Brief Review of the Approaches to Elucidate the Mechanism of the Radiation-induced Demagnetization

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Radiation-induced demagnetization

The magnetic field intensity decreases when the magnets are put in a strong radiation environment.



Several approaches to clarify the mechanism of the radiation-induced demagnetization that have been done so far.

- (1) The approach to examine the relations between the field degradation and the environmental factors such as magnet shape, temperature and so on.
- (2) The approach to examine the changes of the properties and the microstructures of the magnet after irradiation.
- (3) The approach to compare and examine the experimental results and the computer simulations.

(1) The approach to examine the relations between the field degradation and the environmental factors.

Radiation-induced demagnetization shows the following dependencies:



(1) The approach to examine the relations between the field degradation and the environmental factors.

Radiation-induced demagnetization shows the following dependencies:

- 1 material chemical component microstructure manufacture
- 2 coercivity
- 3 temperature
- 4 permeance coefficient (P_c) magnet shape outer magnetic field inflection point (B-H curve)

These factors are similar to the factors that influence the demagnetization originated from the reversal magnetization. Especially, 2, 3, and 4 are related to the coercivity decrease caused by the internal magnetic field that is described by permeance coefficient.

(2) The approach to examine the changes of the properties and the microstructures of the magnet after irradiation.

Remagnetization

The damaged magnets by irradiation were remagnetized and compared the magnetic properties before irradiation.

	Particle	Remanence	Coercivity
Cost 1988 [84]	Fast neutron	No change	20 percent increased
Kähkönen 1990 [3]	Proton 20 MeV	No change	
Okuda 1994 [90]	Electron 17 MeV	No change	
lto 2002 [66]	Proton 200 MeV	No change	
Chen 2005 [264]	Neutron ~ 10 MeV	Decreased (flux dependence)	
Klaffky 2006 [266]	Thermal neutron	No change	No change
Qiu 2008 [270]	Electron 2.5 GeV	Decreased	No change

Microstructures change

	Particle	Methods	Results
Talvite	Proton	Positron annihilation measurement	no detectable vacancy concentration
1991 [8]	20 MeV 20 Mrad		
Gao	γ (⁶⁰ Co)	X-ray diffraction	no phase changes
2008 [262]	200 Mrad	Positron annihilation spectroscopy	the positron lifetime variations are all little
Yang	Proton	Soft X-ray absorption spectrometry (XAS)	the destruction of atomic symmetry and changes of bonding environment induce the change of atomic magnetic moment
2009 [290]	9 MeV 7 × 10 ¹⁴ p/cm ²	Mössbauer spectrometry	
2011 [245]		X-ray diffraction (XRD) XAFS	Proton irradiation has no effect on the long-range structure, but significantly affects the atomic local structure
			The alignment degree of the magnet decreases and the internal stress of the lattice increases.
			The coordination number of Fe-Nd in the first neighboring coordination shell of the Fe atoms decreases and the disorder degree increases.

Demagnetization occurs before clear structural changes.



Demagnetization occurs by magnetization reversal in early stages.

(3) The approach to compare and examine the experimental results and the computer simulations.

	Code	Simulation	Particle
Qiu 2008 #270	FLUKA	Absorbed dose 1 MeV equivalent neutron fluence*	Electron 2.5 GeV
Asano 2009	FLUKA	Star density**	Electron 2, 4, 6, 8 GeV
Leitner 2010 #282	FLUKA	Total dose Non-electromagnetic dose Neutron fluence	Electron 13.7 GeV

They suggested the important role of neutron.

FLUKA is a Monte Carlo simulation code for the interaction and transport of particles and nuclei in matter.

* 1MeV equivalent neutron fluence, is widely used to characterize the displacement damage of the electronic devices in which the main material is Si when they are irradiated by neutrons.

** A star was defined by a hadronic inelastic interaction (spallation reaction) at energy higher than the threshold and excludes the spallation due to annihilating particles.

Basic model of magnetization reversal

Process of magnetization reversal by heat and field



	Model
Blackmore	It appears to be similar to a thermal heating of the sample at elevated temperature.
Cost 1988 #84	Collision cascade has a higher probability of nucleating a reverse domain when the temperature is closer to the Curie temperature.
Brown 1988 #52	
Kähkönen ^{1990 #3}	
Zeller 1990 #139	
Makita 2004 #208	
Gao 2006 #246	

Brown's model

Brown focused on the magnetic interaction.

The decay of magnetic remanence during neutron irradiation is presumably caused by a combination of nucleation of reverse magnetic domains and depinning of domain walls, leading to domain wall motion and demagnetization.

Magnetic interaction of the neutron's magnetic moment with the magnetization of the material



Magnetic excitation in the magnetization of a grain

Nucleating a reverse domain

Excitation at a magnetic domain wall pinning site

Depinning of the domain wall, which is then free to move

Inelastic collisions with the atoms

Local disruption crystal structure magnetic anisotropy



Creation of additional pinning sites

Increasing the coerecivity



Part of the energy of the incoming particle is transferred to the primary knock-on atom.

The energy is then diffused into the lattice raising the temperature of a spherical region.

Thermal spike

If the temperature rises above the Curie temperature and if this sphere is large enough the demagnetizing field can turn the spins and nucleation of a new domain occurs.

The domain immediately grows to the size of the grain.

* J. Phys. Condens. Mater 4 (1992) 1007-1014

** Graphic image is added by Bizen.

The sensitivity of NdFeB permanent magnet materials to radiation induced demagnetization is shown to be the result of loss of coercivity. This allows the magnet to demagnetize at locations which are subjected to the largest external and internal demagnetization fields.

Makita's model

Since all of the magnets have the same Curie temperature, the difference in the demagnetization rate can be attributed to the probability of nucleation of a reverse domain in a locally heated region located around a knock-on atom, which is analogous to the thermal demagnetization of the magnets that usually starts below their Curie points.

Makita made clearer explanation of the origin of the nucleation by focusing on the "decrease of the magnetic anisotropy" in place of the "loss of the coercivity". This is because the magnetic anisotropy is the origin of the coercivity.



(1) Temperature rise (below Curie temperature)

(2) Magnetic anisotropy decreases

(3) Nucreation and reversal magnetization

* J. Magn. Soc. Jpn.28, (2004), 326-329

Gao's model

As an external energy source, γ -ray irradiation will *decrease the ordering of magnetic moments by disturbing the electron spin* of Fe and Co atoms in the ferromagnetic phase.

γ-ray (⁶⁰Co)

Fe-Cr-Co and $Nd_2Fe_{14}B$ magnet

	Fe-Cr-Co		$Nd_2Fe_{14}B$	The Curie temperature should be taken into
Curie temperature	945 K	>	700 K	account in high-energy particle irradiation conditions, but for γ -ray irradiation, <i>coercivity</i>
Demagnetization	16.5 %	>	2.5 %	mechanism is the dominant factor

J. Magn. Magn. Mater. 302 (2006)





Two points of view to consider the mechanism

Wide effect and local effect

- 1. Wide unstable region (magnetic moment instability)
- 2. Local hot spot (quasi-thermal spike)

1. Wide unstable region (magnetic moment instability) model

Similar process of the magnetization reversal caused by heat and field would occur in the *low energy particle irradiation*.





This technique is a stabilization against the after-effect that is associated with the fluctuations of magnetic moment in thermal or magnetic energy.



2. Local hot spot (quasi-thermal-spike) model

Local hot spot is generated by high energy photoneutron

Thermal spike generates very high temperature in very small area.

The energy of the knock-on atom transfers to the magnet atom by the process similar to the thermal spike (quasi-thermal-spike*).



* Temperature generated by the thermal spike is higher than the melting point, however, as Makita stated, demagnetization can occur below Curie temperature much lower than the thermal spike temperature.

Inverse magnetic domain is formed in local hot spot

(1) Core of energy release

Melting and structural changes make the magnetic properties different.

(2) Nucleation of magnetization reversal

Heat and intense properties change generates the nucleus of inverse domain.

(3) Low anisotropy region caused by temperature rise

Inverse domain wall easily expands in the low anisotropy region.

(4) Higher anisotropy region

Expansion of the domain wall is limited depending on the magnitude of the anisotropy.

Inverse domain expands against anisotropy field

This is the reason of coercivity dependence though the nucleation is inevitable.



High-energy electron irradiation will cause the local hot spot.

The experimental results of the demagnetization caused by the high-energy electron irradiation were compared to the computer simulations.

Neutron collision density distribution [Energy and target dependence]





Energy and target dependence (Experiment and calculation)



The calculation results are normalized to the 4 GeV experimental results.



Thermally stabilized magnet Calculated by FLUKA Y. Asano et. al. J. Synchrotron Rad. (2009)

Energy and target dependence (Experiment and calculation)



Thermally stabilized magnet Calculated by FLUKA Y. Asano et. al. J. Synchrotron Rad. (2009)

Application case

Fig.1 Geometry of simulations. (unit: mm)

density (cm⁻³

Stainless steel, aluminium, and YAG (Y3Al5O12: density 4.6 g cm³) are usually use for the material of OTR. Figure 4 shows the results of the star density distributions for each OTR materials with 0.1 mm thick. The OTR made of aluminium produce the lowest star density in the same thickness of 0.1 mm. In the case of aluminium OTR with 1 mm thick, the peak position of the star density moves to almost 7.2 m distance from the OTR and the density increases from about 2.4 x 10-7 to about 7.8x107 star cm3 electron so that the thickness of the OTR is also important to reduce the star density.

Table 1 Comparison of maximum star density and peak position between the 2 mm and 4 mm ID gap width. (*Density unit: star · cm⁻³ · electron⁻¹) ID gap width

> ID gap 2mm • : Al :YAG

: SUS

~ F F C C C C

Distance from OTR (m)



The estimation of the demagnetization has been performed quantitatively as functions of the electron energy, the gap width of the ID, and the dependence on material of the OTR.

> Y. Asano, Proceedings of 2nd International Particle Accelerator Conference 2011, San Sebastian, Spain, THPC165, (2011).

Effective protection

The efficient methods to increase the radiation resistance of the magnet is following:

- 1. Designing the magnetic circuit of higher permeance coefficient
- 2. Selecting the magnets with high coercivity
- 3. Selecting the magnets with high temperature-stability
- 4. Applying the stabilization technique to the magnets

5. Using the magnets at very low temperature

Using the magnets at very low temperature



To increase the coercivity in the grain, the best way is to decrease the magnet temperature.

Resistance against the radiation increases at low temperature.

T. Bizen et. al. Radiat. Meas. (2007)

Summary

Typical past experiments and models of radiation-induced demagnetization are reviewed and summarized.

If the nucleation is the origin of the radiation-induced demagnetization, two types of demagnetization should be considered.

The demagnetization cannot be estimated sufficiently by a simulation of the dose alone, because the mechanism of the demagnetization depends on the particle energy.

The simulation of the star density was in good agreement with the demagnetization. The star density can be used for estimation of the radiation-induced demagnetization.

Thank you for your attention.