

22/08/11

**High harmonics from gas,
a suitable source for seeding FEL
from vacuum-ultraviolet
to soft X-ray region (XUV)**

G. LAMBERT

guillaume.lambert@ensta.fr



M.E. Couprie, M. Labat, O. Chubar **Synchrotron Soleil, Gif-sur-Yvette, France**

D. Garzella, B. Carré **CEA, DSM/SPAM, Gif-sur-Yvette, France**

T. Hara, H. Kitamura, T. Shintake, Y. Tanaka, T. Tanikawa **SPring-8/RIKEN Harima Institute, Hyogo, Japan**

B. Vodungbo, J. Gautier, A. Sardinha, F. Tissandier, Ph. Zeitoun, S. Sebban, V. Malka **LOA ENSTA-Paristech, Palaiseau, France**

J. Luning **LCPMR, Paris**

C.P. Hauri **Paul Scherrer Institute, Villigen, Switzerland**

M. Fajardo **Centro de Física dos Plasmas, Lisboa, Portugal**

Thanks to **JSPS**

Laserlab Integrated Infrastructures Initiative RII-CT-2003-506350

TUIXS European project (Table top Ultra Intense XUV Sources)

FP6 NEST-Adventure n.012843

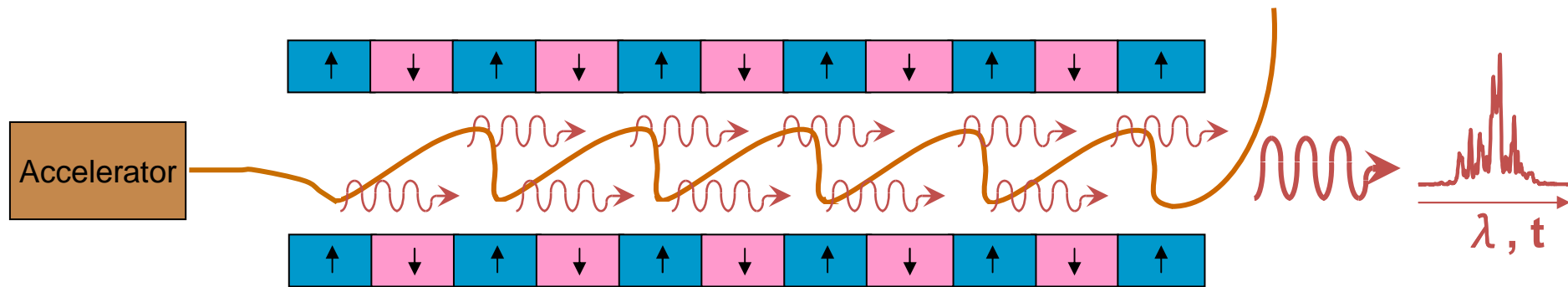
European Research Council Paris ERC project 226424



XUV Free Electron Lasers (FEL) : SASE

FLASH (2004, down to 4.1 nm), SCSS (2007, 50 nm), SPARC (2009, 160 nm)...

SASE: Self Amplified Spontaneous Emission



- High number of photon : 10^{12} photons
- Short pulse duration (sub ps)
peak power (GW)
- Relatively high repetition rate (tens Hz)
- High wavelength tuning
- Variable polarization
- Good wavefront (*Bachelard, PRL 106, 2011*)
- Good spatial coherence (*Ischebeck, NIMA 507, 2003*)

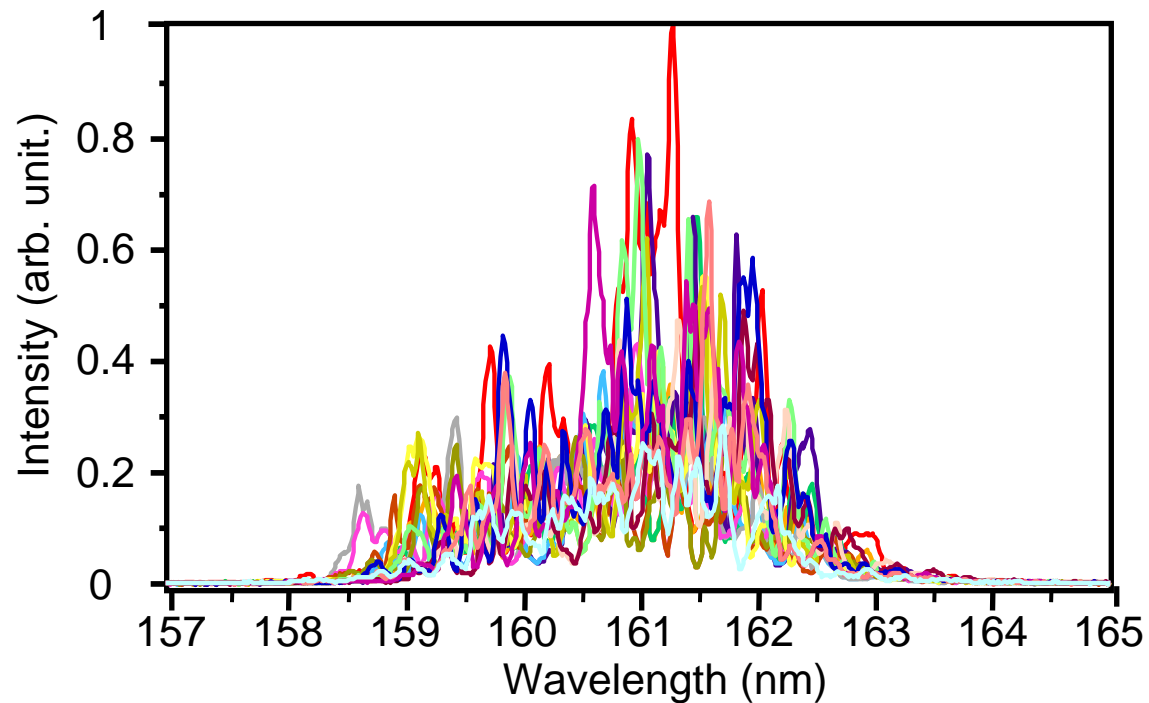
very performing tool for user experiment but...

SASE limitations

-Weak gain at short wavelength, single pass
long undulator (tens m)

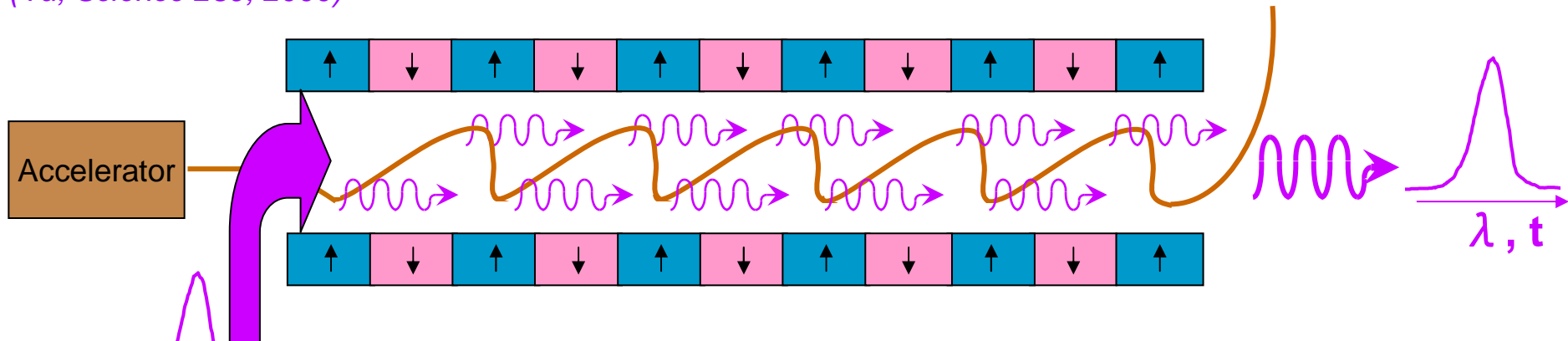
-Relatively important shot to shot variations: intensity, temporal/spectral profile
jitter for pump-probe experiments

limited temporal coherence (*Saldin, Opt. Commun. 202, 2002*)



How to reduce/supress SASE limitations: “seeding fully coherent light” in XUV ?

Previous demonstrations in IR with CO₂ and Ti: Sa lasers, then in UV with crystals
(Yu, Science 289, 2000)



External source
 λ, t

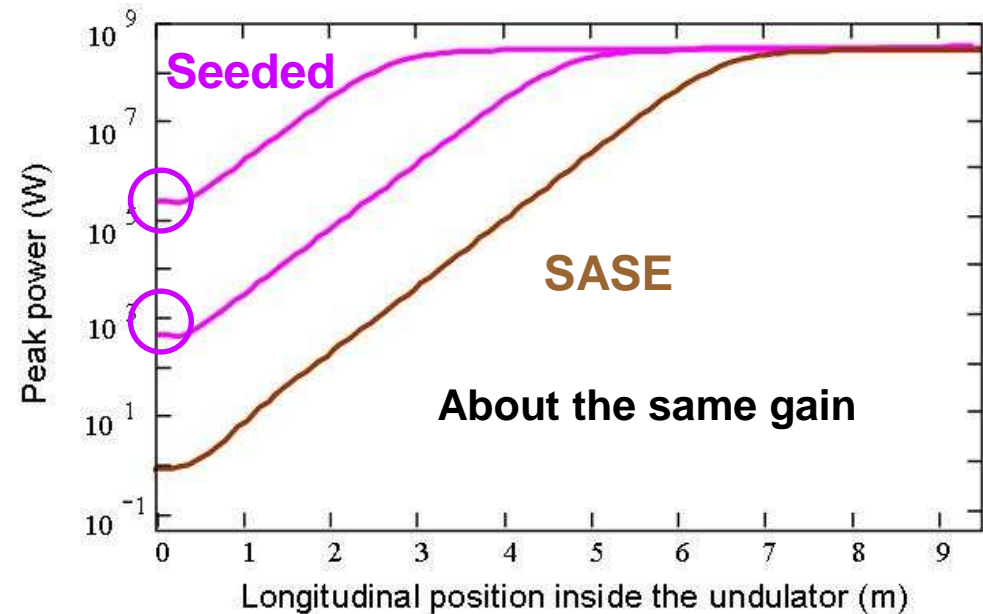
-Coherent

Improvement of the temporal coherence

-Intense

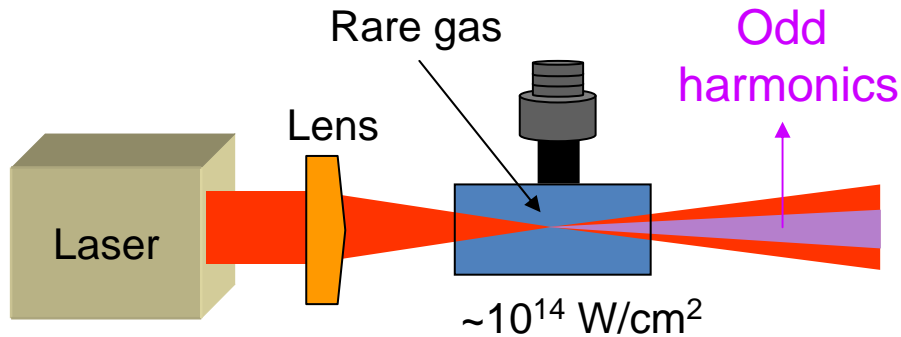
Decreasing of the saturation length

=> shorter undulator

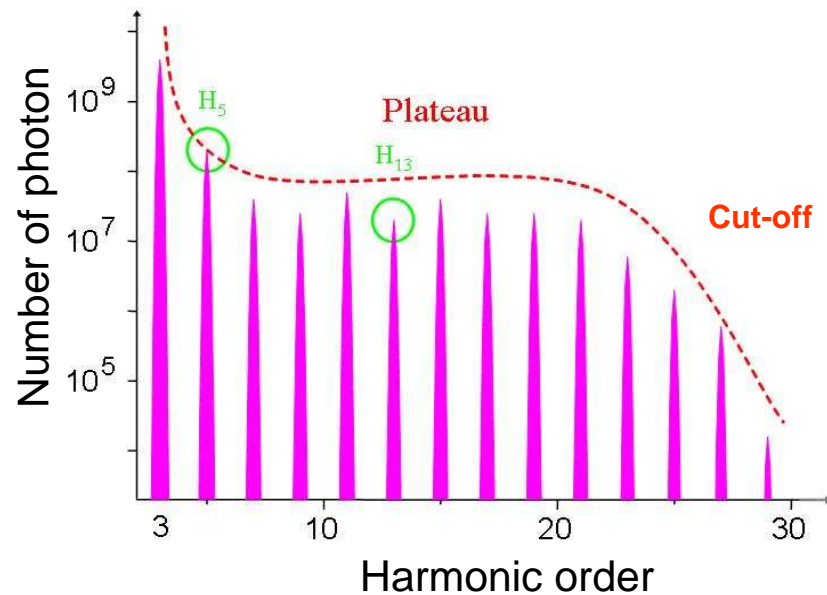


XUV radiation : Harmonics produced from gas (HHG)

Why proposing High Harmonics Generated in gas (HHG) ?



- Spatially and temporally coherent
- Relatively tunable
- Short pulse duration (10-100 fs)
- Conversion efficiency (10^{-5} to 10^{-7})
~ 100 nJ - nJ

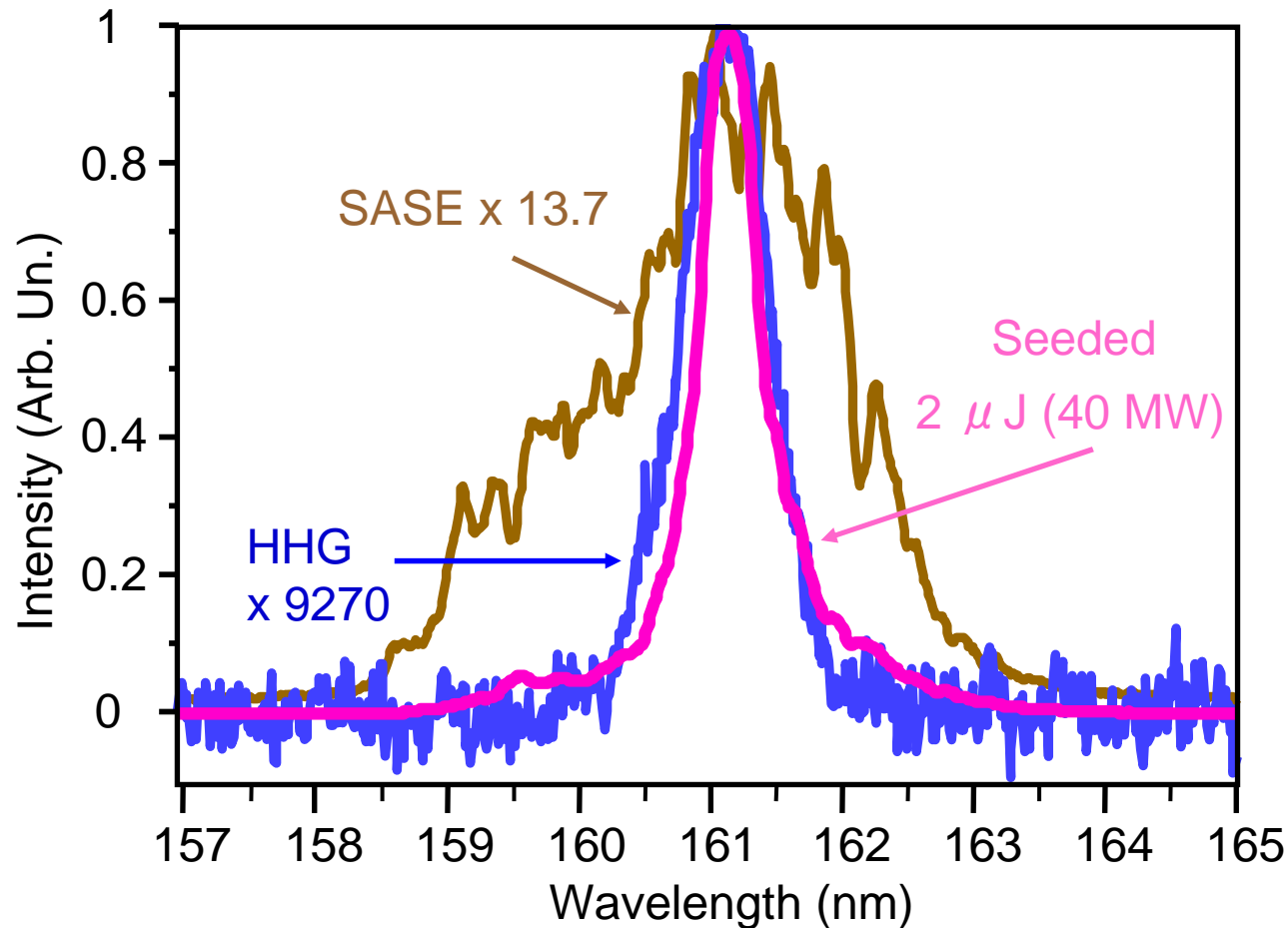


-Limited energy per pulse
Compared to SASE shot noise?

Direct seeding with HHG at 160 nm at SCSS: fundamental spectrum

SCSS Test Accelerator (Japan) at 150 MeV with 4.5 m long undulator sections

(Lambert, Nature Physics 4, 2008)



-High amplification

-Regular and quasi-perfect Gaussian shape

-Spectral width / ~5

-Pulse duration / ~20

Unseeded: 1 ps

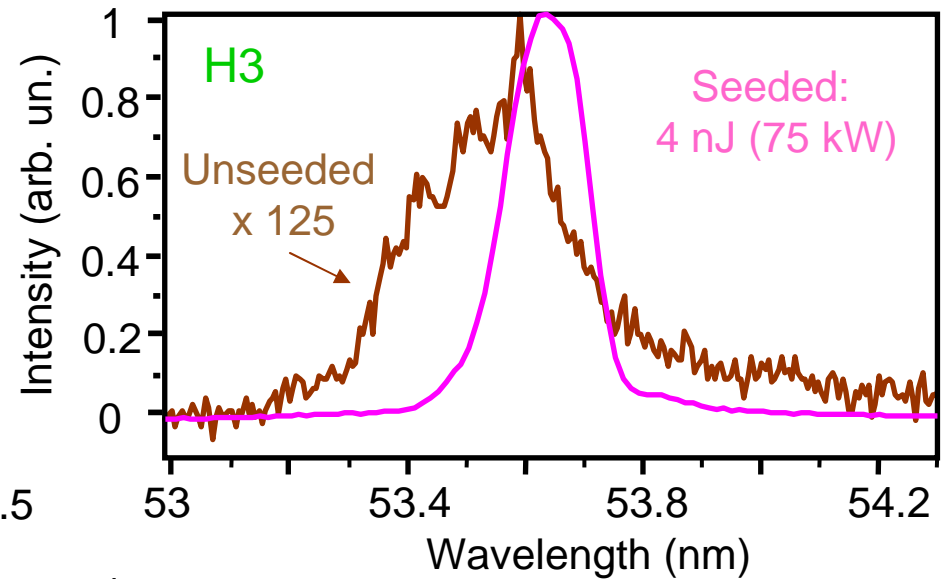
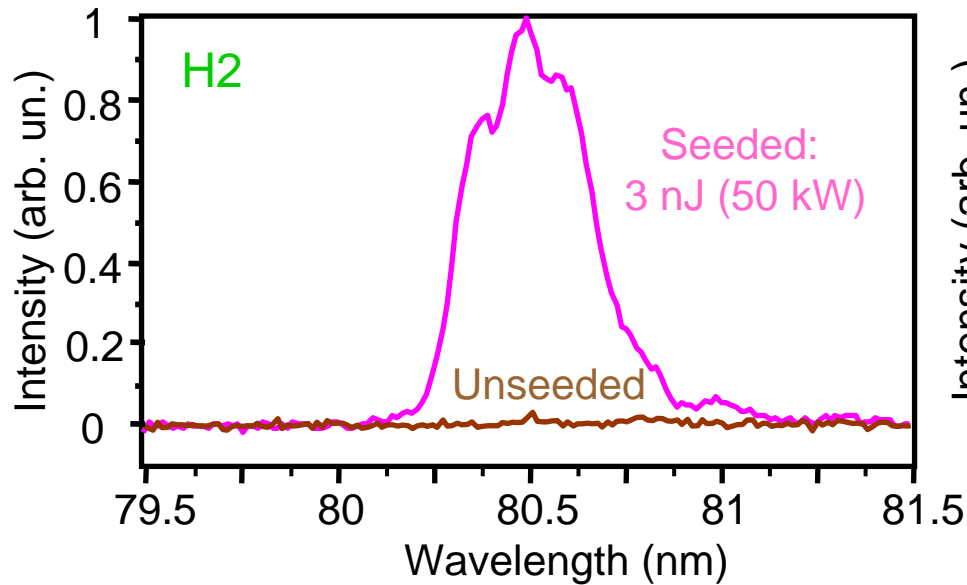
HHG: 50 fs

Seeded: 50 fs



Strong improvement of the temporal coherence/SASE

Non Linear Harmonic spectra



(Tanikawa, EPL 94, 2011)

-Odd and even harmonics from 2nd (80 nm) to 7th (23 nm)

-Clear amplification

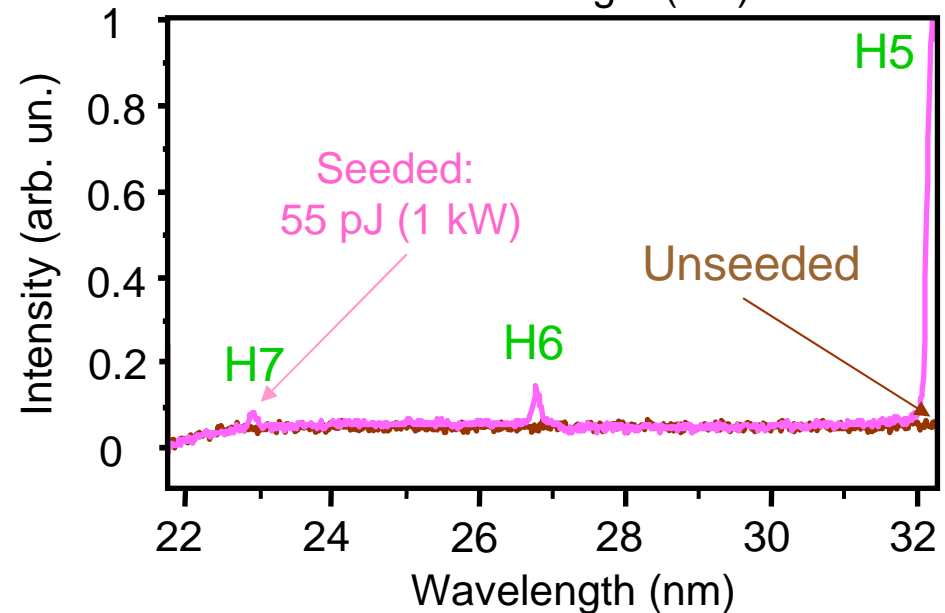
-Regular and quasi Gaussian spectral shape

-Spectral narrowing

-Shorter pulse duration

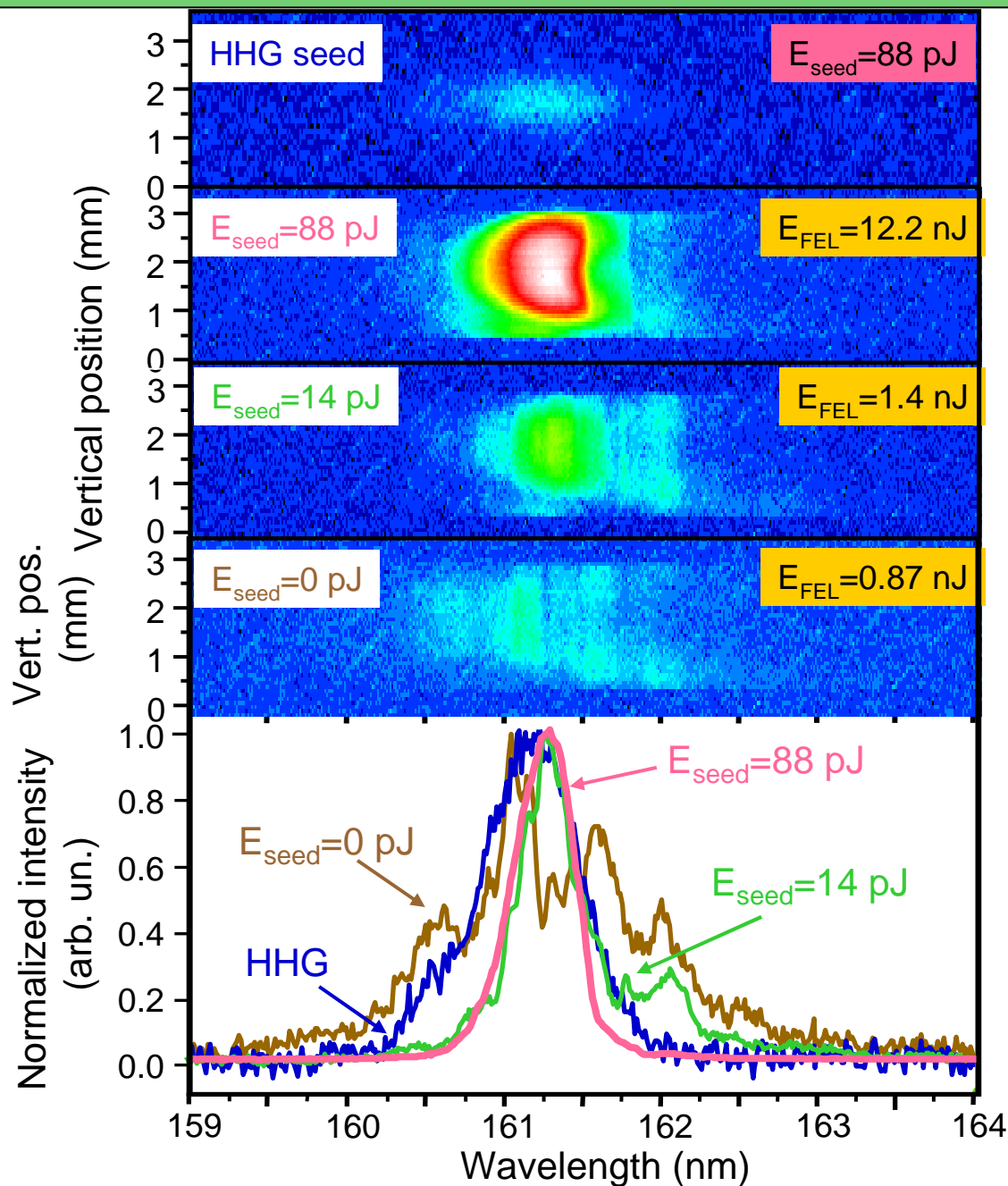


Improvement of the temporal coherence



Intense and coherent emission at short wavelength while $E=150$ MeV

Evaluation of the seed level requirement for observing coherent emission

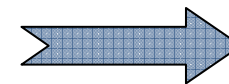


For $E_{seed} \geq 88$ pJ:
 -Stable emission
 -Regular and quasi Gaussian spectral profiles

Very weak level of injection

88 pJ (160 nm) \rightarrow 10 pJ (spatial overlapping) or $\sim 200 \times P_{e, shot\ noise}$

At 13 nm ($P_{e, shot\ noise} < 10$ W) \rightarrow
 $E_{seed} < 0.7$ nJ ($200 \times 10 \text{ W} \times 10 \times 35 \text{ fs}$)



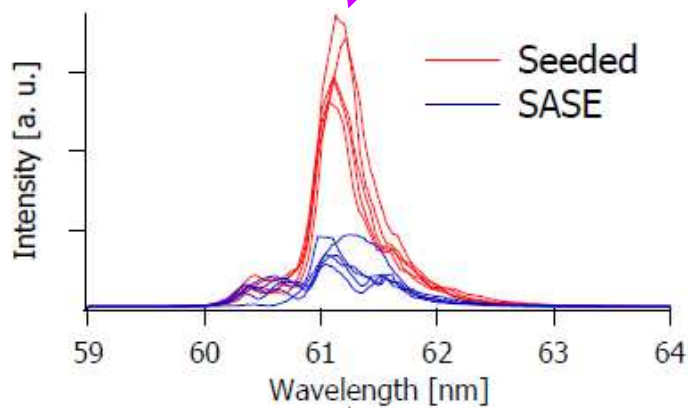
Seeding in XUV !

Below 13 nm is a challenge!

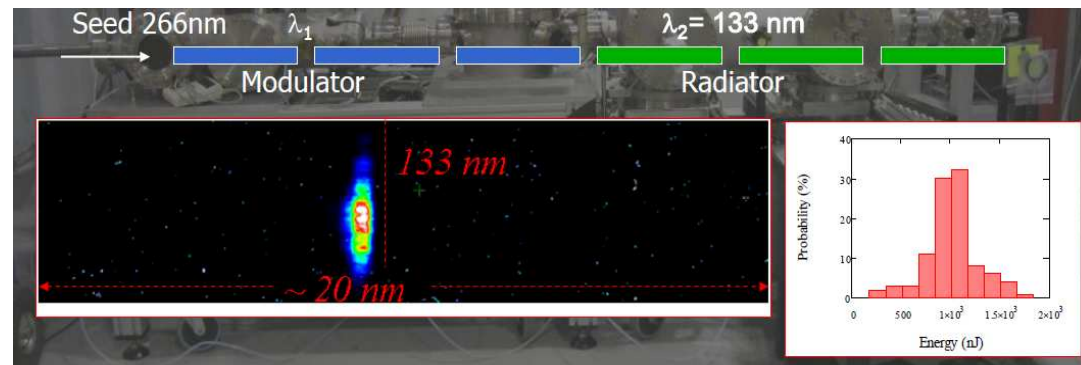
(Lambert, EPL 88, 2009)

Overview of the HHG seeded FELs

Name	Wavelength (nm)	state	Seeding process	Country
SCSS Test Accelerator	60 / 160	demonstrated	Direct HHG	Japan
SPARC	160 / 266	demonstrated	Direct HHG / HGHG	Italy
sFlash	38-13	Due 2010-2011	Direct HHG	Germany
Fermi	100-10	Due 2011-2012	HGHG, direct HHG?	Italy
SwissFEL	5	Due 2018	EEHG, direct HHG	Switzerland
SPARX	15	Due ?	HGHG, direct HHG?	Italy



(Togashi, Optics Express 19, 2011)



(Giannessi, Proceedings FEL10 JACOW (2011) Labat WE0B3)

What has to be improved on HHG for seeding FEL in future?

Harmonics properties relevant for seeding

Already obtained

- fs pulse duration
- Full coherence
- High repetition rate:
kHz HH currently
First MHz HH in xenon (J. Boulet et al.
optics letters, 34, 1489 (2009))

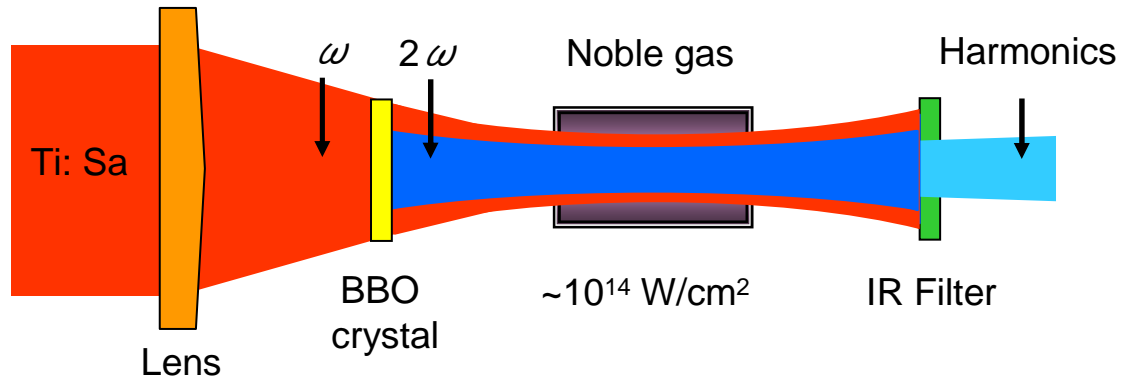
To be improved

- Intensity at short wavelengths
- Tuneability (only odd HH):
=> need to considerably chirp the driving laser and/or change the gap of the undulator
- Two colour mixing
- Wavefront: diffraction limited beam (aberration-free)
=> need to drastically clip the IR or HH beam or use adaptive optics for IR or HH
- Variability of the polarization
- Stability of the shot to shot intensity

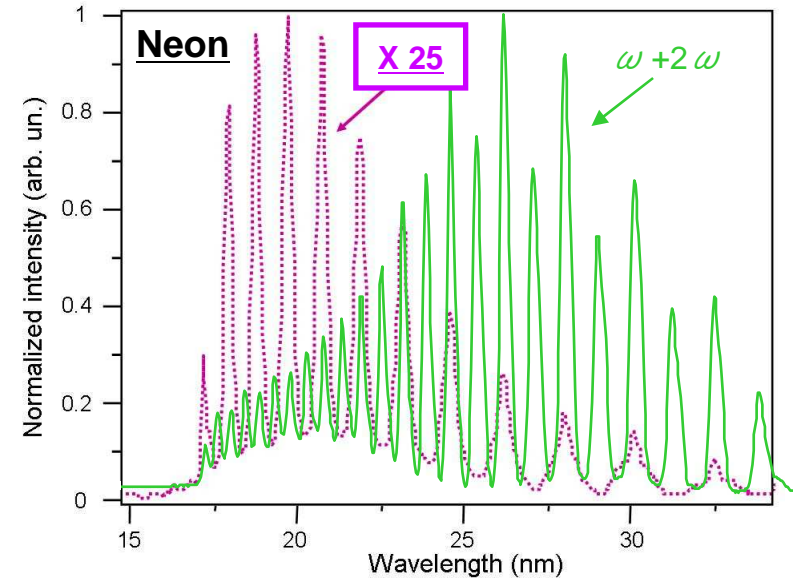
Keep the simplicity of the classical HHG setup

High order harmonics generated with a two-colour field

Technical principle:



(Lambert, NJP 11, 2009)



-double harmonic content
even types:
 $2 \times (2n+1)$ from 2ω
 $2 \times (2n)$ from the mixing

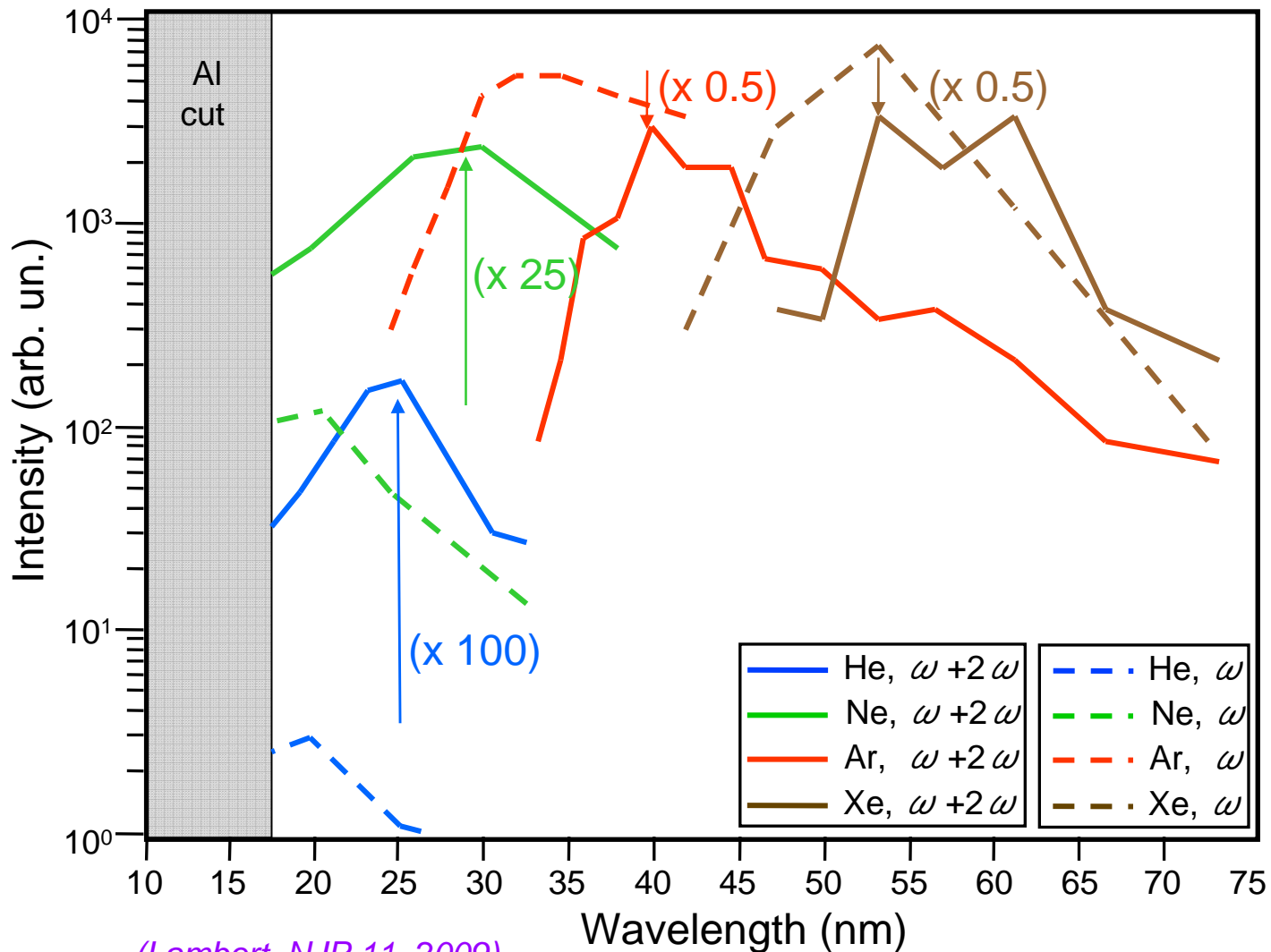
-redshift: $E_{\text{Cut-off}} = I_P + 3.2U_P$

$$U_P \propto I_{\text{Laser}} \lambda_{\text{Laser}}^2$$

-increase of the number of photons

Harmonic spectra obtained with either ω or $\omega+2\omega$ technique

100 μm thick BBO crystal, and with the optimization parameters corresponded to ω :
 $E_\omega < 6 \text{ mJ}$, $L_C = 7\text{-}9 \text{ mm}$ and $P_G = 30\text{-}35 \text{ mbar}$



-Flat spectra (same intensity level for odd and even harmonics)

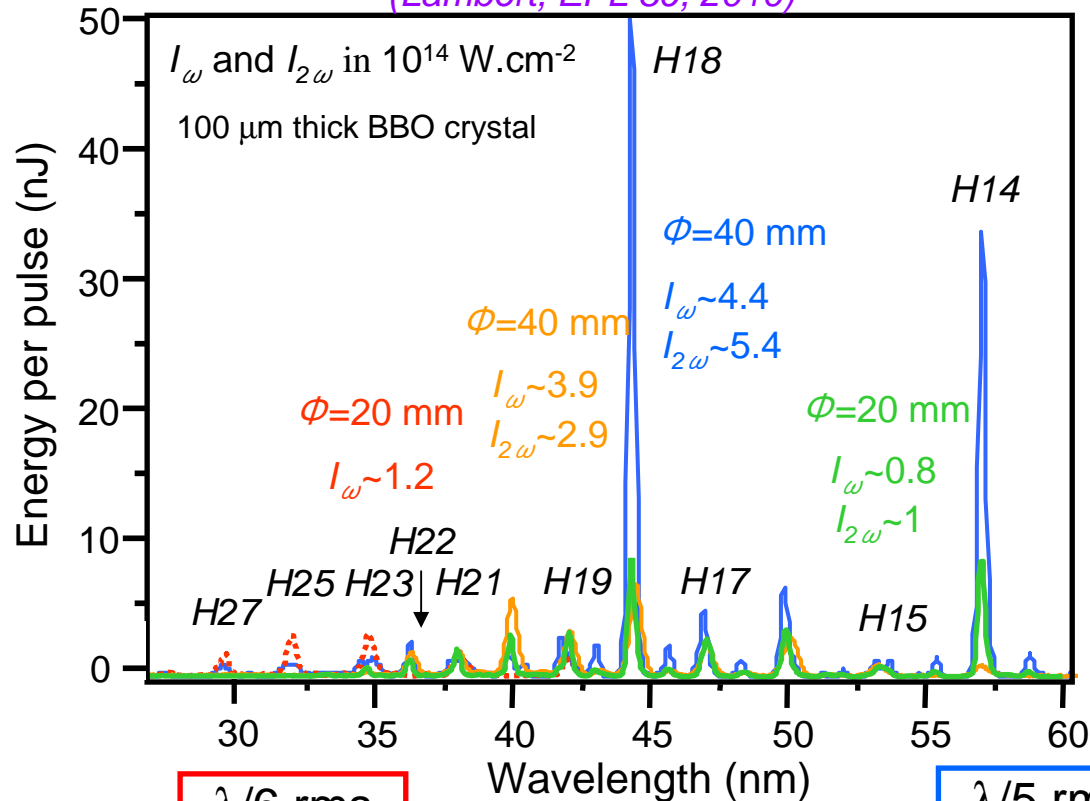
-Increase limited at high wavelengths due to an already relatively high efficiency for Xe and Ar

=> $\omega+2\omega$ technique compensates the weak efficiency at short wavelengths

(Lambert, NJP 11, 2009)

Optimization of both flux and wavefront (Ar gas)

(Lambert, EPL 89, 2010)



-iris clipping technique:

change the focusing geometry/energy

clean the major part of the distortions in the outer part of the beam: λ to $\lambda/6$ rms

- ω ($L_C=8$ mm and $P_G=30$ mbar) to

$\omega+2\omega$ ($L_C=4$ mm and $P_G=16$ mbar)

-very high increase on $2x(2n+1)$ type of even harmonics (50 nJ) due to strong blue/IR and distortions limited: $\lambda/5$ rms

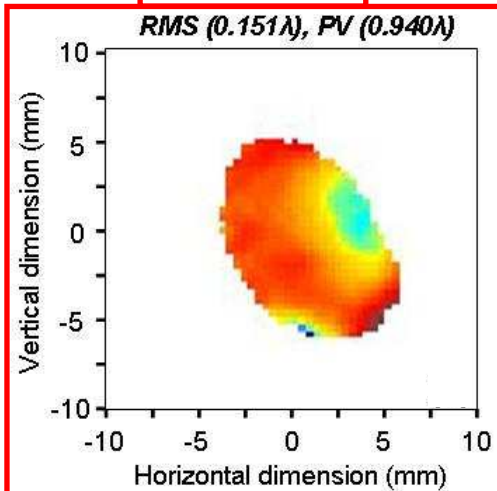
-iris clipping: limited decrease of intensity

But distortions about $\lambda/17$ rms:

First aberration-free high harmonic beam

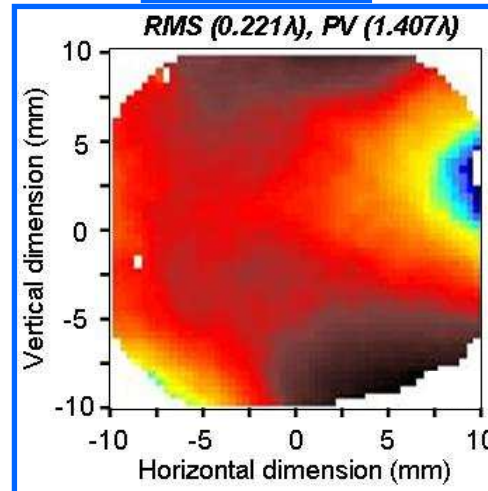
$\lambda/6$ rms

RMS (0.151 λ), PV (0.940 λ)



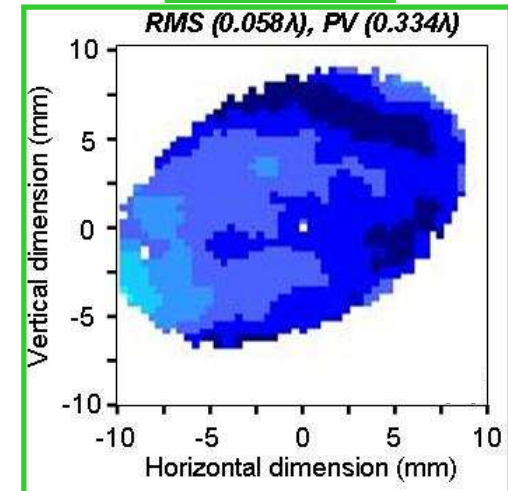
$\lambda/5$ rms

RMS (0.221 λ), PV (1.407 λ)



$\lambda/17$ rms

RMS (0.058 λ), PV (0.334 λ)



Then next?

Already obtained

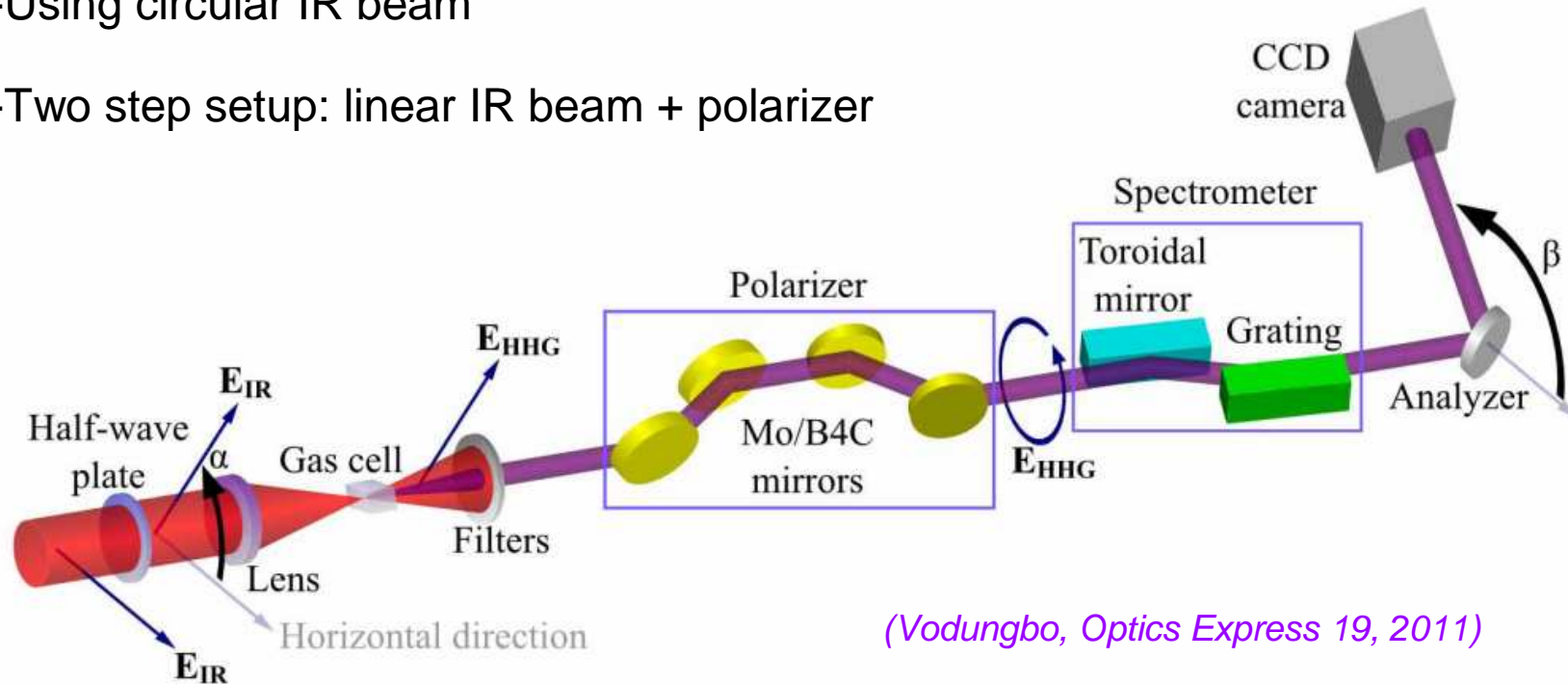
- fs pulse duration
 - Full coherence
 - High repetition rate: kHz to MHz soon
 - Intensity at short wavelengths
 - Tuneability: both odd and even harmonics
- Use parametric amplifier (1.2-1.5 μm)
- Wavefront: aberration-free beam
 - Simple system

To be improved

- Variability of the polarization
- Stability of the shot to shot intensity?

Circularly polarized HH

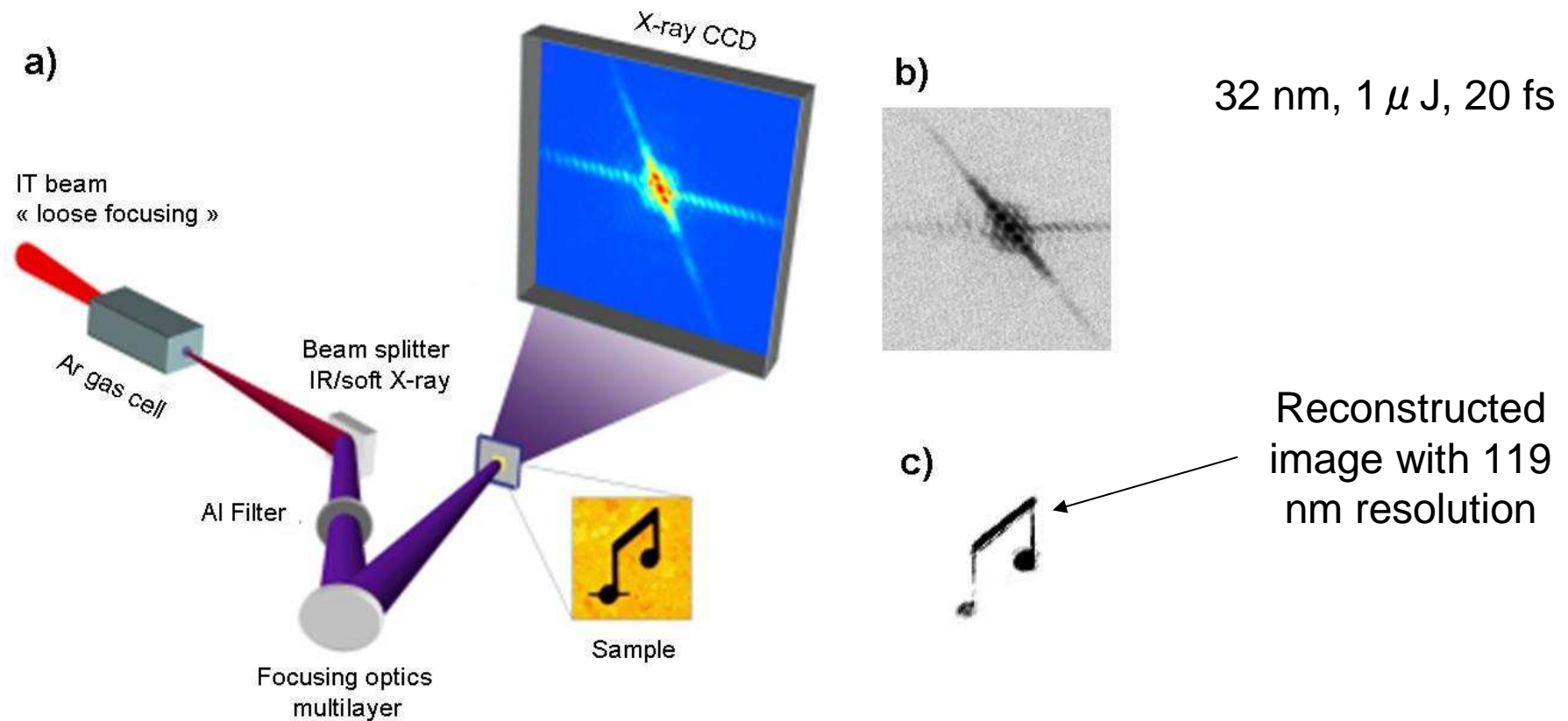
- Using circular IR beam
- Two step setup: linear IR beam + polarizer



Harmonic order	31 st	33 rd	35 th	37 th	39 th	41 st	43 rd	45 th
λ (nm)	26.3	24.7	23.3	22	20.9	19.9	19	18.1
P_{Cmax} (%)	100	100	97	91	85	76	66	61
T_C (%)	2.6	2.7	3.7	3.6	4.4	4.3	4.0	3.7

In XUV performances are close: weak flux harmonics
weak temporal coherence

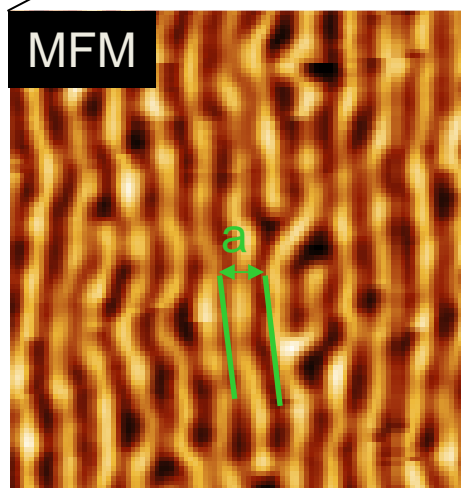
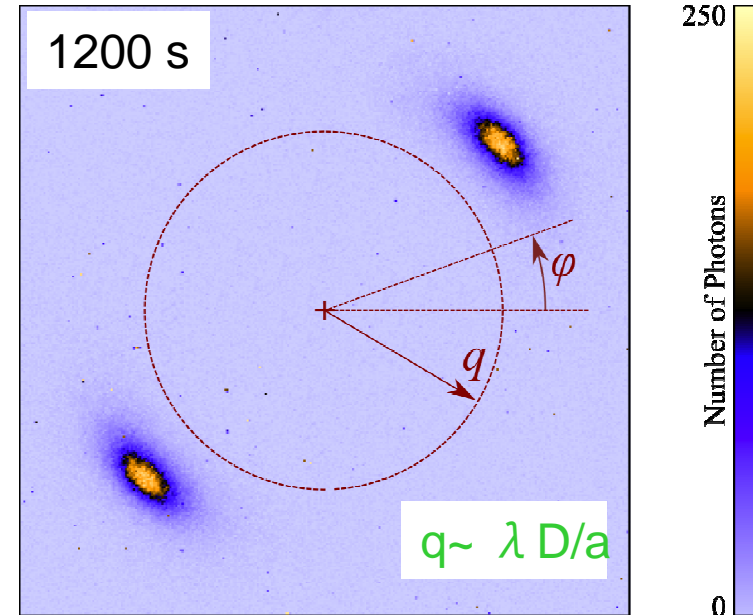
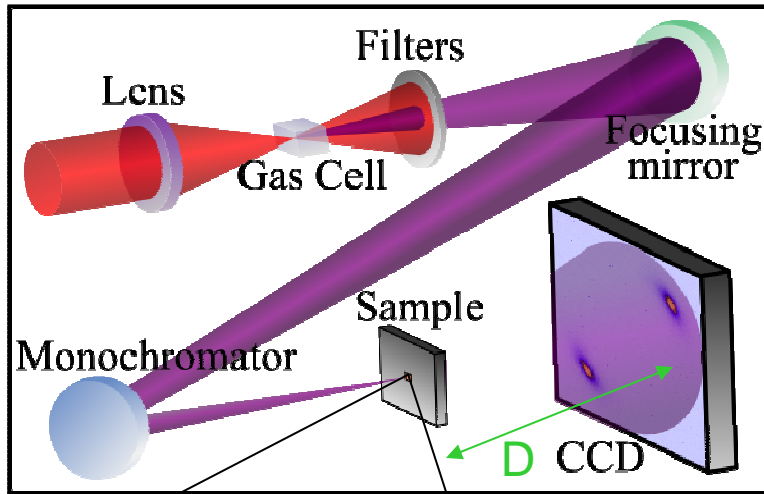
Single shot coherent diffraction imaging of nano-object (*Ravasio et al. PRL 103 (2009)*)



Pump-probe experiments

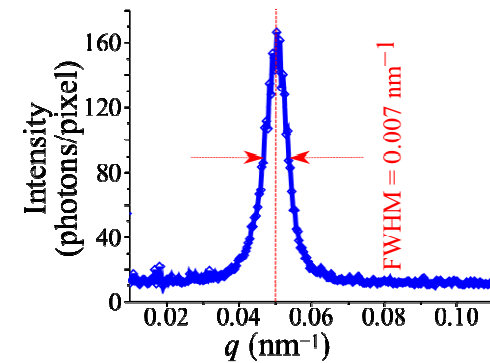
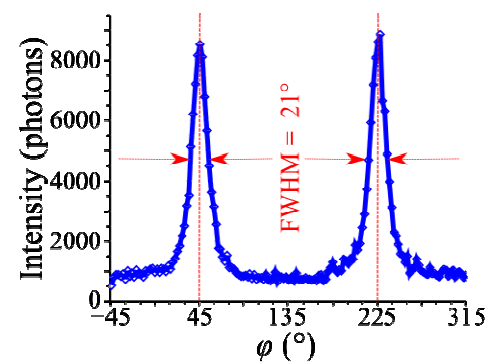
Natural synchronization between Laser and HH

Coherent diffraction imaging of magnetic domains (Vodungbo, EPL 94, 2011):



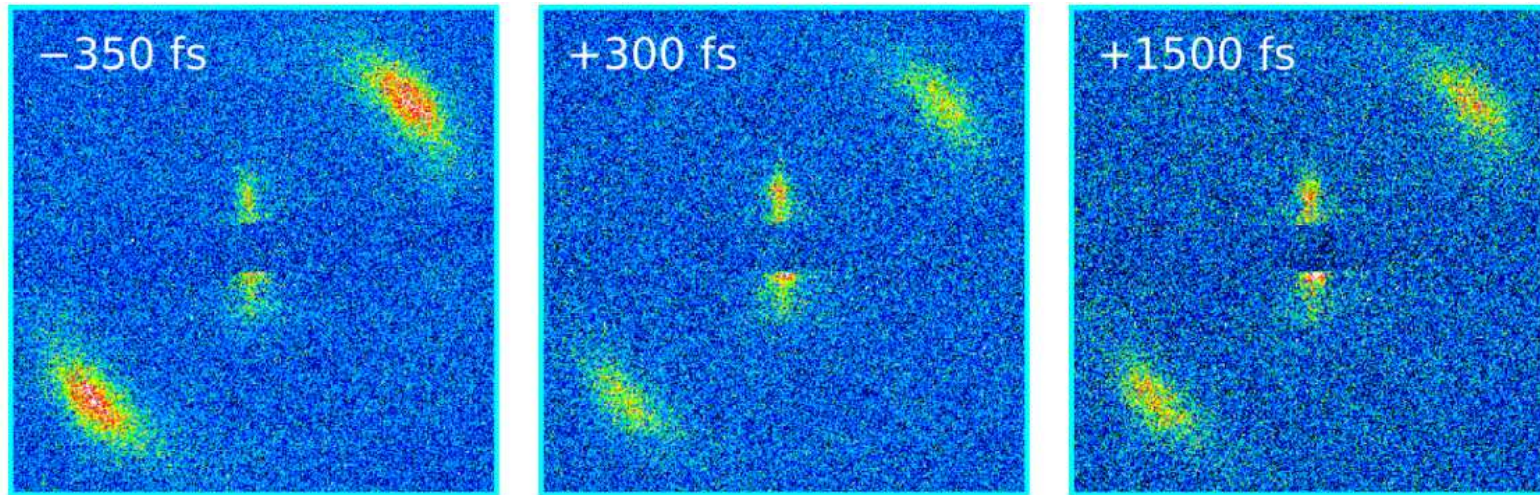
Co $M_{2,3}$ edge
at 60 eV

Aligned
magnetic
domains

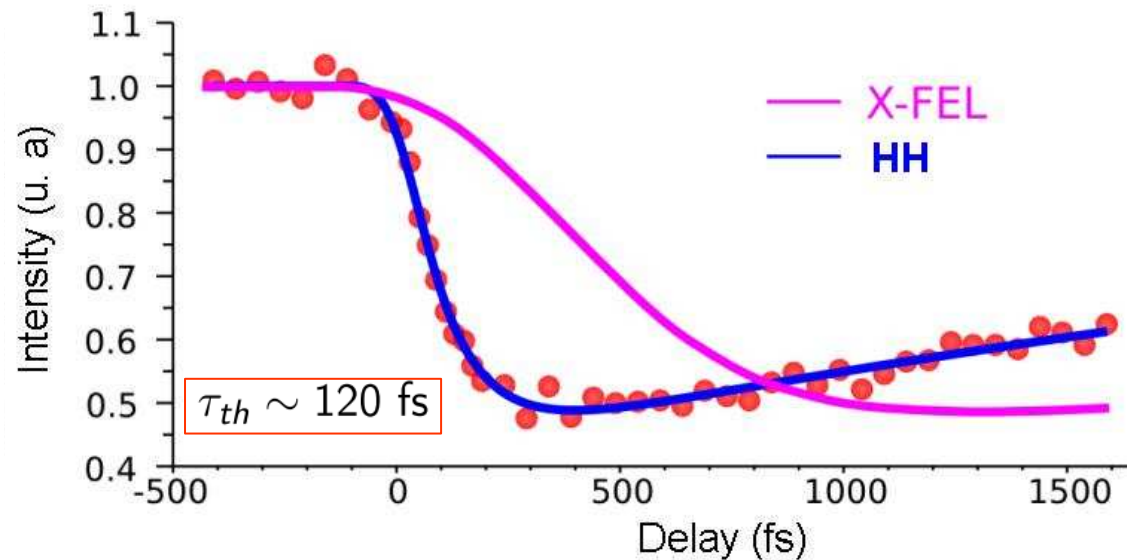


- Magnetic domain orientation: $45^\circ \pm 10^\circ$
- Width distribution of magnetic domains: $65 \text{ nm} \pm 5 \text{ nm}$

Ultra-fast demagnetization



Integrated density gives magnetization : $I \sim M^2$



- much shorter time scale
- better temporal resolution for HH?

And/or due to XFEL at 800 eV (L edge)

fitted with $G(t) * \left\{ C_0 + C_1 H(t) \left[1 - \exp\left(\frac{-t}{\tau_{th}}\right) \right] \exp\left(\frac{-t}{\tau_{s-p}}\right) \right\}$ (Boeglin, Nature 465, 2010)

Thank you
for your attention