right]\right].$$
 (5)

 \rightarrow Allows for the selection of a high density core !



Amplitude reconstruction

In the case of non-linear beams, special care must be taken in the reconstruction of amplitude as tails significantly bias the covariance matrix

Optimal procedure for amplitude reconstruction:

- $\circ~$ Compute $\Sigma~$ and $\mu~$ for the whole sample;
- 1 Calculate all the particle amplitudes A^i_{\perp} ;
- 2 Register the highest amplitude in the distribution;
- 3 Update Σ and μ by removing the highest amplitude point;
- 4 Iterate from 1.



Amplitudes at TKD station 5





Amplitude distribution evolution





Subemittance definition and properties

The α -subemittance, e_{α} , is defined as the emittance of the core fraction α of the parent beam. For a truncated 4D Gaussian beam of covariance matrix S, it satisfies



The statistical uncertainty carried by this measurement is identical to that of the emittance, scaled by the fraction α as

$$\frac{\sigma_{e_{\alpha}}}{e_{\alpha}} = \frac{1}{\sqrt{\alpha}} \frac{\sigma_{\epsilon_n}}{\epsilon_n} = \sqrt{\frac{2}{\alpha N d}}.$$
(8)

Subemittance evolution



Fractional emittance definition and properties

The α -fractional emittance, ϵ_{α} , is defined as the phase-space volume occupied by the core fraction α of the parent beam. For a truncated 4D Gaussian beam of covariance matrix S, it satisfies



In 4D, a fraction α of $\mathbf{9\%}$ yields the volume of the RMS ellipsoid, $V_{\rm RMS}$

The **convex hull** is a prime candidate for volume reconstruction. It computes the smallest volume that contains the core αN points.



Phase-space volume evolution



Toy analysis of fractional quantities

A toy analysis (Gaussian input beam, toy absorber) shows:

- The **same relative change** is seen in the RMS emittance and all of the fractional quantities, for any fraction
- $\circ~$ The change in fractional quantities exhibit the same relation with β_{\perp} and the input emittance, ϵ_i
- The fractional quantities are **more robust** against losses and non-linearities as the tails do not influence their measurement



 \rightarrow Plots produces for a core 9 % selection, i.e. size of the RMS ellipse

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Non-parametric density estimation: k-Nearest Neighbours

For a given point x, find the k closest points in the input cloud. Find the distance R_k to the k^{th} point and compute the 4D local density estimate as

$$\rho(\boldsymbol{x}) = \frac{k}{\mathcal{V}_k} = \frac{2k}{\pi^2 R_k^4},\tag{11}$$

with \mathcal{V}_k the volume of the 4-ball centred in \boldsymbol{x} of radius R_k .

The rule of thumb choice of $k = \sqrt{N}$ yields quasi-optimal results for a broad array of distributions.

Right plot shows great agreement between theoretical Cauchy distribution (red) and estimation (blue).





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Density estimation at TKD station 5



Phase-space volume evolution



Conclusions

Status of the amplitude-based analysis:

- Selecting the low amplitude core gets rid of apparent emittance reduction due to scraping and apparent emittance growth due to beam filamentation in the downstream section;
- A toy MC shows that the **exact same behaviour** can be observed for the subsample and fractional emittance as for the RMS definition;
- Method shows a **clean cooling signal** in a realistic MC.

Status of the non-parametric analysis:

- Systematic study well advanced, *k***NN robust in 4D**, low error and no bias for large samples with the rule-of-thumb *k* selection;
- Method applied to the toy MC to study its behaviour, identical trend as with the amplitude-based fractional emittance;
- Method also shows **cooling signal** in a realistic MC.