





### Wir schaffen Wissen – heute für morgen

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**Overview of Seeding Methods for FELs** 



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- General Features of Seeded FEL
- SASE FEL as a Background Process
- Direct Seeding Methods
- Electron Beam Manipulations
- Cascades and Hybrid Configurations
- Self-Seeding Methods
- Summary



In Comparison to SASE FELs seeded FELs can offer the improvement:

- **1.** Control/Improvement of the Longitudinal Coherence
- 2. Improved Brilliance
- 3. Energy Stability of FEL Output Pulse
- 4. Spectral Stability at Selected Frequency
- 5. Synchronization with External Source (Pump-Probe)
- 6. Ability to Increase FEL Efficiency with Taper
- 7. FEL becomes shorter



### **Disadvantage of Seeded FELs**

FEL Performance gets more sensitive to electron beam fluctuation: Energy Jitter SASE SASE FEL Seeded FEL P(E) [a.u.] Energy Jitter shifts central Huge fluctuation in output wavelength but keeps photon power when energy jitter number almost unchanged becomes comparable to FEL bandwidth 12.35 12.36 12.37 12.38 12.39 124 12.41 12.42 12.43 12.44 12.45 E [keV] FEL performance unchanged Temporal overlap problems Seeded Energy Jitter from arrival time unless seed signal is longer Power growth flattens at Post Saturation Oscillation of Green's function of FEL process <sup>o</sup>(E) [a.u.] saturation with a slight growth FEL power FEL Pulse Jitter





12.35 12.36 12.37 12.38 12.39 124 12.41 12.42 12.43 12.44 12.45 E [keV]

# Synchronization to External Seed Signal

#### Seed pulse must be shorter than electron bunch length

Otherwise FEL pulse length is defined by electron bunch length, including bunch arrival jitter



Seed pulse must be longer than cooperation length

Pulse will be stretched by FEL process. Identical performance than single spike SASE operation

Goal is mutual exclusive to maximum brilliance Maximum brilliance is given by bunch length and requires a seed signal longer than bunch length

Arrival time jitter must be less than bunch length

Otherwise there is a chance of missing overlap. Bunch will laser in SASE mode



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- · Well-defined input signals allows to optimize the tapering profile
- However side-band instability will modulate profile envelope
- Slippage effects stops the trapping of the electrons

#### Example: LCLS (self-seeded)





$$P_{sn} = \frac{3\sqrt{4\pi}\rho^2 P_{beam}}{N_{\lambda}\sqrt{\ln(N_{\lambda}/\rho)}} \propto \frac{\gamma}{\sqrt{a-\ln\gamma}}$$

Towards shorter wavelengths the typical beam energy increases and the shot noise signal gets larger Example: 5 nm @ 2 GeV  $\rightarrow$  P<sub>sn</sub> ~ 100 W (N<sub> $\lambda$ </sub>: #electrons/wavelength,  $\rho$ : FEL Parameter)



 $P_{seed} > 2000 P_{sn}$ 

# High Harmonic Generation (HHG) Seeding



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### **Proof-of-Principle Experiment**

#### 1.0 G. Lambert et al, Nature Physics Vol. 4 (2008) 296 0.8 Intensity (arb. units) Test Injector at SCSS (Spring 8) 0.6 -Drive Laser at 800 nm Seeded Unseeded 0.4 -× 2.600 Infrared beam from Ti:sapphirelaser install ation HHG Harmonics: 160 nm (n=5) 0.2 Laser hutch Vertical position (mm) Intensity (arb. units) Delay line 0.2 0.4 0.6 0.8 1.0 Achromatic lens system Streak carnera 3 Seeded 2 · Focusing lens = 7 mShielding wall Telescope Accelerator tunnel Vertical position (mm) 0 HHGgascell 3 **Electron beam** Unseeded dump FEL 2 OTR Undulator Undulator Electron beam Fifth harmonic radiation screens section 2 section 1 Grating from LINAC **Periscope**4 Vertical position (mm) 3 Second C-band Dipole acceleration 2 · section CCD camera Accelerator Magnetic chicane Diagnostic and radiating systems Spectrometer system (150 MeV) 158 159 160 161 162 163 164 Wavelength (nm)

HHG Seed is amplified by a factor of about 500

SASE signal is 2600 times weaker than seeded FEL(good contrast of seed to shot noise)



## **Comments on HHG Seeding**

HHG sources as seeds for FEL have been demonstrated at various experiments with the current record of 39 nm [*C. Lechner, et al, Proc of FEL Conference 2012*]

However further progress requires significant R&D in the source development:

- Decrease the wavelength (extending the plateau of the HHG process)
- Increase the efficiency of HHG process to overcome increasing shot noise power
- Control/preserve the phase front and mode content of the HHG source
- Control the bandwidth of a given harmonic to match FEL bandwidth

Wavelengths below 20 nm difficult to achieve

Best for sync FEL pulse to external signal

Very little increase in brightness (Single-spike SASE might be better alternative)



- Induced energy modulation at longer wavelength is changed into rich harmonic current content after compression with a chicane.
- To avoid smearing out the energy modulation must be larger than intrinsic energy spread
- A selected harmonic is picked up with a succeeding undulator.



Theory: L.-H. Yu, Science 289 (2000) 932



#### **Proof-of Principle Experiment (SDL)**

A. Doyuran et al, PRL 86 (2001) 5902



L.-H. Yu et al, NIMA 528 (2004) 436



#### FERMI @ Elettra – User Facility based on HGHG

E. Allaria et al, Nature Photonics. 6 (2012) 699





#### Fresh Bunch Technique (e.g. FEL II at FERMI@Elettra)





Strong progress in the last year mostly due to the success of FERMI. Wavelengths down to 4 nm have been achieved.

However an HGHG FEL cannot be optimized much for pulse energy:

- Only fresh bunch feasible for cascades due to the tremendous sensitivity of same bunch cascades
- Long bunches reduces the current and thus the saturation power
- Only a subsection of the bunch contribute to the final radiation stage (similar to HHG seeds)
- Energy spread of initial beam has to be less than in SASE case, limiting the use of laser heaters in the injector and machine.

Wavelengths down to 1 nm seems feasible

Good for sync FEL pulse to external signal

Not optimized for pulse energy.

# **Echo-Enabled Harmonic Generation**



Basic Idea [D. Xiang and G. Stupakov, PR STAB 12 (2009) 030702]:

- First stage: Modulation and overcompression to generate energy bands
- Second stage: HGHG principle but spacing of bands defines harmonics

High efficiency for bunching

$$b_{\rm max} = \frac{0.39}{m^{1/3}}$$



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### **Experiments for EEHG**

[D. Xiang et al PRL 105 (2010) 114801]

 $\lambda_1$ = 759 nm,  $\lambda_2$  = 1590 nm

Little Chirp



[Z.T. Zhao et al Nature Photonics 6 (2012) 360]

 $\lambda_1$ = 1047 nm,  $\lambda_2$  = 1047 nm





# Wavelength Limit of EEHG

Although scaling towards shorter wavelength is promising, there are practical reasons for the higher harmonic numbers:

- Total width of energy modulation is limited by FEL process ( $\Delta\gamma/\gamma < \rho$ )
- Number of lines defines harmonic *m* with an average line spacing of  $\delta \gamma / \gamma < \rho / m$
- Hyperfine structure can be blurred out by:
  - Quantum Fluctuation of the incoherent emission in modulator and chicane
    - Favors low magnetic field and long chicanes
  - Intrabeam scattering [G. Stupakov, FEL 2011]
    - Favors compact chicanes

Wavelength limit at about 1 nm Good control of electron chirp

1<sup>st</sup> chicane can lengthen bunch





Basic Idea:

- 1<sup>st</sup> stage operates as SASE FEL, but stopped before saturation
- Radiation is filtered, introducing longitudinal coherence
- Delay of radiation field is matched with delay electron beam with a magnetic chicane. The chicane removes also any induced bunch, removing the imprint of SASE in the bunch (quasi fresh bunch)
- Beam and radiation are overlapped in a second stage, operating as an FEL amplifier.

First proposed for soft X-ray FEL FLASH [J. Feldhaus et al, Opt. Comm 140 (1997) 341] but never realized due to the strong delays in the photon and electron path.

Idea brought up again for hard X-ray [*G. Geloni, Jour. Of Modern Optic 58:16 (2011) 1391*], using the transmission around the stop band of a Bragg reflection (see next slide).

More compact design for soft X-ray [Y. Feng, LCLS] makes self-seeding feasible for longer wavelength



### **Example Performance for SwissFEL at 1 nm**





### **Self-Seeding for Hard X-rays**





Hard X-ray Self-seeding done at LCLS for 8 keV [J. Amann et al, Nature Photonics 6 (2012) 693]

- Wavelength: 1.5 nm
- Diamond Crystal, using (400) reflection
- Reduction of Bandwidth by factor 40 observed
- Output energy very sensitive to electron energy jitter



# Slippage Enhancing Schemes I

#### Mode Locked FEL (N.R. Thompson et al, PRL 100 (2008) 203901)

Modulator



- Increase Slippage by delay lines
- Undulator modules shorter than one gain length
- Lock with a modulation synched to slippage per stage
- Single mode selectable by increasing delay (1,2,4,8 etc) → iSASE [J. Wu et al, FEL Conference 2012]



- For high K-values, the FEL is operated at the higher harmonics *n*
- Phase shifter disrupts fundamental by Δφ=m 2π / n with m is integer, optimized for best suppression of fundamental.
- Requires rather short undulator modules less than two gain length.

Enhanced slippage and thus narrow bandwidth

Full Undulator: [E. Schneidmiller, DESY 12-070]

Sub-section with high harmonic (pSASE): [D. Xiang et al, PR STAB 16 (2013) 010703]



Example for SwissFEL Soft X-ray at 2 nm

$$L_g \propto \lambda_u / \left(\lambda_u \mathbf{x} K \mathbf{x} J J_n\right)^{\frac{2}{3}}$$

$\lambda_{u}$	4 cm	5 cm
K	1	2.3
n	1	3
JJ <sub>n</sub>	0.8	0.4
L <sub>g</sub>	2.5 m	2.5 m



## **Comparison and Summary**

Method	Direct Seeding (HHG)	HGHG Cas. or EEHG	Self-Seeding
Wave Length Limit	>20 nm	> 1nm	> 0.1 Å
Synchronization	Good	Good	None
Brilliance	Similar to SASE (penalty from seed BW)	Slightly better (penalty from lower current)	Much better than SASE
Pulse Length	~10 fs	10 – 100 fs	As electron bunch
Signal-to- Background	Poor	Moderate - Good	Excellent
Complexity	Moderate (excluding source)	High	Moderate
Electron Beam Requirement	Arrival time and energy stability	Arrival time and energy stability, lower energy spread	Energy stability
Undulator Length	Slightly less than SASE FEL	Comparable and longer than SASE FEL	50% longer than SASE FEL



### **Summary**

Seeding is very promising to improve the quality of FEL as a user facility.

Several methods are proposed for seeding and sucessfully demonstrated down to 1 Ångtrom.

Except for the synchronization with external signal, self-seeding is most promising and robust method with no inherent limitation below 1 nm.

Seeding at very short wavelength are very limited. Novel ideas emerging to improve performance of seeded FELs

