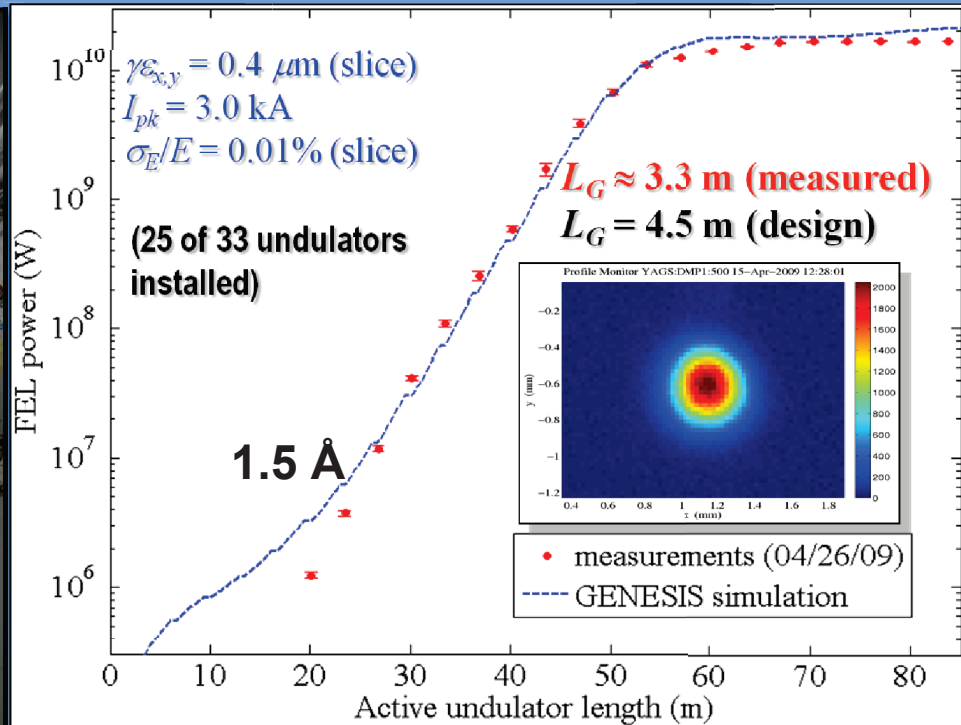
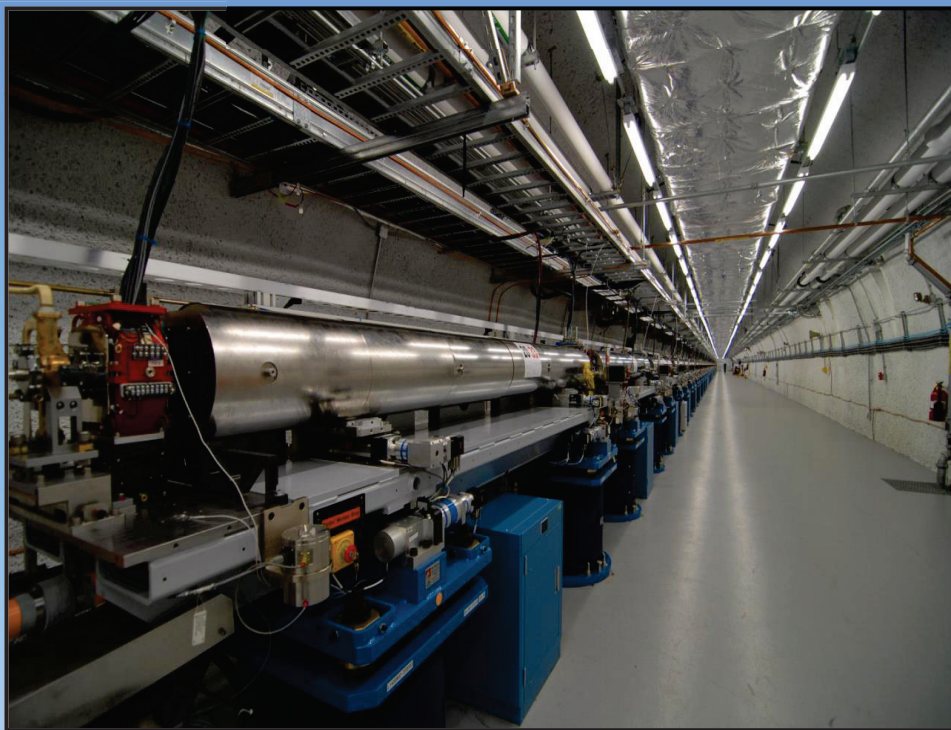


X-Ray FEL R&D Brighter, Better, Cheaper

Zhirong Huang



LCLS: A new era of x-ray sources



- SASE wavelength range: **40 – 1 Å**
- Photon energy range: **0.3 – 12.8 keV** *F.-J. Decker, WEP022*
- Pulse length FWHM **5 - 500 fs** (SXR only)
- Pulse energy up to **5 mJ**, rep. rate **120 Hz**
- User operation since **2010**

A growth spurt in XFELs

SACLA 2011
8.5 GeV, 60 Hz NC



XFEL 2016
17.5 GeV, 2800 x 10 Hz SC



PAL XFEL 2016
10 GeV, 60 Hz NC



SWISS FEL 2017
5.8 GeV, 100 Hz NC

- EUV/soft X-rays: FLASH (2005), FERMI (2012), SXFEL (2017)
- LCLS-II (2020): a new SC CW FEL at 1 MHz rep. rate

Introduction

- Despite the early success, it is widely recognized that XFELs continue to have significant potential for improvement.
- New methods have been rapidly developing to provide FEL seeding, extremely short x-ray pulses, variable double pulses, two-color FEL generation, and polarization control.
- Many of the proposals were implemented in the LCLS since 2011 through FEL R&D program that I have the privilege to contribute. Here I present a personal (incomplete) overview.
- I also like to discuss dreams/progress towards compact XFELs and conclude with some final remarks.

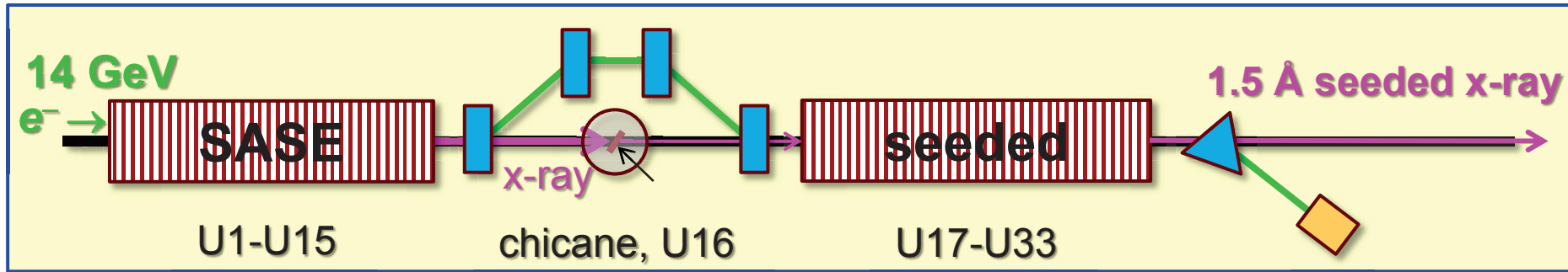
Outline

- *Introduction*
- *Improving temporal coherence (“**brighter**”)*
- *X-ray pulse controls (“**better**”)*
- *Compact coherent sources (“**cheaper**”)*
- *Summary*

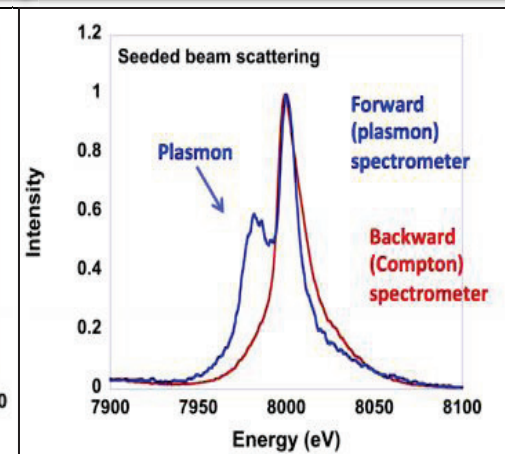
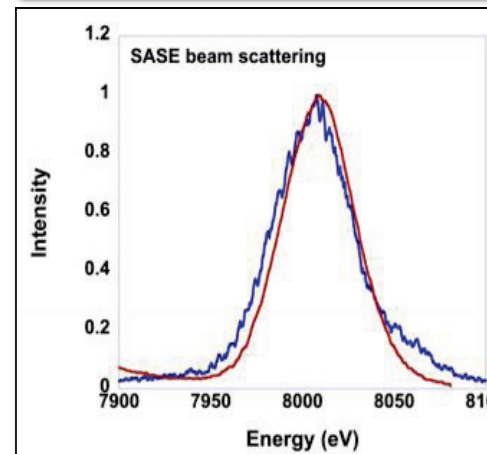
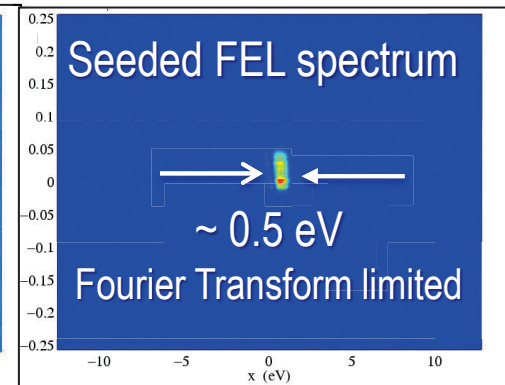
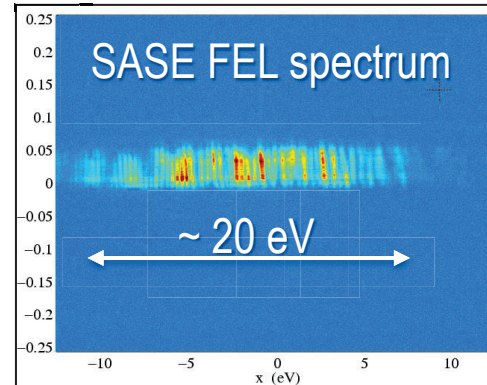
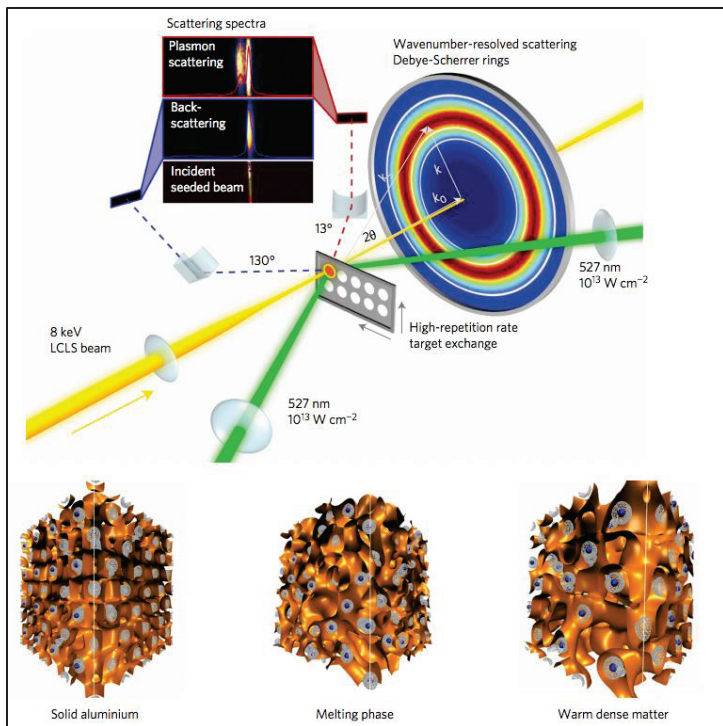
Hard x-ray self-seeding @ LCLS

J. Amann et al., Nature Photon. 6, 693 (2012)

G. Geloni et al., DESY 10-133, 2010.



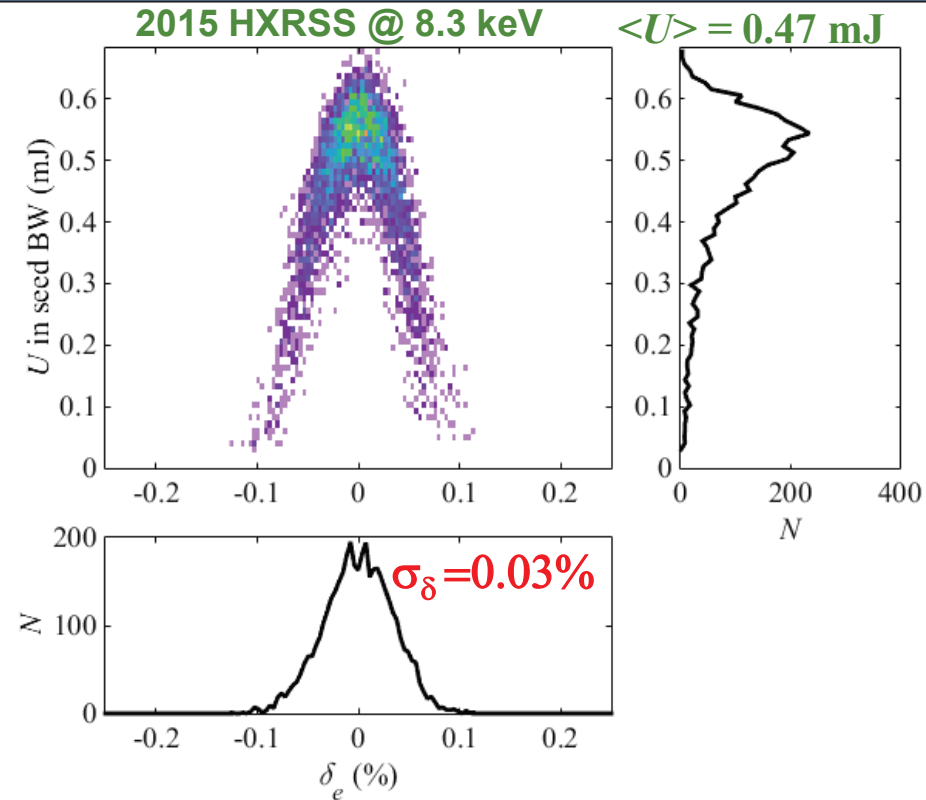
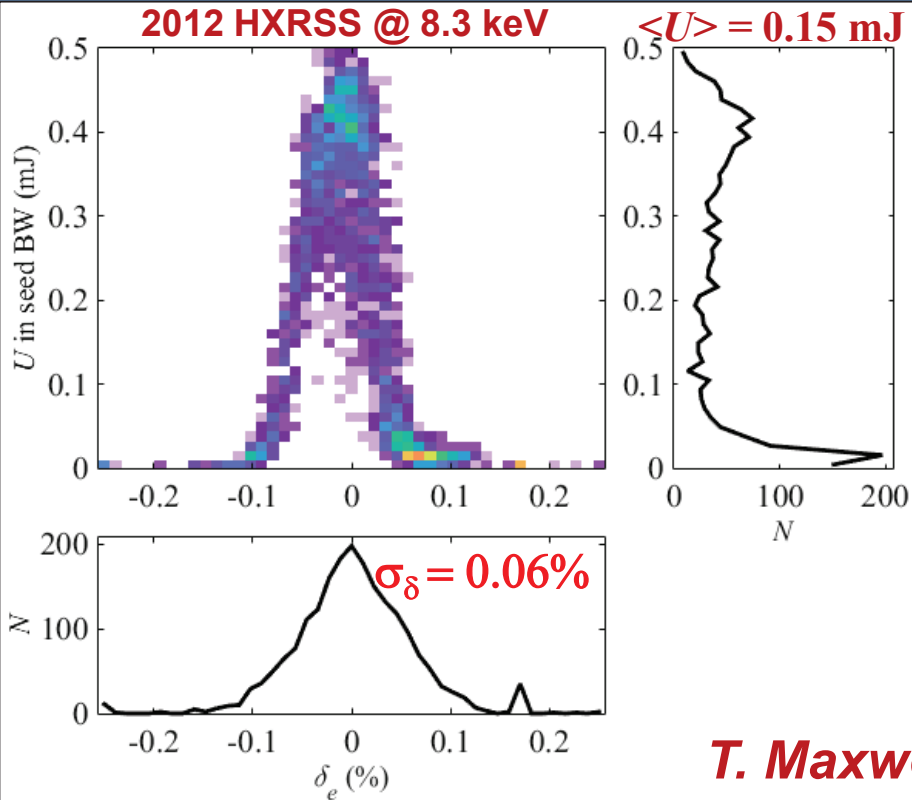
User operation started in 2013



L. Fletcher et al., J. Instrum. 8, C11014 (2013); Nature Photon. 9, 274 (2015)

Improving energy jitter for seeding

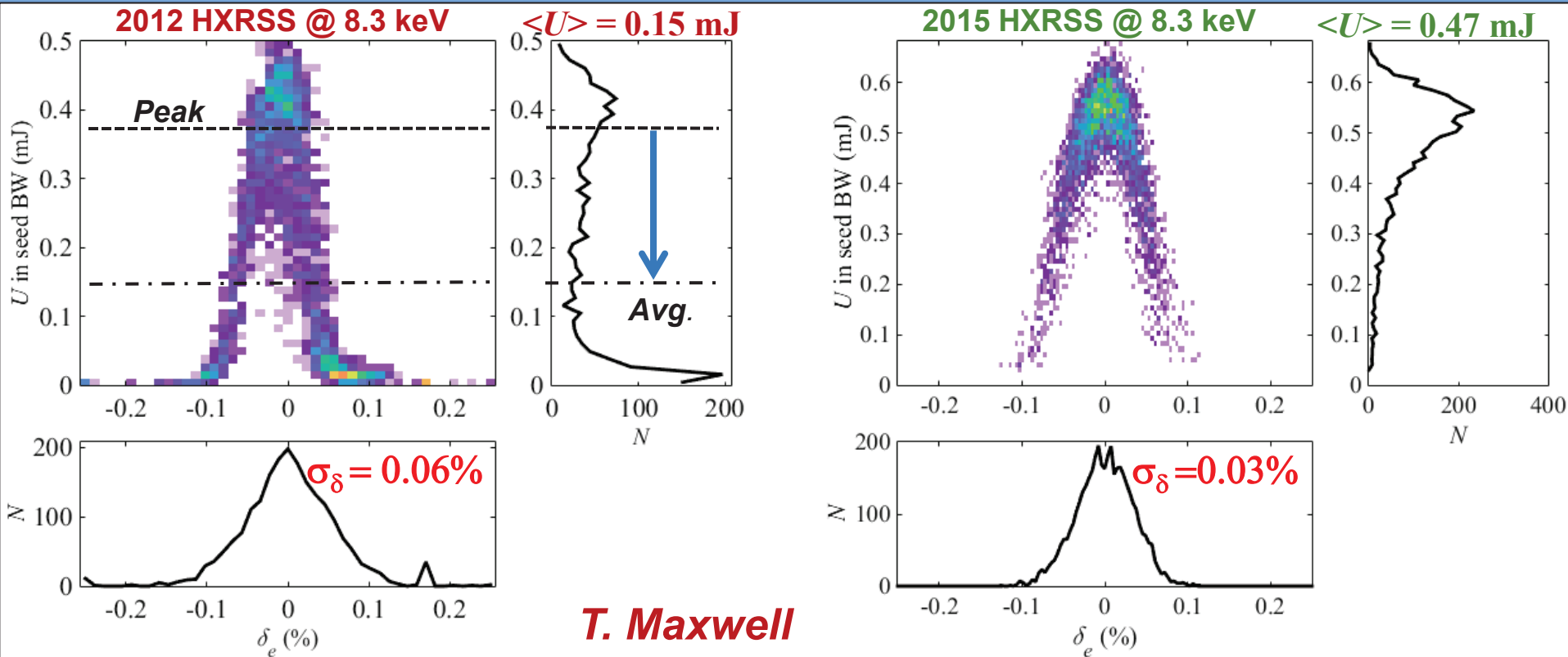
■ Seeded FEL intensity is sensitive to beam energy jitters (need $\sigma_\delta < \rho/2$).



T. Maxwell

Improving energy jitter for seeding

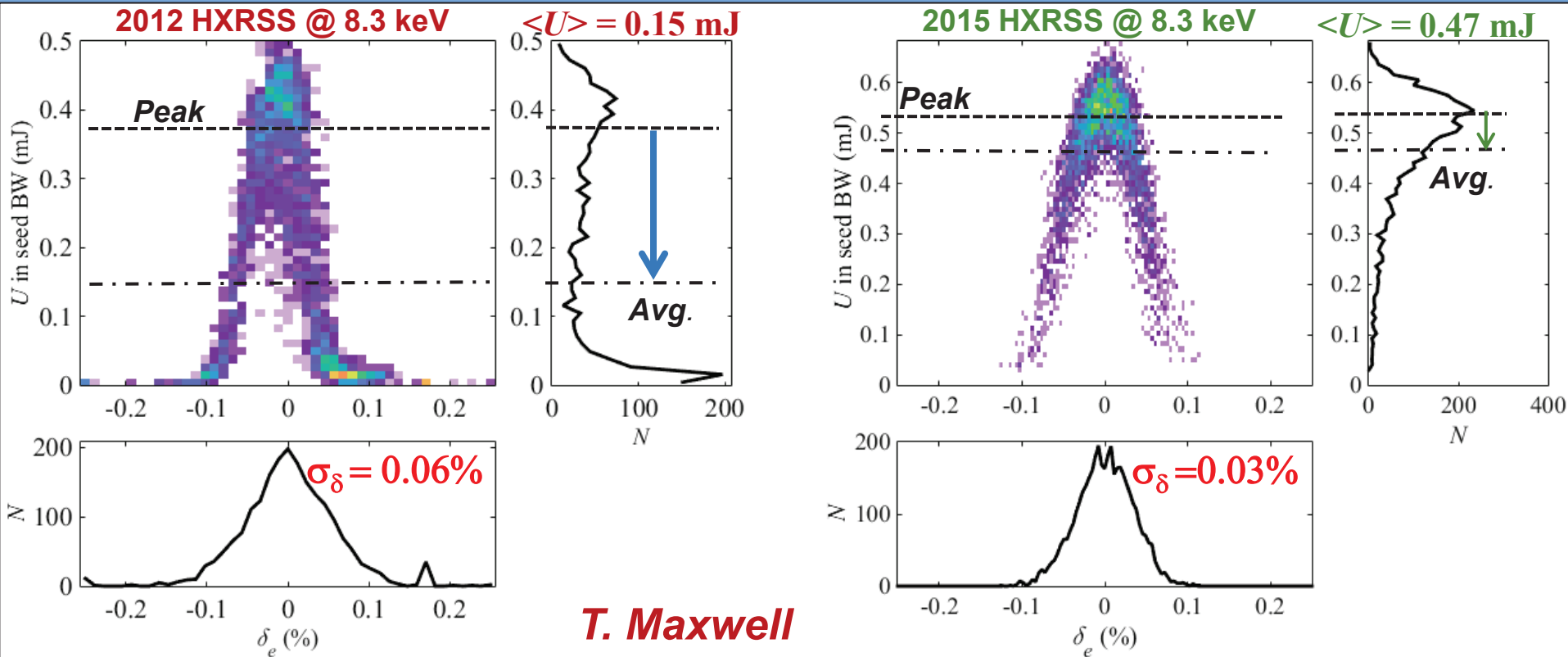
- Seeded FEL intensity is sensitive to beam energy jitters (need $\sigma_\delta < \rho/2$).



- HXR and SXR E jitter reduced by factor 2 since 2012
 - Injector RF tuned up, compression scheme optimized

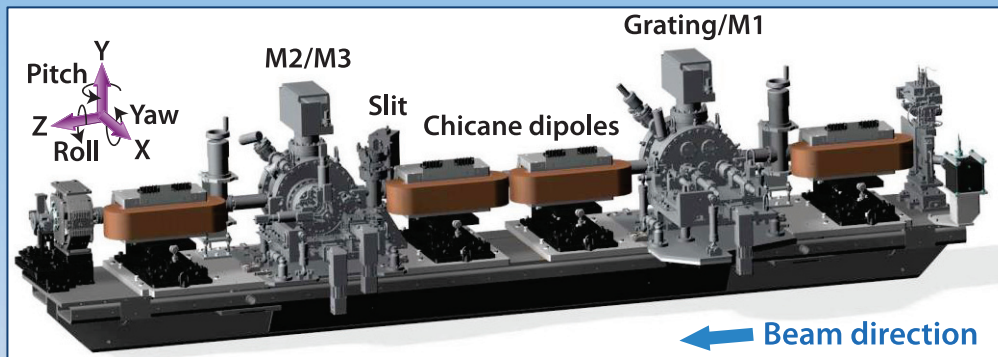
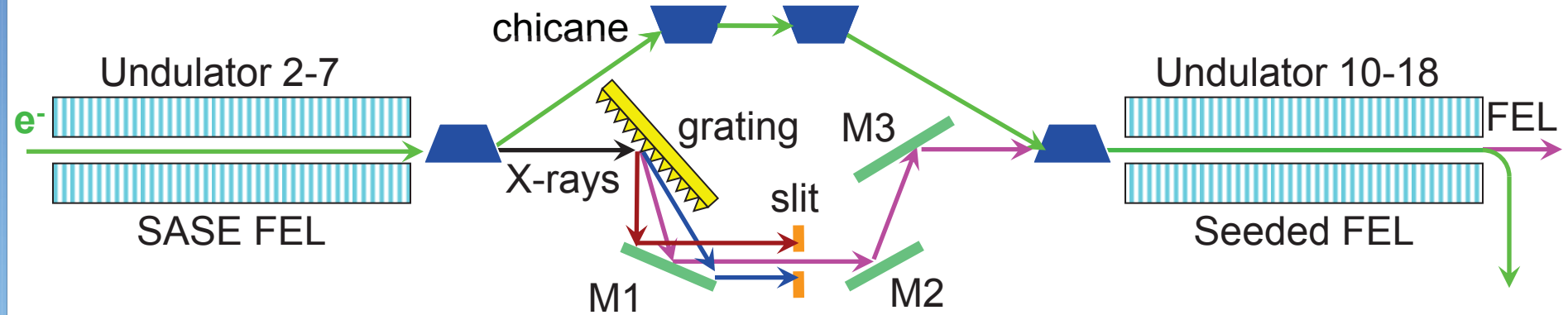
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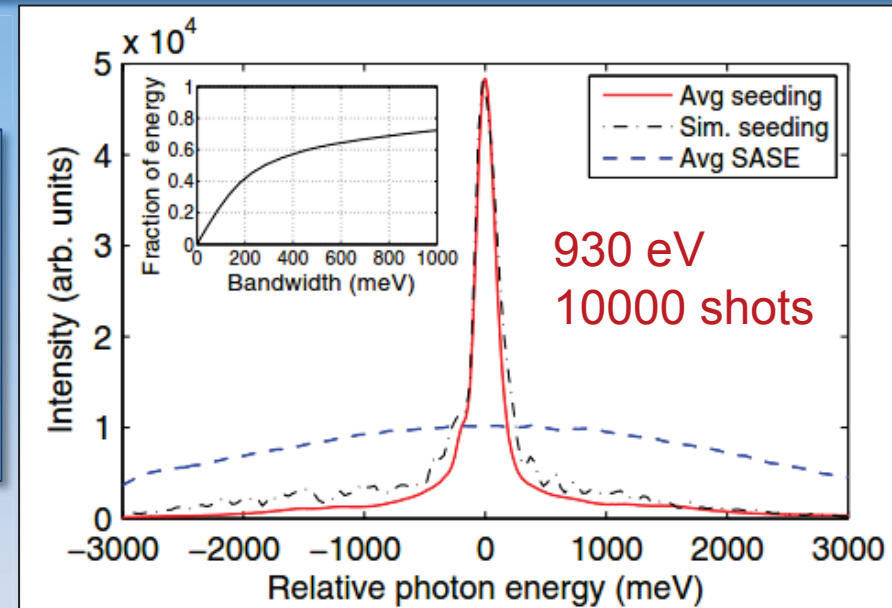
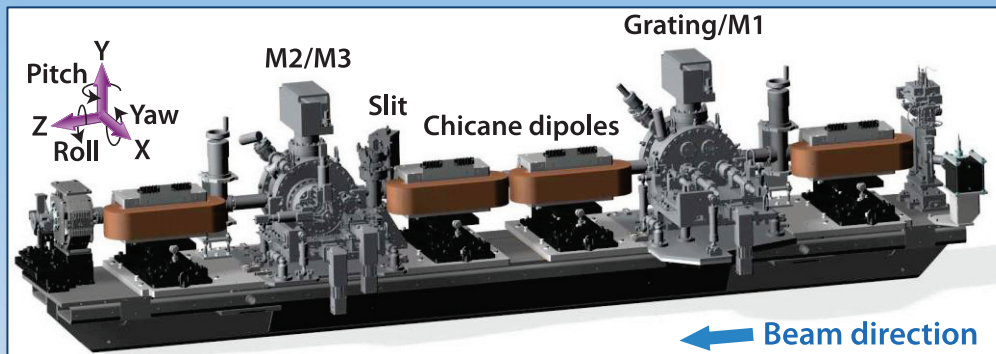
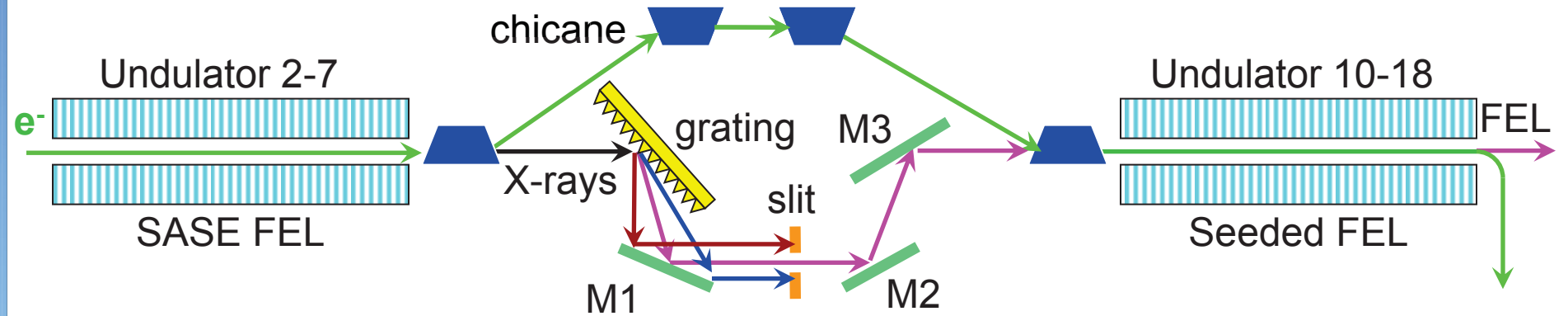


- HXR and SXR E jitter reduced by factor 2 since 2012
 - Injector RF tuned up, compression scheme optimized
- Hardware improvements incoming for further reduction
 - High-power RF terminating loads + higher-rated deuterium thyratrons

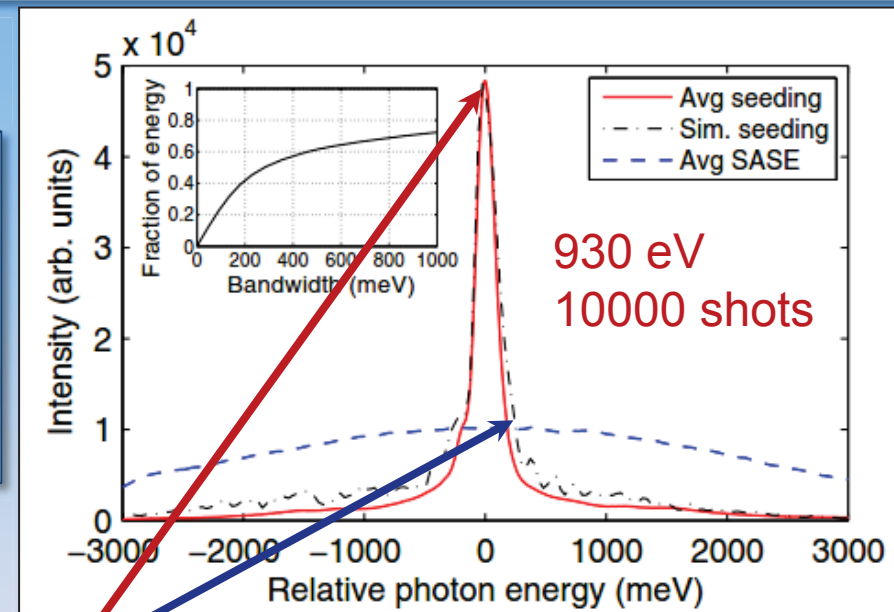
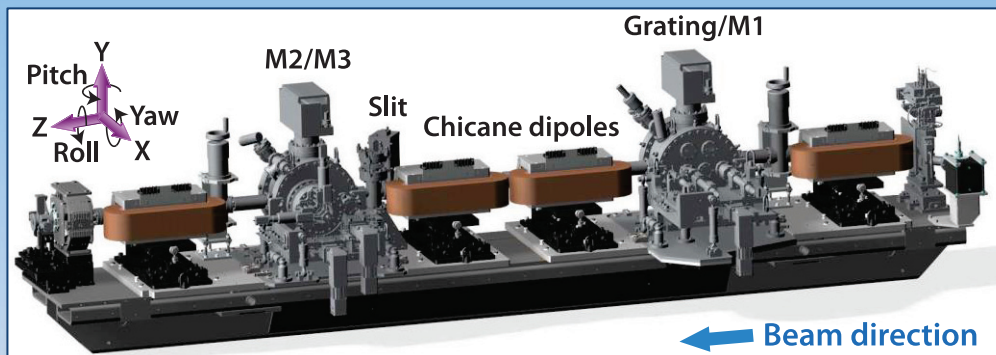
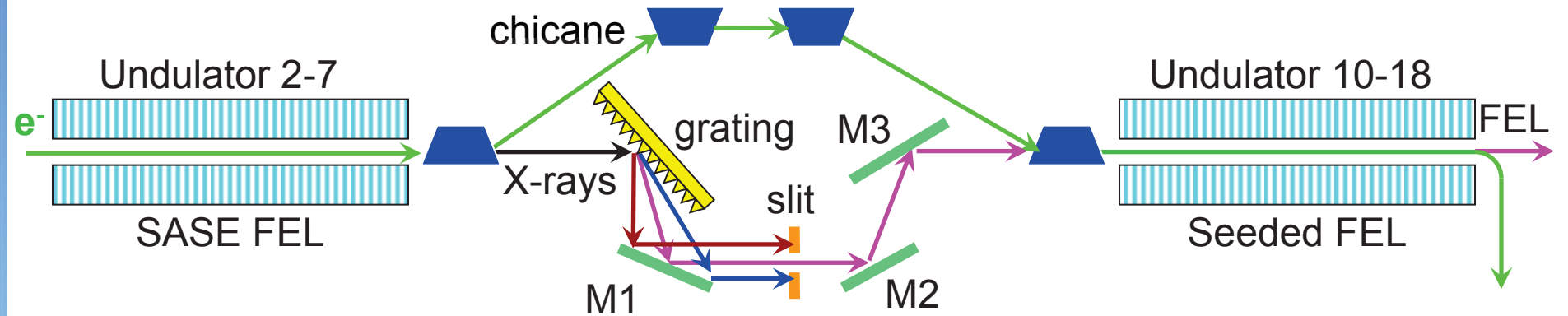
Soft X-ray Self-Seeding (SXRSS)



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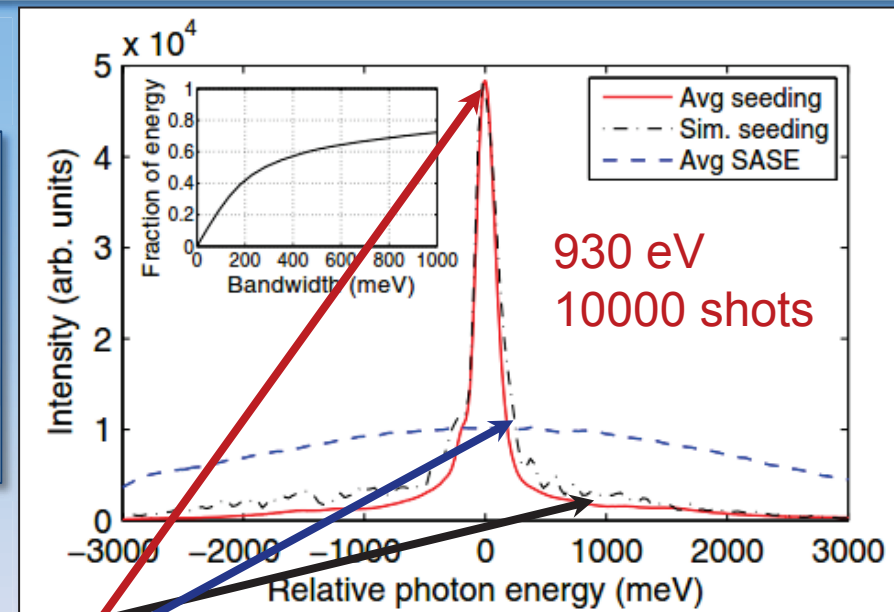
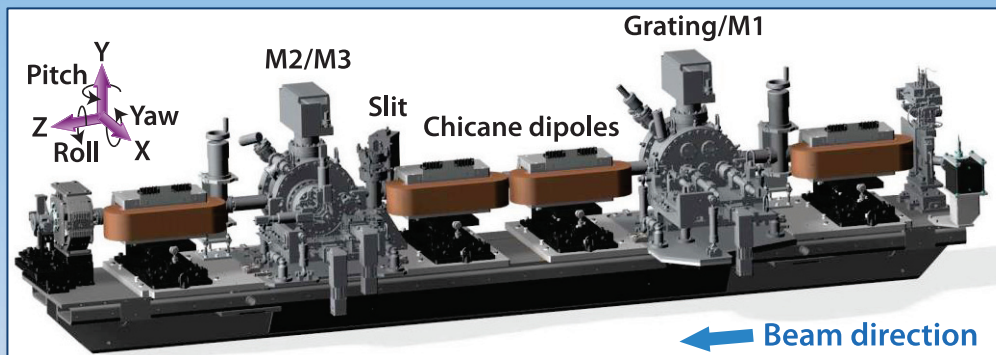
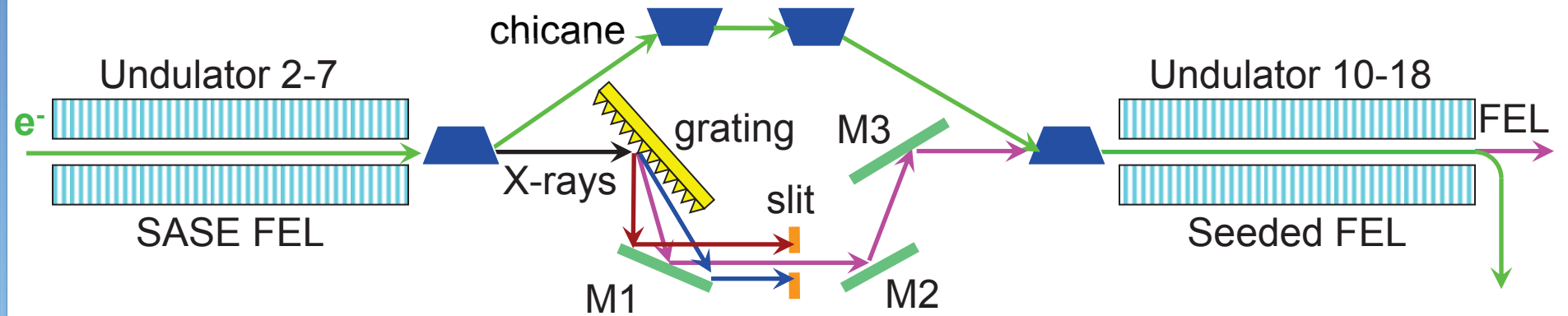


Soft X-ray Self-Seeding (SXRSS)



■ Spectral brightness $\sim 5\times$ higher than SASE

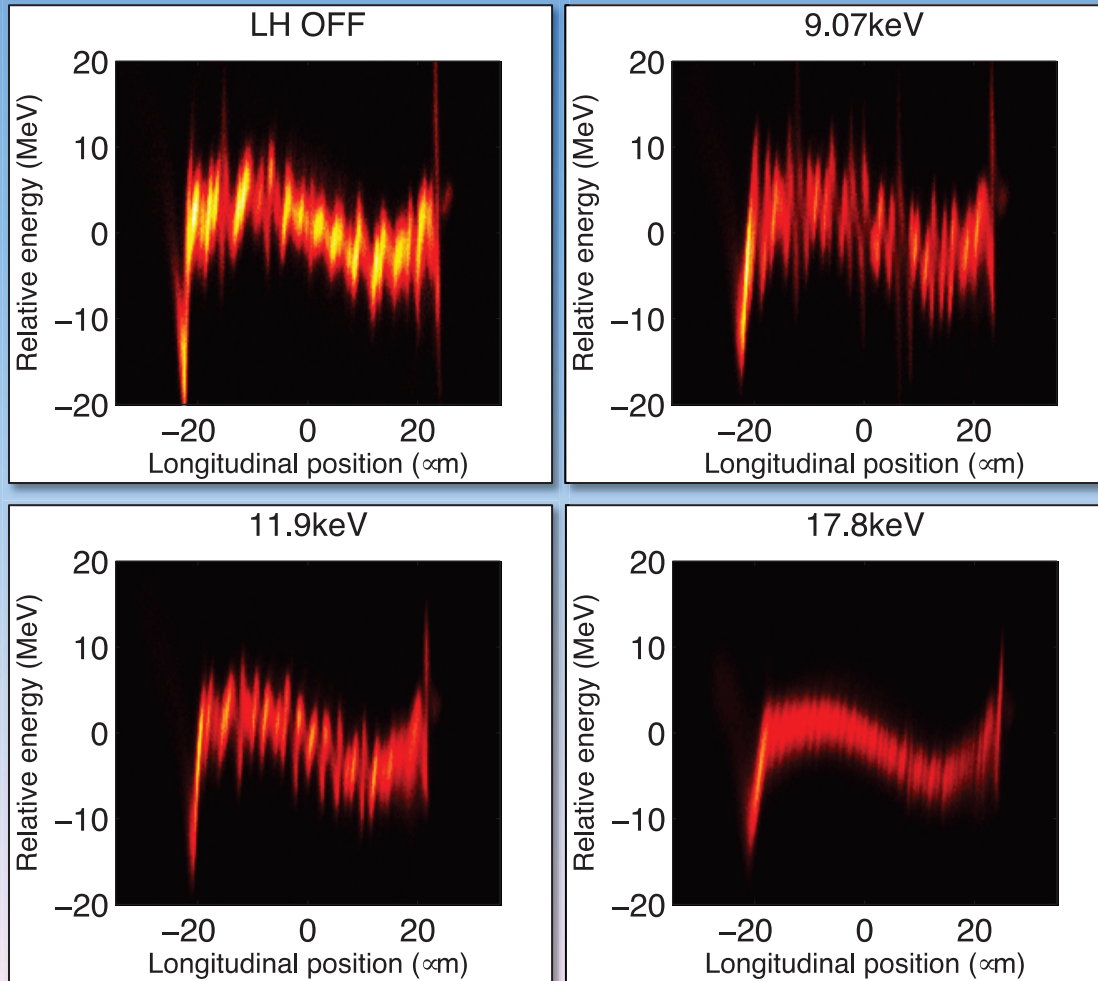
Soft X-ray Self-Seeding (SXRSS)



- Spectral brightness $\sim 5\times$ higher than SASE
- Observe spectral pedestals (may degrade certain applications w/o a mono).

Where are SXRSS pedestals from?

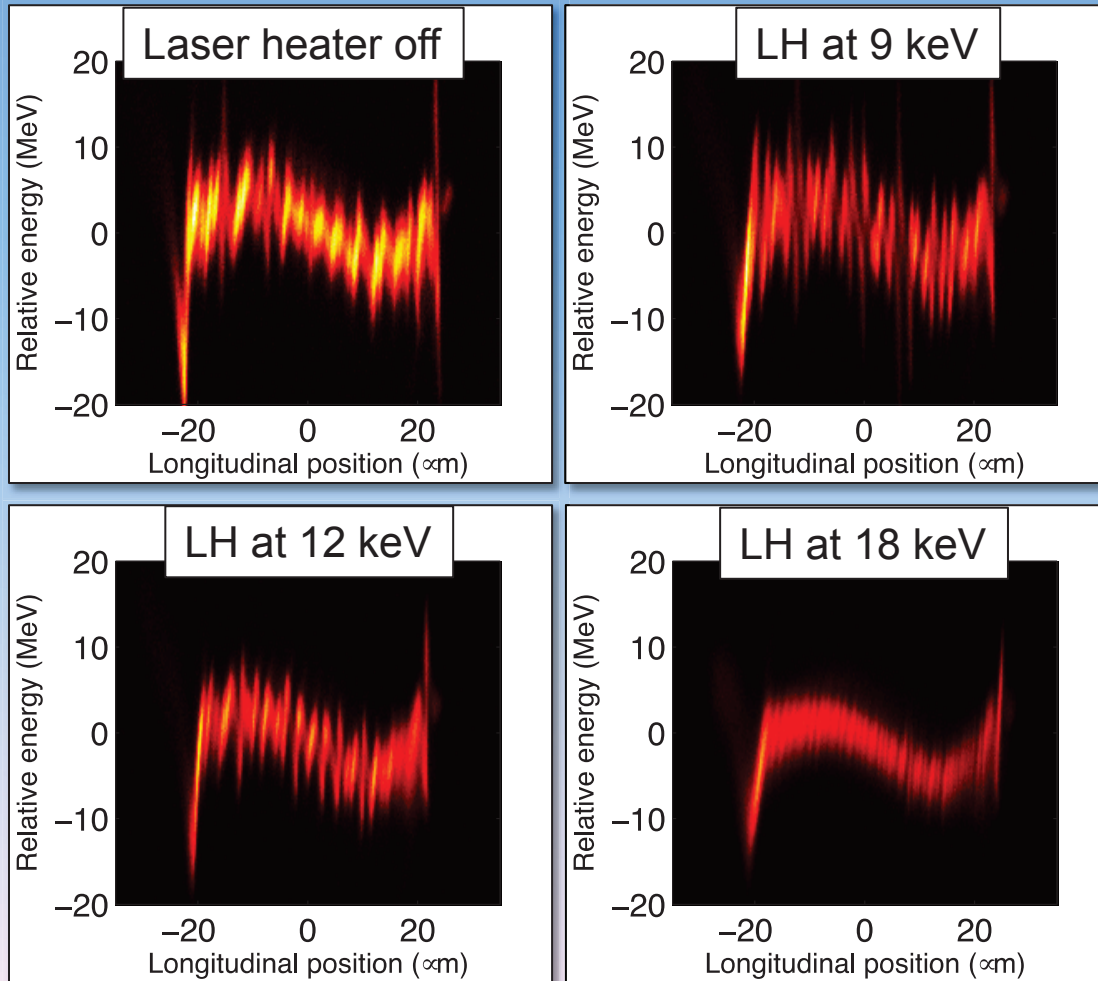
- Studies ruled out various sources (spectrometer noise, seeding monochromator optics)
- Focus on beam microbunching instability (μ BI) that drives sideband growth



Observation of μ BI at 4 GeV with X-band Transverse Deflector

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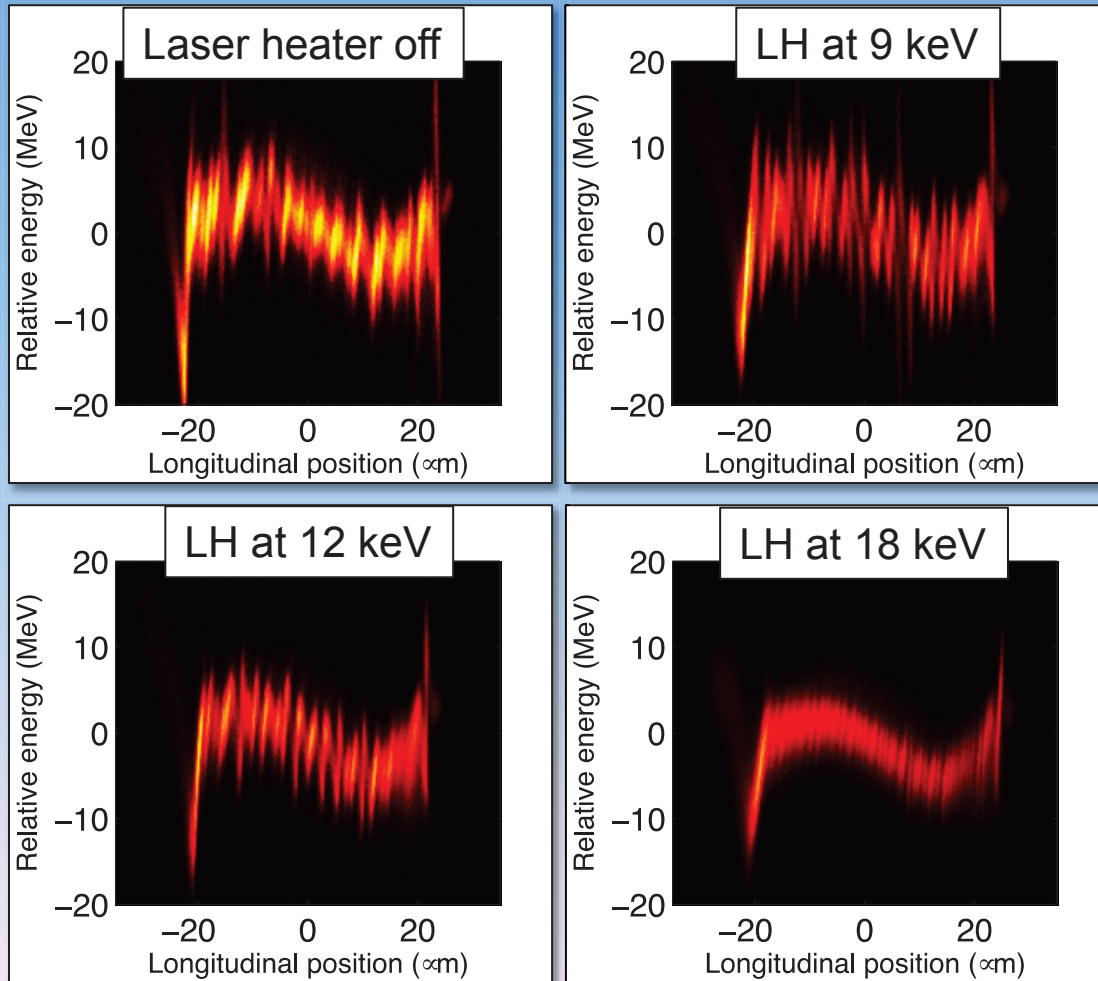
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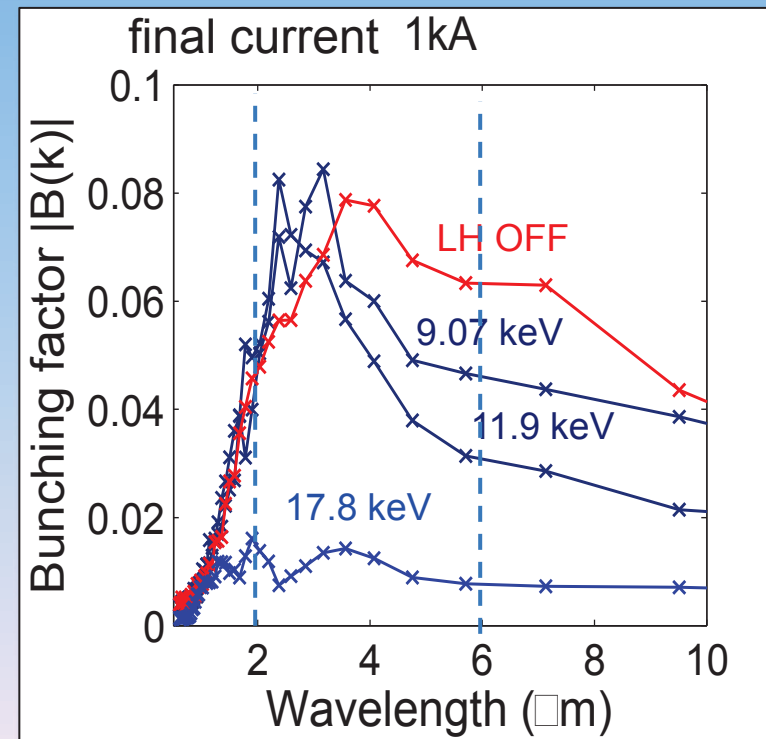
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Observation of μ BI at 4 GeV with X-band Transverse Deflector



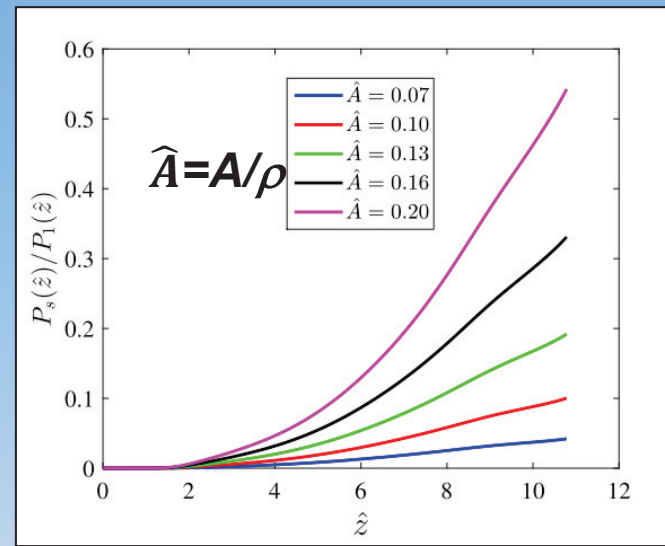
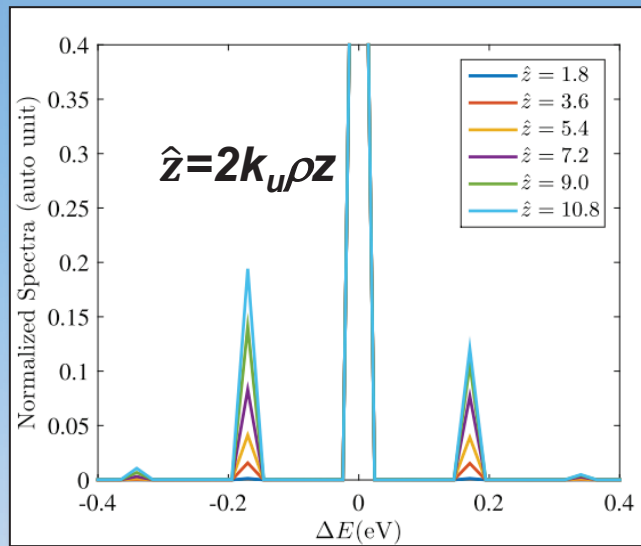
μ BI-induced sidebands in a seeded FEL

- Sideband ($1\text{nm}/2\mu\text{m} = 5 \times 10^{-4}$ within FEL bandwidth)
- Sideband power vs. seed signal power

Z. Zhang, WEP024

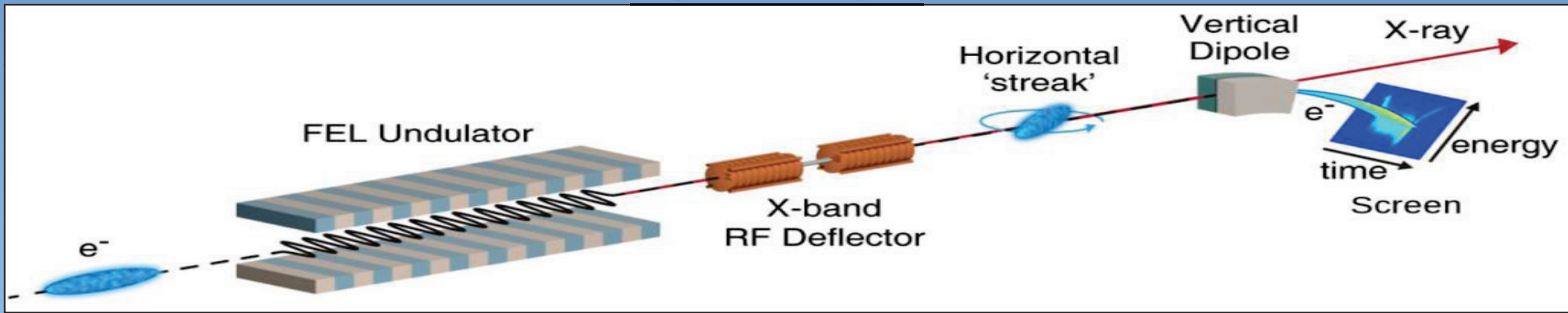
$$\frac{P_s(\hat{z})}{P_1(\hat{z})} = \frac{\hat{A}^2}{9} \hat{z}^2 = \frac{(2k_u \rho z)^2}{9} \frac{A^2}{\rho^2}$$

Sideband grows quadratically with undulator length and E mod. amplitude

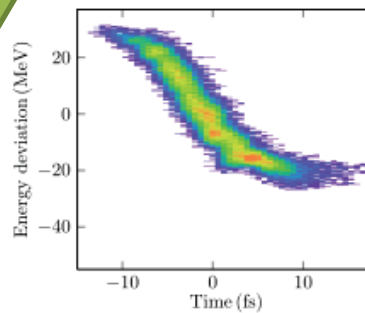
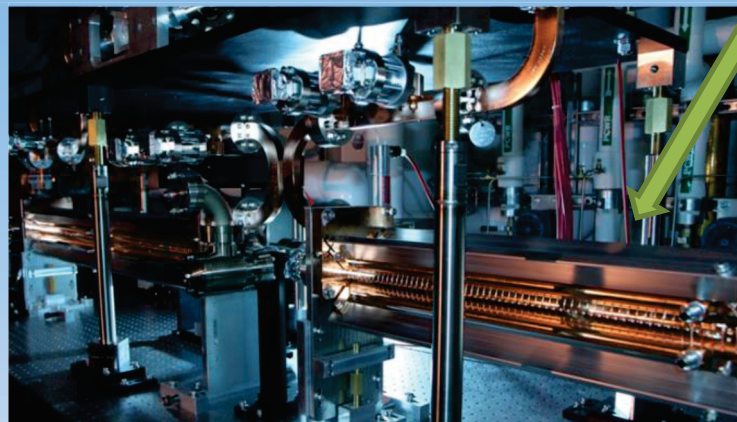
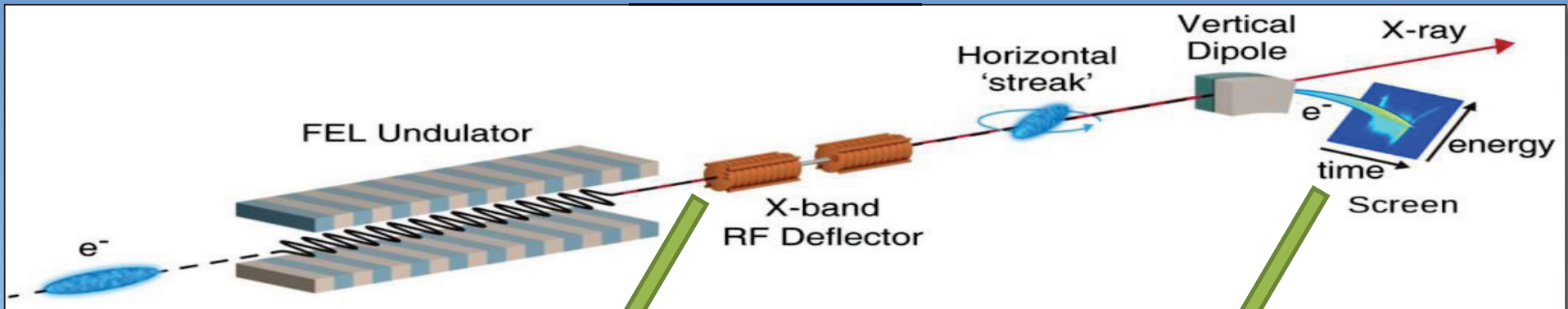


- Take $2\rho k_u z \sim 9$ for seeding saturation,
In order for sideband power < seed power $\rightarrow A < 0.3 \rho$
Take LCLS E = 4 GeV, $\rho \sim 10^{-3}$, $A < 1$ MeV (very stringent requirement!)
- Needs very uniform beams in longitudinal phase space
 - Laser heater shaping for more effective heating **S. Li, WEP005**
 - Single BC instead of 2 BCs to reduce μ BI (similar to FERMI)

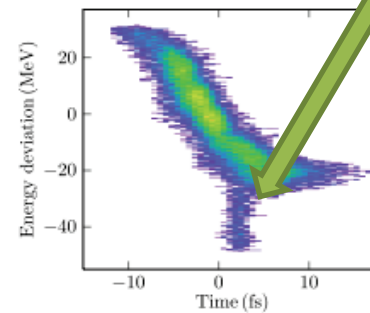
fs X-ray diagnostics with an X-band transverse deflector (XTCAV)



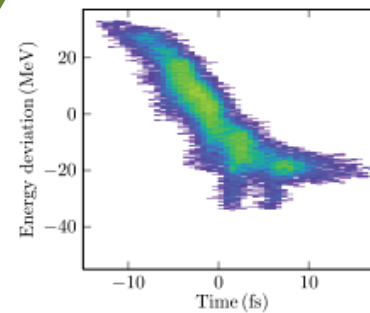
fs X-ray diagnostics with an X-band transverse deflector (XTCAV)



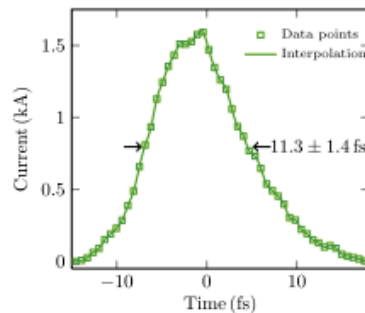
(a) Lasing-off



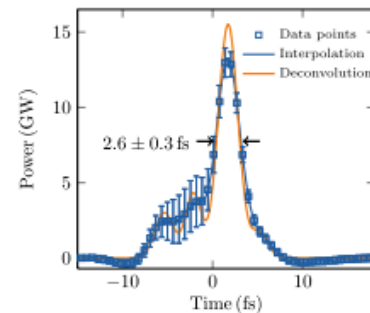
(b) Lasing-on, shot-1



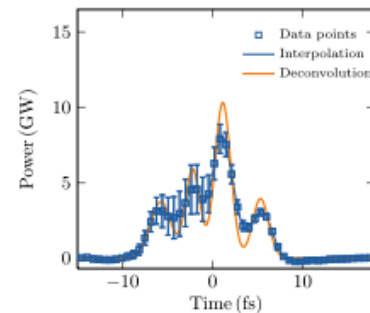
(c) Lasing-on, shot-2



(d) Electron current

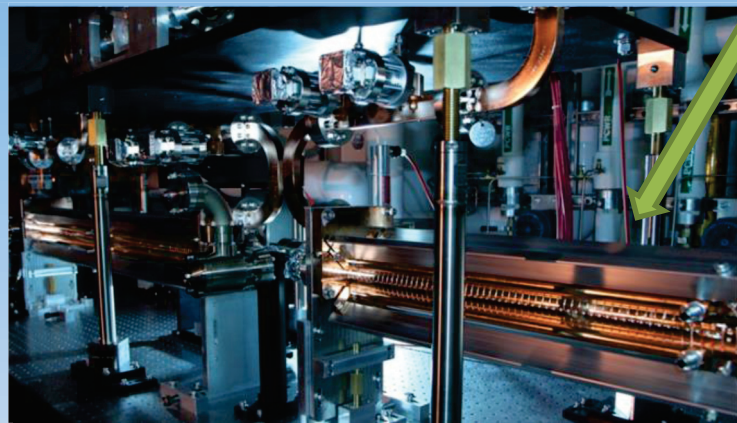
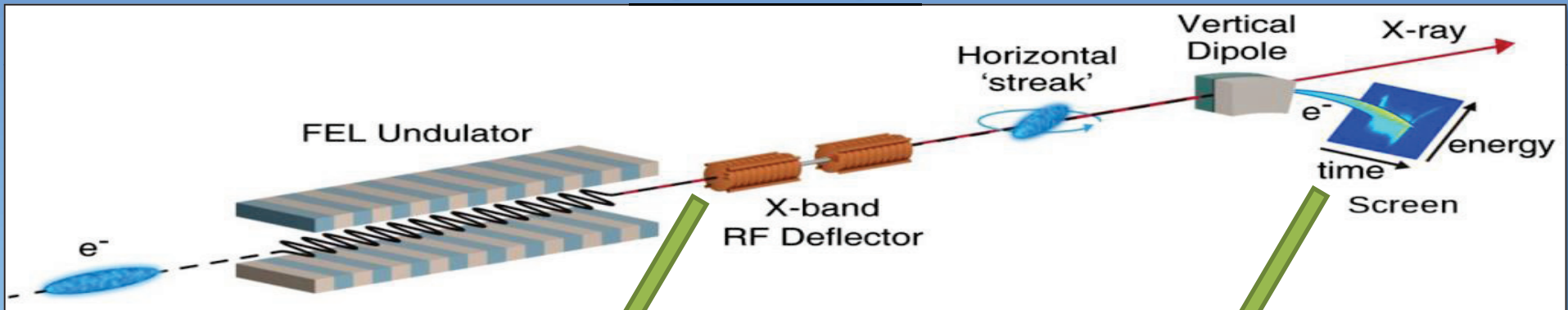


(e) X-ray power, shot-1

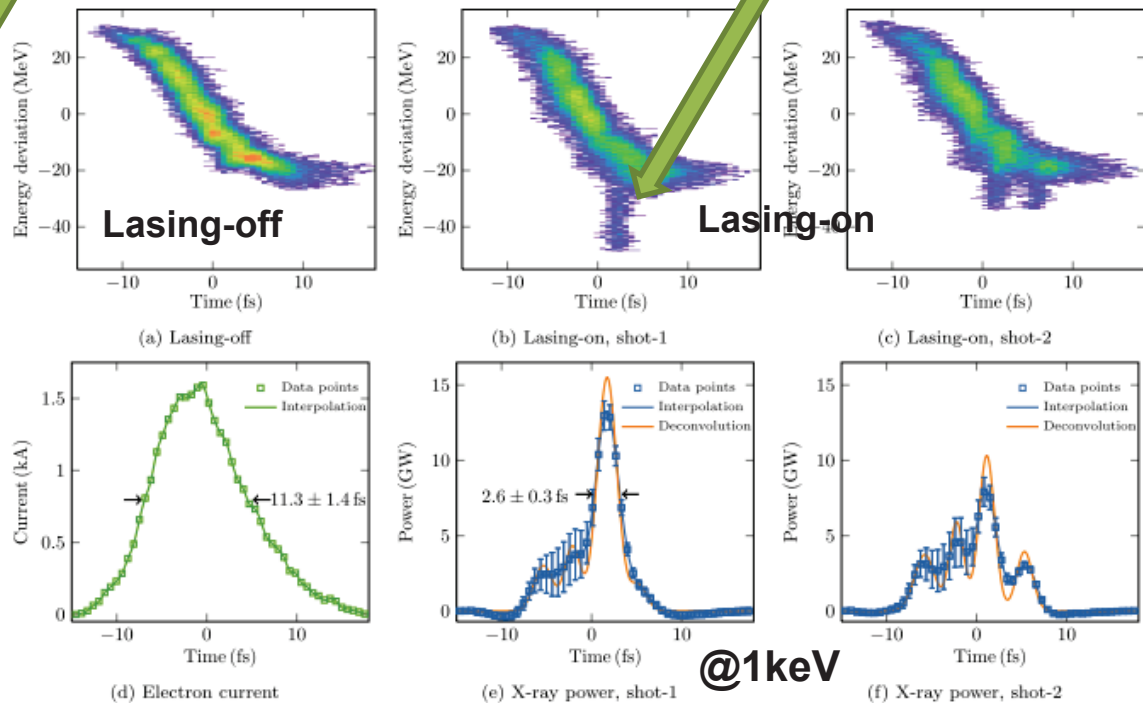


(f) X-ray power, shot-2

fs X-ray diagnostics with an X-band transverse deflector (XTCAV)



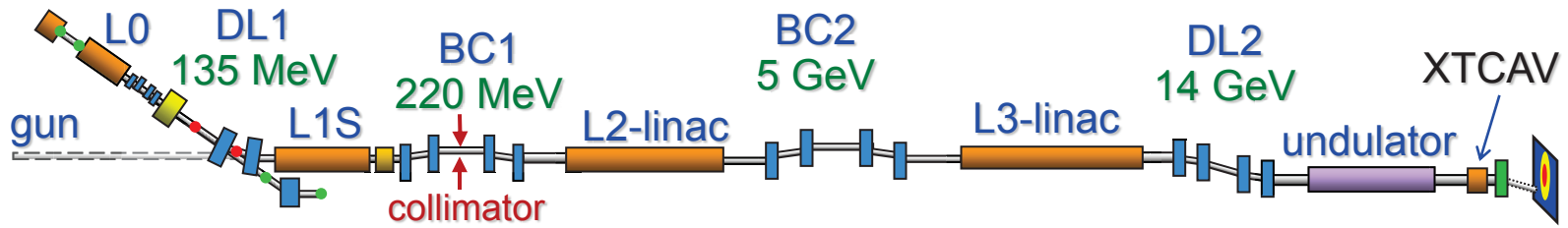
*C. Behrens, et al., Nature Comm.,
5, 3762 (2014)*



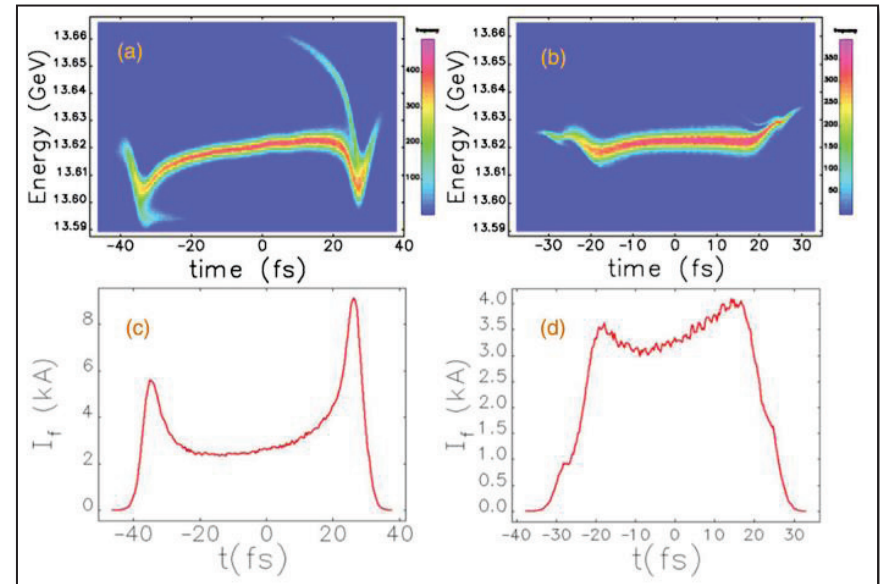
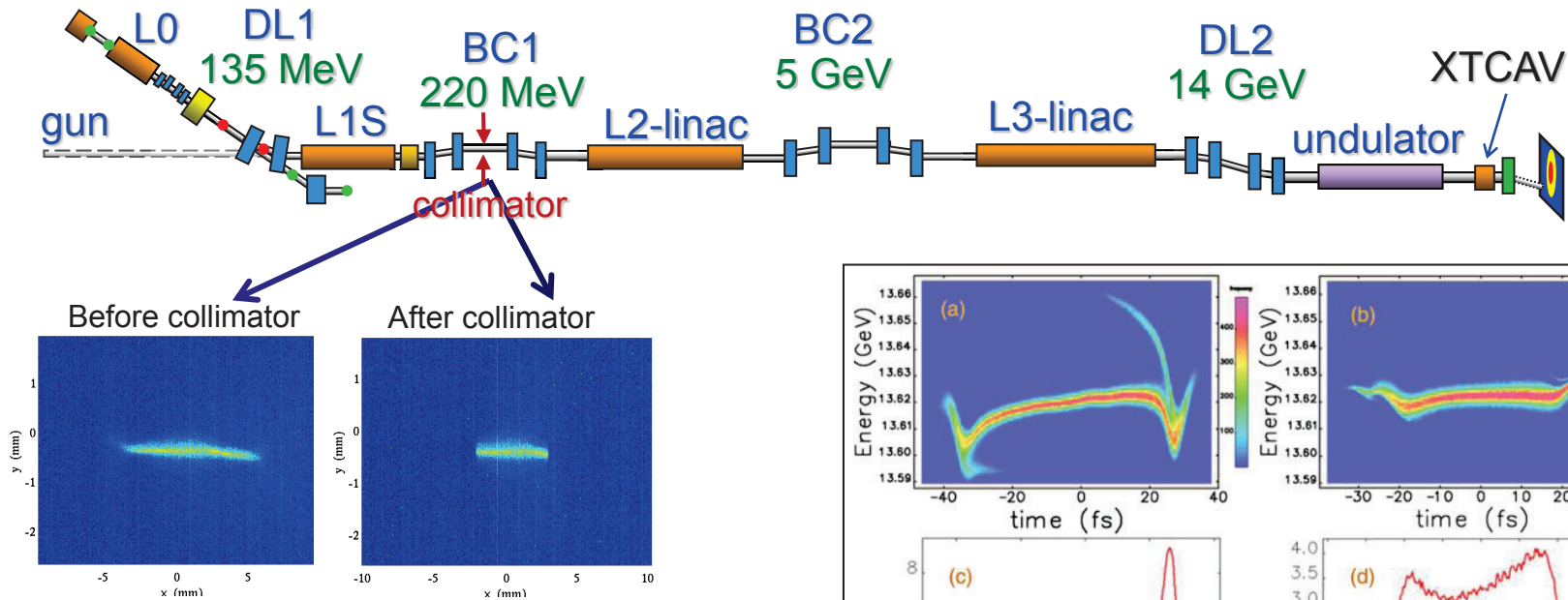
Reconstructed X-ray profile 2.6 fs (FWHM)

Y. Ding, TUA01

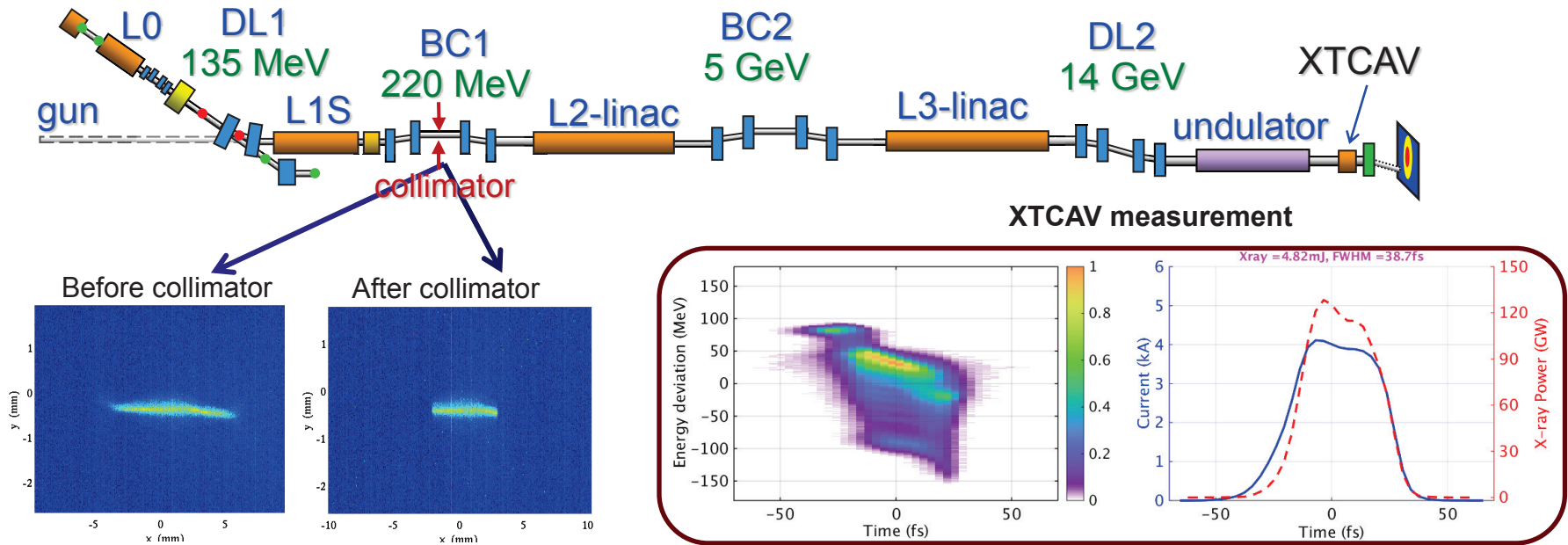
Generating high power FEL with a horn-collimated beam



Generating high power FEL with a horn-collimated beam

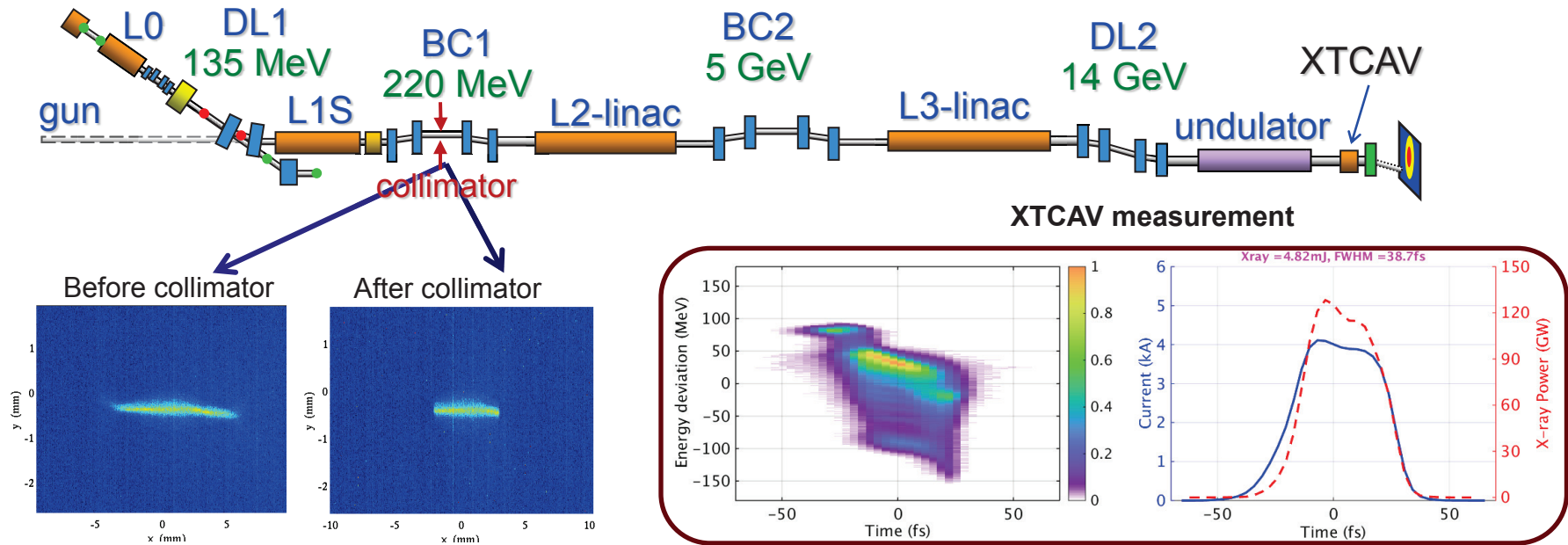


Generating high power FEL with a horn-collimated beam

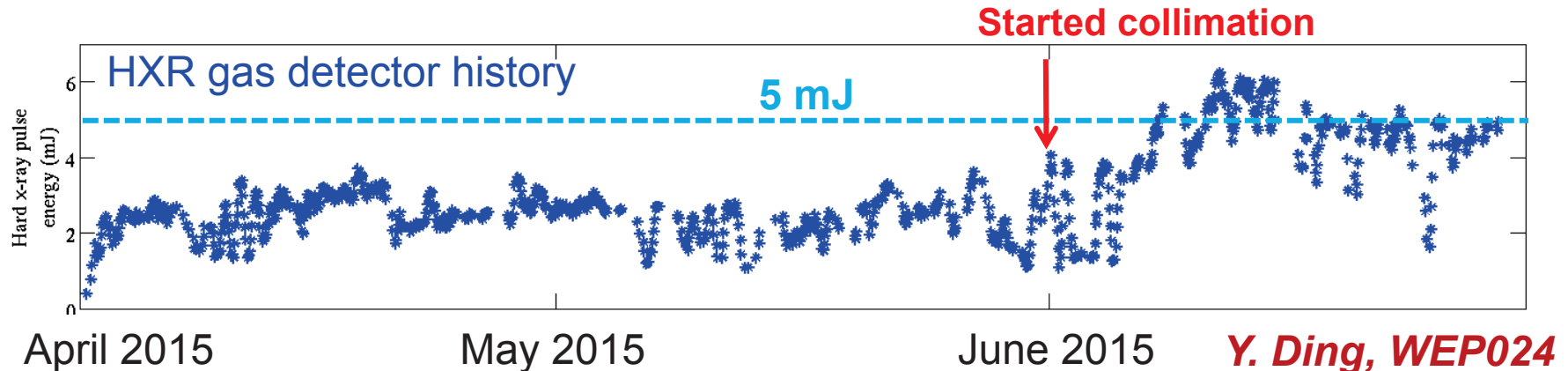


simulations

Generating high power FEL with a horn-collimated beam

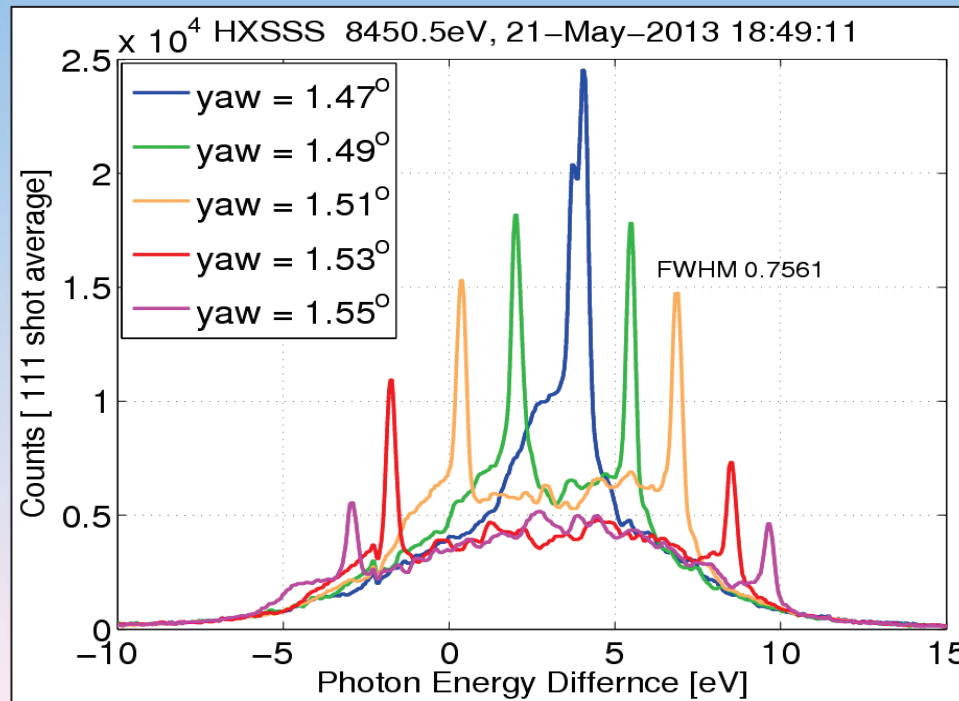
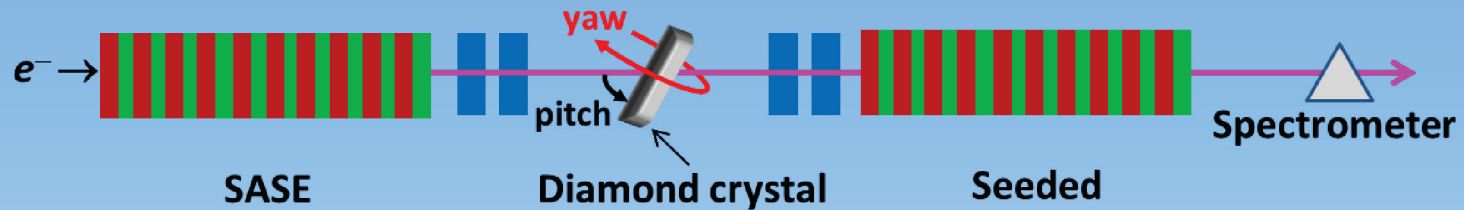


- **Routinely** achieved hard x-ray pulse energy $\sim 4\text{--}5$ mJ and ~ 100 GW (from XTCAV).
- Further improvement on this mode, such as compensating BC2 CSR effect for shorter bunches and a few hundred GW power, will be studied.



2-color seeded FEL

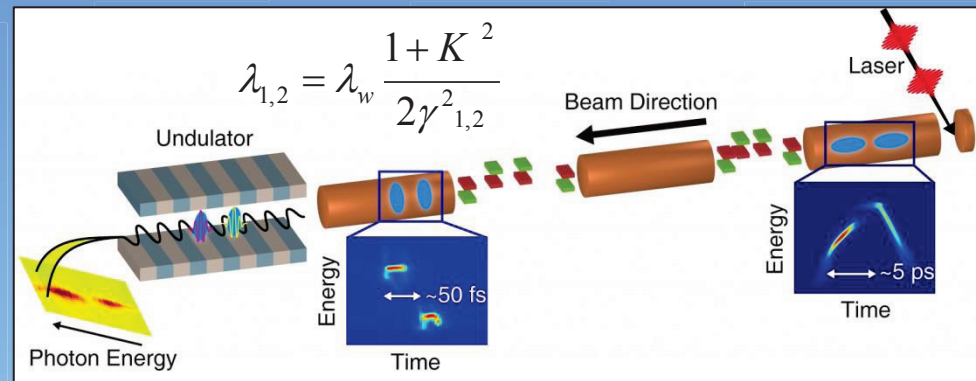
- Exciting and probing samples above and below absorption edges, imaging at different wavelengths.
- By adjusting the yaw angle in addition to the usual pitch angle of the seeding crystal, two-color hard X-ray self-seeded FEL with two different crystal diffraction planes were demonstrated.



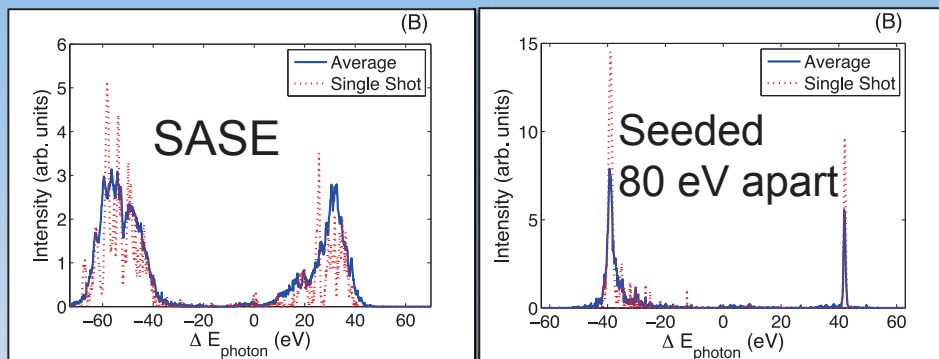
A.A. Lutman et al.
PRL 113, 254801, 2014

Twin-bunch, high-power 2-color FEL

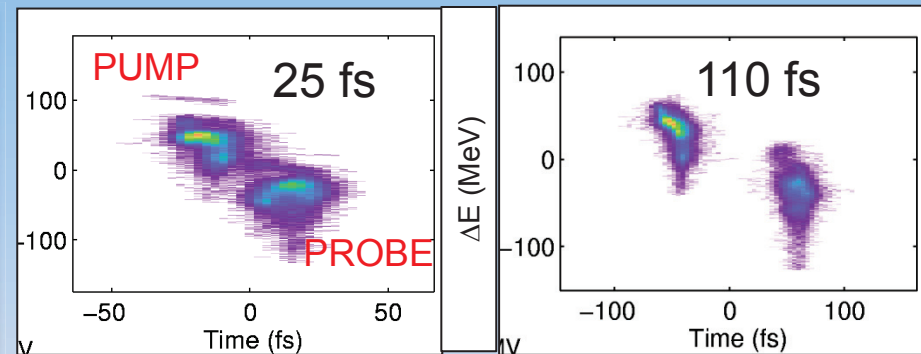
- Two bunches (in same RF bucket) are accelerated and compressed
- Each bunch lases at its own FEL color, hence improving the peak power over previous schemes by one order of magnitude



- Allows for independent control of energy and time delay



Hard X-ray FEL spectra



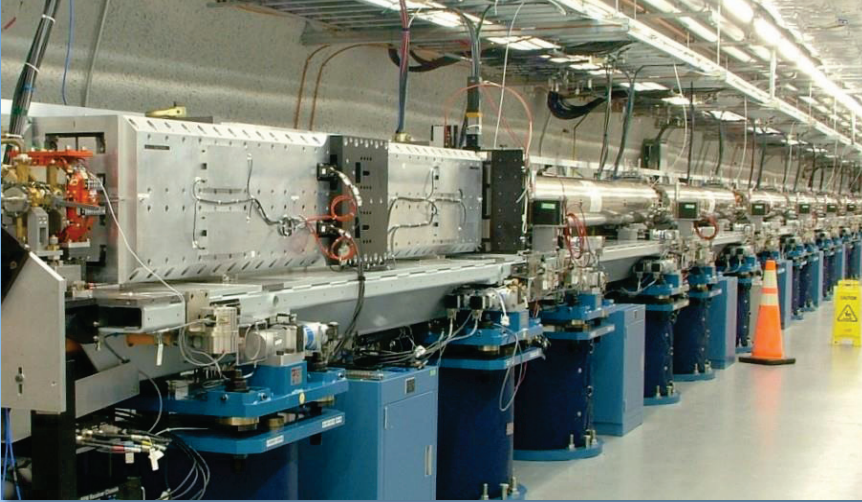
Time delays (measured w/ XTCAV)

- Extend to soft X-rays by combining with double-slotted foil to generate shorter bunches and negative to positive delays

A. Marinelli, WEP080

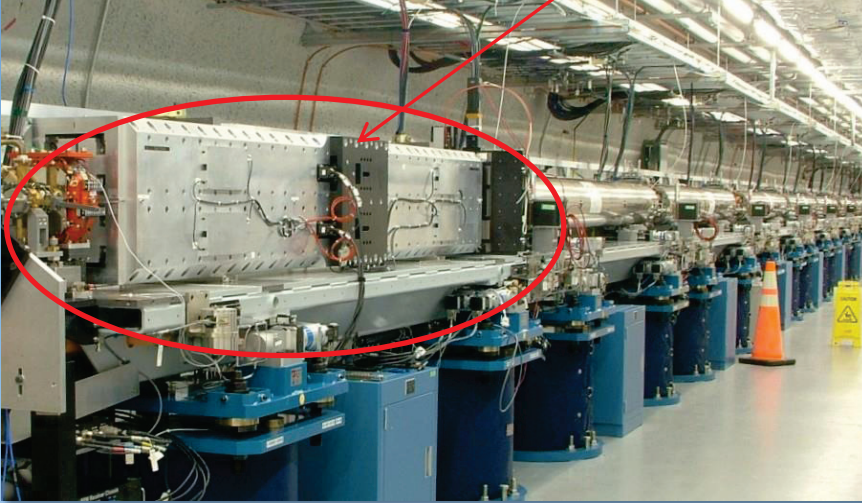
Novel Undulators

- A new type of undulator (“Delta”) has been developed and commissioned at LCLS for polarization control. *(H.-D. Nuhn, WED01)*



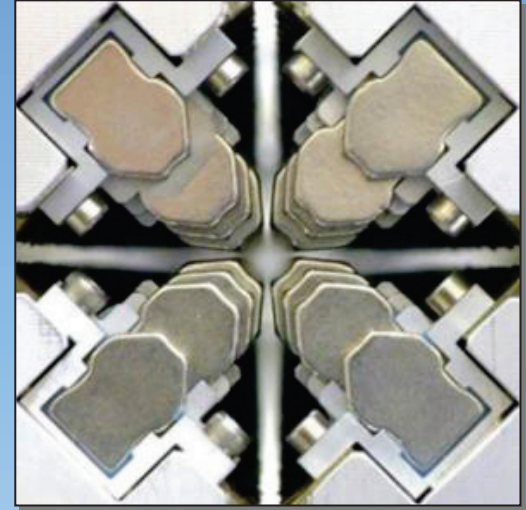
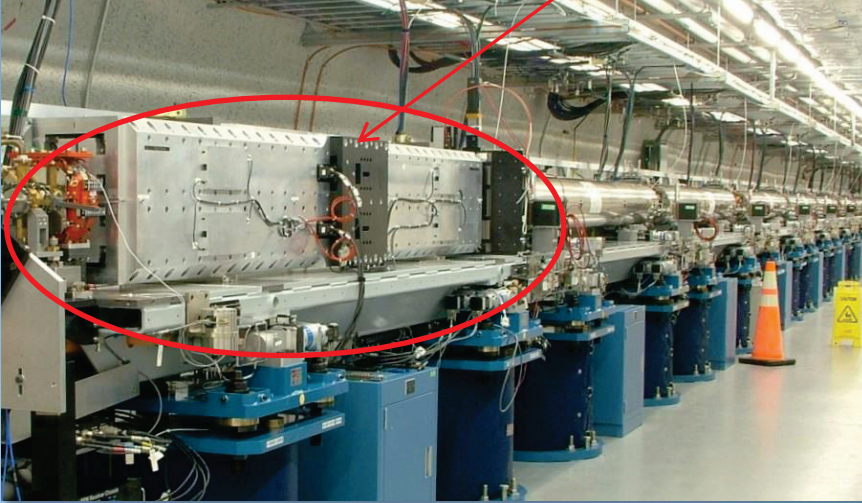
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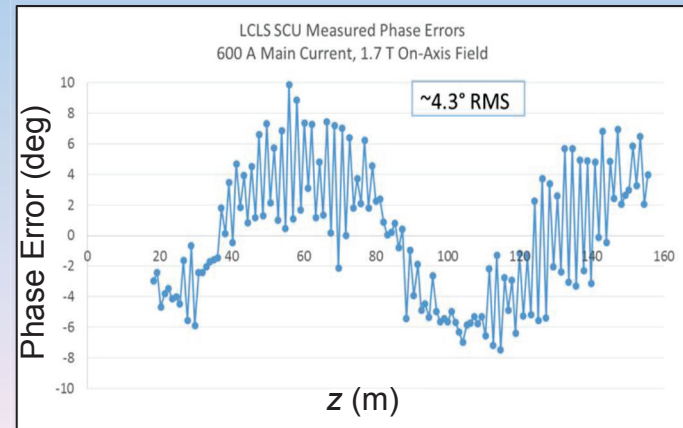


Novel Undulators

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- Superconducting undulator improves FEL efficiency and extends spectral range for LCLS-II (*P. Emma et al., Proc. of FEL 2014*)



ANL prototype (NbTi) phase error controlled to < 5 deg rms

Compact XFELs

- Synchrotrons can provide 30-60 beamlines per machine, much more cost effective than XFELs (1-2 beamlines at present)
- Strong motivation to develop compact coherent x-ray sources at a small fraction of cost and size of big XFELs
- Search “Compact XFEL” on google

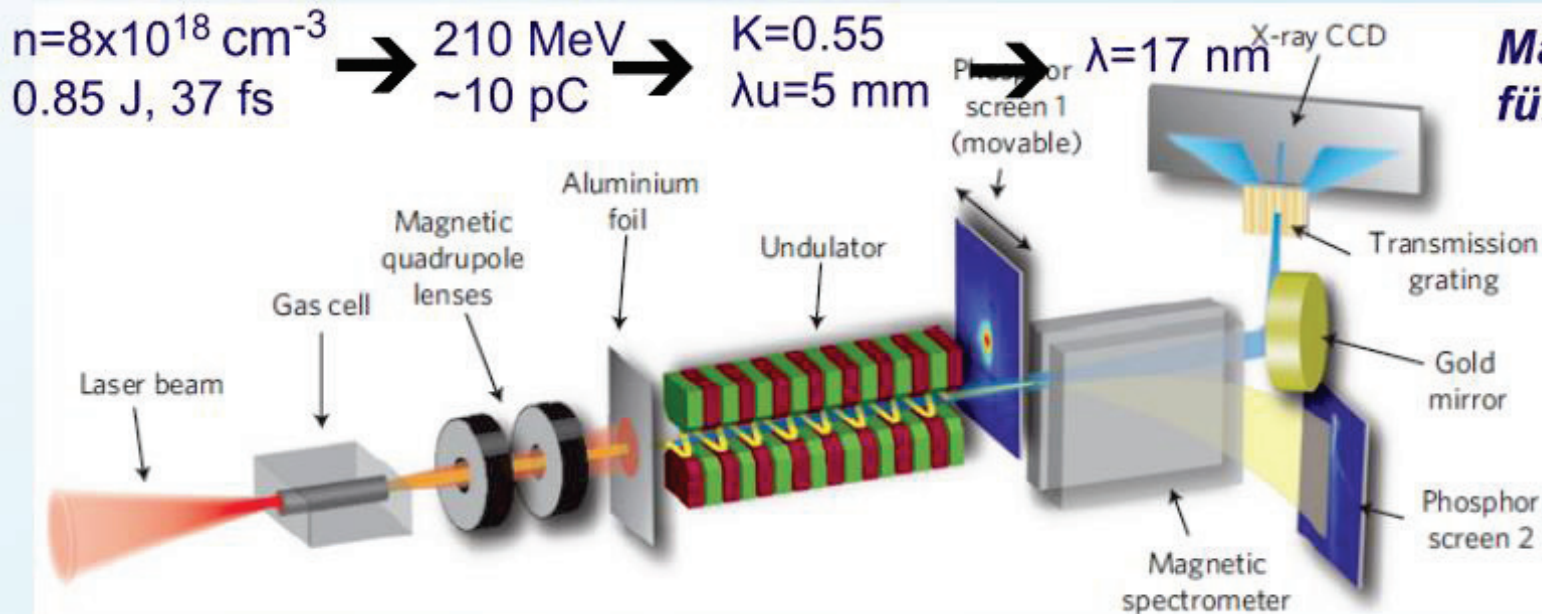


- Many ideas exist to make compact coherent sources
 - Some relies on Inverse Compton Scattering (replacing cm-period undulator with μm wavelength laser $\rightarrow$ lower e-beam energy by 100)
 - Some relies on advanced accelerator methods (e.g., plasma accelerator boosts gradient by 100x or more) *M. Hogan, WEC01*

Laser plasma accelerators (LPAs)

C. Schroeder

- Energy: $\sim 100 \text{ MeV} - 4 \text{ GeV}$
 - Obtained 100TW-1PW laser pulses in mm - cm long plasmas
- Charge: $\sim 1 - 100 \text{ pC}$
 - Depends on tuning, energy spread due to beam loading
- Energy spread: $\sim 1 - 10\%$ level
 - Depends on amount of charge, trapping physics
- Normalized Emittance: $\sim 0.1 \text{ micron}$
 - Based on divergence measurements ($\sim 1 \text{ mrad}$) and e-beam spot ($\sim 0.1 \text{ micron}$)
 - Improved measurements needed



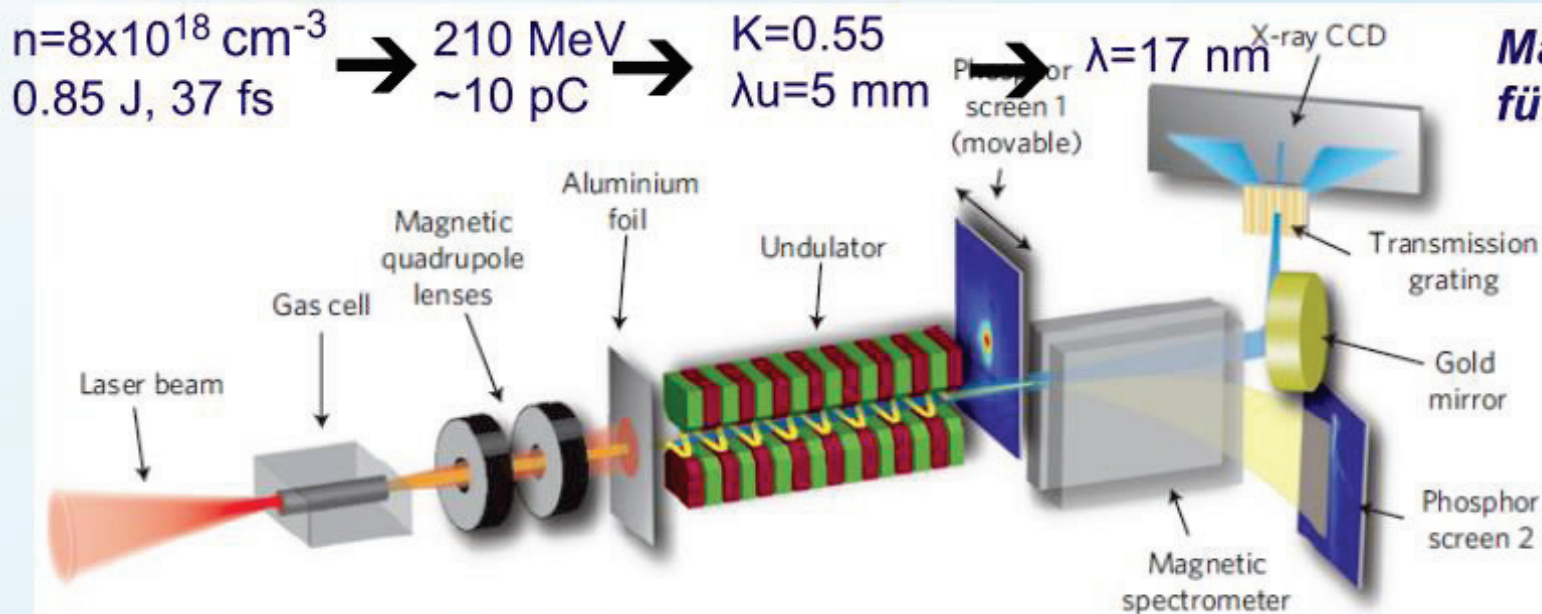
*Max-Planck-Institut
für Quantenoptik*

M. Fuchs et.
al., Nature
Physics 5, 826
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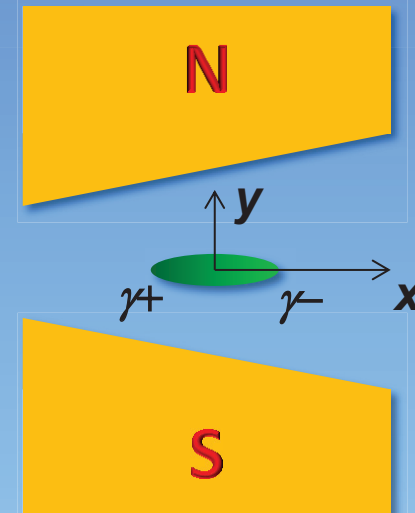
Compact XFELs using Transverse Gradient Undulator

- TGU can be used to compensate large energy spread of laser plasma accelerator beams for a compact FEL

$$\alpha = \frac{dK/K_0}{dx}$$

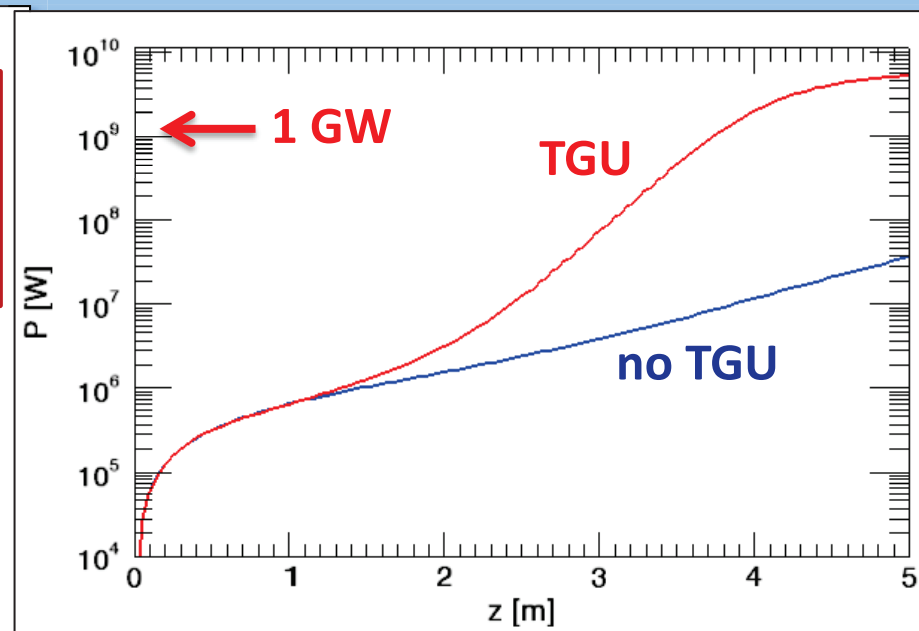
- FEL resonance can be satisfied for all energies if dispersion at undulator is chosen as

$$\eta = \frac{2 + K_0^2}{\alpha K_0^2}$$



T. Smith et al., J. Appl. Phys. 1979

Parameter	Symbol	EUV	X-ray
Beam energy	$\gamma_0 mc^2$	500 MeV	1 GeV
Norm. transv. emittance	$\gamma_0 \epsilon_x$	0.1 μm	0.1 μm
Peak current	I_0	5 kA	10 kA
Flat-top bunch duration	T	10 fs	5 fs
Rel. rms energy spread	σ_δ	2%	1%
Undulator type		Hybrid	SC
Undulator period	λ_u	2.18 cm	1 cm
Undulator length	L_u	5 m	5 m
Undulator parameter	K_0	1.85	2
Transverse gradient	α	43 m^{-1}	150 m^{-1}
Horizontal dispersion	η	3.7 cm	1 cm
Resonant wavelength	λ_r	31 nm	3.9 nm

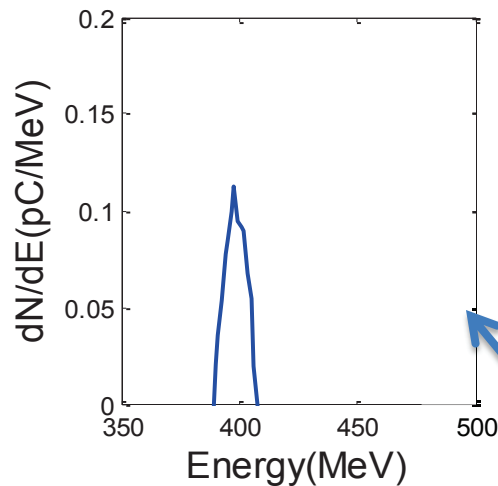


Z. Huang, Y. Ding, C. Schroeder, PRL 109, 204801 (2012)

High-quality high-energy electron beams from a cascaded LPA

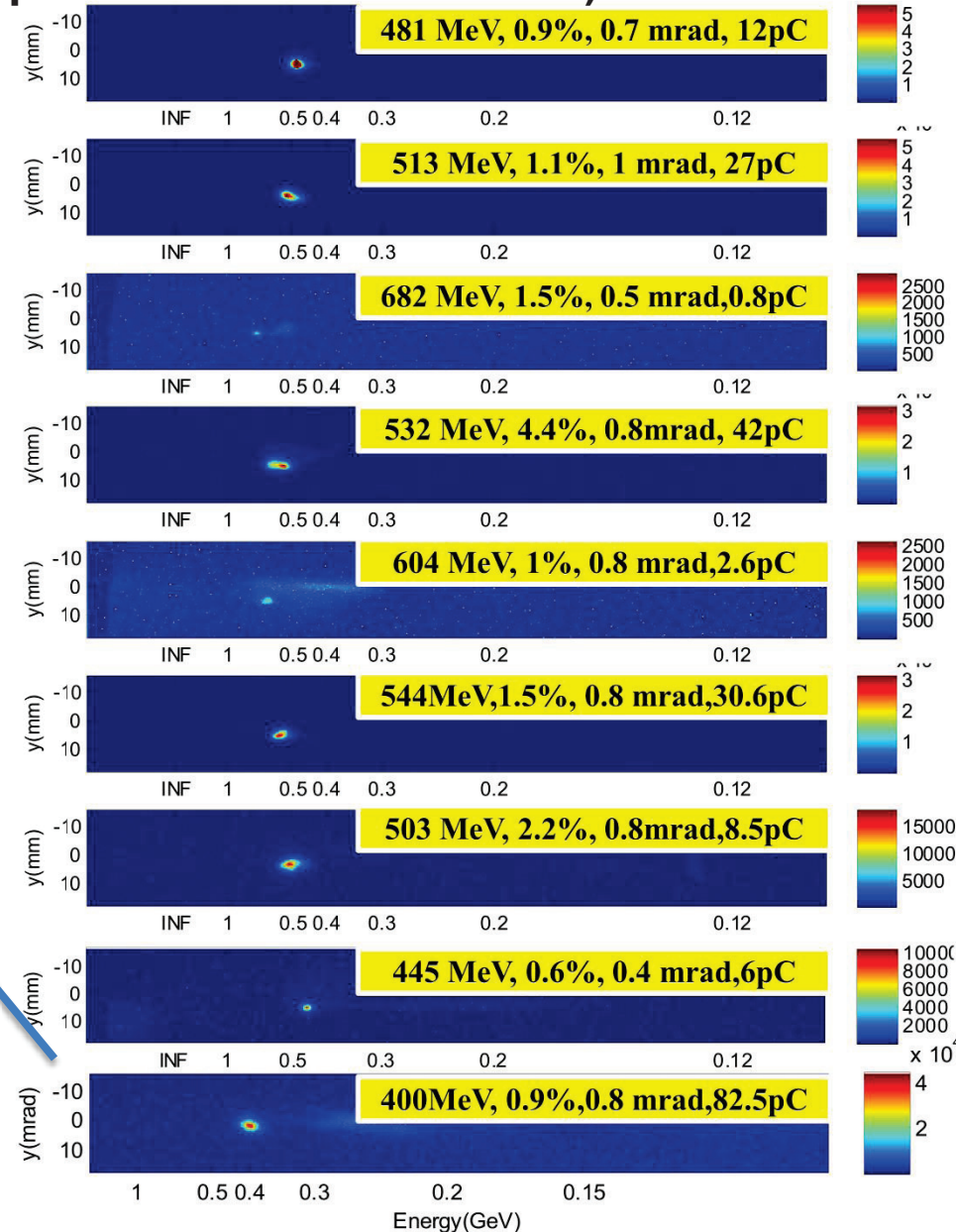
Courtesy J.S. Liu (Shanghai Institute of Optics and Fine Mechanics)

Peak energy: **0.4-0.6 GeV**
Energy spread: **$\sim 1\%$**
Beam charge : **up to 82.5 pC**
Divergence: **0.4-1.0 mrad**



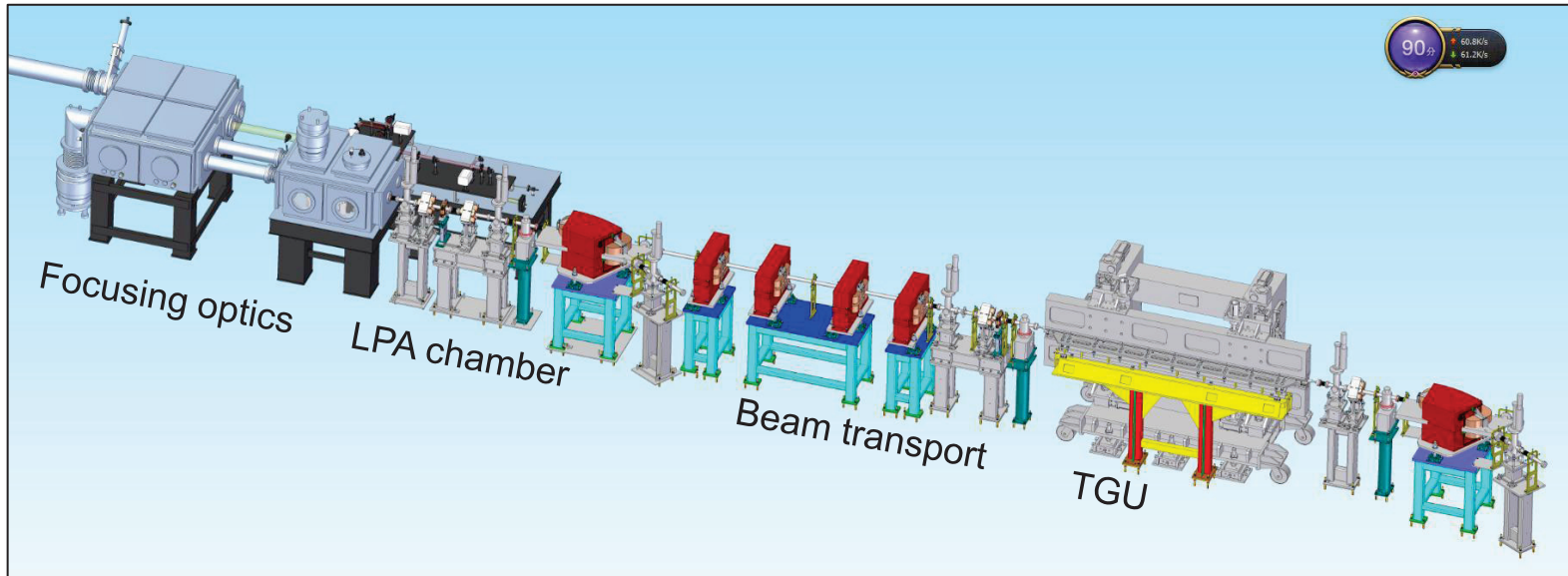
Peak energy	398 MeV
Energy spread (rms)	0.8%
Divergence (rms)	0.8 mrad
Beam charge	82.5 pC

J.S. Liu et al., *Phys. Rev. Lett.* 107, 035001 (2011).



LPA FEL with TGU (SIOM/SINAP)

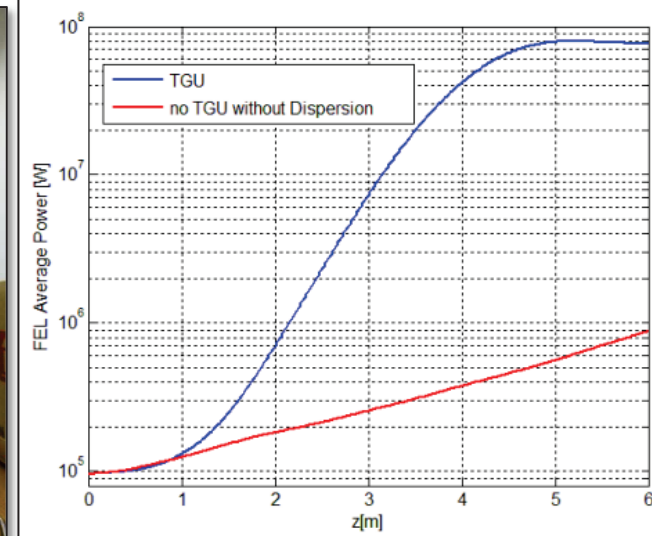
- Plan a demonstration experiment at 30 nm (400 MeV, 6 m TGU)



SIOM LPA setup (J.-S. Liu)



SINAP TGU assembly (D. Wang)



Summary

- XFEL is a breakthrough in light source development, enabling atomic-scale imaging at femtosecond time resolution.
- Vigorous accelerator and FEL R&D can drastically improve FEL coherence and brightness, which in turn demands better control of electron beams and advanced diagnostics and undulator development.
- Many challenges exist for realization of compact XFELs, but the efforts are worth pursuing and will be complementary to large facilities.

Acknowledgement

- I would like to thank Kwang-Je Kim, who introduced me to the beautiful and intricate FEL world, and thank Argonne LEUTL team that first demonstrated SASE saturation!
- I wish to thank the entire LCLS team and talented SLAC accelerator research staff and engineers. They made all these possible and are so much fun to work with!
- Thank you all for your attention!