



Distributed Seeding for Narrow-band X-ray Free-Electron Lasers

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Outline

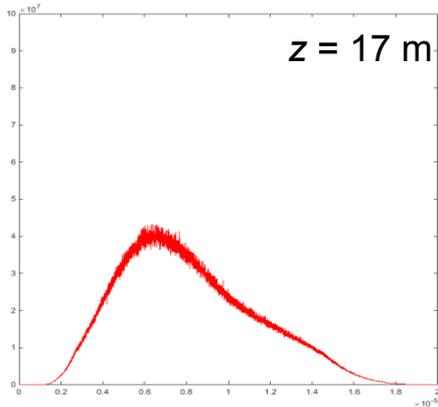
- **SASE Self-Seeding**
- **Distributed Seeding (DS)**
- **MaRIE X-ray FEL**
- **Time-dependent Genesis Simulations**
- **Summary**



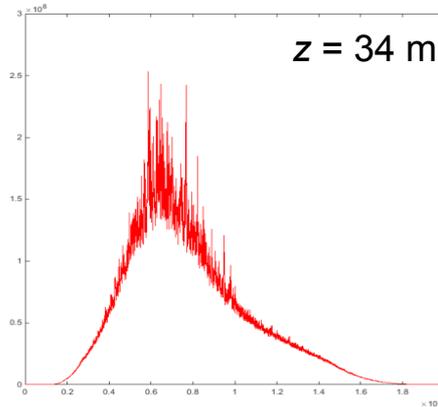
SASE spikes form during exponential growth



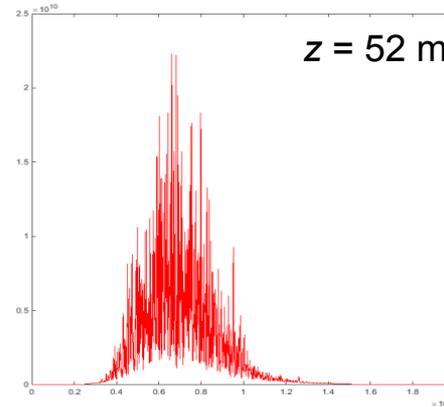
Power versus s plots



Radiation pulse resembles electron current profile. Radiation spectra peak at wavelength longer than λ_0

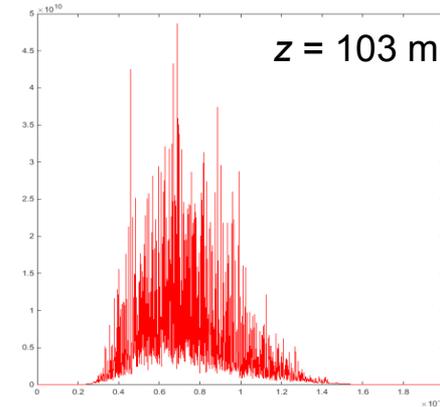


SASE spikes appear in exponential growth regime. Spectral peak shifts toward resonance wavelength λ_0



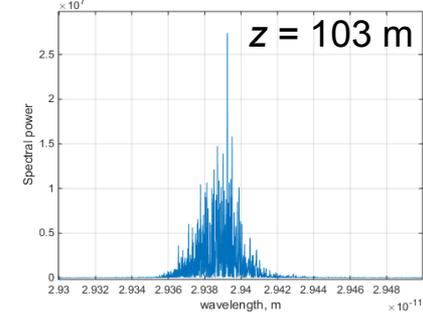
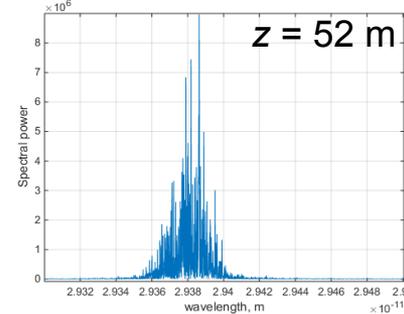
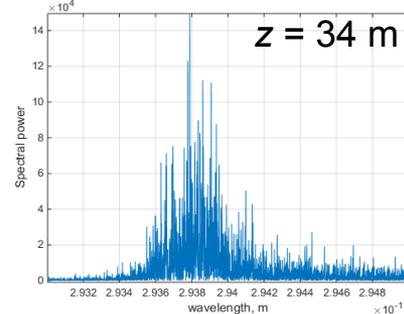
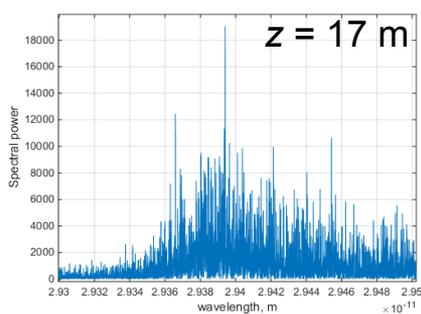
At saturation, SASE spikes have a coherence length of

$$l_c = \frac{\lambda}{2\sqrt{\pi\rho}}$$



More SASE spikes are added beyond saturation. Spectra broaden and shift to slightly longer wavelength.

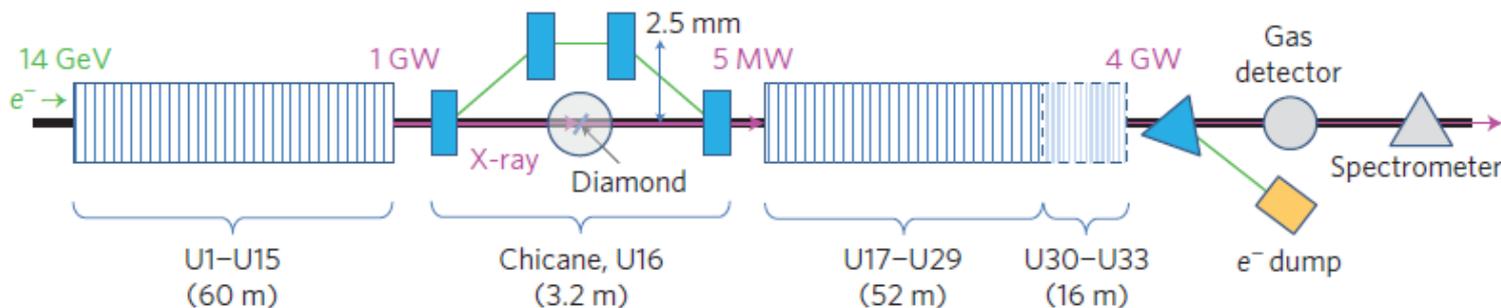
Power versus wavelength plots



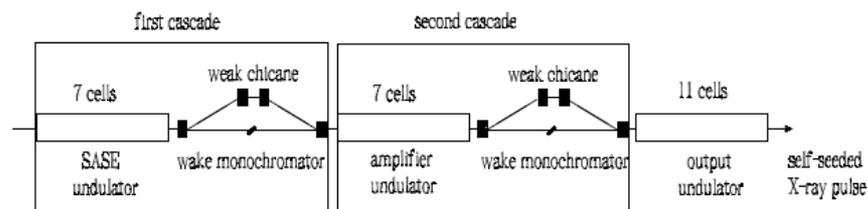
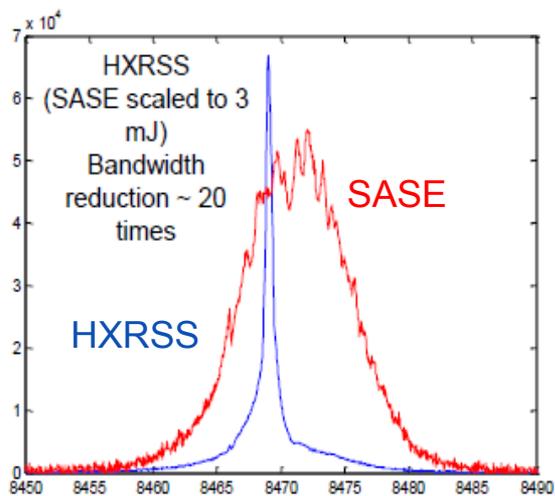


SASE Self-Seeding

Hard X-ray Self-Seeding (HXRSS) has been demonstrated at LCLS and SACLA



Ref: J. Amann *et al.*, *Nature Photonics* **6**, 693-698 (2012).

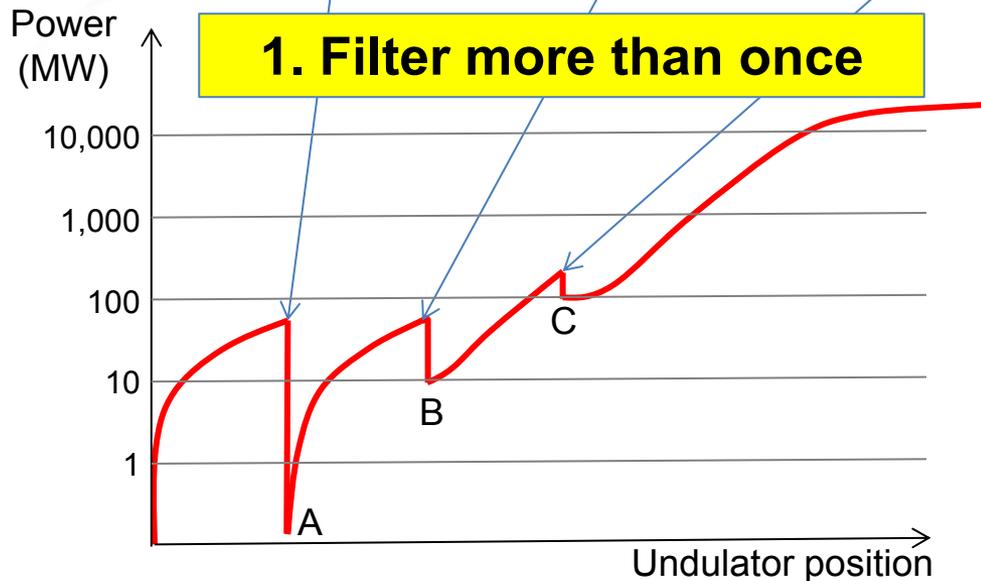


Ref: L. Geloni *et al.*, DESY **10-080**, (2010).

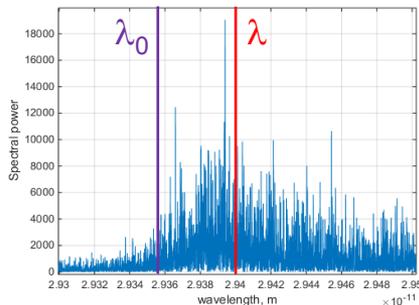
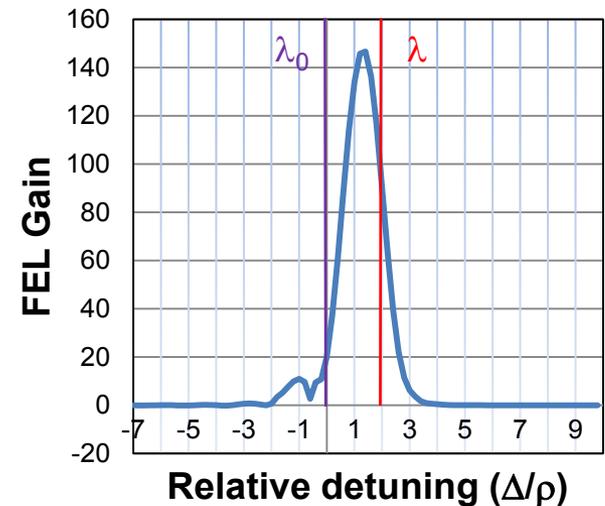
SASE background contributes to about one-half of HXRSS radiation energy. Geloni *et al.* proposed cascade self-seeding to improve HXRSS contrast.



Distributed Seeding (DS)



Gain curve of 2 FODO periods



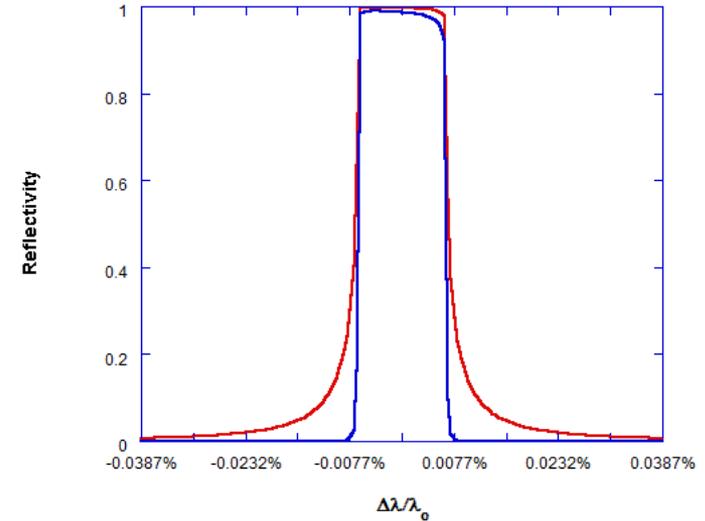
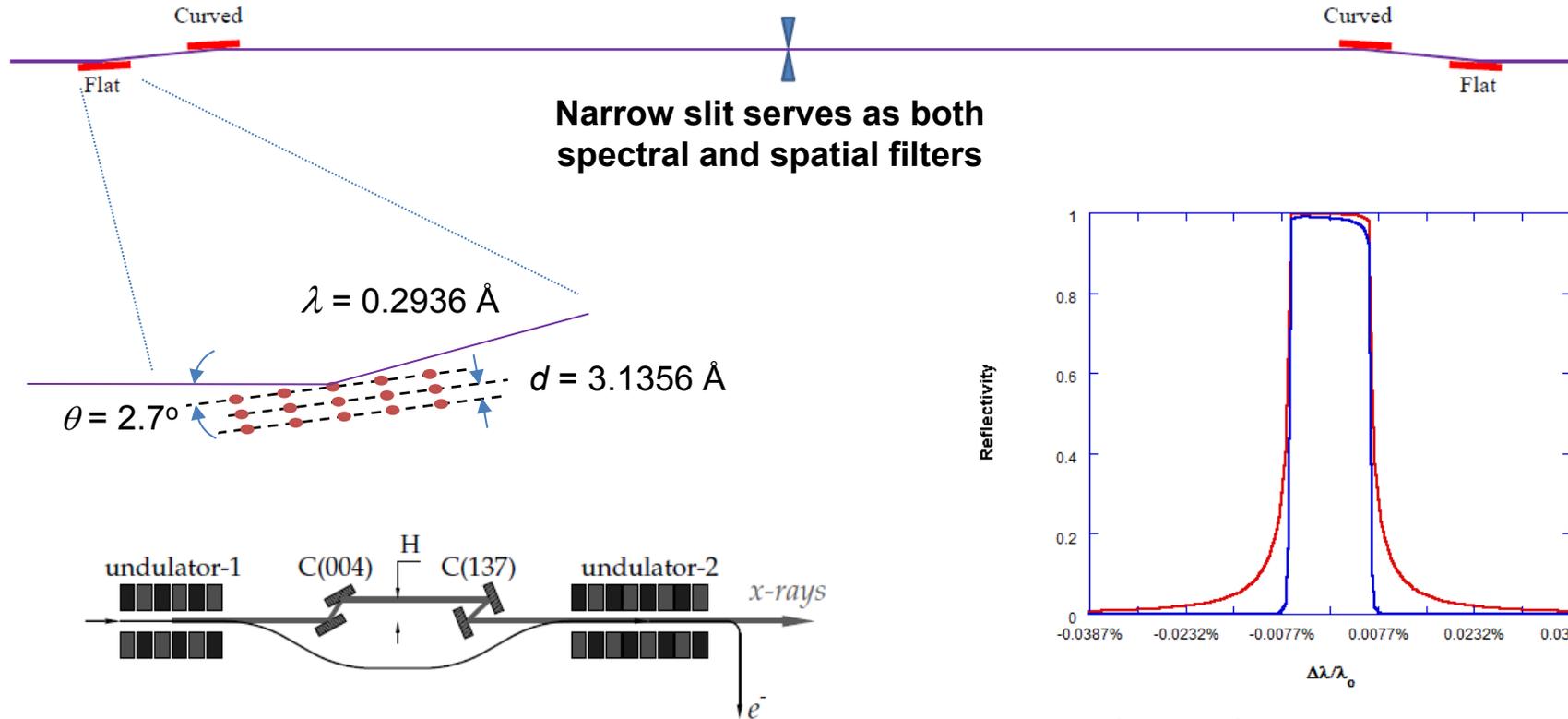
2. Filter before SASE spikes appear

3. Filter at a wavelength longer than λ_0

Four-bounce Bragg crystal monochromator



Each Si(111) crystal deflects X-ray beam by $\pm 5.4^\circ$; total deflection is zero.



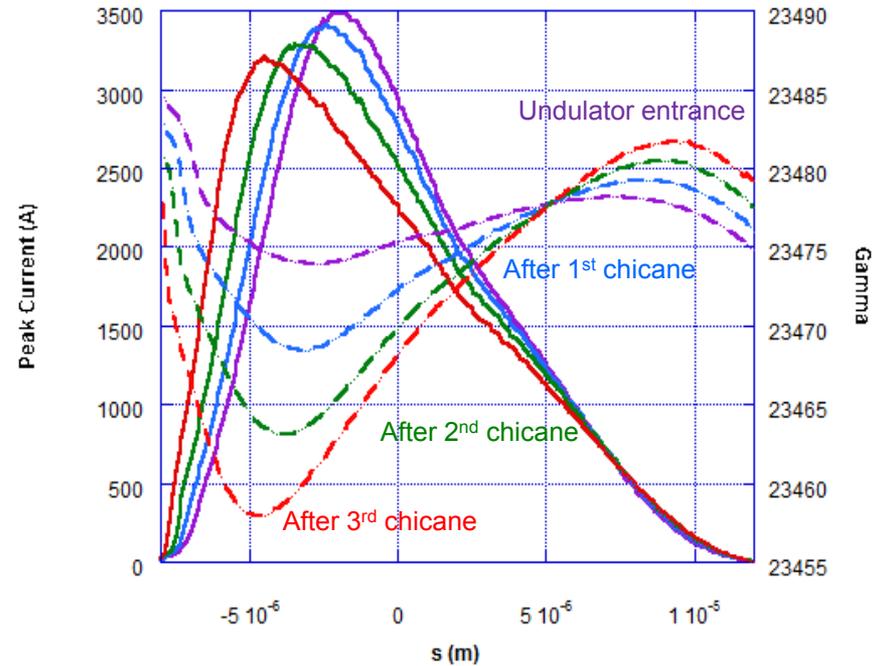
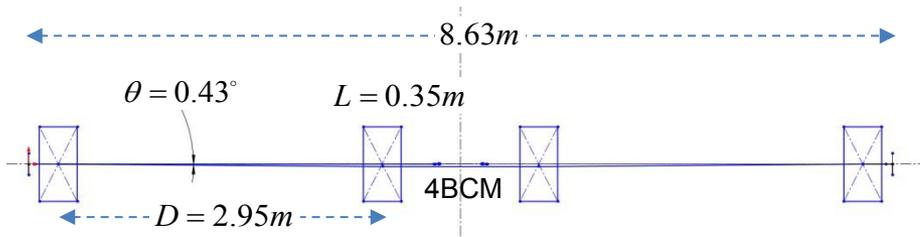
Si(111) Spectral Response

Four-bounce Bragg crystal monochromators (4BCM) have been proposed for SASE self-seeding

Ref: E.L. Saldin *et al.*, Nucl. Instr. Met. Phys. Res. **475** (2001) 357-362.



Small-angle chicanes introduce 1-ps delay in electron path, matching the delay of 4BCM



Delay in electron beam path

$$\Delta S_e = R_{56} = -\theta^2 \left(\frac{4}{3} L + 2D \right)$$

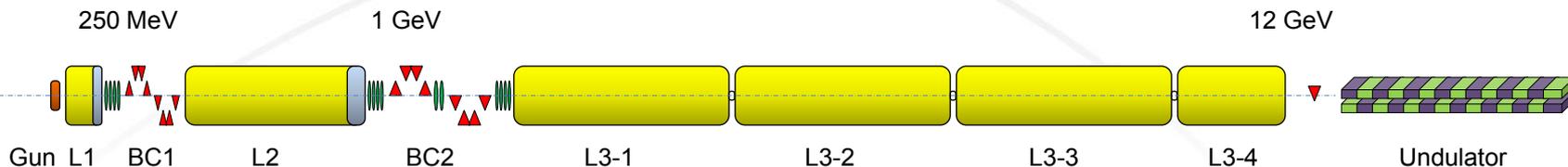
Chicane R_{56} is sufficiently large to erase SASE-induced microbunching.

$$\lambda \ll \left| R_{56} \right| \frac{\sigma_\gamma}{\gamma}$$

The main effects are CSR-induced reduction in the slice beam energy and peak current.

Spoiling SASE microbunching = fresh bunch re-injection

MaRIE X-ray FEL (Matter-Radiation Interactions at Extremes)



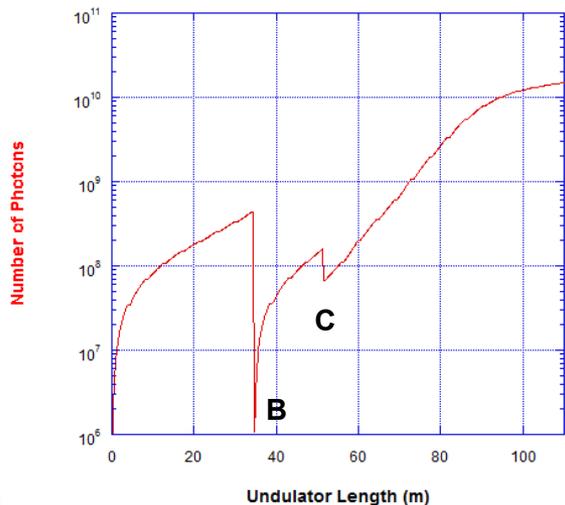
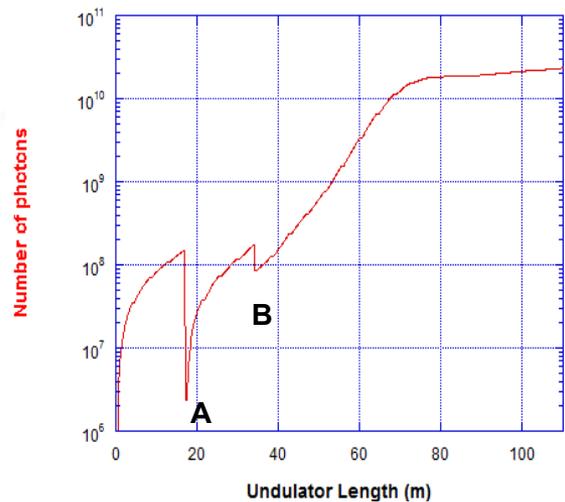
Parameter	Symbol	Value	Unit
Electron beam energy	E_b	12	GeV
Peak current	I_{pk}	3	kA
Bunch charge	Q	100	pC
Slice normalized emittance	ε_n	0.2	μm
Slice energy spread	σ_γ/γ	0.015%	
Undulator period	λ_u	18.6	mm
Peak undulator parameter	K	1.22	
Wavelength	λ	0.2936	\AA
1-D FEL gain parameter	ρ	0.05%	
3-D gain length	L_{G3D}	2.5	m

Details presented in Poster MOP045

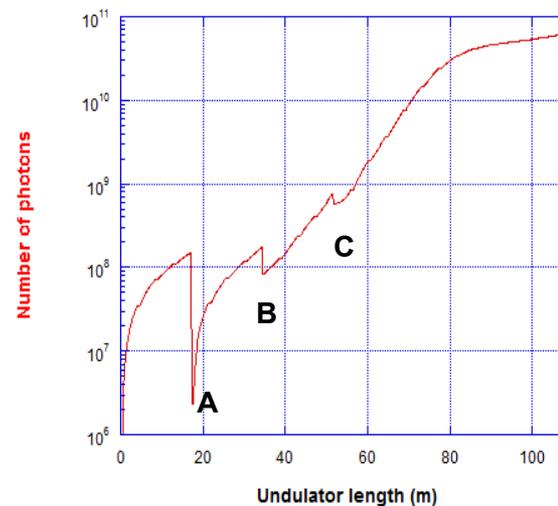
Plots of number of photons versus z for DS



Two-stage DS



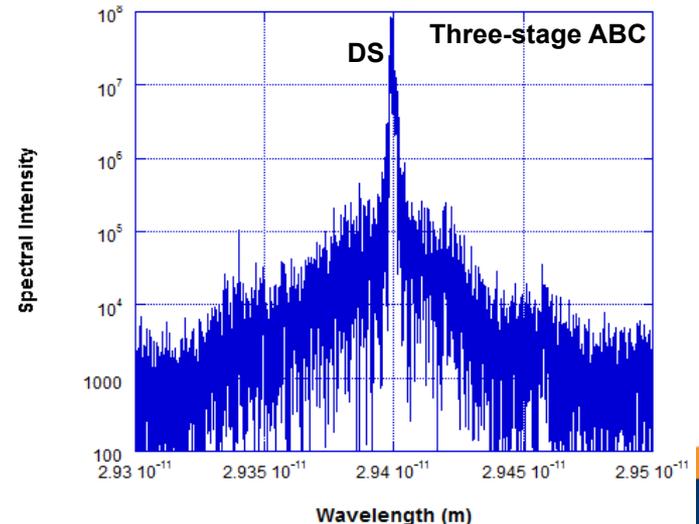
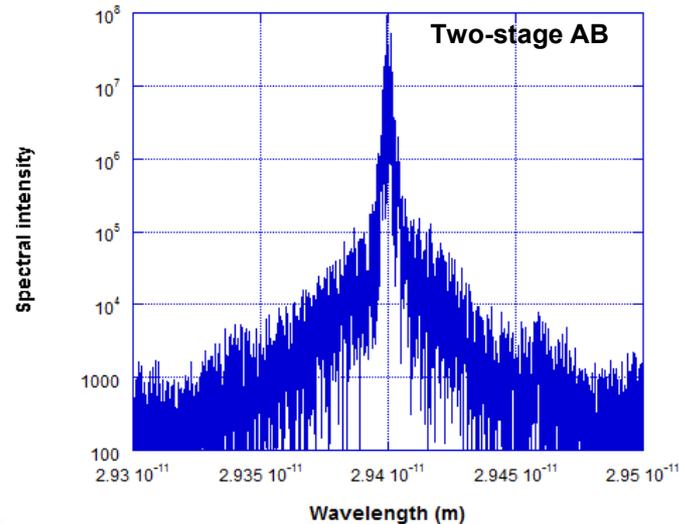
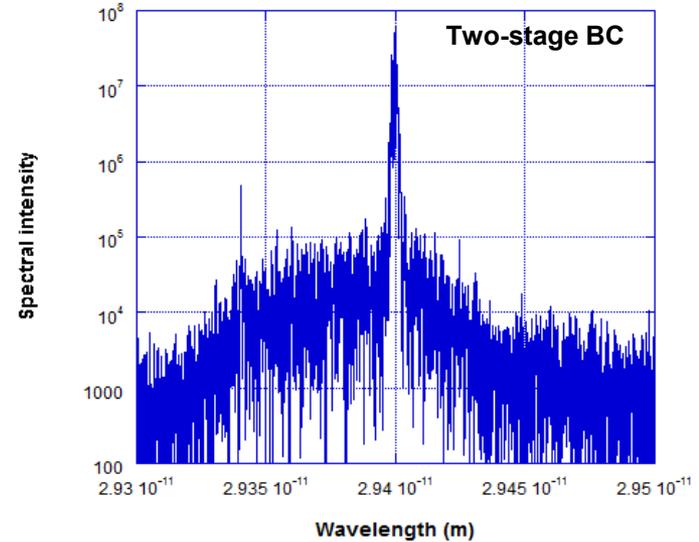
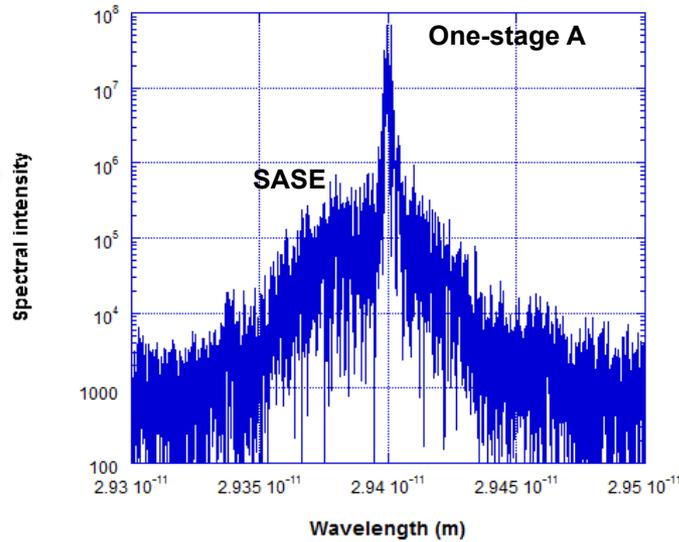
Three-stage DS



Number of photons increases exponentially with z immediately after the 2nd and 3rd filters.

DS yields higher contrast over SASE than SS

Three-stage DS achieves 0.008% bandwidth





Summary

- **DS differs from SASE Self-Seeding in three aspects: Filter more than once, filter before SASE spikes appear, and filter at a longer wavelength.**
- **Filtering at a wavelength longer than the resonance wavelength improves the contrast between the narrow-line DS signal and broadband SASE.**
- **Both two-stage and three-stage DS produce <0.01% relative bandwidth.**
- **Time-dependent Genesis simulations show the three-stage DS can deliver $>2 \times 10^{10}$ photons/bunch at 42 keV for the MaRIE XFEL.**

But will it work in the real world?

DS has been running for years!



THANK YOU

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