Prototype of White Rabbit Network in LHAASO



Hongming Li, Guanghua Gong Tsinghua University, Beijing Qiang Du LBNL, USA



Synchronization is a crucial concern in distributed measurement and control systems. White Rabbit provides subnanosecond accuracy and picoseconds precision for large distributed systems. In the Large High Altitude Air Shower Observatory project, to guarantee the angular resolution of reconstructed air shower event, a 500 ps overall synchronization precision must be achieved among thousands of detectors. A small prototype built at Yangbajin, Tibet, China has been working well for a whole year. A portable calibration node directly synced with the grandmaster switch and a simple detectors stack named Telescope are used to verify the overall synchronization precision of the whole prototype. The preliminary experiment results show that the long term synchronization of the White-Rabbit network is promising and 500 ps overall synchronization precision is achievable with individual calibration and temperature correction.

LHAASO

Tracing galactic cosmic rays sources > 30 TeVWith angular resolution $< 0.5^{\circ}$

KM2A sub-detector array: Covering 1.2km²

5632 electron detectors

1221 muon detectors Timestamps Synchronization < 500 ps (rms) Jitter of Synchronous ADC clock < 100 ps

High data throughput (26 Gbps) with minimum loss

MEAN of timestamps offset

Prototype

Locations & Components:

ARGO Experiment hall, Yangbajin, Tibet, China (4300m) Rubidium clock constrained GPS

4 White-Rabbit Switches V4.0.1

50 Compact Universal Timing Endpoint(CUTE), WRPCV2.1

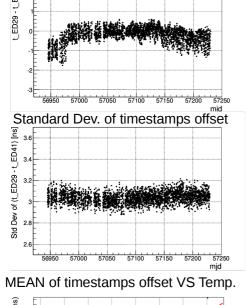
Feature:

4-layer heirachy Node by node calibration Dynamic Temperature correction

Portable Calibration Node



- Normal CUTE WR Node inside box, water sealed
- Armored optical power composite cable
- Rugged composite connector
- ➤ Directly connected to/synced with Grand Master Switch (GMS)
- reference PPS for field measurement



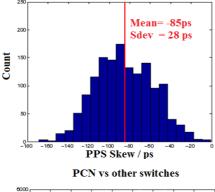
MEAN of timestamps of	fset VS Temp.
LED29 - t ED41 (ns)	
ED4	
239-1	
-0.5	
-1	Entries 2672
	χ ² / ndf 234.9 / 56 p0 0.5262 ± 0.0034
	p1 0.06113 ± 0.00024
-1.5	
	-10 -5 0 :D41 - Temp_ED29 (K)
icinp_c	p_LD20 (it)

Telescope

- ▶ High possibility of Synchronizing Detection
- ► Sync Deviation = Timestamps offset

Simple Detectors Stack

Large Standard Deviation (~3ns) and the temperature dependency (61ps / °C) is mainly caused by the detectors.



GMS vs PCN

			I CIV	v 5 U t	псі	SWILL	псз		
6000		•	V	/RS3	WR	.S2	WR	RS1	
5000									
Count				П	ı		L		
Ö 3000	<u>'</u>			Ш	·	П			
2000	<u>'</u>			Ш	Ш	Н			
1000									
-9	350 -	-300 -2	50 –200	-150	-100	-50	0	50	1
			P	PS S	kew	/ ps			

- ▶ WRPC as IP core integrated with detector electronics (ED & MD)
- ▶ Power noise causes bad precision
- ▶ Remote Status Monitor

	Nada	Node vs PCN (ps)		
	Node index	Accuracy	precision	
	2	-138	100	
	5	-56	28	
	8	28	28	
	12	-27	150	
	16	5	28	
	17	-15	162	
	24	97	156	
	32	-40	34	
	39	-28	41	
	42	-157	135	
	44	-102	26	
	50	-45	34	

