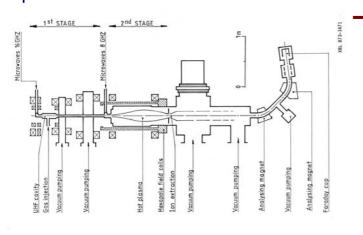


ECR ion sources have made remarkable improvements over the last few decades

Supermatios (Geller, 1974) 15 eµA of O⁶⁺



Factor 200 increase

VENUS (2007) 28 GHz 2850 eµA of O⁶⁺

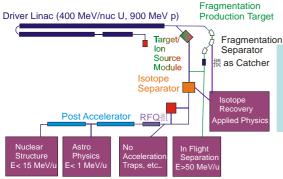


Generation 1.0 f≤10 GHz

Generation 2 10 < f ≤ 20 GHz Generation 3.0 20<f<40 GHz

Demand for increased intensities of highly charged heavy ions continues to grow

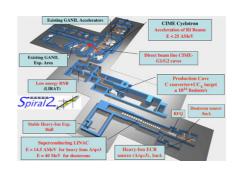
FRIB MSU, USA



270 eµA U³³⁺ and 270 eµA U³⁴⁺

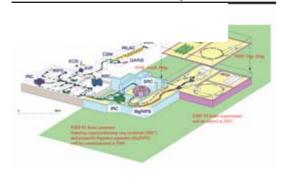
Experimental Areas

SPIRAL 2, GANIL, France



1mA Ar¹²⁺

RIKEN, Japan



525 eµA U³⁵⁺



750 μA Bi³⁵⁺



Present Performance of 3rd Generation Sources

Ion	Intensity eµA	Source
Ar ¹²⁺	860	VENUS
Xe ²⁷⁺	455	SECRAL
Bi ³⁰⁺	225	VENUS
Bi ⁴¹⁺	22	SECRAL
U ³³⁺	205	VENUS

Projected Requirements

Ion	Intensity eµA	Project
Ar ¹²⁺	1000	Spiral 2-GANIL
Bi ³⁵⁺	750	IMP HIRFL
U ³³⁺	270	FRIB
U ³⁵⁺	525	HRIBF- RIKEN



Third Generation Superconducting ECR Sources

- SuSI MSU
- SC-ECR RIKEN
- -----
- SECRAL IMP
- VENUS LBNL

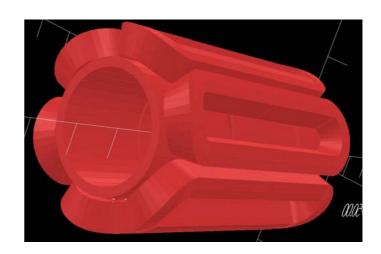


SECRAL*, IMP, Lanzhou, China



In operation at 18 and 24 GHz

3.7 T axial, 2 Tesla radial

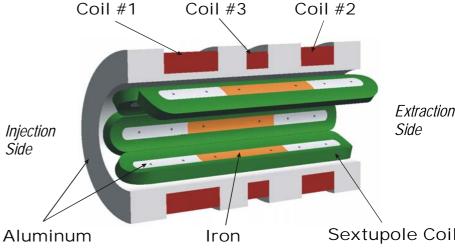


Solenoid in sextupole



VENUS 28 GHz





Sextupole-in-Solenoid

Achieved magnetic fields Binj ≤4 T, Bext ≤ 3 T, Brad≤2.2 T ECREVIS, SERSE, SUSI, MS-ECRIS, RIKEN SC-ECR

lacktriangle



Standard Model for ECR ion sources

Frequency scaling $n_e \alpha \omega_{rf}^2$

$$B_{ecr} = \underline{m_e \, \omega_{rf}}$$

Confinement criterion at 28 GHz

Bconf ≥2 B_{ecr} at walls

$$B_{inj} \sim 3 - 4 B_{ecr}$$
 on axis

$$B_{rad} \ge 2 B_{ecr}$$
 on the walls

$$B_{min} \sim 0.5-0.8 B_{ecr}$$
 on axis

$$I \propto \omega_{rf}^2/m$$
 $I \propto n_{ion} / \tau_{ion}$

Enhancements

Electron sources (Bias Probe, Electron Gun, Plasma First Stage, Wall Coatings)

2 Frequency Heating

Solids Ovens, Direct Insertion, Sputtering



Fourth Generation ECR Ion Sources



- Model calculations for 4th Generation source
- Choose 56 GHz (2 times 28)
- Conventional coil geometry

For a 56 GHz ECR $B_{ecr} = 2 T$

Confinement criterion

 $B_{conf} \ge 2 B_{ecr}$ at walls

 $B_{ini} \sim 3$ B_{ecr} on axis

 $B_{rad} \ge 2 B_{ecr}$ on the walls

ECRIS-56

$$B_{inj} \sim 6 T$$

$$B_{ext} = 4 T$$

$$B_{rad} = 4 T$$

ECRIS-56 Magnetic field is a challenge



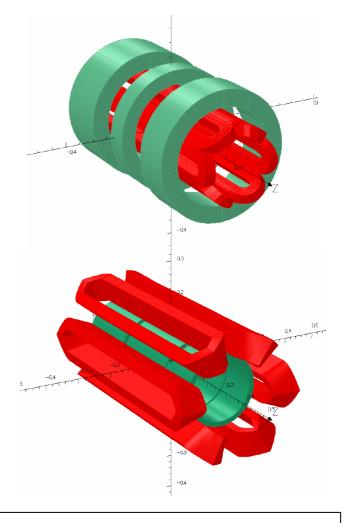
ECR Coil Layouts

Sextupole-in-solenoid:

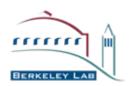
- Minimizes the peak field in the sextupole
- Solenoid field causes strong asymmetric forces on the sextupole coil ends

Solenoid in sextupole:

- Minimizes the influence of the solenoid on the sextupole coil field and forces
- More compact
- Higher field in the sextupole coil (larger radius)
- Strong forces on the solenoid coils
- Iron contribution less effective at high field

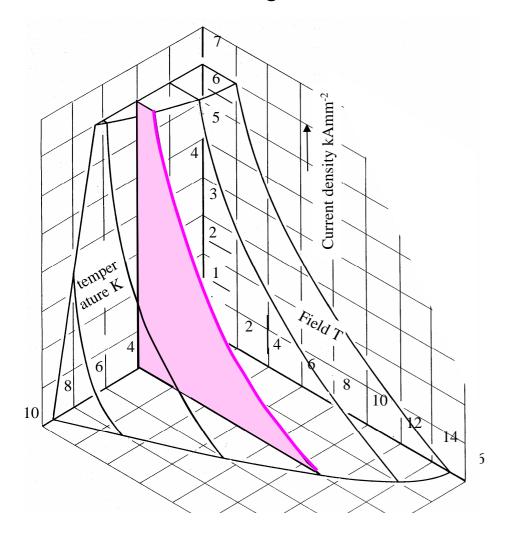


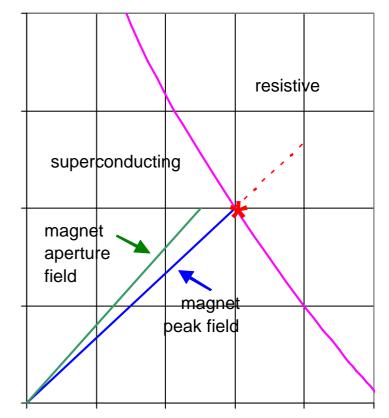
LBNL group chose the Sextupole-in-Solenoid because it has the potential to reach higher magnetic fields



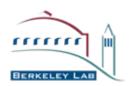
Superconducting Magnets

Critical line and magnet load lines

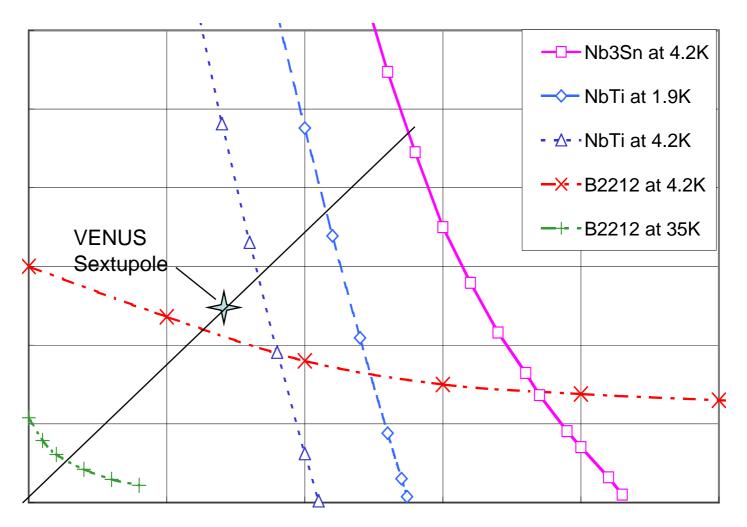




we expect the magnet to go resistive 'quench' where the peak field load line crosses the critical current line *



Engineering Current Densities for various materials



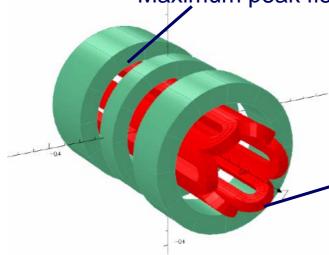


LBNL Design Effort for a 56 GHz ECR

- LBNL Supercon group is developing Nb₃Sn magnets (dipoles and quadrupoles for the LHC Upgrade)
- Supercon and the 88-Inch Cyclotron ECR group developed the VENUS magnet structure
- LBNL R&D funds have supported a preliminary design effort for a Nb₃Sn ECR magnet structure for a 4th Generation ECR ion source

Sextupole-in-Solenoid for 56 GHz: Clamping Structure





There are two limits to the maximum achievable field with this design

Maximum force on the end point (up to 175 MPa)

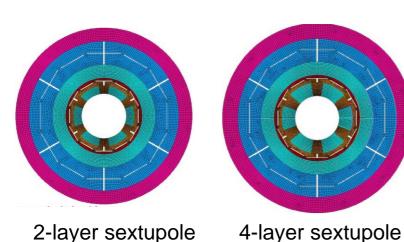


- In the end region each layer is subdivided in two blocks of conductors separated by end-spacers.
- The number of turns per block and the relative axial position of the end spacers were optimized to reduce the peak field in the end region.
- The coils are lengthen to reduce the peak field
- Shell type support structure
 Claude Lyneis Cyclotrons 10 Lanzhou



Magnetic Design

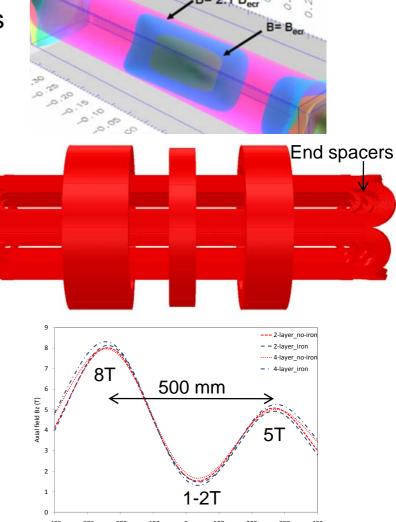
- Cos (3θ) sextupole winding, end spacers
- Keystone Rutherford cable, 15 mm wide
- Two and four layer options:



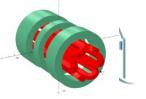
Prototype cable



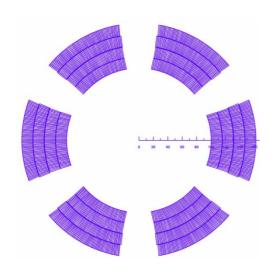




Axial location z (mm)



Sextupole design concepts



Cable properties				
Strand Dia	0.8 mm			
Fill factor	~ 33%			
No strands	35			
Cable	~ 15.2x1.5 mm			

- 4-layer coils using cables (675 conductors/coil)
- The same cable design is currently used by the LARP program to develop high field quadrupoles for future LHC luminosity upgrades (peak fields 15 T)

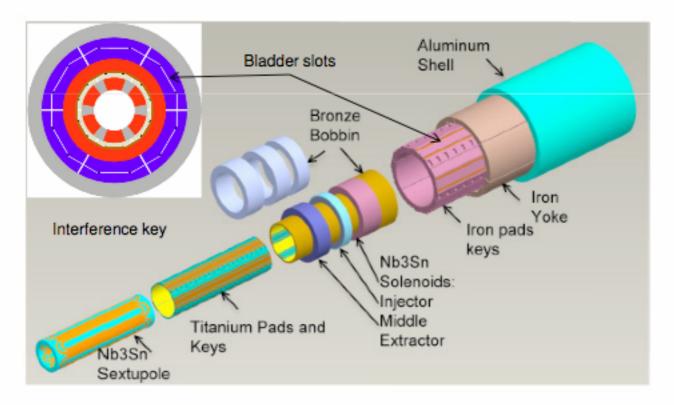


 The cable design requires high 8.2kA current leads, the 56 GHz cryostat will most likely require He filling during operation.



Shell-based Mechanical Structure

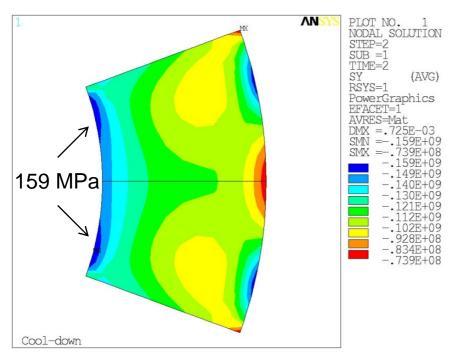
- Primary mechanical support is provided by a thick Aluminum shell
- Assembly (warm) pre-load by pressurized bladders and interference keys
- Pre-load increase at cool-down due to shell-yoke differential contraction
- The coils remain in compression up to the operating point





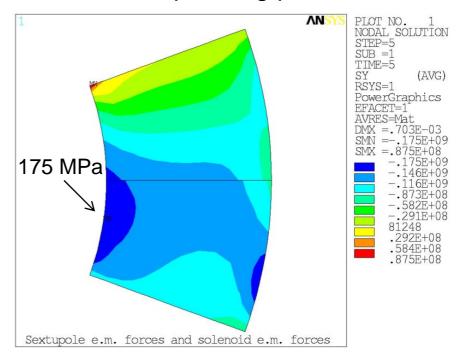
Sextupole coil stresses for 56 GHz

After cool-down



- Solenoid intercepts 50% of compressive the force
- Maximum stress 159 MPa in the "solenoid center" region

At the operating point



- Asymmetric stress profile in the "solenoid end" region
- Maximum stress 175 MPa



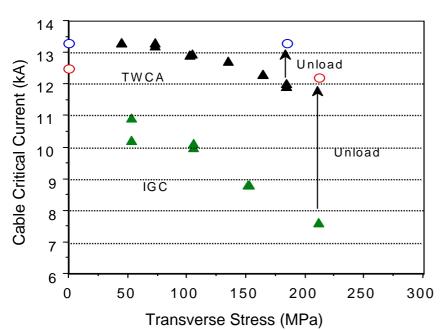
Nb₃Sn Challenges

Brittleness:

- React coils after winding
- Epoxy impregnation

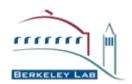
Strain sensitivity:

 Mechanical design and analysis to prevent degradation under high stress



00		
Material	NbTi	Nb₃Sn
Dipole Limit	10-11 T	16-17 T
Reaction	Ductile	~675°C
Insulation	Polymide	S/E Glass
Coil parts	G-10	Stainless
Axial Strain	N/A	< 0.03 %
Transverse stress	N/A	< 200 MPa

LAYER 2

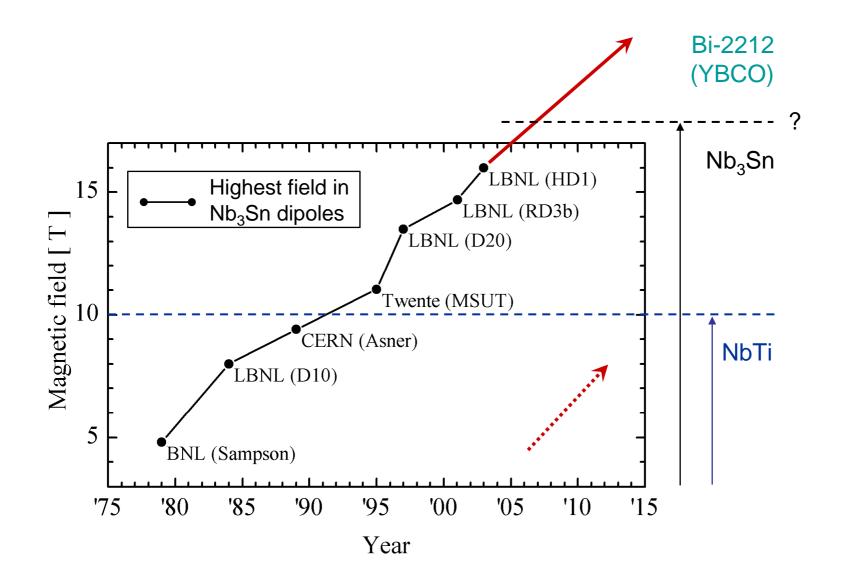


ECRIS-56 -- Other Challenges

- Gyrotrons at 53, 60 and 70 GHz at 200 kW for 100 ms can be run at 30 kW cw. "No problem" to extend to 50 kW cw.
- Power requirements and chamber cooling
 - Total RF power $\sim n_e V$ or $\sim f^2 * V$. VENUS at 1 kW/liter has not reached the saturation power density
 - The heat deposition on the plasma wall is highly non-uniform and 'burnout" is a concern.
- Bremsstrahlung heating of the cryostat will require significantly more cryo-cooling power.



Progress in Maximum Field





• Why is this the time to be developing a 4th Generation ECR Ion Source?

- Heavy ion driver requirements are beyond the reach of 3rd
 Generation Source performance
- The R&D time needed for a new generation source is quite long. Example: VENUS (9 years from proposal to 28 GHz operation)
- High Energy Physics is driving the technology for Nb3Sn magnets—LHC upgrade—Nuclear physics can take advantage of these developments
- While the magnets are the most demanding technical challenge— The design studies show it is feasible to build an 4th Generation source at $f \ge 50$ GHz
- The cost of such a source should only be about 2 or 3% of the cost of a state-of-the-art Rare Isotope Beam facility



4th Generation ECR Ion Source

- As Geller predicted, frequency scaling promises us higher intensity and higher charge states
- There are technical challenges, but there are no "show stoppers"
- The design and construction of a magnet structure for a 4th Generation ECR is the most challenging task
- Next step, construction of a prototype Nb₃Sn ECR ion source magnet structure for 56 GHz

ECRIS-56 A new twist on an old idea!

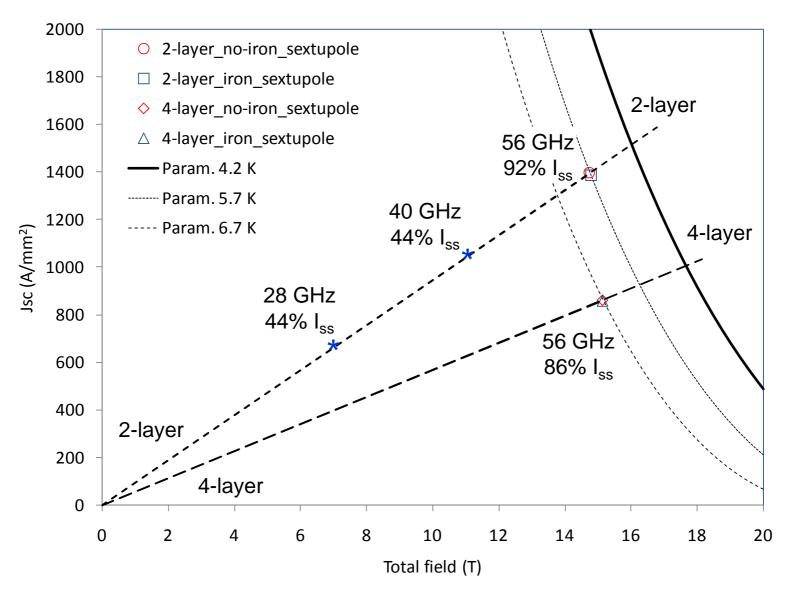
"... we propose a bolder extrapolation.

...With a 56 GHz generator, TRIPLEMAFIOS should furnish up to U⁵⁰⁺ ions!"

Richard Geller, IEEE-Trans NS-23, 1976



Operational Conditions





Quench Protection

Design parameters: 2-layer design: I_{op}=13.2kA; L= 33mH; U=2.9MJ

4-layer design: $I_{op} = 8.2kA$; L=163mH; U=5.5MJ

Active protection and full heater coverage is required:

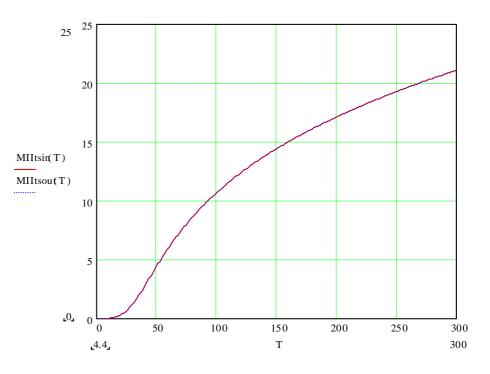
• T_{max} for 100% heater coverage: 390K (2-layer); 260K (4-layer)

• T_{max} for 75% heater coverage: 430K (2-layer); 280K (4-layer)



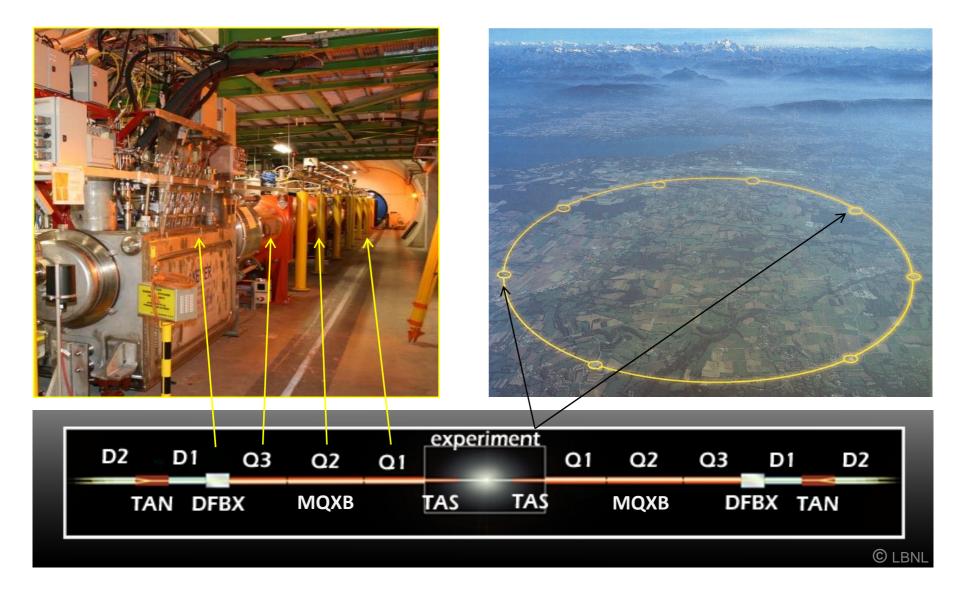








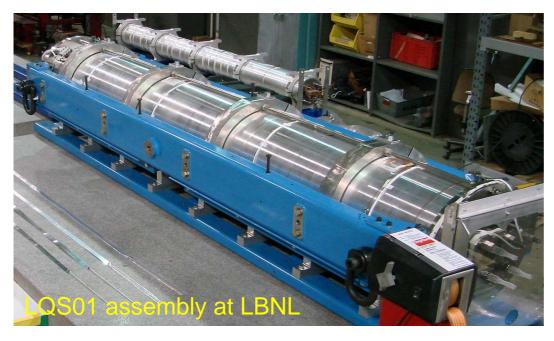
Nb₃Sn Magnets for the LHC Upgrades



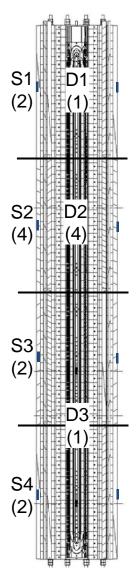


Long Quadrupole Shell (LQS)

- TQ length scale-up from 1 m to 4 m
- Coil Fabrication: BNL+FNAL
- Coil and magnet instrumentation: LBNL
- Mechanical structure and assembly: LBNL
- Test: FNAL (November 2009)
- Target gradient 200 T/m









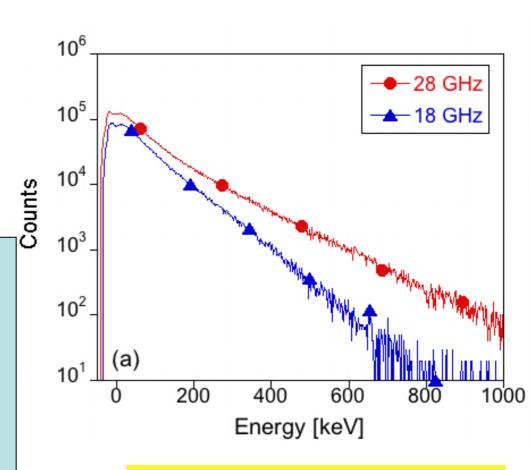
Summary

- The magnet requirements for 56 GHz operation are challenging but feasible using Nb₃Sn conductor
- Design tools and fabrication technology are available
- Sextupole-in-solenoid preferred to achieve the highest field
- Shell based support structures are suitable to provide the required pre-load and prevent conductor motion at all coil locations
- Next step: detailed engineering design and prototype fabrication
- Nb₃Sn properties also provide key advantages in the field range accessible to NbTi



VENUS Bremsstrahlung Measurements

- Measurements of axial bremsstrahlung at 18 and 28 GHz
- B fields are scaled by frequency
- $B_{min}/B_{ecr} = 70\%$
- RF input power 1.5 kW
- Bremsstrahlung is more intense at 28 GHz
- Much larger high energy tail at 28 GHz
- Cryostat shielding is ineffective above 500 keV
- Mean electron energy increases with RF frequency Alain Girard (2000)



More shielding and 4 K cooling will be required for 56 GHz



New VENUS Plasma Chamber with X-ray Shielding and Increased Water Cooling

