



Electron Cloud Mitigation Investigations at CEsrTA

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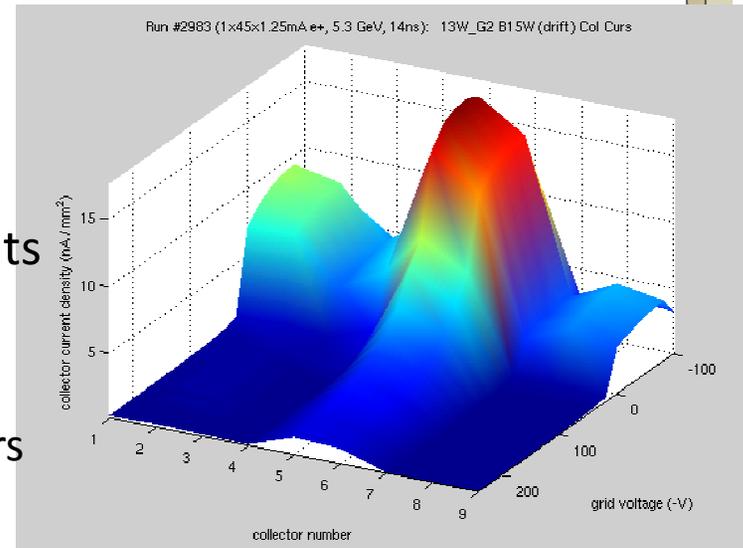
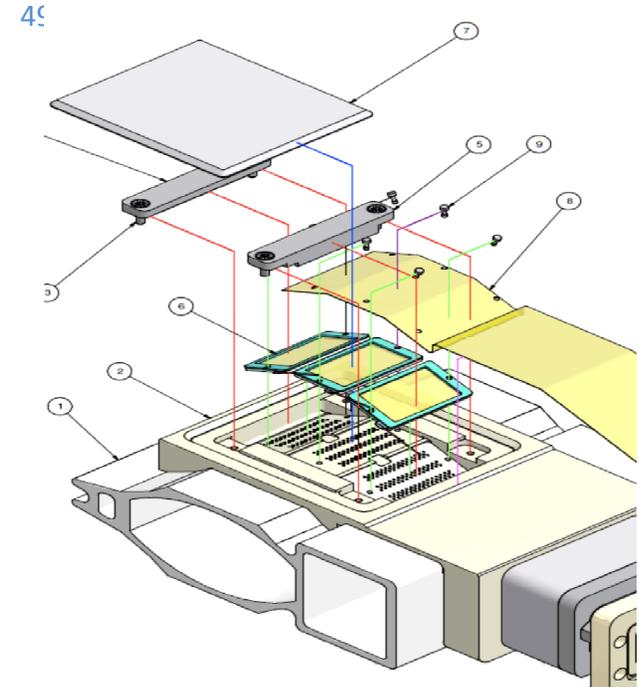


- The density and distribution of the electron cloud can depend strongly on several parameters that can vary substantially throughout an accelerator. These include...
 - Local photon flux
 - Vacuum chamber shape and material
 - Primary and secondary emission properties of the material
 - Magnetic field type and strength
- Therefore it is useful to have a detector that can sample the electron cloud locally. At CEsrTA we have used...
 - Retarding field analyzers (focus of this talk)
 - TE-Wave transmission (see talk by S. DeSantis, poster by J. Sikora)
 - Shielded pickups (poster by J. Crittenden)
- Several EC mitigation techniques have been proposed, many of which have been studied at CESR...
 - Beam pipe coatings (TiN, amorphous Carbon, NEG)
 - Grooved beam pipes (in dipole regions)
 - Solenoids (in drift regions)
 - Clearing electrodes



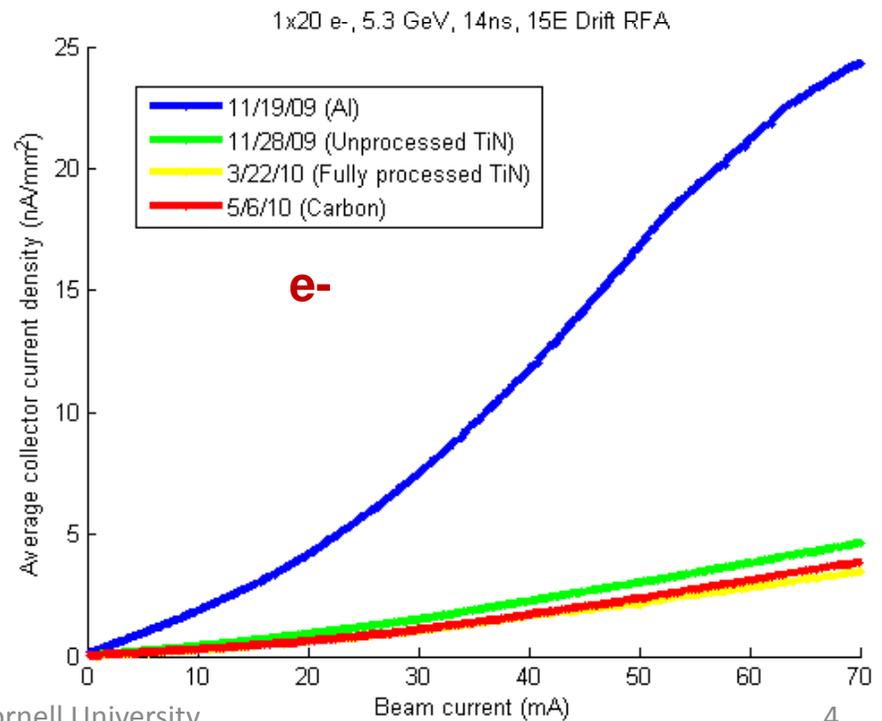
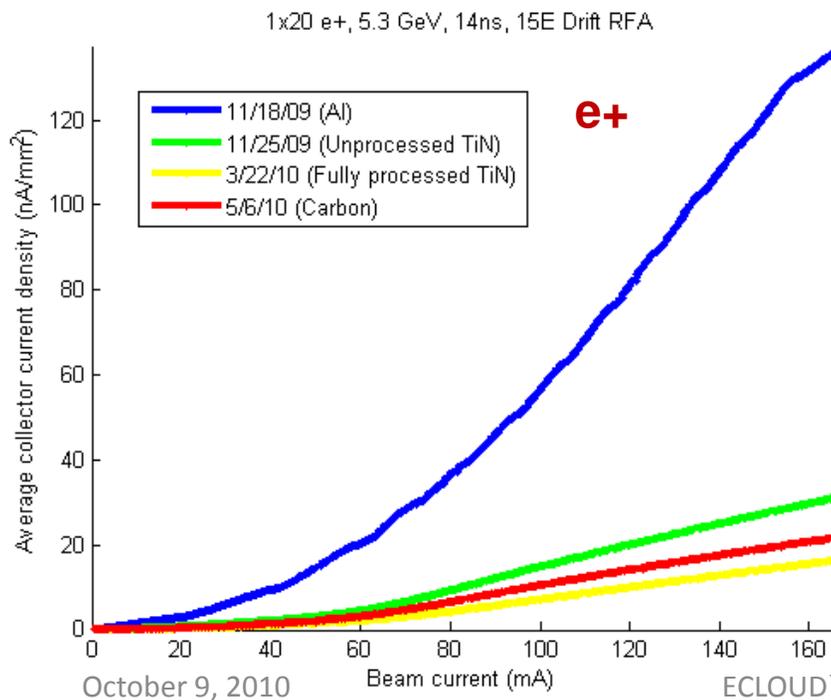


- RFAs consist of...
 - Holes drilled into the beam pipe to allow electrons to pass through
 - A “retarding grid” to which a negative voltage can be applied, rejecting any electrons which have less than a certain energy
 - A collector which captures any electrons that make it past the grid
 - Often there are several collectors arranged transversely across the top of the beam pipe
 - Left: CESR thin drift RFA
- So RFAs provide a local measure of the electron cloud density, energy distribution, and transverse structure
- There are two common types of RFA measurements
 - “Voltage scans,” in which the retarding voltage is varied, typically between +100 and -250V
 - “Current scans,” in which the RFA passively monitors while the beam current is gradually increased

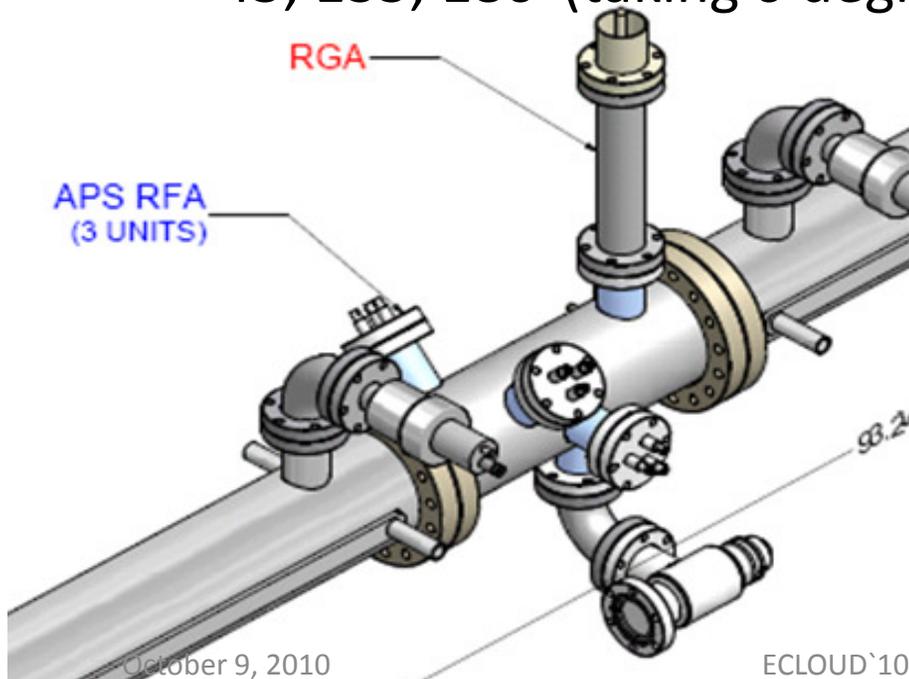




- We have installed chambers with different beam pipe coatings in the same place in CESR, to do as direct a comparison as possible
- Plots show average collector current vs beam current for a 20 bunch train of positions, 5.3 GeV, 14ns spacing
 - Comparing three different chambers (Al – blue, unprocessed TiN – green, processed TiN- yellow, Carbon – red) that were installed in 15E at different times
 - Both coatings show similar performance, much better than Al
 - Carbon chamber did not show significant processing

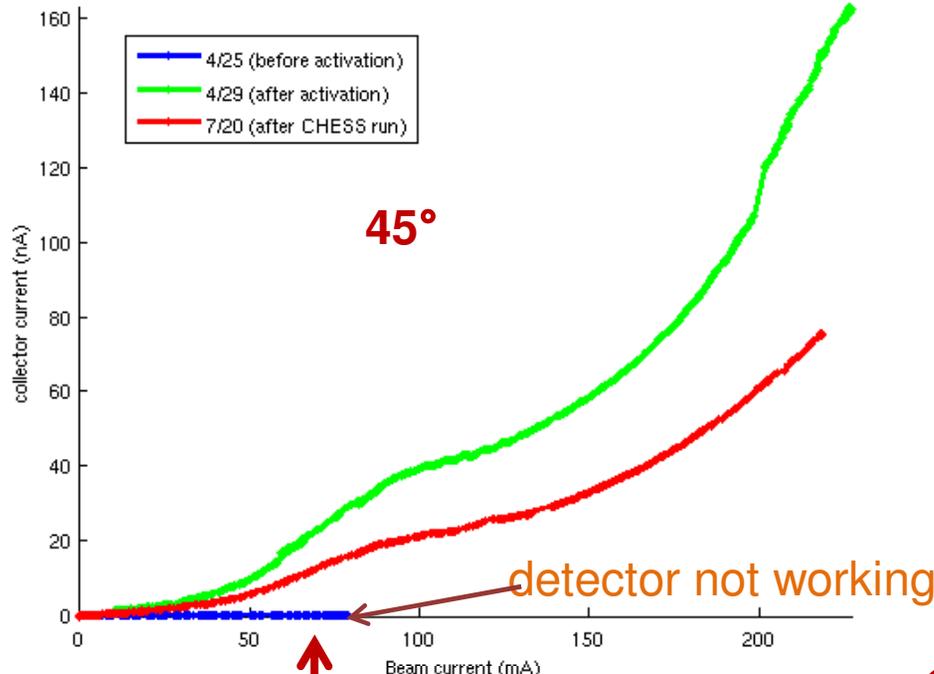


- Installed in L3 straight before April run
 - NEG activated on 4/28
 - Plots compare signal before activation, after activation, and after CHESS run
- 3 single collector (“APS style”) RFAs located at different azimuthal locations in the chamber
 - 45, 135, 180° (taking 0 degrees as source point)

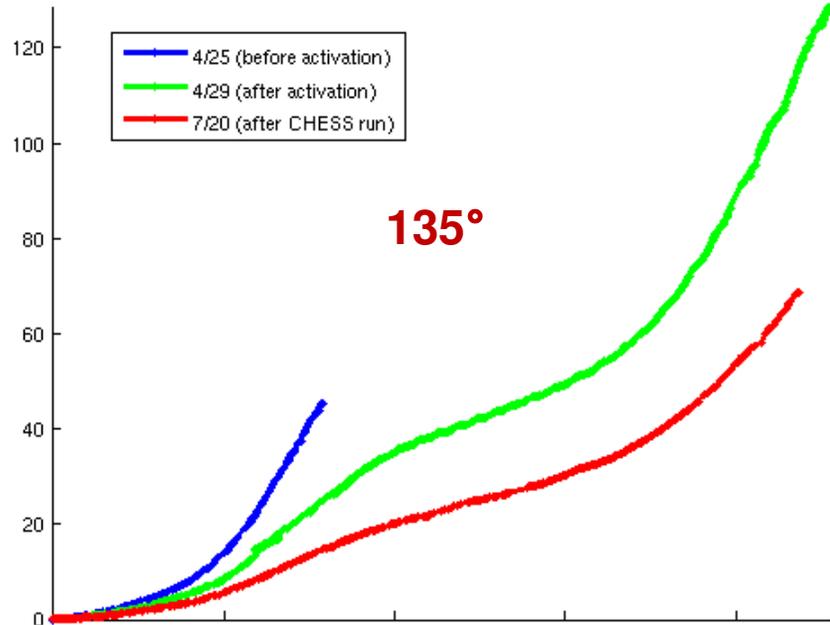


• Signal in all three RFAs was reduced significantly by activating the NEG, and further reduced by processing during the CHESS run.

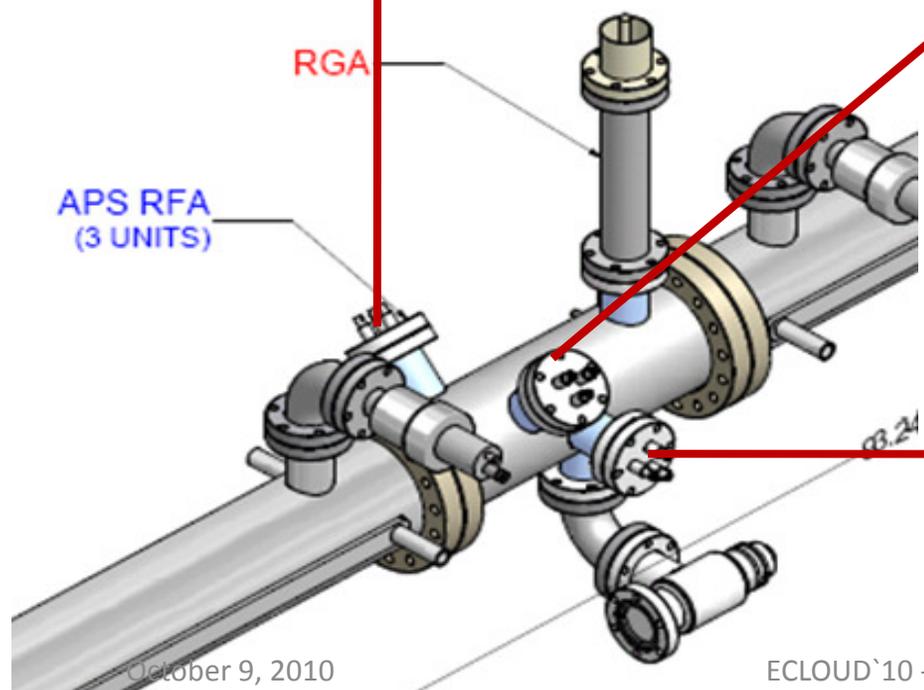
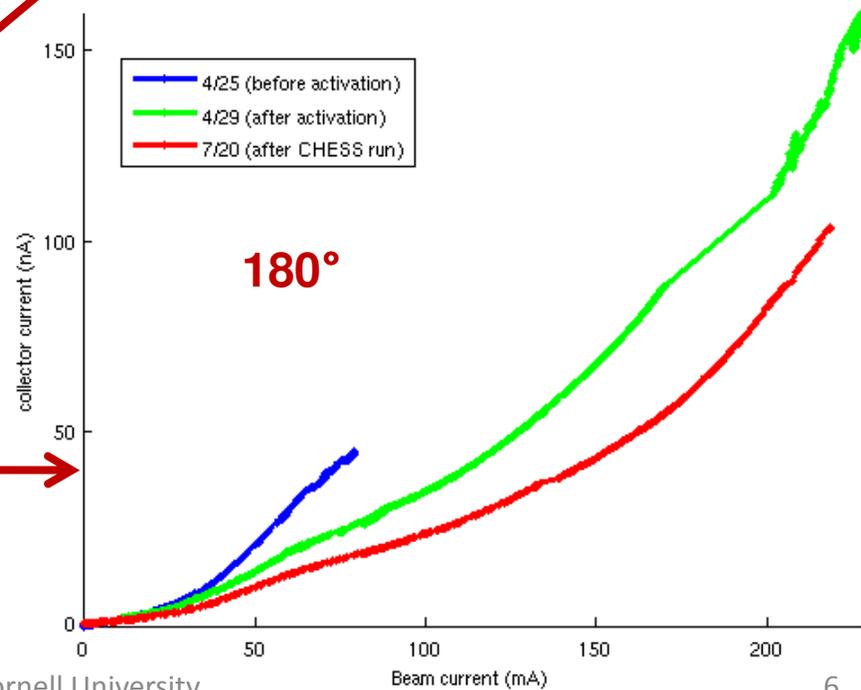
NEG Coated Drift Chamber RFA #4 (45 degrees from outside), 1x20 e+, 14ns, 5.3 GeV



NEG Coated Drift Chamber RFA #3 (45 degrees from inside), 1x20 e+, 14ns, 5.3 GeV

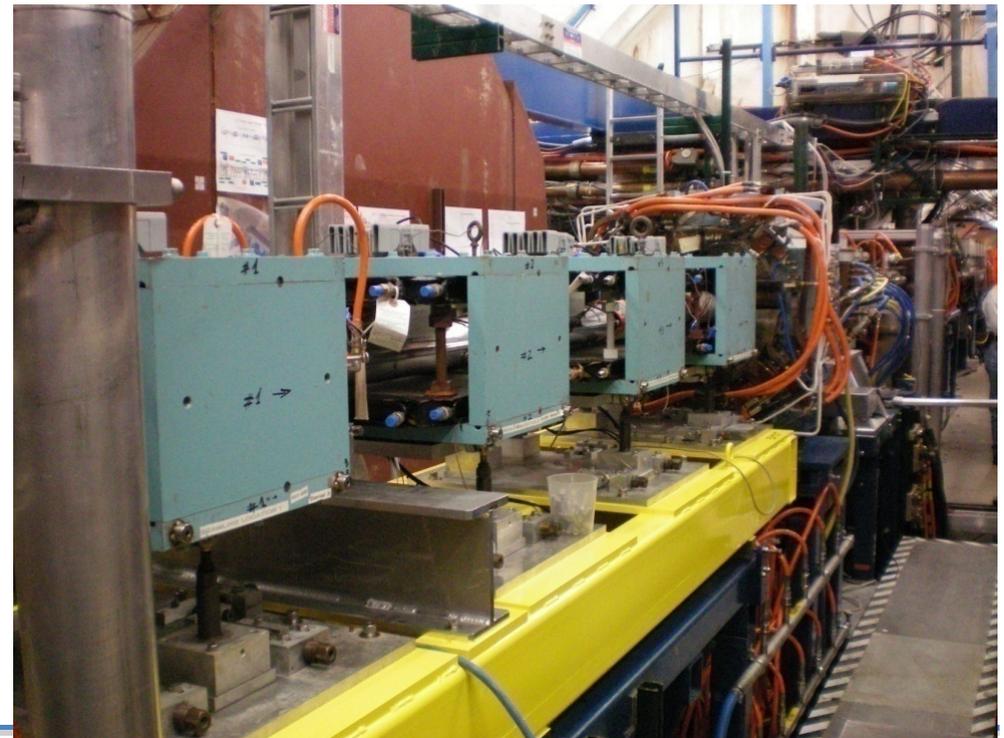


NEG Coated Drift Chamber RFA #2 (Radial Inside), 1x20 e+, 14ns, 5.3 GeV





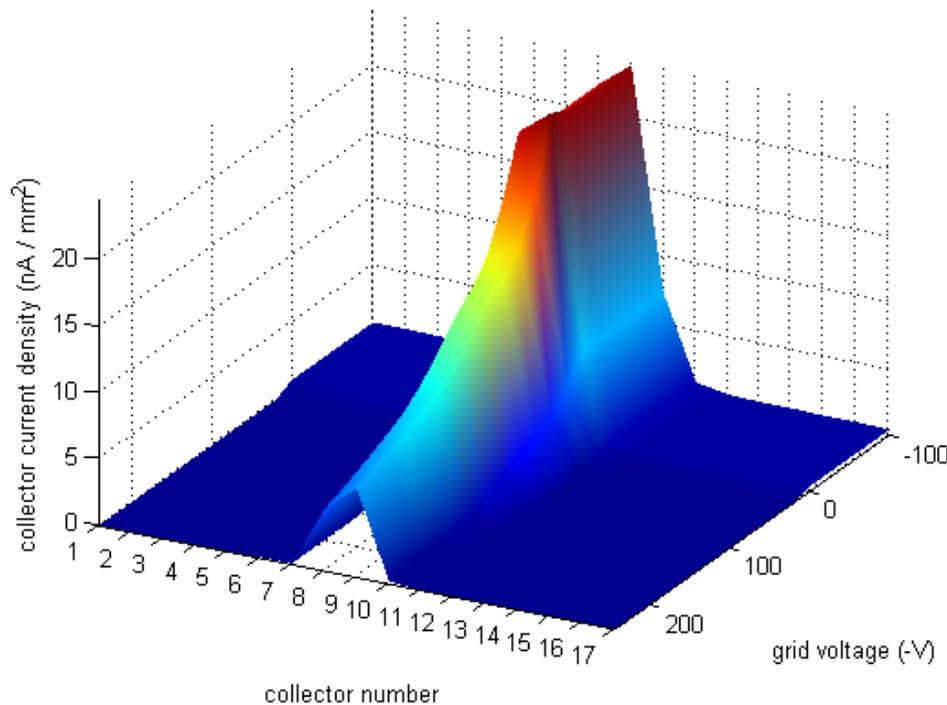
- We have installed the PEP-II chicane in our L3 straight region
 - Each magnet is instrumented with a 17 collector RFA
 - This allows us to investigate the behavior of the cloud as a function of magnetic field
 - Range: ~25 - 1100 Gauss
- Two different mitigation techniques are employed
 - TiN coating (2 magnets)
 - Grooves + TiN coating (1 magnet)
 - The last magnet is bare Aluminum



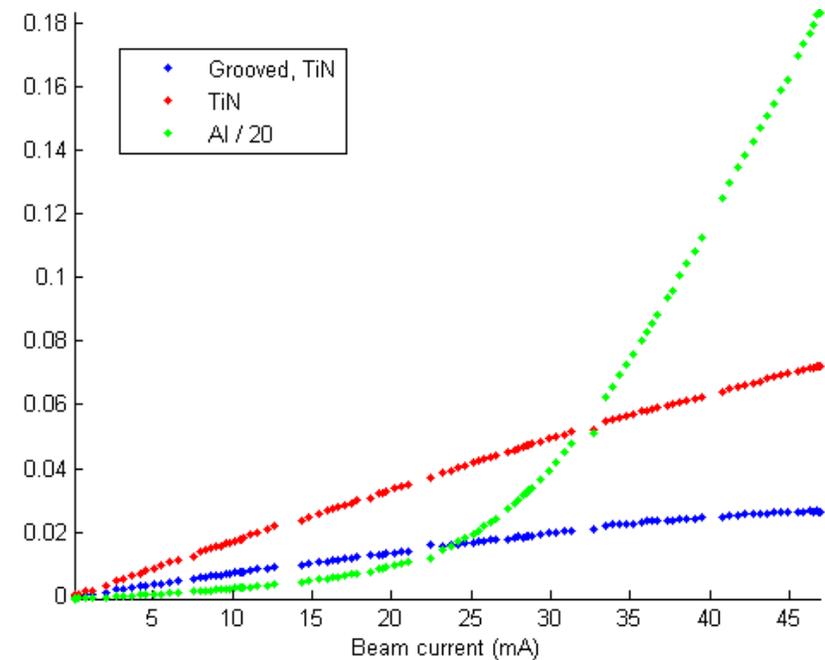


- Left plot is typical voltage scan for Al RFA, 1x45x1.25mA e+, 14ns, 5.3GeV,
- Left plot is current scan, 1x45 e+, 14ns, 5GeV
- Both mitigation techniques show drastic improvement relative to Aluminum
 - Note that Al signal is divided by 20
 - Al shows significant mutipacting
 - TiN actually seems to saturate
 - Groove + TiN is even better than just TiN

Run #2983 (1x45x1.25mA e+, 5.3 GeV, 14ns): L3a_G1 SLAC RFA 4 (Bare Al) Col Curs

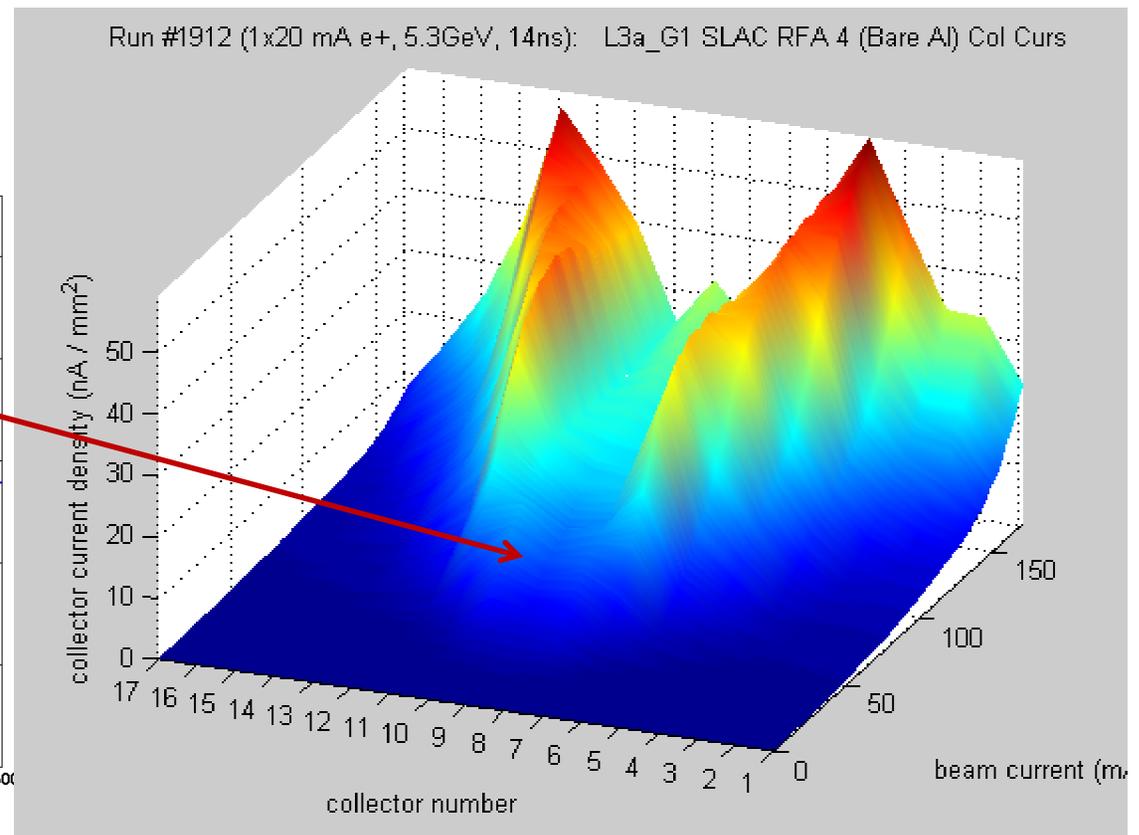
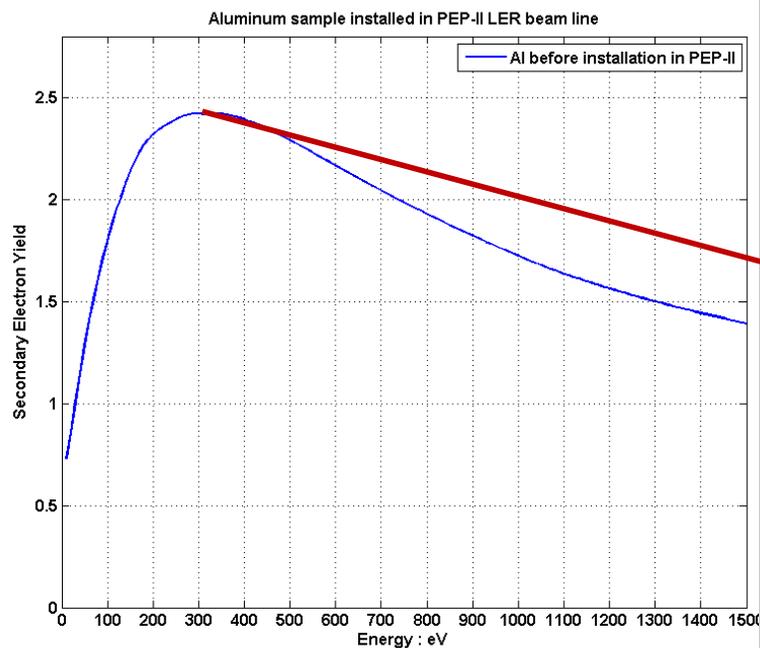


Run 1336: 1x45 e+, 5GeV, 14ns



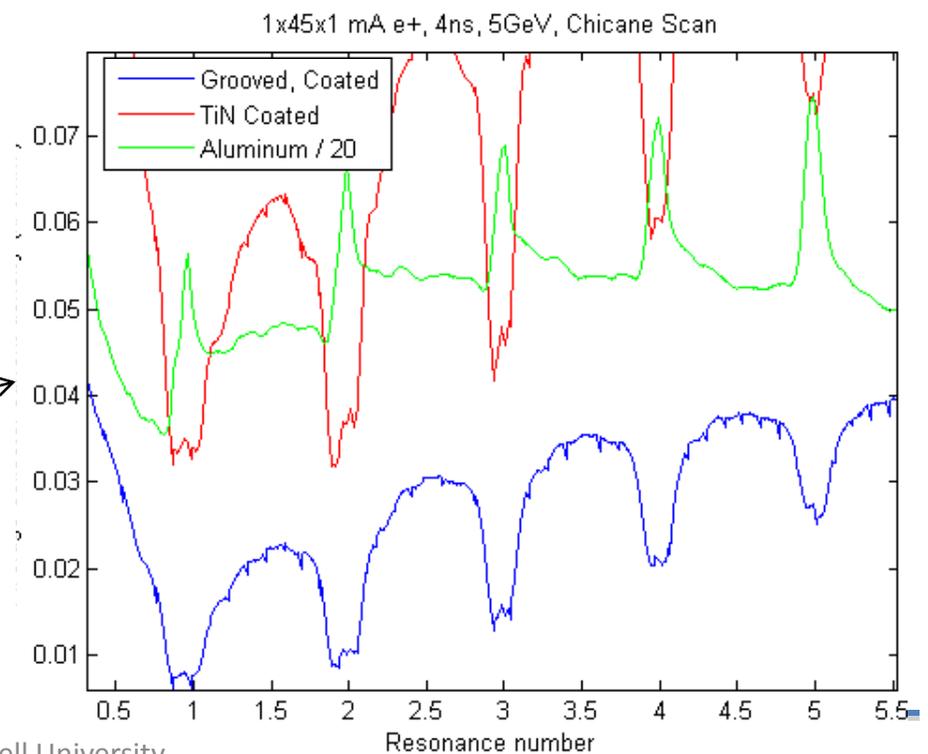
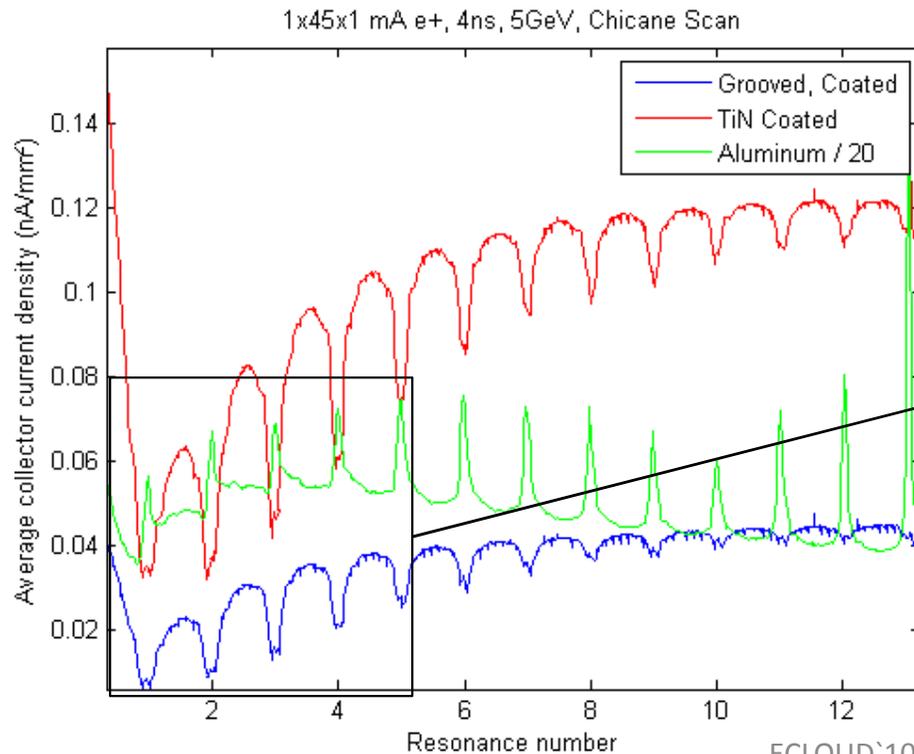


- With sufficient bunch current, one can push the average cloud energy in the center of the pipe past the SEY peak
 - This causes a bifurcation of the peak density
- Conditions: 1x20 e+, 5.3 GeV, 14ns, +50V on grid
- Plot shows collector currents vs beam current (\sim cloud energy) and collector number (horizontal position)
 - Aluminum SLAC RFA
(in chicane), \sim 700G dipole field



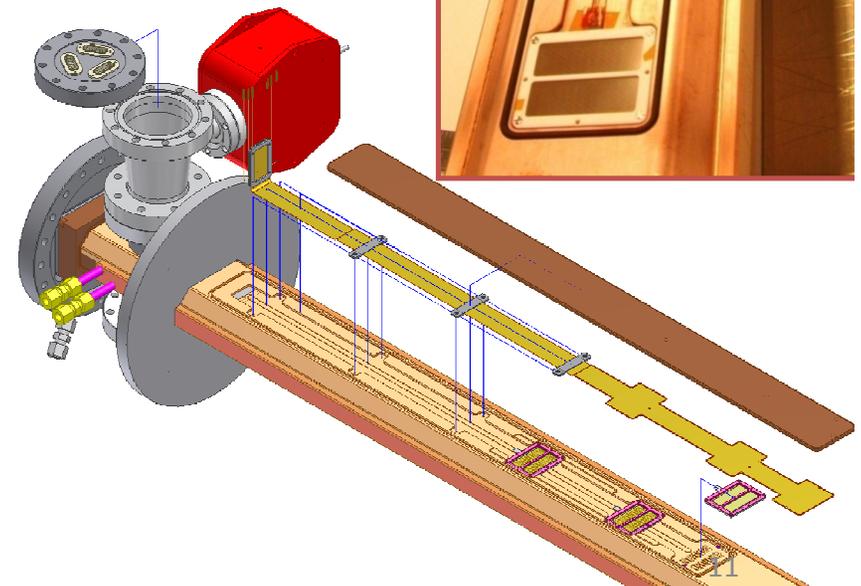
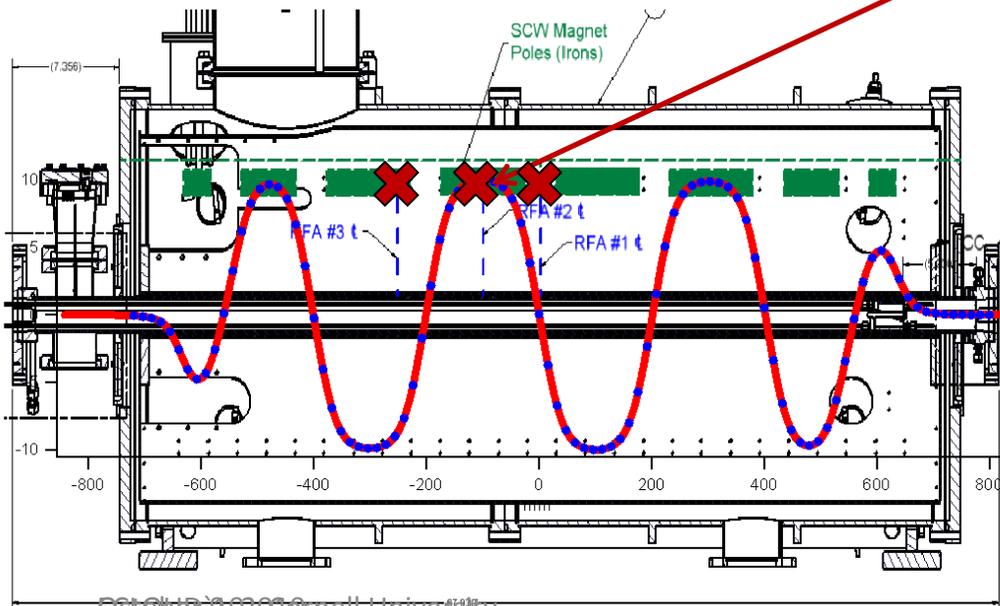
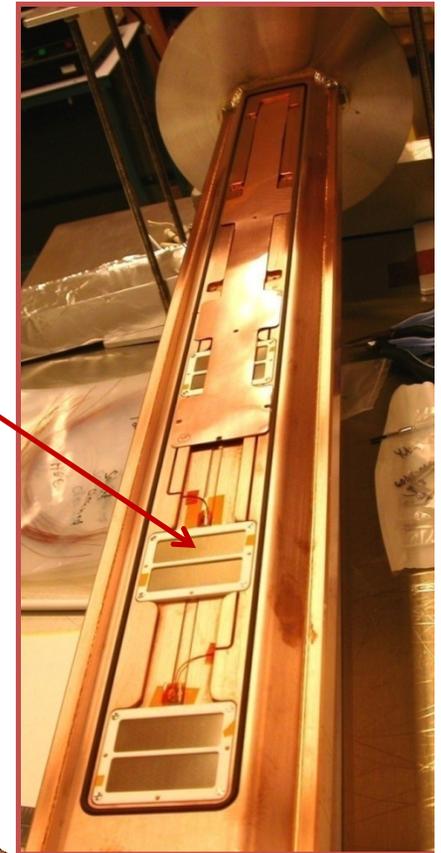


- RFA currents monitored while chicane dipole fields are increased
- We are looking for “cyclotron resonances”
 - When the bunch spacing is an integral multiple of the cyclotron period of an electron
 - Data are plotted against “resonance number” (= bunch spacing / cyclotron period)
- 1x45x1 mA, 4ns, 5GeV, positrons
- On resonance, there are peaks in the Al chamber and dips in the TiN and grooved chambers
 - Both dips and peaks are exactly on resonance
 - Not clear what causes dips vs peaks





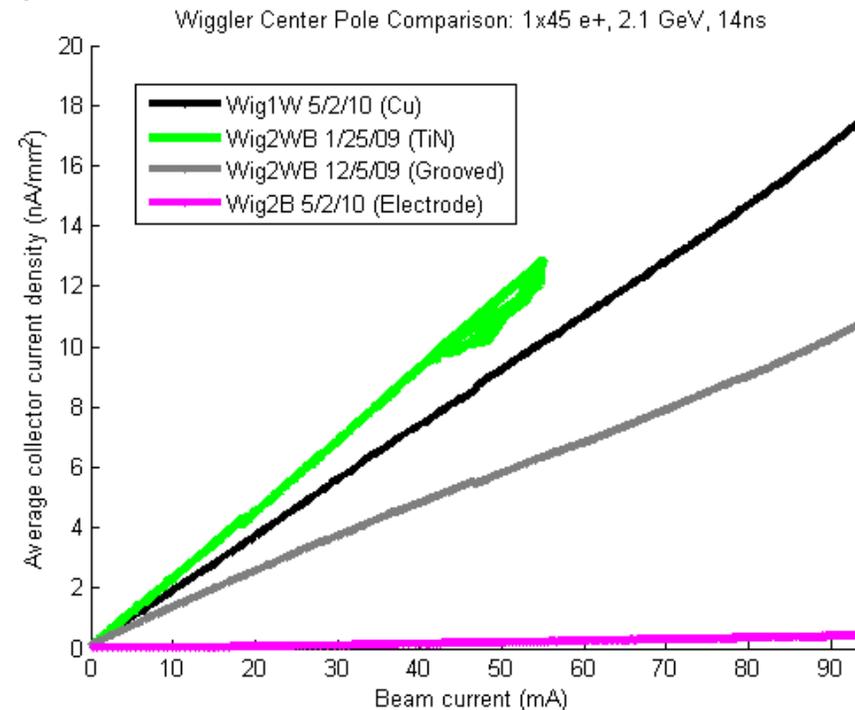
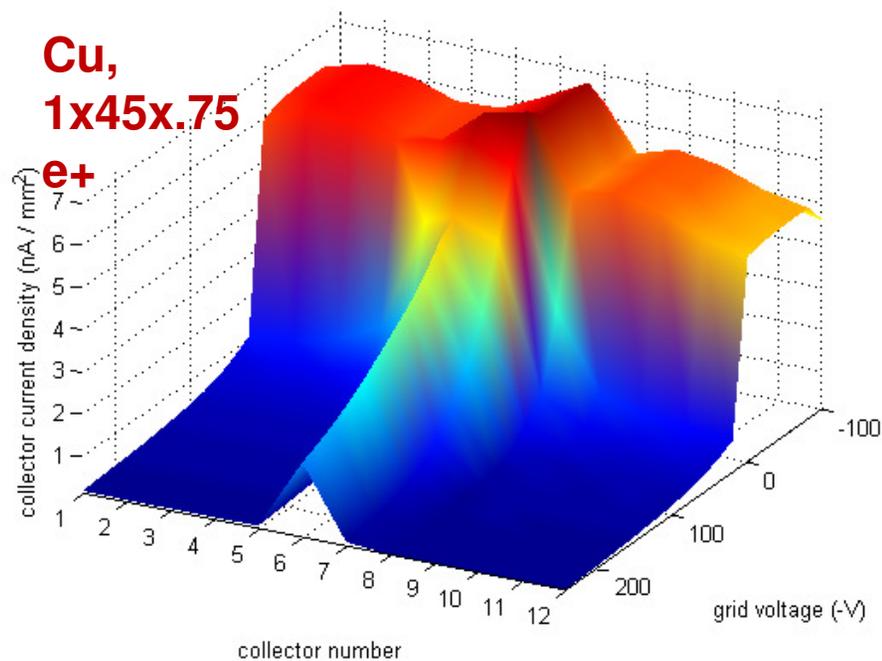
- We have three wigglers instrumented with RFAs
 - Bare Cu
 - TiN coated
 - Clearing electrode
 - Previously installed: grooved
- Each wiggler has three RFAs
 - Plots shown will be for an RFA in the center of a wiggler pole
 - There are also RFAs in a longitudinal and intermediate field
 - RFAs have 12 collectors and are built into the beam pipe





- Left plot shows typical voltage scan in Cu center pole wiggler
- Right plot shows average collector current density vs beam current
 - 1x45 e+, 2.1 GeV, 14ns
 - TiN, Grooved, Electrode chamber all in same location at different times
- Cu, TiN, and grooved chambers all within a factor of two
 - Electrode chamber does significantly better

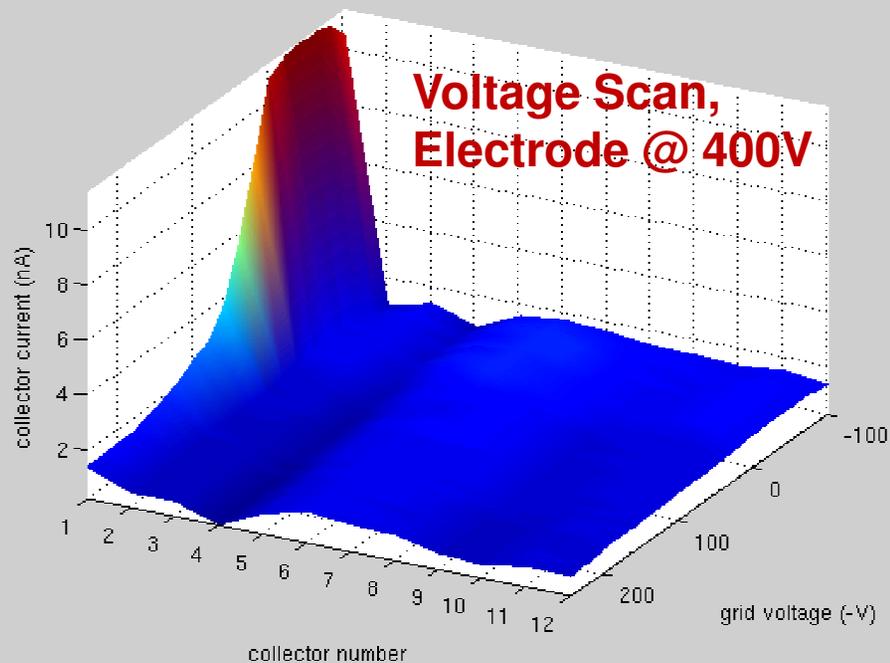
Run #2955 (1x45x.75mA e+, 2.1 GeV, 14ns): 01W_G1 Wig1W Center pole Col Curs



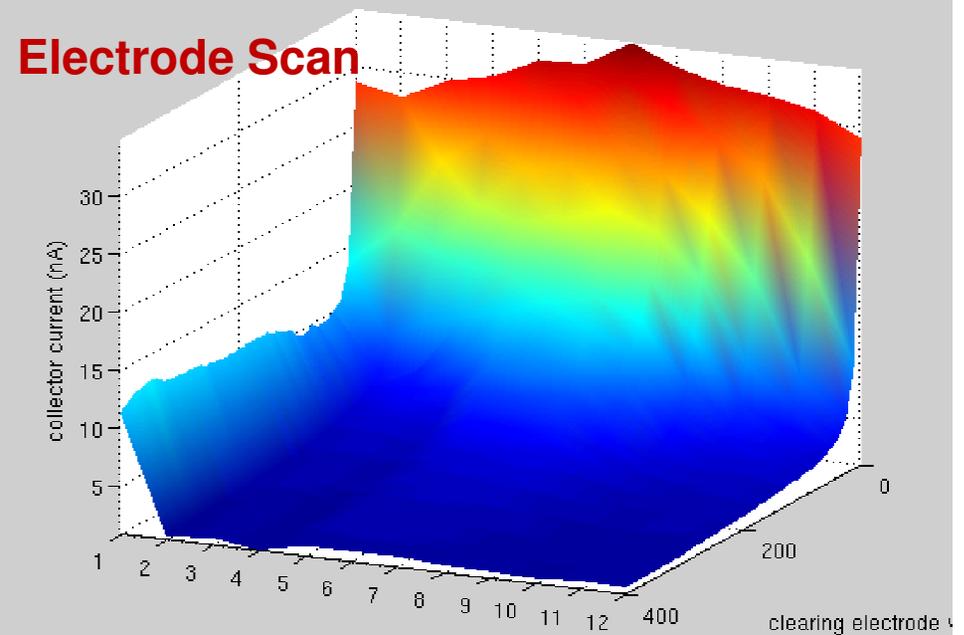


- Goes up to 400V
- 1x20x2.8 mA e+, 14ns, 4 GeV, wigglers ON
- Cloud suppression is very strong, except on collector 1
 - Electrode is exactly the width of the RFA
 - In other collectors, signal is essentially gone by 100V

Run #2569 (Electrode:400V, 1x20x2.8mA e+, 4GeV, 14ns): 01W_G2 Center pole Col Curs

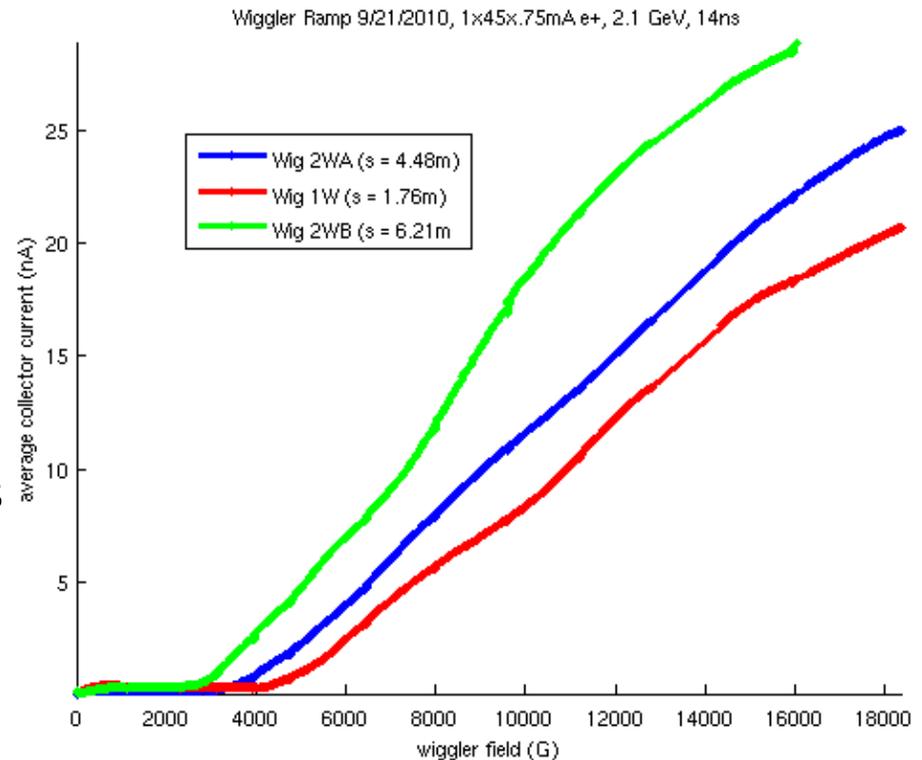


Run #2568 (Electrode Scan: 1x20x2.8mA e+, 4GeV, 14ns): 01W_G2 Center pole Col Curs





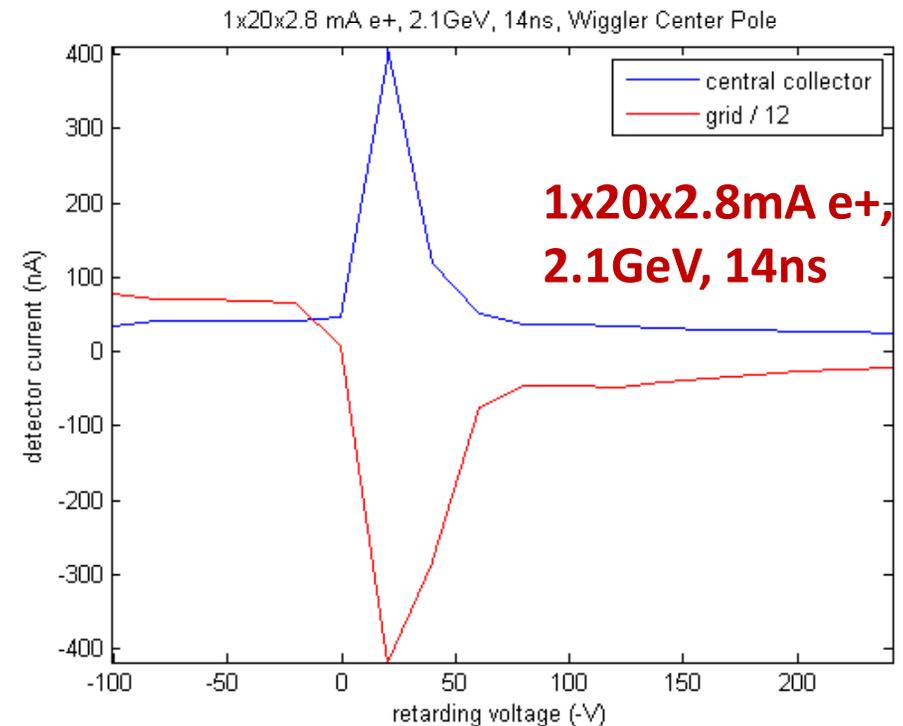
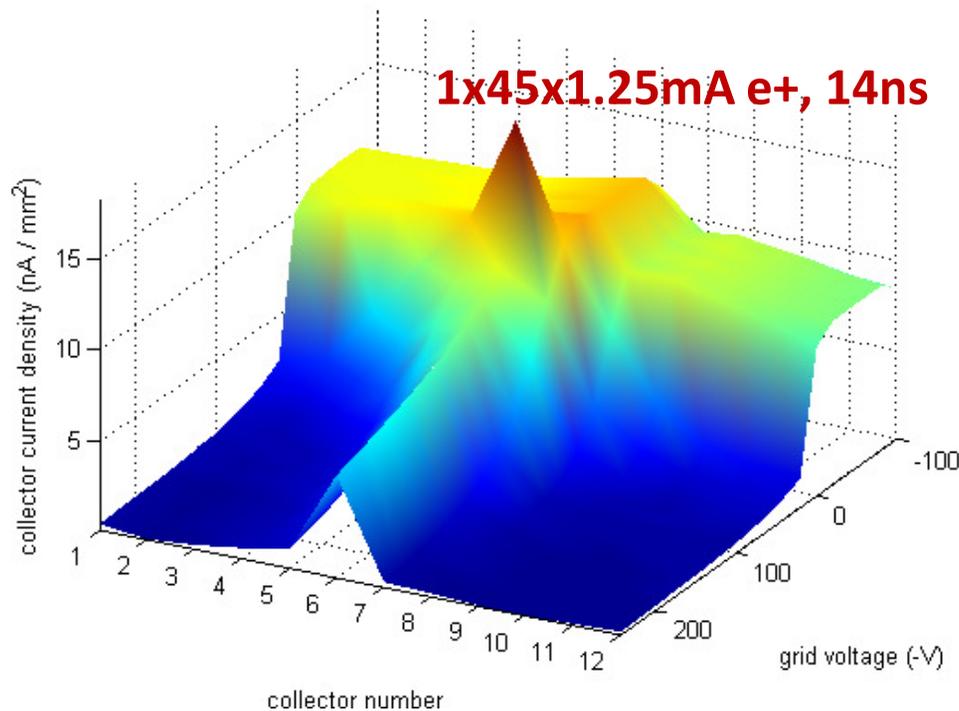
- LO RFA currents were monitored while wigglers were ramped down
- Plot shows average collector current in wiggler center pole RFAs as a function of wiggler field strength
 - Note “turn on” of signal at each RFA, presumably as photons from upstream wigglers hit the beam pipe at that location
 - Further downstream wigglers turn on sooner
 - Beam conditions:
 - $1 \times 45 \times \sim .75 \text{ mA } e^+$,
 - Normalized to beam current
 - 2.1 GeV, 14ns
- Helpful, since photon flux is difficult to calculate in straight sections (depends strongly on reflections)





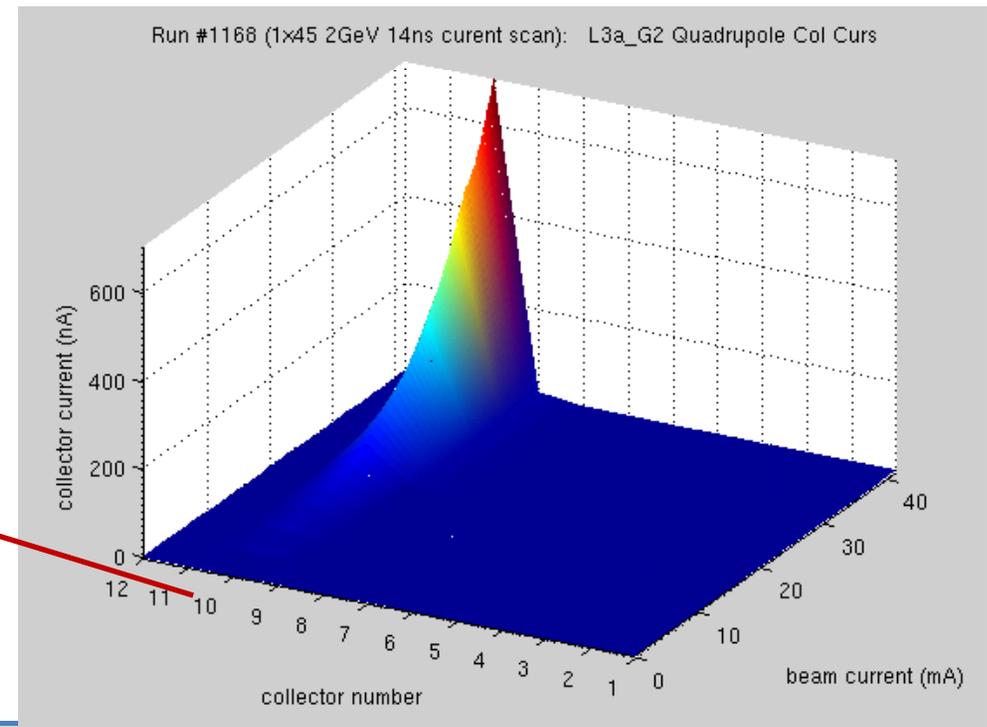
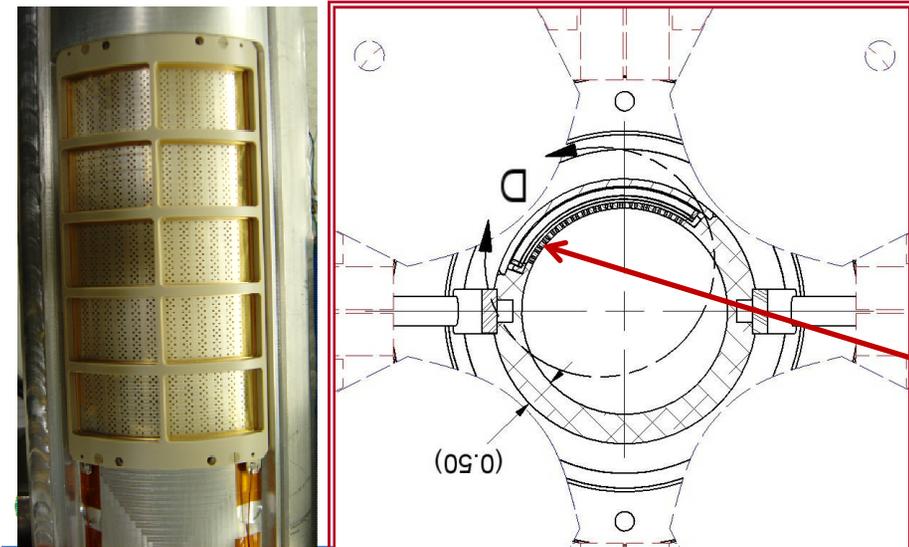
- In a high magnetic field (e.g. wiggler pole center), electrons are strongly pinned to the field lines
- Secondary electrons produced on grid can be accelerated through retarding voltage back out into vacuum chamber
- End result is a resonant condition between retarding voltage and bunch spacing
 - Leads to an enhancement in signal at low (but nonzero) retarding voltage

Run #2585 (1x45x1.25mA e+, 14ns, 2.1GeV): 01W_G1 Wig1W Center pole Col Curs



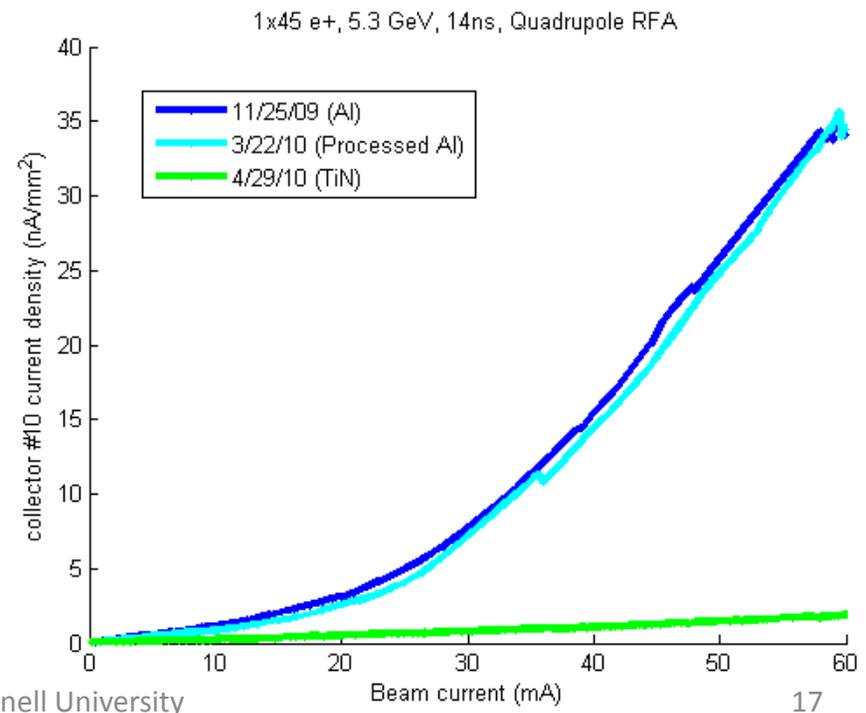
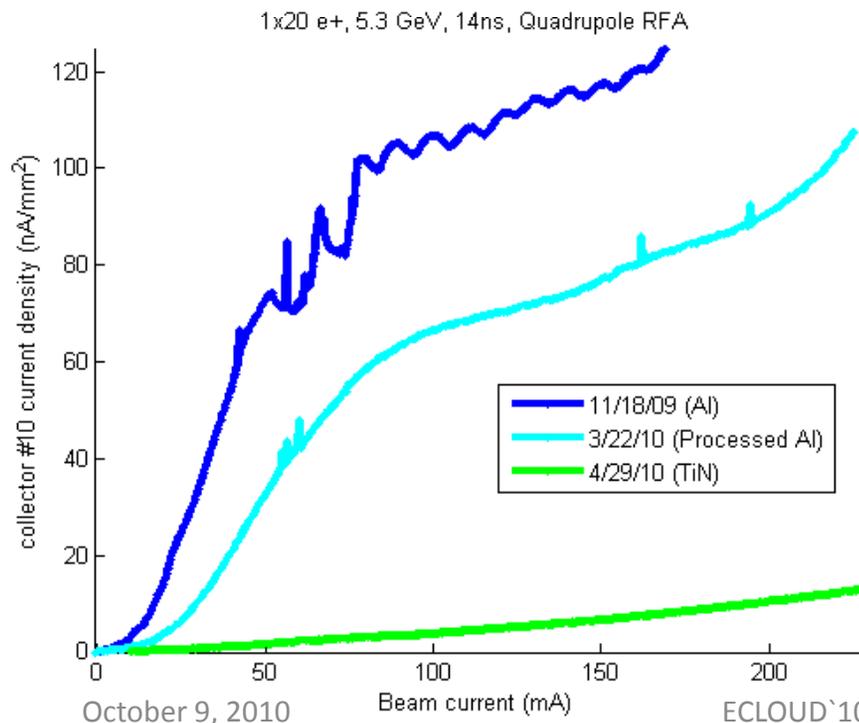


- We have instrumented a quadrupole chamber with an RFA
- One collector sees a huge amount of current
 - This is where the electrons are guided by the quad field lines
- There have been both bare Al and TiN coated chambers installed in the same location





- Plotting current in collector #10 (the one that sees a large signal)
- TiN shows improvement of well over an order of magnitude

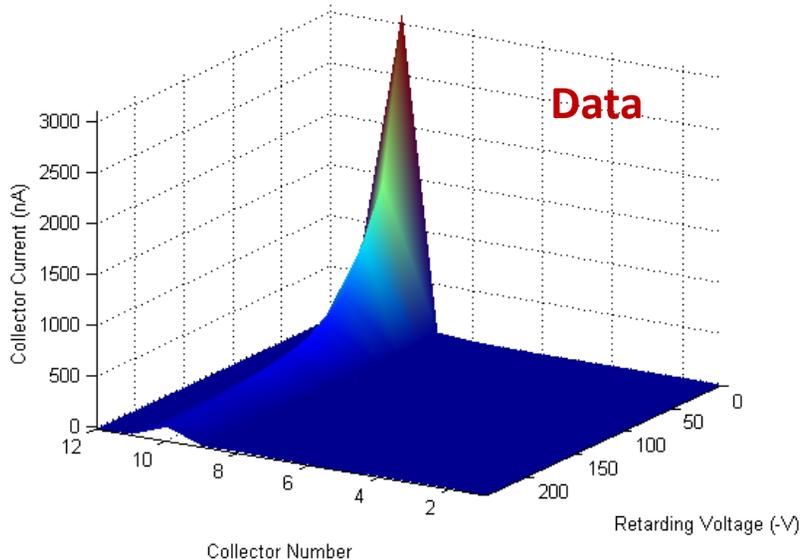




Slow Buildup in Quadrupole

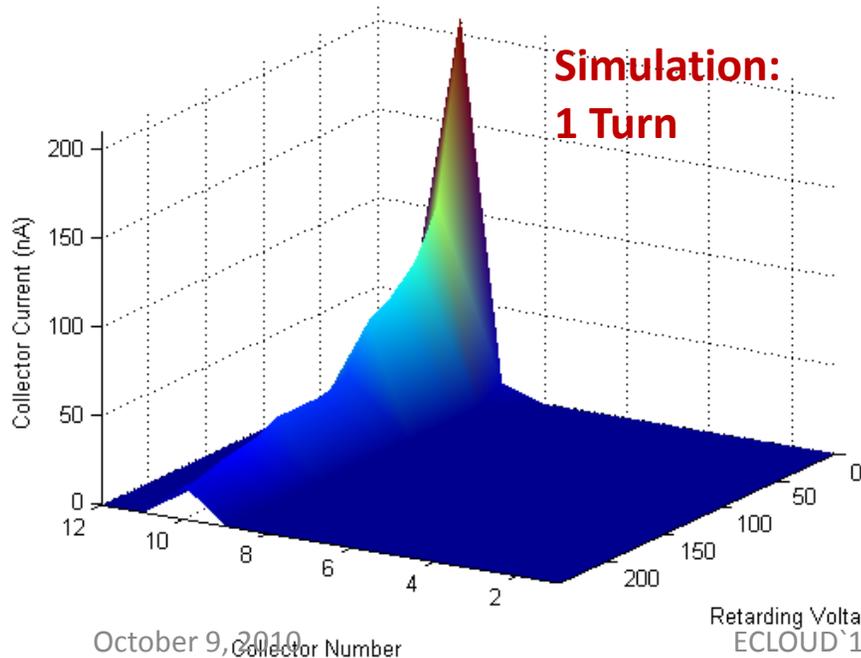


1x45x1 mA e+, 5.3GeV, 14ns, 9.2 T/m Quad, 1 Turn- Data

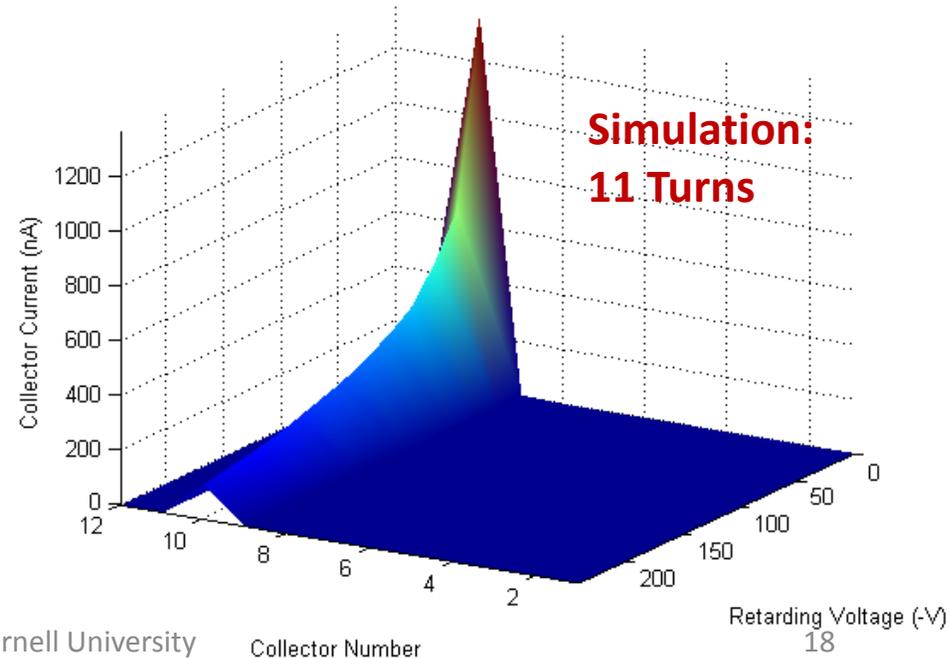


- 1x45x1 mA e+, 5.3GeV, 9.2 T/m
- 1 turn simulation underestimates data by more than an order of magnitude
- 11 turn simulation is quite close at high energy, within a factor of 2 at low energy
- This indicates cloud is building up over several turns before it reaches equilibrium
 - So it must be persisting over the $\sim 2\mu\text{s}$ between trains

1x45x1 mA e+, 5.3GeV, 14ns, 9.2 T/m Quad, 1 Turn- Simulation



1x45x1 mA e+, 5.3GeV, 14ns, 9.2 T/m Quad, 11 Turns- Simulation





- Goal: Use RFA data to provide constraints on the surface parameters of the chamber --> a challenging exercise
- Requires cloud simulation program (e.g. POSINST or ELOUD)
- Also need a model of the RFA itself
 - Method 1: post-processing
 - Perform a series of calculations on the output of a simulation program to determine what the RFA would have seen had it been there
 - Relatively easy, can perform an entire “voltage scan” on the output of one simulation
 - Method 2: integrated model
 - Put a model for the RFA in the actual simulation code
 - More self-consistent, can model effects of the RFA on the development of the cloud
 - Need to do a separate simulation for each retarding voltage



- **Beam pipe hole secondaries**

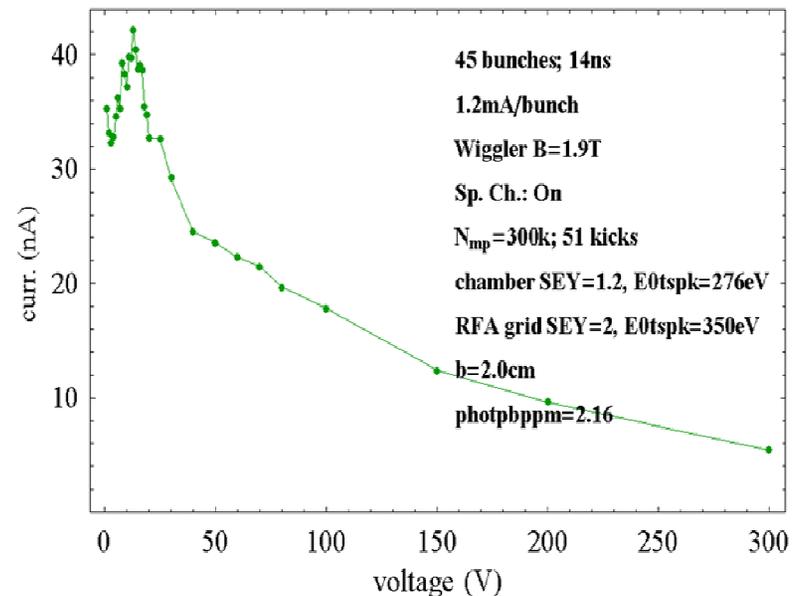
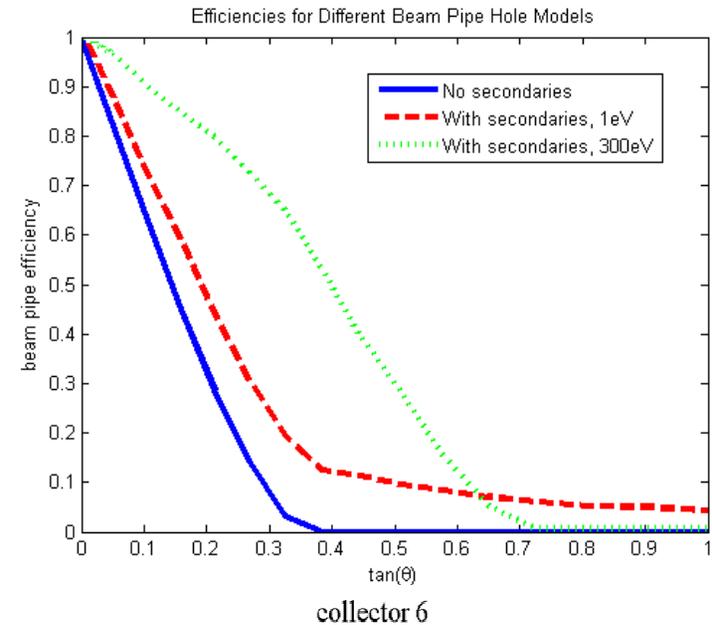
- Secondary electrons can be generated in the beam pipe holes in front of the RFA, leading to a low energy enhancement in the RFA signal.
- We have developed a specialized particle tracking code to quantify this effect.
- This code indicates low energy electrons maintain some probability of a successful passage even at high incident angle (due to elastic scattering)
- High energy electrons have a higher efficiency at intermediate angles (due to the production of "true secondaries."

- **Photoelectron model:**

- The traditionally used low energy photoelectrons do not provide sufficient signal for electron beam data with high bunch current.
- A Lorentzian photoelectron energy distribution with a wide width (~150 eV) has been added to POSINST.

- **Interaction with cloud:**

- The "resonant enhancement" has been observed qualitatively with integrated models in ECLLOUD in POSINST





- Need a systematic method to extract best fit simulation parameters from large amount of data.
- 1. Choose a set of (related) voltage scans
- 2. Choose a set of simulation parameters
- 3. Do a simulation with the nominal values for each parameter
- 4. Postprocess the output of simulations to obtain a predicted RFA signal
- 5. For each data set and each parameter, do a simulation with a high and low value of the parameter, and determine the predicted RFA signal
- 6. For each data point in the simulated voltage scan, do a best linear fit to the curve of RFA signal vs parameter value. The slope of this line determines how strongly this point depends on the parameter
- 7. Try to find a set of parameters that minimizes the difference between data and simulation, assuming linear dependence of each voltage scan point on each parameter.
- 8. Repeat the process until fits stop getting better

• Simulations have been done for
beam conditions shown in table

Condx #	Run #	Bunches	Spacing (ns)	Energy (GeV)	Bunch Current (mA)	Species
20	2615	20	14	5.3	2.8	e+
21	2619	20	14	5.3	10.75	e+
22	2624	45	14	5.3	0.75	e+
23	2626	45	14	5.3	1.25	e+
24	2628	45	14	5.3	2.67	e+
25	2632	9	280	5.3	4.11	e+
26	2635	20	14	5.3	2.8	e-
27	2642	20	14	5.3	10.75	e-
28	2647	45	14	5.3	0.8	e-
29	2651	45	14	5.3	1.25	e-
30	2655	9	280	5.3	3.78	e-





Parameter "Domains"



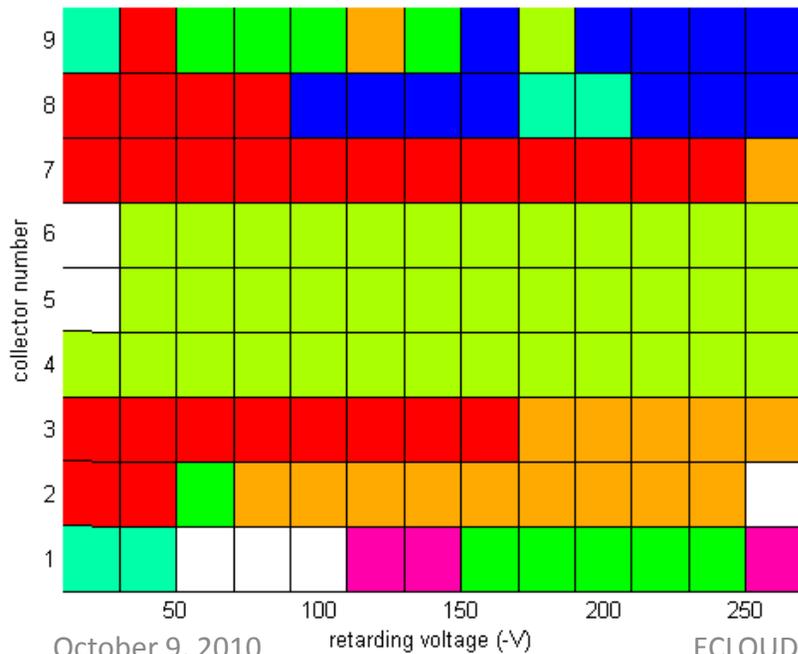
49th ICFA Advanced Beam Dynamics Workshop

	dtspk (true secondaries)
	P1rinf (rediffused)
	delta0 (yield at E=0)
	E _{max} (peak energy)
	e ⁺ quantum efficiency
	e ⁻ quantum efficiency
	primary energy width, e ⁺
	primary energy width, e ⁻
	beam displacement

- We want to understand where each parameter matters the most
 - Plots show the "strongest" (i.e. highest slope) parameter, as a function of retarding voltage and collector number, for various conditions
 - Color coded according to legend to the left
- Examples shown are for Aluminum chamber

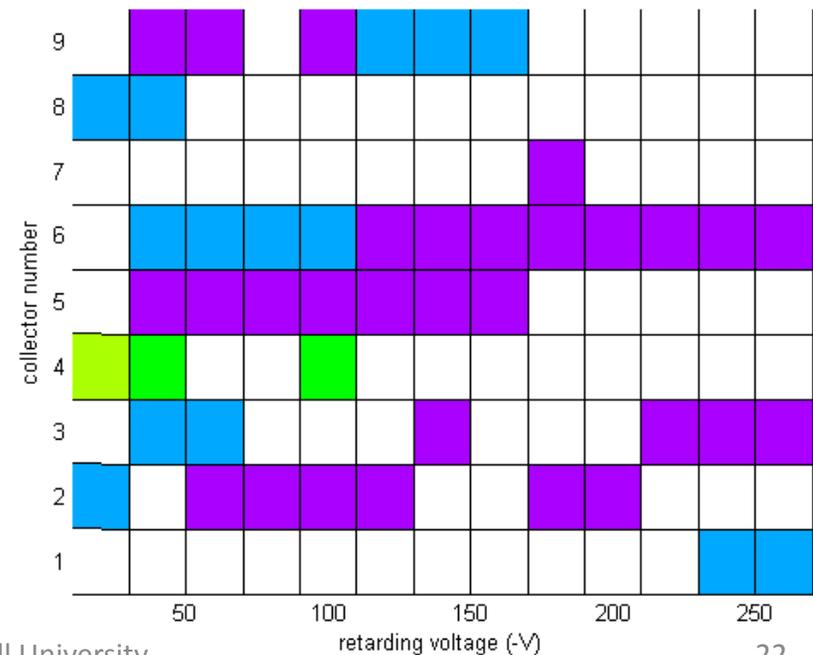
1x20x10.75mA, e⁺, 14ns

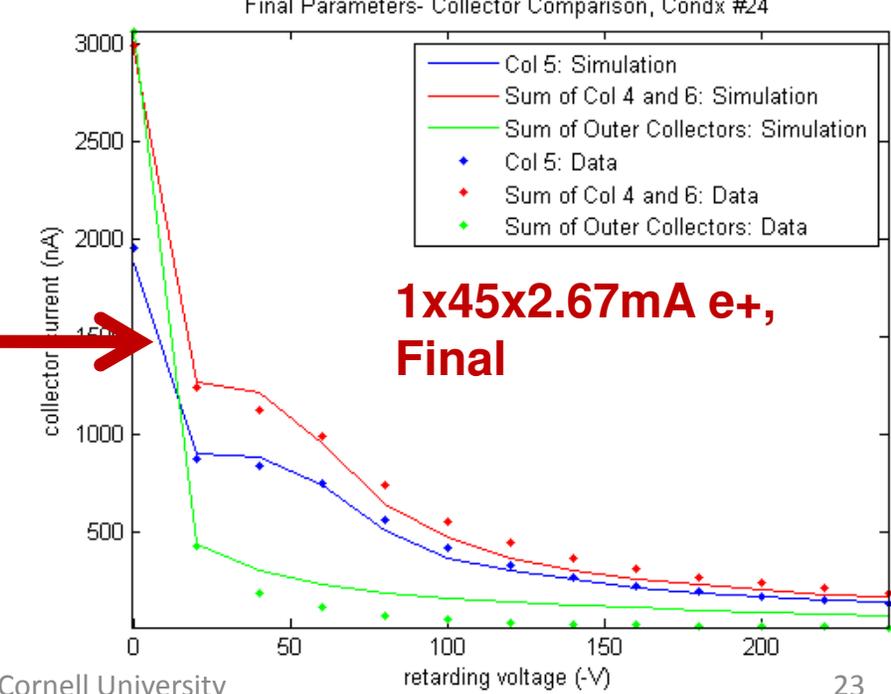
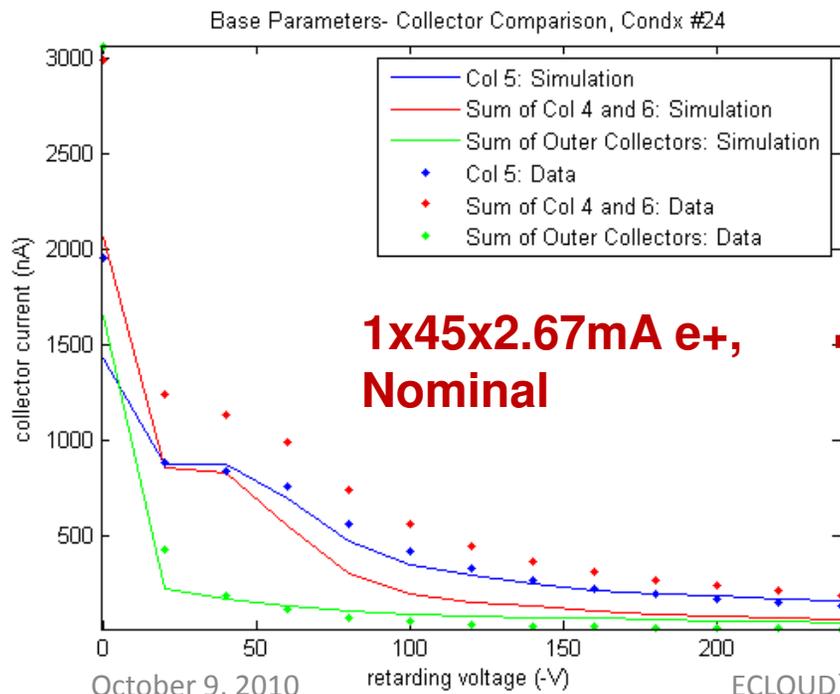
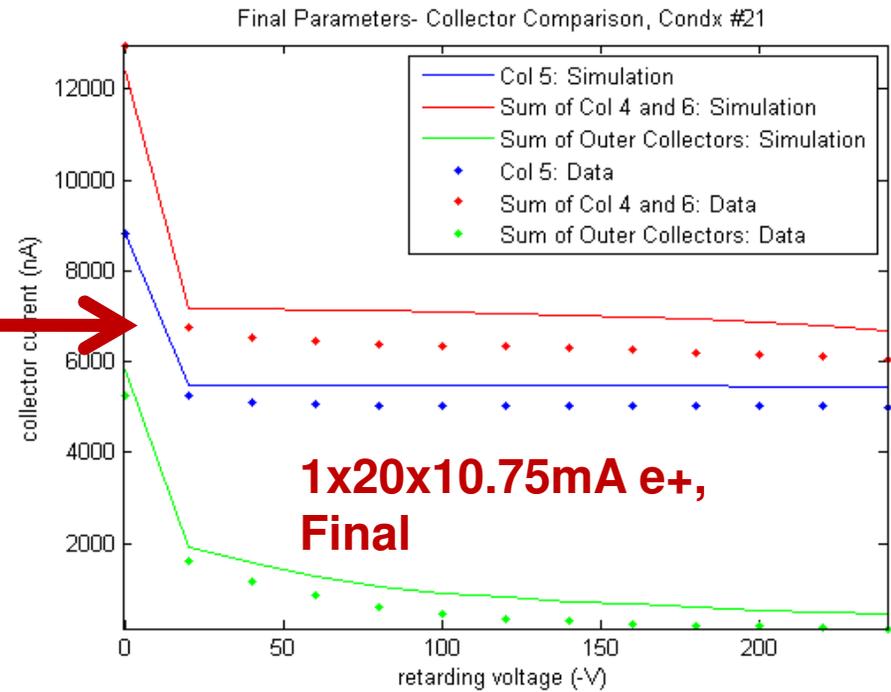
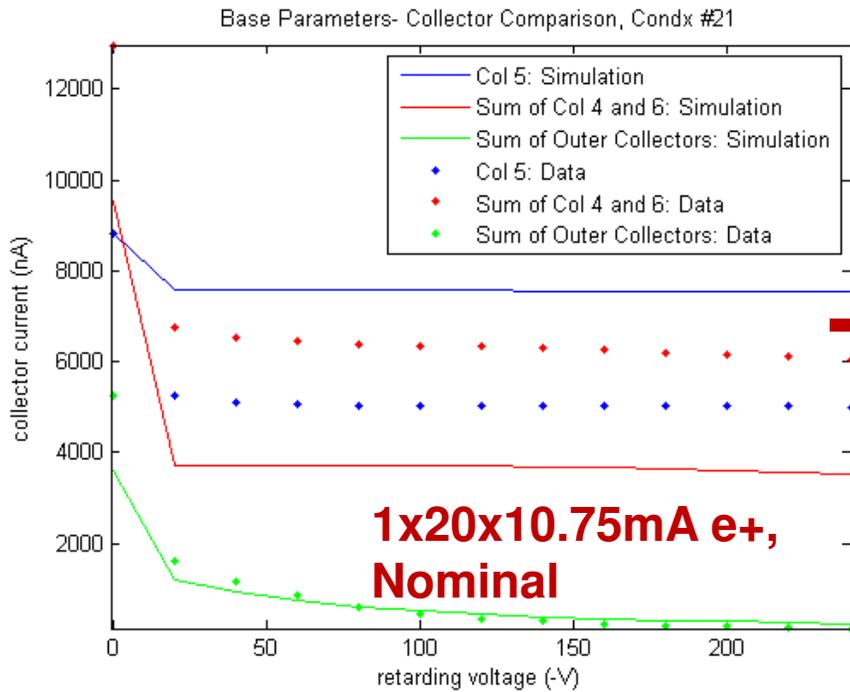
Strongest Parameter, Condx #21

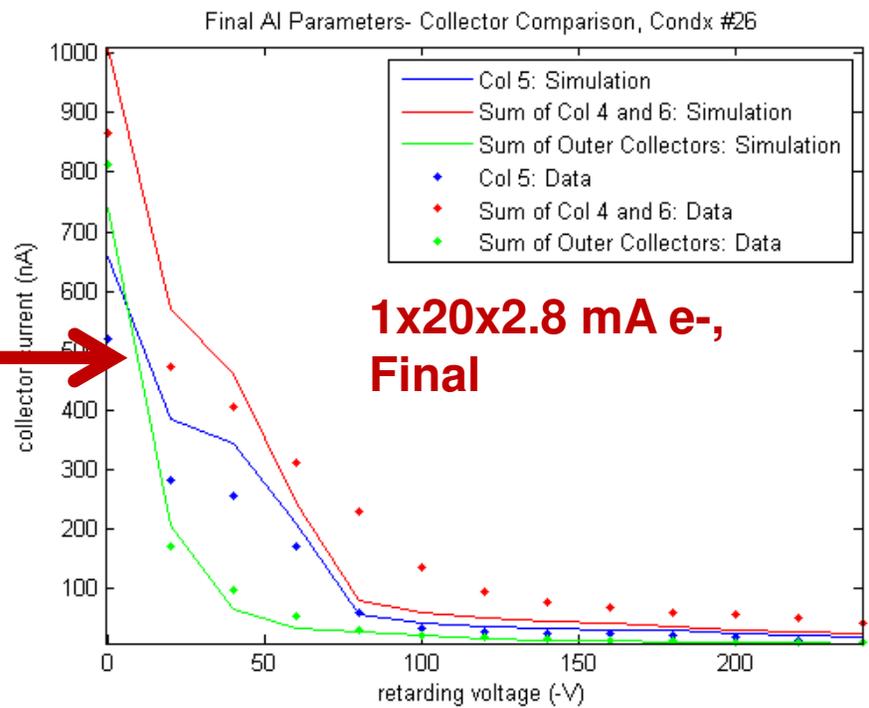
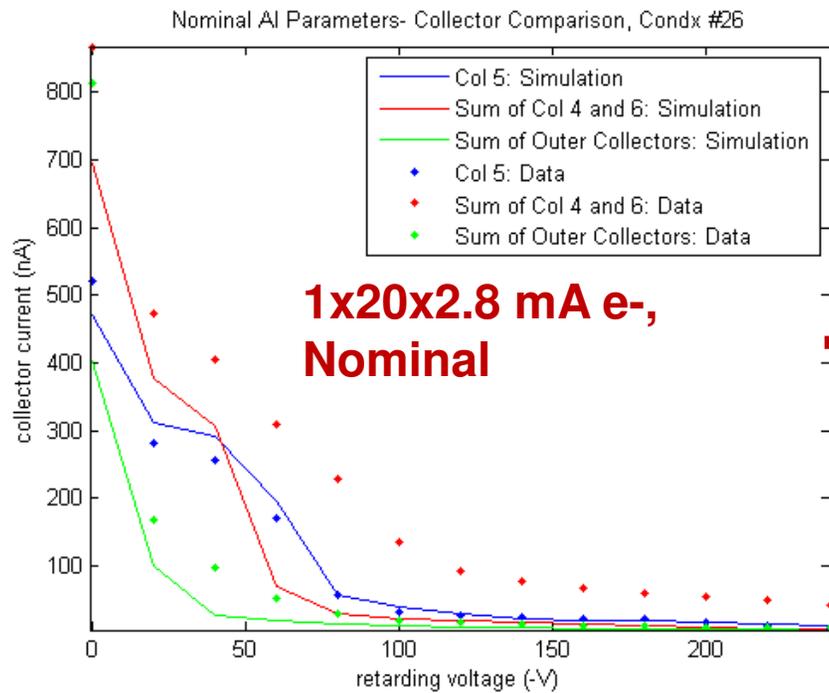
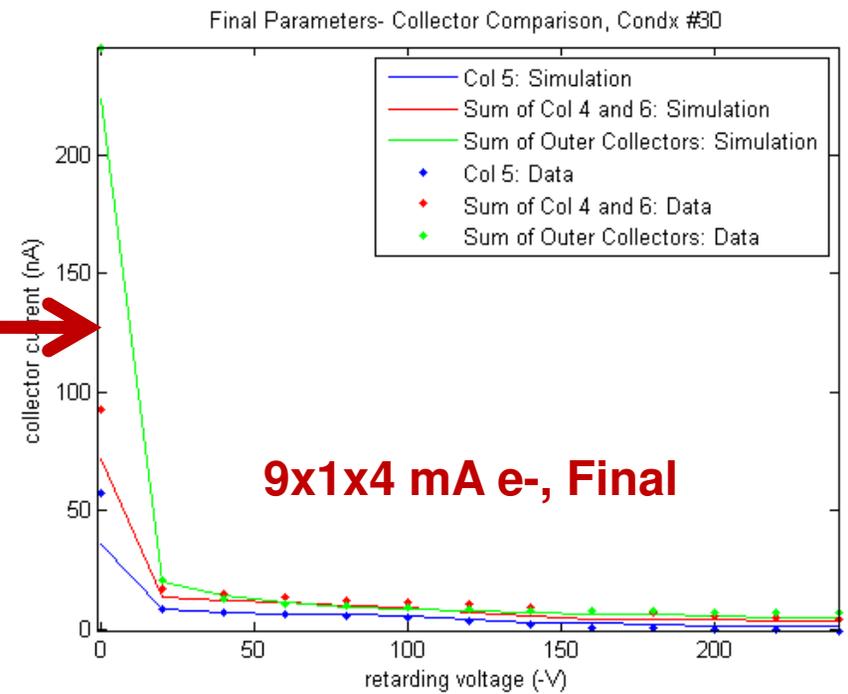
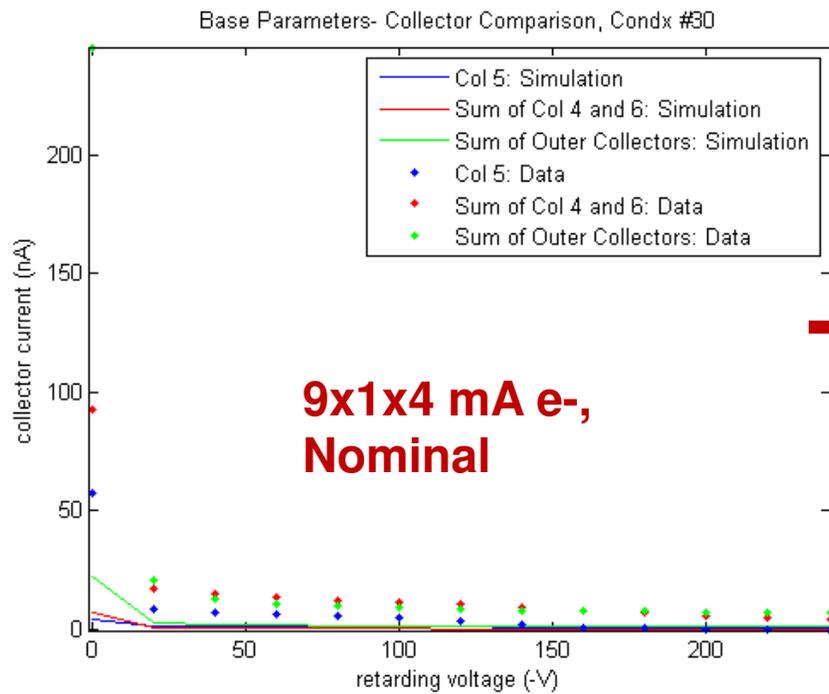


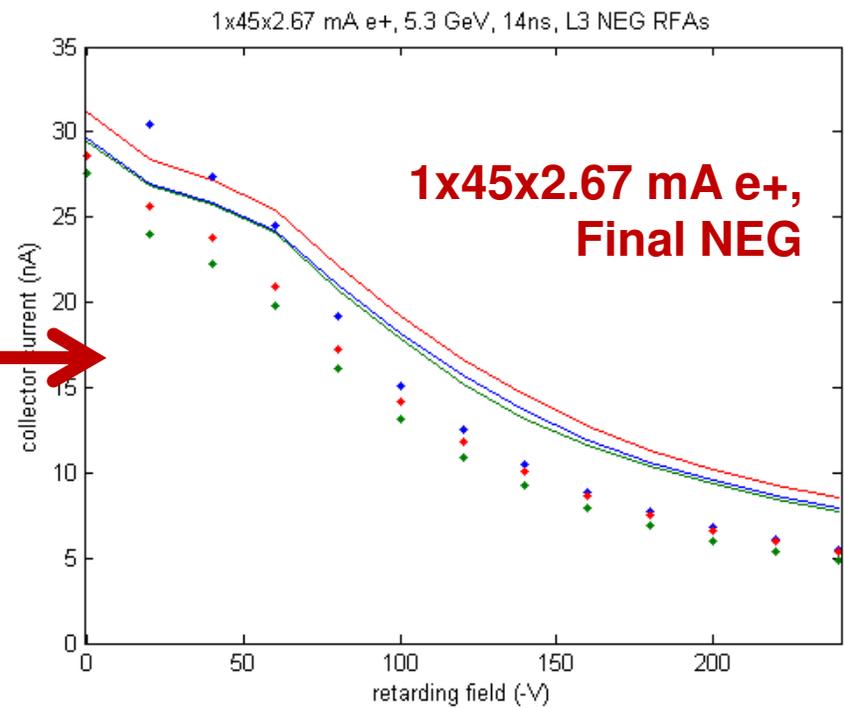
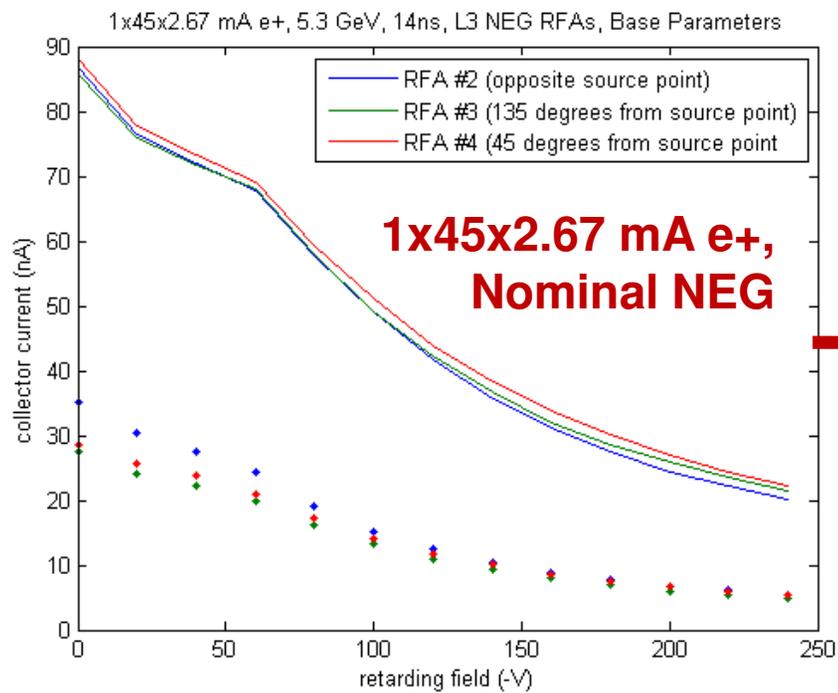
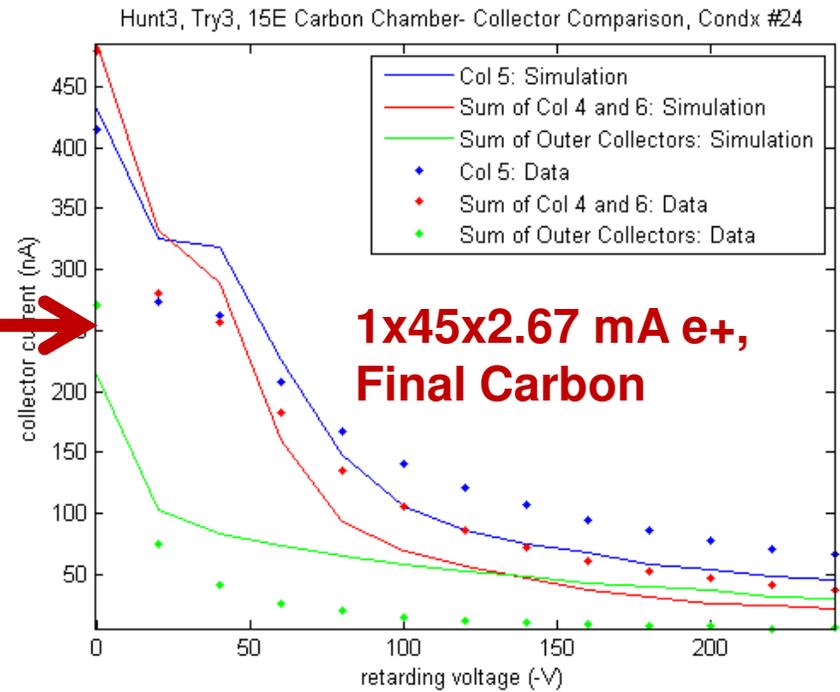
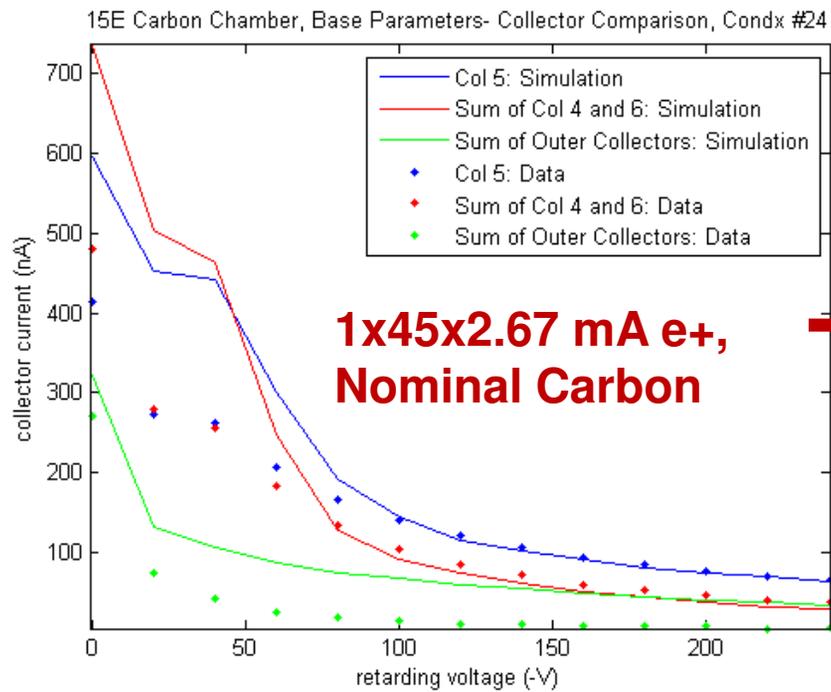
9x1x4 mA, e⁻, 280ns

Strongest Parameter, Condx #30











- Best fit parameters shown below
 - Note very low peak SEY ($\sim .9$) for Carbon and NEG coatings
 - Very low quantum efficiency for NEG is probably due to overestimation of photon flux
 - NEG chamber is in a straight section, far from any dipoles, so flux is difficult to estimate

Parameter	Description	Nominal Value(s)	Final Value: Al	Final Value: Carbon	Final Value: NEG
dtspk	Peak "true secondary" yield	1.8 (Al), .8 (C, NEG)	2.18	0.618	0.715
P1rinf	"Rediffused" yield at infinity	0.2	0.227	0.221	0.173
dt0pk	Total peak yield (δ_{max})	2.0 (Al), 1.0 (C, NEG)	2.447	0.879	0.928
P1epk	Low energy elastic yield ($\delta(0)$)	0.5	0.416	0.26	0.452
E0tspk	Peak yield energy (E_{max})	310 (Al), 500 (C, NEG)	314	486	500
queffp	Quantum efficiency	0.1	0.106	0.096	0.027





- A great deal of RFA data has been taken throughout the CEsrTA program
 - RFAs have been installed in drifts, dipoles, quadrupoles, and wigglers
- Several mitigation techniques have been investigated
 - In drifts, beam pipe coatings (TiN, Carbon, and NEG) all seem quite effective in suppressing secondary yield
 - Primary electrons could still be an issue
 - In dipoles, TiN coating was found to be very effective
 - Grooves + TiN is even better
 - TiN also suppresses the cloud in quadrupoles
 - A clearing electrode was found to be most effective in a wiggler chamber
 - Also gets rid of primary electrons
- A systematic method has been used to improve agreement between RFA data and simulation, and best fit simulation parameters have been obtained.
- Future work includes:
 - Quantifying errors and correlations in best fit parameters
 - Repeating analysis for RFAs in magnetic fields
 - Continuing development of integrated RFA models
- Detailed comparisons of RFA, SPU, and TE-Wave measurements

