## Tapering Enhanced Stimulated Superradiant Amplifier

#### **IFEL decelerator**

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# Outline

- IFEL background
- Tapering Enhanced Stimulated Superradiant amplifier
- Low gain regime
- High gain regime
- Few case studies
  - 10.6 micron proof of principle experiment at ATF
  - X-rays amplification at LCLS
- Conclusion

#### Inverse FEL Background

• Rubicon IFEL experiment recently demonstrated high quality acceleration of 50 MeV e-beam at BNL ATF in a strongly tapered helical undulator



# What happens if we do the same experiment in reverse ?

- 10.3 um-driven IFEL decelerator.
- Decelerate 90 MeV to 50 MeV using 100 GW input power.
- Potentially demonstrate ~40 % (!!!) energy extraction from a relativistic electron beam
- For comparison, FEL's typically get efficiencies of ~0.1%



## **TESSA** concept

- Reversing the laser-acceleration process, we can extract most of the energy from an electron beam provided:
  - A microbunched input e-beam
  - An intense input seed

# Tapering-Enhanced Stimulated Superradiant



# Applications

- High efficiency conversion of electron beam energy to coherent radiation opens door to very high average power light sources.
- Wavelength set by e-beam energy and resonant condition -> wide tunability
  - High average power IR and visible lasers.
  - X-rays.
  - EUV-L applications.
- LLNL Paladin experiment in the '80s showed large conversion efficiency using tapered undulator and waveguide at 250 GHz
- Differences from current optimization tapering schemes for short wavelength FEL
  - Strong tapering (both period and amplitude)
  - large seed intensity much above FEL saturation level –
  - Higher initial input energy (so that at the decelerated output energy undulator is still feasible)

# **TESSA** theory: tapering optimization

- First approximation. Frozen field, small gain regime
  - Do not compensate for radiation emission.
  - For constant period there is an analytical solution varying only the magnetic field amplitude (larger gap, weaker permanent magnets, smaller current, etc.)
  - Varying also period allows more flexibility and might be technologically simpler.
- Optimum tapering is obtained by matching resonant energy gradient to the available ponderomotive gradient



• Efficiency proportional to number of periods and laser normalized vector potential.

$$\eta \cong 2\pi N_u K_l \sin \psi_r$$

- Including diffraction
  - higher intensity smaller spot size
  - shorter interaction length longer Rayleigh range



# **Nocibur Genesis simulation**

- Input power: 100 GW
- Output power: 127 GW
- 1 kA input beam @ 70 MeV
- Decelerated to ~40 MeV

80

70

60

50

40

30

0.0

1D sim

0.2

0.3

Position (m)

0.4

120

100 1 18

meld-1

0.5

6 22

00

20

0.1

Energy (MeV)



#### Application to short wavelengths: Where to get the high **intensity** seed?

- Oscillator configuration
  - Low gain regime
  - Build-up is complex problem
  - Could use low rep-rate igniter pulse
- Afterburner following FEL amplifier
  - Simpler
  - Use mirrors to refocus radiation (issues of damage on mirrors?)
  - Efficiency limited by seed power.... but





# Low gain vs high gain regime

- Low gain
- Neglect radiation power increase along the undulator.
- Trapping and deceleration will work for any beam current.
- Beam-independent efficiency.
- Proportional to sqrt (input seed power)



#### • High gain regime

- As the radiation power increases we can taper and extract energy more efficiently ! Power grows along the undulator as the particles are decelerated.
- $\circ$  Tapering depends on amplitude of ponderomotive potential  $\propto \sqrt{I}$
- But some of the generated power diffracts away (3D effects)
- How to optimize?

#### High Gain TESSA Genesis Informed Tapering optimization Scheme

Due to strong diffraction, and external seed laser, gain guiding singlemode formulas not sufficient to describe laser driving intensity ⇒ Solve numerically with help of 3D FEL code -- Genesis!

 $\Rightarrow$  Genesis Informed Tapering Scheme (GITS) optimization

Solve tapering period-by-period

- Run Genesis on a period
- Measure min intensity seen by particles => threshold for capture
- Calculate new period and undulator parameter
- Saves tapering as well as simulated data

GITS offers options to dynamically optimize different simulated e-beam and radiation parameters: maximize power, minimize detrapping, etc.

# Single stage TESSA 14 m long @ 13 nm

enable >10kW

50 100 150 200 250 x-plane

- Input energy: 750 MeV
- Input current: 3 kA ٠
- Gap const 7.5 mm ٠
- Beam size 20 um (emittance 0.5 um) •
- Resonant phase: 0.5 to 1.0 ٠
- Lu = 14 m, zw = 3 m, zr=1 m٠

0 0

Input power: 5 GW

@14 m

250

200

170 um rms

Output power: 1.0 TW (44% conversion)





## Time dependent simulation



## Conclusion

#### • TESSA - tapering enhanced stimulated superradiant amplifier

- Strongly tapered helical undulator
- Refocus FEL radiation to drive beyond FEL saturation
- 3D simulation guided tapering
- "Nocibur" low gain demonstration at ATF will be first step.
  - Measure energy beam spectrum, CO2 output power, mode quality, spectrum.
  - All hardware required for the experiment already in hand !
- Strongly driven system: effects of energy spread and emittance reduced compared to SASE FEL.
- Applications in various spectral ranges
- Many interesting points
  - Mirrors to refocus radiation
  - Start-up in oscillator configuration
  - Side-band supression

# Efficiency estimate: Low gain regime

- Define  $\eta = \frac{\gamma_0 \gamma}{\gamma_0}$
- For constant *K*<sub>1</sub> and small efficiency
- Taking into account diffraction, we need to find a compromise between:
  - higher intensity smaller spot size
  - shorter interaction length longer Rayleigh range

$$\frac{\partial}{\partial z_r} \left[ \int_0^{L_u} \sqrt{\frac{1}{z_r} \frac{1}{1 + (\frac{z - \frac{L_u}{2}}{z_r})^2}} dz \right] = 0$$

$$\int_0^1 \frac{1}{1 + (\frac{z - \frac{L_u}{2}}{z_r})^2} dz$$
Peak for  $z_r = L_u/6$ 

Plug-in numbers for Nocibur

η	0.45	estimate !
N <sub>u</sub>	11	Reasonable
K	0.016	
w <sub>0</sub>	1 mm	
λ	10.3 um	
Р	100 GW	



 $\eta \cong 2\pi N_{\nu}K_{l}\sin\psi_{r}$ 

Х

0.4 0.6

0.2

08